

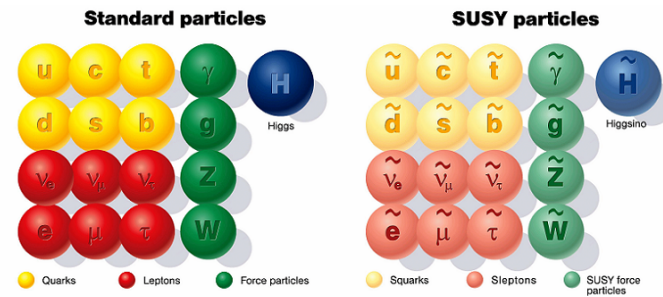
SUSY after the Higgs discovery

G. Ross, Madrid, September 2013



Low scale SUSY

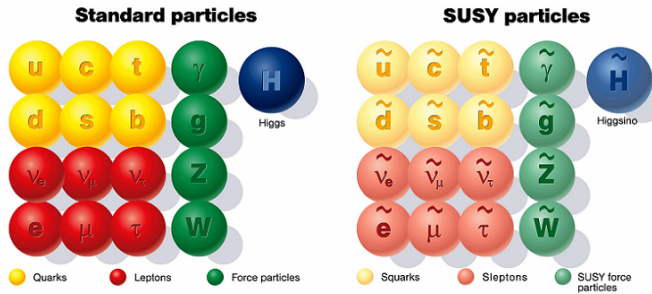
MSSM:



...motivation?

Low scale SUSY

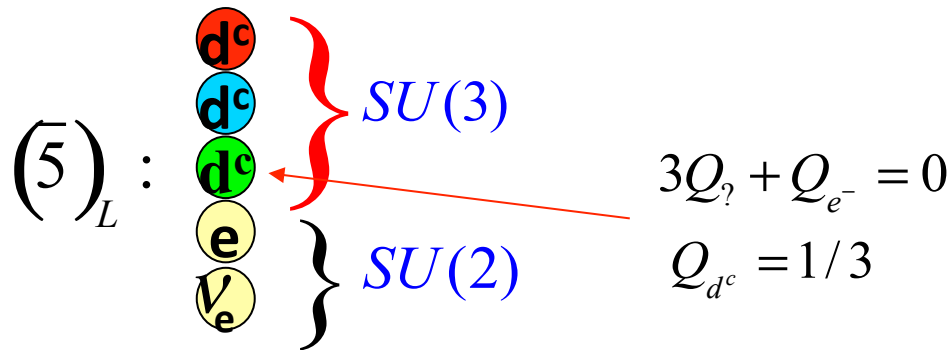
MSSM:



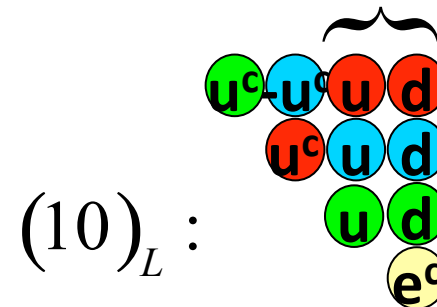
GUTS:

e.g. $SO(10) \supset SU(5) \supset SU(3) \otimes SU(2) \otimes U(1)$
 $\qquad\qquad\qquad g_5 \qquad\qquad\qquad g_3 \qquad\qquad\qquad g_2 \qquad\qquad\qquad g_1$

Georgi Glashow 1974



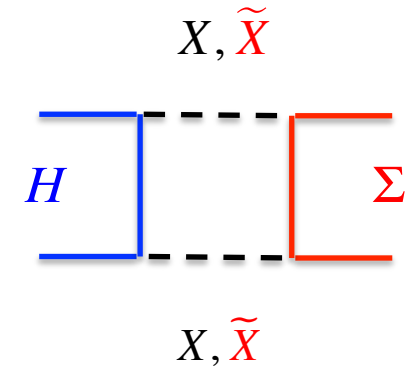
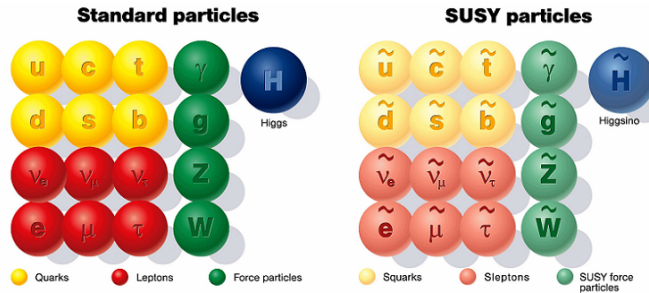
LH states SU(2) doublets



$(16)_L = (10)_L + (\bar{5})_L + (1)_L$ $\leftarrow \nu_{e,L}^c \equiv \nu_{e,R}$

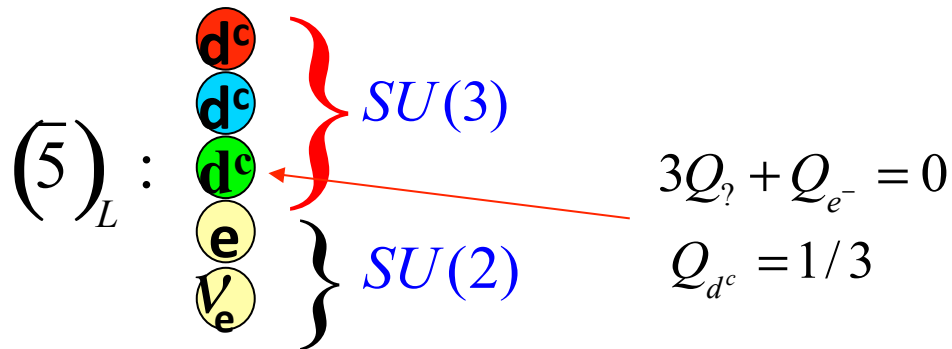
Low scale SUSY

MSSM:

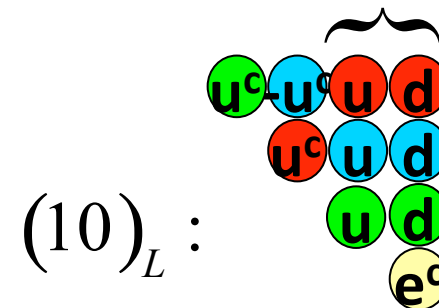


SUSY GUTS: the hierarchy problem

e.g. $SO(10) \supset SU(5) \supset SU(3) \otimes SU(2) \otimes U(1)$
 $g_5 \quad g_3 \quad g_2 \quad g_1$



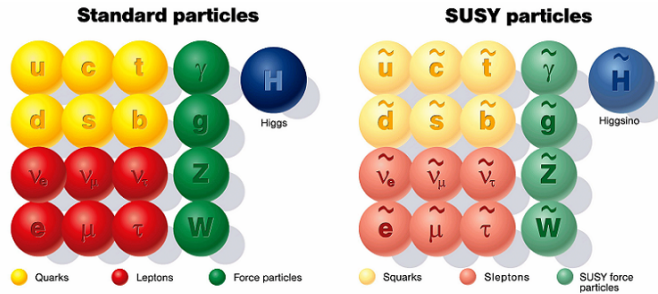
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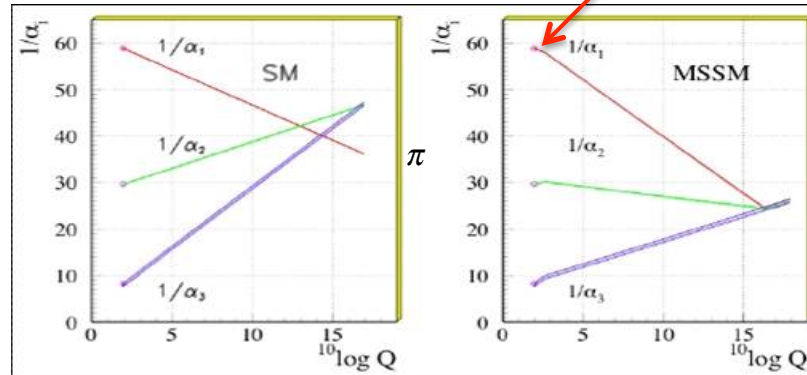
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Low scale SUSY

MSSM:



SUSY GUTS:



The (SUSY) Standard Model as an EFT:

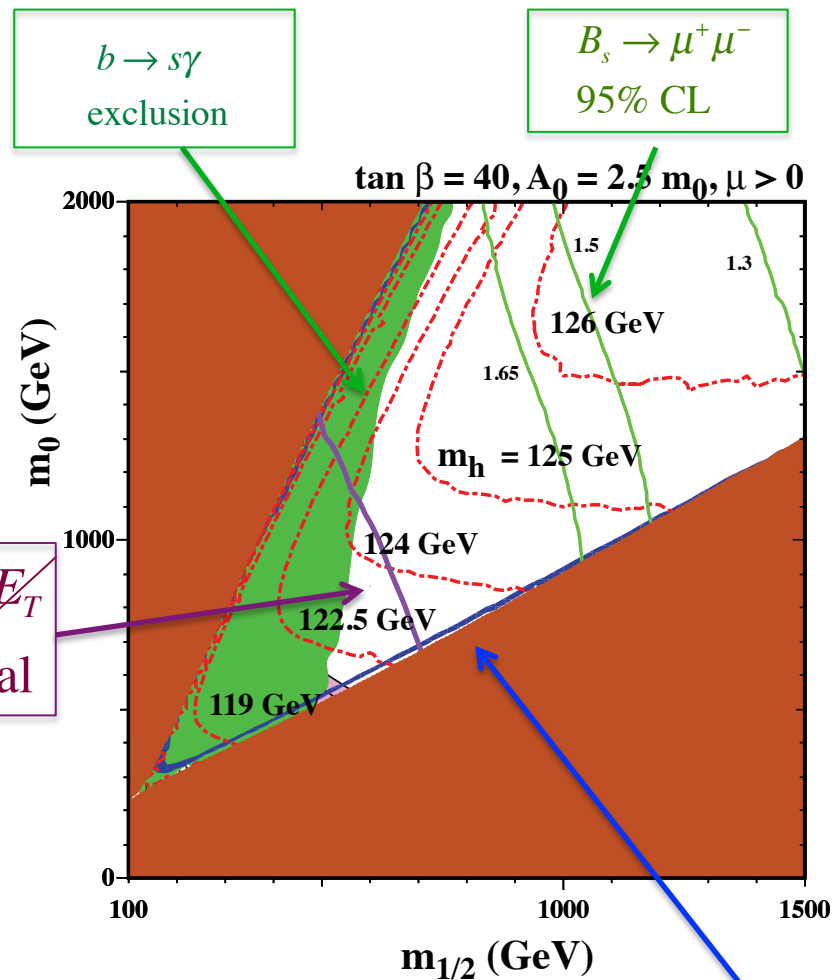
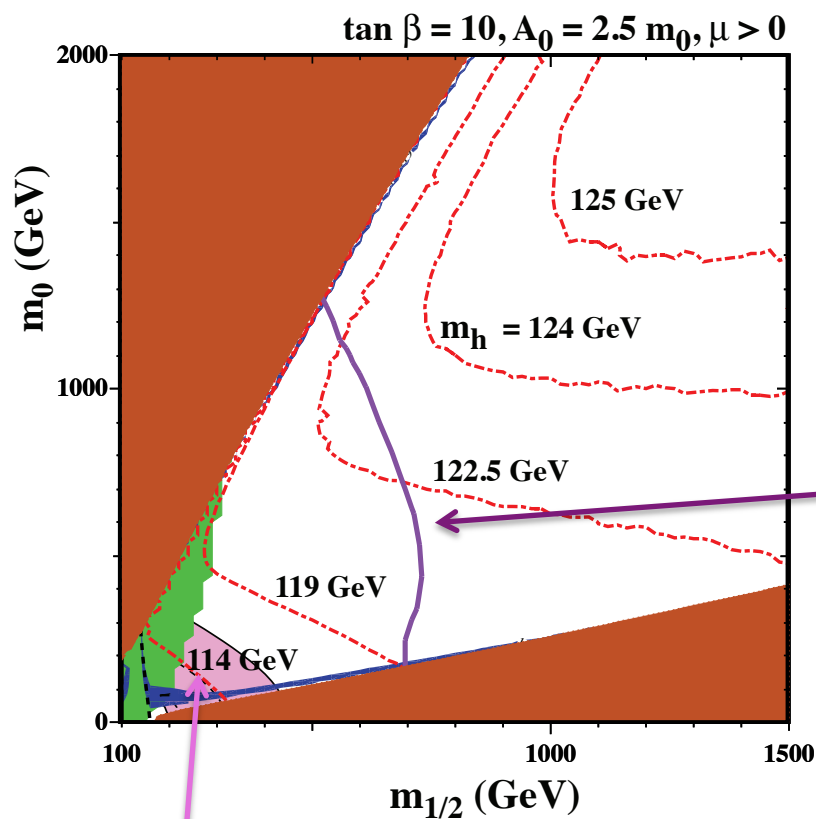
$$A_\mu \checkmark, \Psi \checkmark, H \checkmark ?$$

$$M_{Higgs}, M_{W,Z} \ll M_{Planck}, M_{GUT}, \dots \checkmark$$

CMSSM fits after the Higgs: $\gamma_i = \mu_0, m_0, m_{1/2}, A_0, B_0$

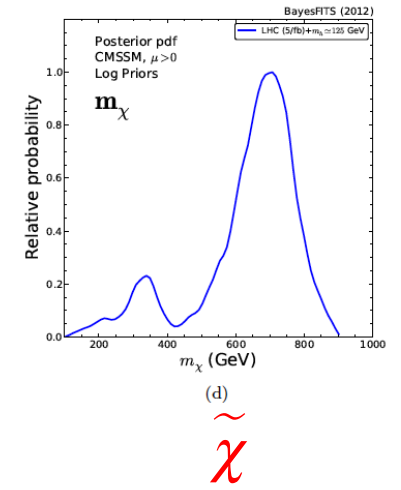
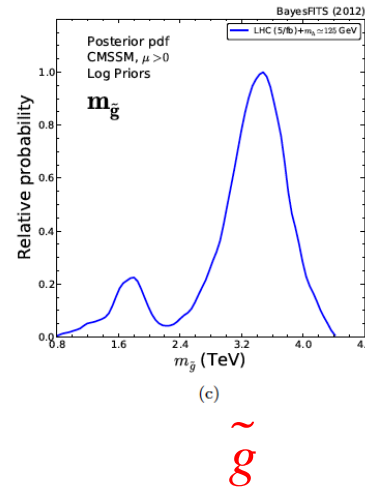
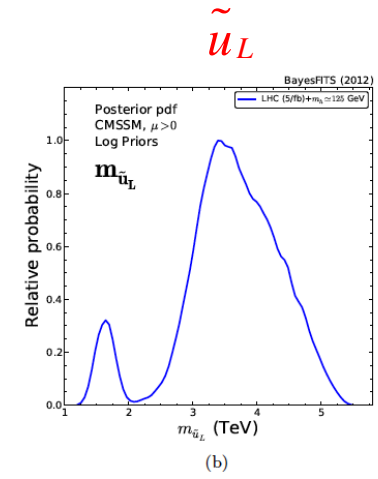
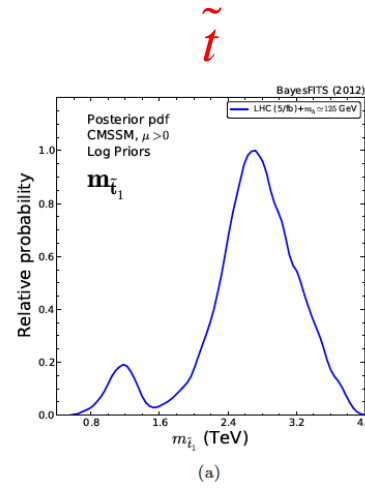
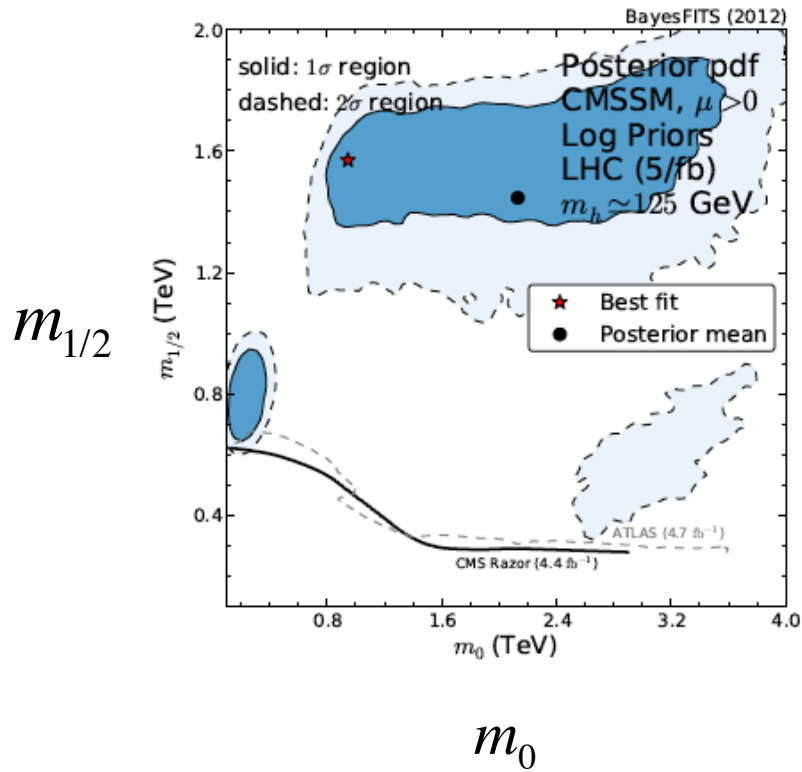
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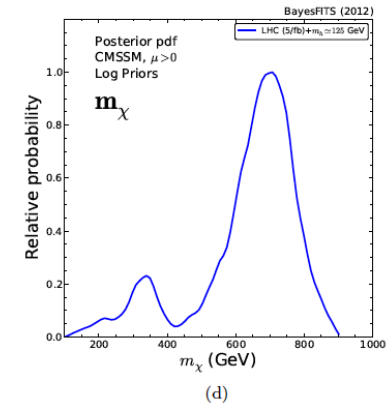
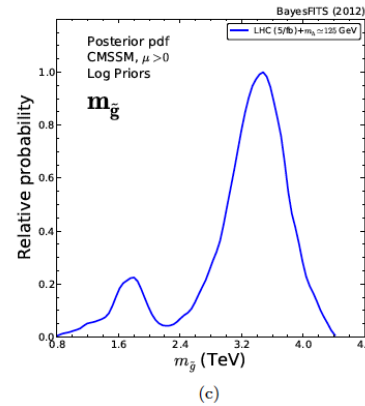
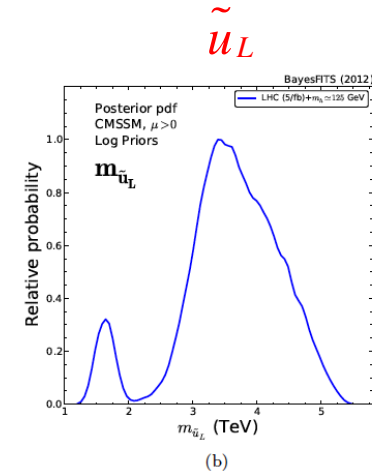
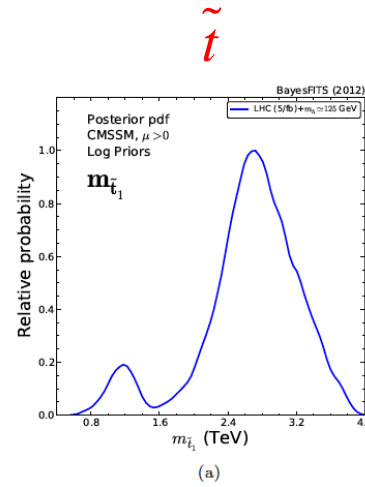
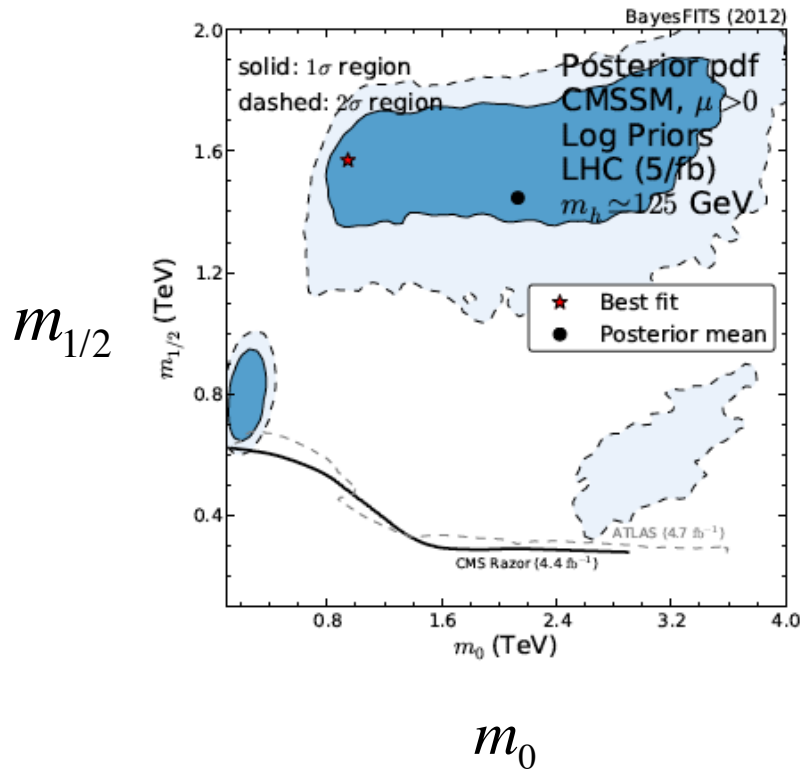


Under pressure!

SUSY spectrum : CMSSM



SUSY spectrum : CMSSM



Little hierarchy problem:

$$v^2 \sim \delta m_{H_u}^2 \simeq -\frac{3y_t^2}{4\pi^2} \left(m_{stop}^2 + \frac{g_s^2}{3\pi^2} m_{gluino}^2 \log\left(\frac{\Lambda}{m_{gluino}}\right) \right) \log\left(\frac{\Lambda}{m_{stop}}\right) \quad ?$$

This talk $\Lambda \sim M_{GUT}$

breaking

Little hierarchy problem \Rightarrow definite SUSY structure

MSSM: 105 +(19) Parameters

$$M_Z^2 = \sum_{\tilde{q}, \tilde{l}} a_i \tilde{m}_i^2 + \sum_{\tilde{g}, \tilde{W}, \tilde{B}} b_i \tilde{M}_i^2 + \dots$$

$$M_{\tilde{g}} > 1\text{TeV} \Rightarrow \Delta > b \frac{\tilde{M}^2}{M_Z^2} \sim 100$$

\Rightarrow Correlations between SUSY breaking parameters and/or additional low-scale states

breaking

Little hierarchy problem \Rightarrow definite SUSY structure

MSSM: 105 +(19) Parameters

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\Rightarrow Correlations between SUSY breaking parameters and/or additional low-scale states

Fine Tuning measure:

$$\Delta(a_i) = \left| \frac{a_i}{M_Z} \frac{\partial M_Z}{\partial a_i} \right|,$$

$$\Delta_m = \text{Max}_{a_i} \Delta(a_i), \quad \Delta_q = \left(\sum \Delta_{\gamma_i}^2 \right)^{1/2}$$

Ellis, Enquist, Nanopoulos, Zwirner
Barbieri, Giudice

Fine tuning from a likelihood fit:

If v included as a "Nuisance" variable

$$L(\text{data} \mid \gamma_i) \propto \int dv \delta(m_Z - m_Z^0) \delta\left(v - \left(-\frac{m^2}{\lambda}\right)^{1/2}\right) L(\text{data} \mid \gamma_i; v)$$

$$= \frac{1}{\Delta_q} \delta(n_q (\ln \gamma_i - \ln \gamma_i^S)) L(\text{data} \mid \gamma_i; v_0)$$

Fine tuning

Ghilenca, GGR
Cabrera, Casas, de Austri

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Fine tuning

Probabilistic interpretation:

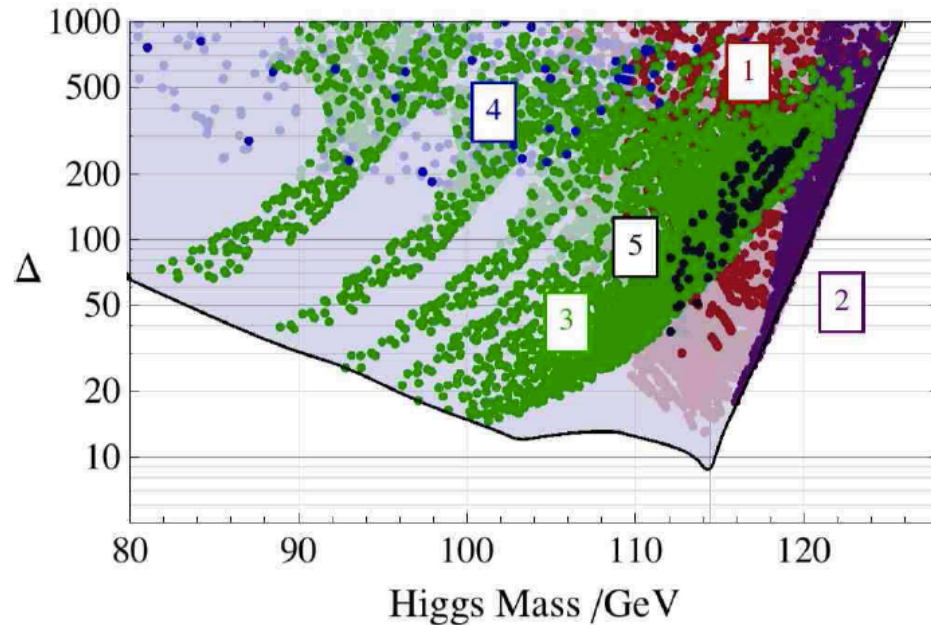
$$\chi_{new}^2 = \chi_{old}^2 + 2 \ln \Delta_q$$

$$\Delta_q \ll 100$$

● The CMSSM - before LHC

$$\gamma_i = \mu_0, m_0, m_{1/2}, A_0, B_0$$

$$v^2 = -\frac{m_{eff}^2}{\lambda_{eff}}$$



Relic density restricted

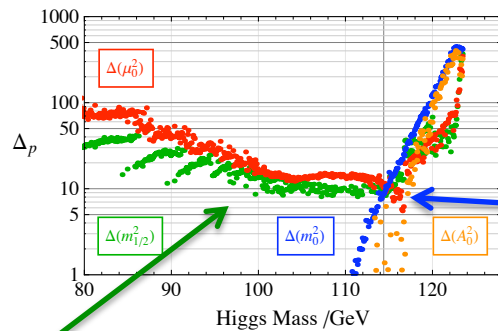
- 1 h^0 resonant annihilation
- 2 \tilde{h} t-channel exchange
- 3 $\tilde{\tau}$ co-annihilation
- 4 \tilde{t} co-annihilation
- 5 A^0 / H^0 resonant annihilation

Within 3σ WMAP:

$$\Delta_{Min}^{EW} = 15, \quad m_h = 114.7 \pm 2 GeV$$

< 3σ WMAP:

$$\Delta_{Min}^{EW} = 18, \quad m_h = 115.9 \pm 2 GeV$$



$\Delta(m_0^2)$
Limit of focus point

$$\Delta^\Omega = \max \left| \frac{\partial \ln \Omega h^2}{\partial \ln q} \right|_{q=m_0, m_{1/2}, A_0, B_0}$$

$$\Delta_{Min}^{EW+\Omega} = 29, \quad m_h = 117 \pm 2 GeV$$

λ_{eff} increases with m_H

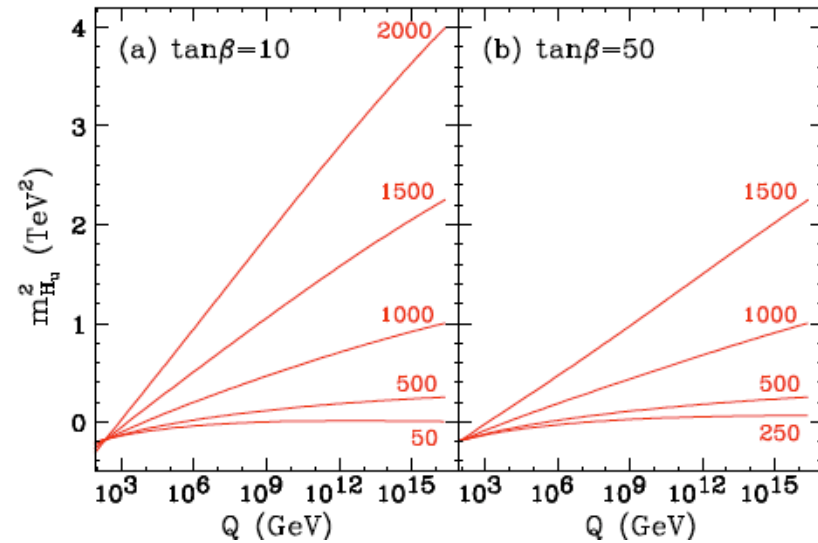
Focus Point

$$2|y_t|^2 (m_{H_u}^2 + m_{Q_3}^2 + m_{u_3}^2) + 2|a_t|^2$$

$$16\pi^2 \frac{d}{dt} m_{H_u}^2 = 3X_t - 6g_2^2 |M_2|^2 - \frac{6}{5}g_1^2 |M_1|^2$$

$$16\pi^2 \frac{d}{dt} m_{Q_3}^2 = X_t + X_b - \frac{32}{3}g_3^2 |M_3|^2 - 6g_2^2 |M_2|^2 - \frac{2}{15}g_1^2 |M_1|^2$$

$$16\pi^2 \frac{d}{dt} m_{u_3}^2 = 2X_t - \frac{32}{3}g_3^2 |M_3|^2 - \frac{32}{15}g_1^2 |M_1|^2$$



$$m_{H_u}^2(Q^2) = m_{H_u}^2(M_P^2) + \frac{1}{2} \left(m_{H_u}^2(M_P^2) + m_{Q_3}^2(M_P^2) + m_{u_3}^2(M_P^2) \right) \left[\left(\frac{Q^2}{M_P^2} \right)^{\frac{3y_t^2}{4\pi^2}} - 1 \right]$$

$$m_0^2$$

$$3m_0^2$$

$$\approx -\frac{2}{3}, Q^2 \approx M_Z^2$$

“Focus point”: $m_{H_u}^2(0) = m_{Q_3}^2(0) = m_{u_3}^2(0) \equiv m^2$

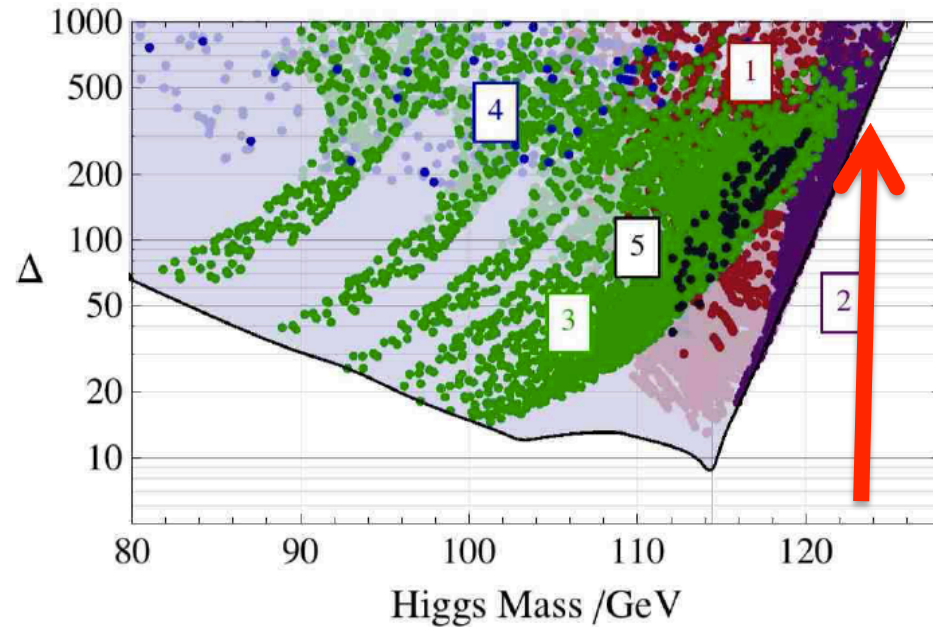
$$m_{H_u}^2(t_0) = a_0 m^2 + \dots, a_0 \leq 0.1$$

i.e. $m_{Q_3}^2, m_{u_3}^2 \gg M_Z^2$ possible

Natural choice

Feng, Matchev, Moroi
 Chan, Chattopadhyay, Nath
 Barbieri, Giudice
 Feng, Sanford

- The CMSSM - after Higgs discovery



$$M_{h^0}^2 = M_Z^2 \cos^2 2\beta + \frac{3M_t^2 h_t^2}{4\pi^2} \left(\ln\left(\frac{M_S^2}{M_t^2}\right) + \delta_t \right) + \dots$$

$M_S^2 = m_{q_3} m_{U_3}$

$\Delta_{Min} > 350, \quad m_h = 125.6 \pm 3 GeV$

Reduced fine tuning (c.f. CMSSM)

- New focus points?

Gauginos: $M_{\tilde{g}, \tilde{W}, \tilde{B}}$ Non-universal gaugino correlations

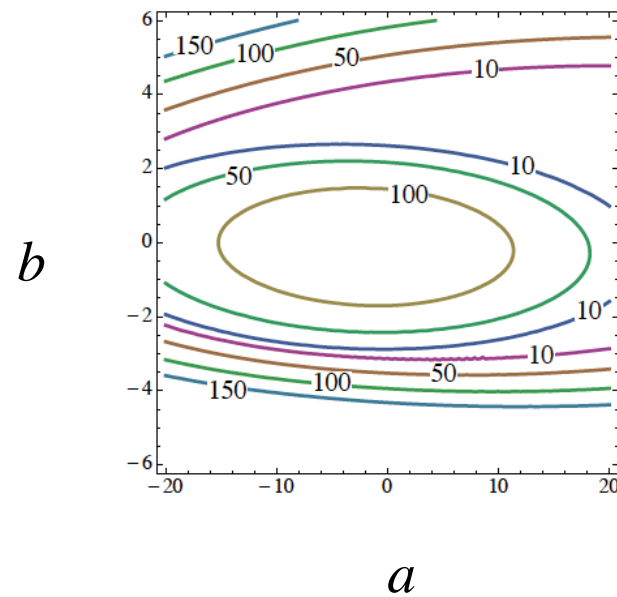
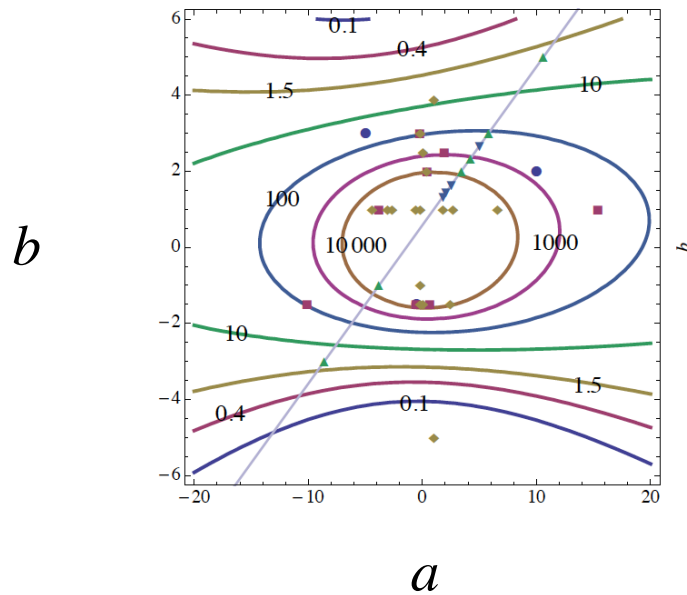
- New degrees of freedom

I. Reduced fine tuning : nonuniversal gaugino masses

$$16\pi^2 \frac{d}{dt} m_{H_u}^2 = 3 \left(2 |y_t|^2 (m_{H_u}^2 + m_{Q_3}^2 + m_{u_3}^2) + 2 |a_t|^2 \right) - 6g_2^2 |M_2|^2 - \frac{6}{5} g_1^2 |M_1|^2$$

New focus point: cancellation between M_3 and M_2 contributions if $|M_2|^2 \simeq |M_3|^2$ at M_{SUSY}

Horton, GGR
Choi et al...



$$M_3 : M_2 : M_1 = 1 : b : a$$

I. Reduced fine tuning : nonuniversal gaugino masses

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New focus point: cancellation between M_3 and M_2 contributions if $|M_2|^2 \approx |M_3|^2$ at M_{SUSY}

$$\Delta_{Min}^{(C)MSSM} = 60 (500), \quad m_h = 125.6 \pm 3 GeV$$

LHC8 SUSY bounds ✓

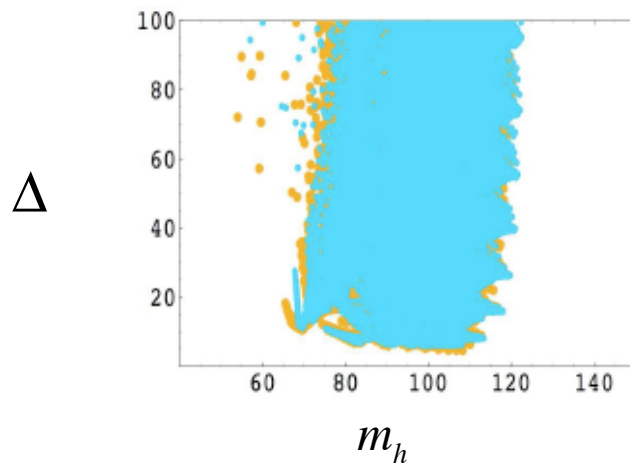
DM relic abundance ✓

DM searches ✗

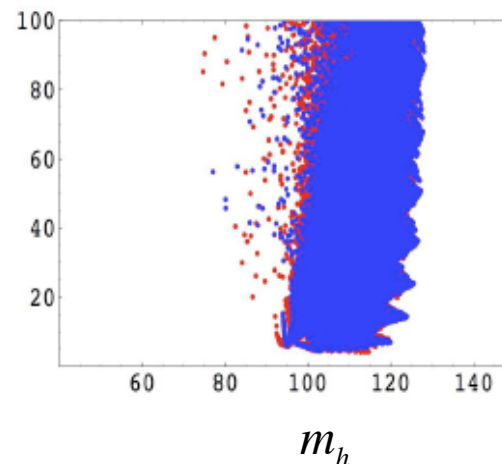
II. Reduced fine tuning : New heavy states - higher dimension operators

$$\delta L = \int d^2\theta \frac{1}{M_*} (\mu_0 + c_0 S) (H_u H_d)^2, \quad S = m_0 \theta\theta \quad \text{Dimension 5}$$

$$\delta V = \varsigma_1 (|h_u|^2 + |h_d|^2) h_u h_d + \varsigma_2 (h_u h_d)^2; \quad \varsigma_1 = \frac{\mu_0}{M_*}, \quad \varsigma_2 = \frac{c_0 m_0}{M_*}$$



MSSM



+ dim 5 operators

Cassel, Ghilencea, GGR
Casas, Espinosa, Hidalgo
Dine, Seiberg, Thomas
Batra, Delgado, Tait
Kaplan,

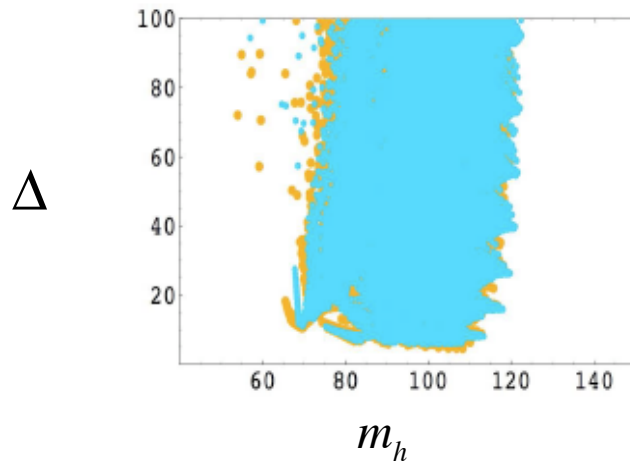
Even for $M_* = 65 \mu_0$ a significant shift of m_h for constant Δ

...effect mainly comes from ς_1 term

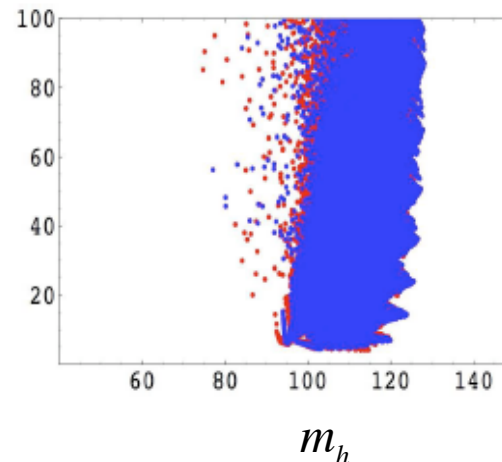
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MSSM



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Even for $M_* = 65 \mu_0$ a significant shift of m_h for constant Δ

...effect mainly comes from ζ_1 term ... origin?

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Singlet extensions

$$W = W_{\text{Yukawa}} + \lambda S H_u H_d + \frac{\kappa}{3} S^3 \quad \text{NMSSM}$$

$$W = W_{\text{Yukawa}} + (\mu + \lambda S) H_u H_d + \frac{\mu S}{2} S^2 + \frac{\kappa}{3} S^3 + \xi S \quad \text{GNMSSM}$$

$$\mu_S \gg m_{3/2} : W_{\text{eff}}^{\text{GNMSSM}} = (H_u H_d)^2 / \mu_s + \mu H_u H_d$$

$$\delta V = \frac{\mu}{\mu_s} (|H_u|^2 + |H_d|^2) H_u H_d \quad \checkmark$$

$$\zeta_2 \propto \frac{m_0^2}{M_*^2}$$

but see Lu et al



II. Reduced fine tuning : New heavy states - higher dimension operators

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$$\delta V = \frac{\mu}{\mu_s} (|H_u|^2 + |H_d|^2) H_u H_d \quad \checkmark$$

but are μ, μ_s naturally small?

SUSY extensions of the Standard Model

$$\begin{aligned}
 W = & h^E LH_d \bar{E} + h^D QH_d \bar{D} + h^U QH_u \bar{U} + \mu H_d H_u \\
 & + \lambda LL\bar{E} + \lambda' LQ\bar{D} + \kappa LH_u + \lambda'' \bar{U}\bar{D}\bar{D} \\
 & + \frac{1}{M} (QQQL + QQQH_d + Q\bar{U}\bar{E}H_d + LLH_u H_u)
 \end{aligned}$$

R-parity:

Z_2

$H_u, H_d +1$

$L, \bar{E}, Q, \bar{D}, \bar{U}, \theta -1$

SUSY states odd

Weinberg, Sakai

SUSY extensions of the Standard Model

$$\begin{aligned}
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 \end{aligned}$$

R-parity: Z_2

Z_N^R R-symmetry

$N=4,6,8,12,24$

Discrete gauge symmetry
-anomaly free
Ibanez, GGR

N	q_{10}	$q_{\bar{5}}$	q_{H_u}	q_{H_d}	q_S
4	1	1	0	0	2
8	1	5	0	4	6

$SU(5), SO(10)$
compatible

 R-symmetry ensures singlets light

SUSY extensions of the Standard Model

$$\begin{aligned}
 W = & h^E LH_d \bar{E} + h^D QH_d \bar{D} + h^U QH_u \bar{U} + \mu H_d H_u + \mu_s S^2 \\
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 \end{aligned}$$

R-parity: Z_2

Z_N^R R-symmetry $N=4,6,8,12,24$

SUSY breaking:

Domain walls safe

$\langle W \rangle, \langle \lambda \lambda \rangle$ $R=2$, non-perturbative breaking

$Z_{4R} \rightarrow Z_2^R$ R -parity

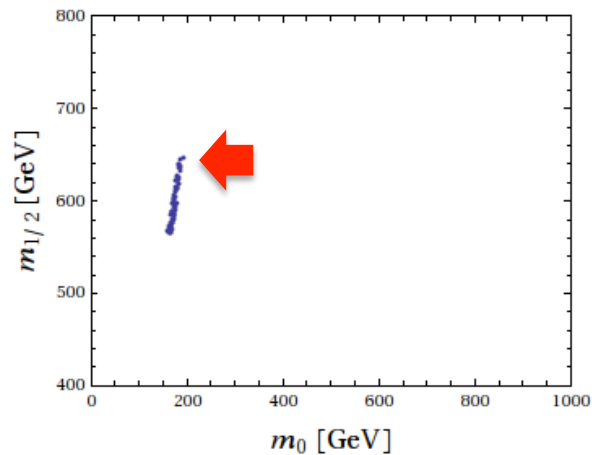
LSP stable

$$\mu, \mu_s \sim m_{3/2}, \quad O\left(\frac{m_{3/2}}{M^2} QQQL\right)$$

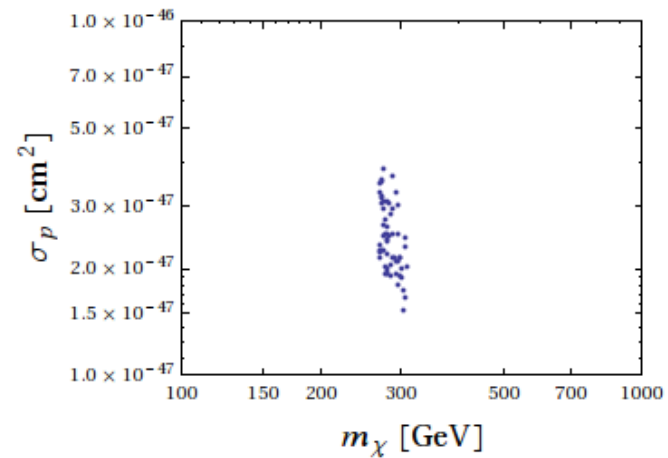
Fine tuning in the CGNMSSM $(\lambda \leq 0.7)$

$$\Delta_{Min} = 60 (500), \quad m_h = 125.6 \pm 3 \text{ GeV}$$

LHC8 SUSY bounds ✗
DM relic abundance ✓
DM searches ✓



Stau co-annihilation



DM searches insensitive

Fine tuning in the (C)GNMSSM ($\lambda \leq 0.7^\dagger$)

Non-universal gaugino masses

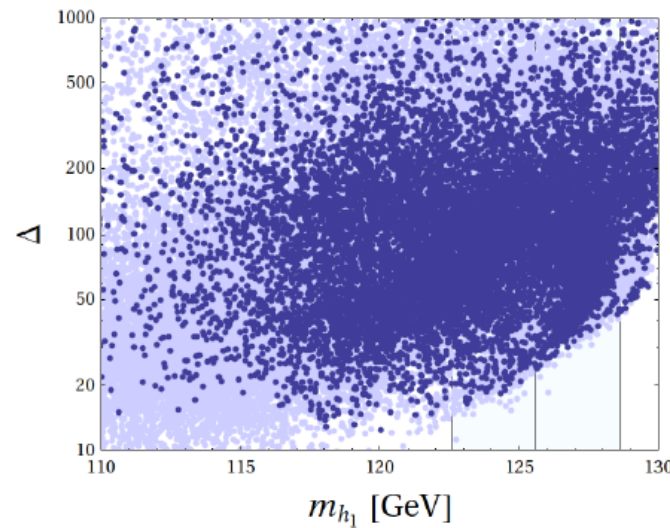
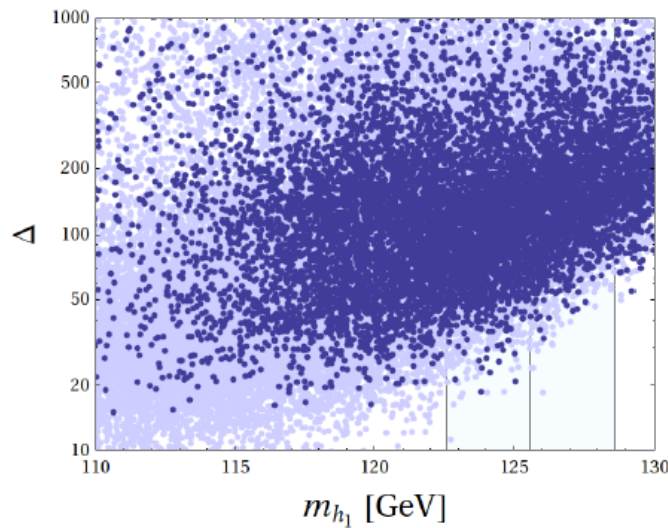
$$\Delta_{Min} = 20, \quad m_h = 125.6 \pm 3 \text{ GeV}$$

LHC8 SUSY bounds ✓

DM relic abundance ✓

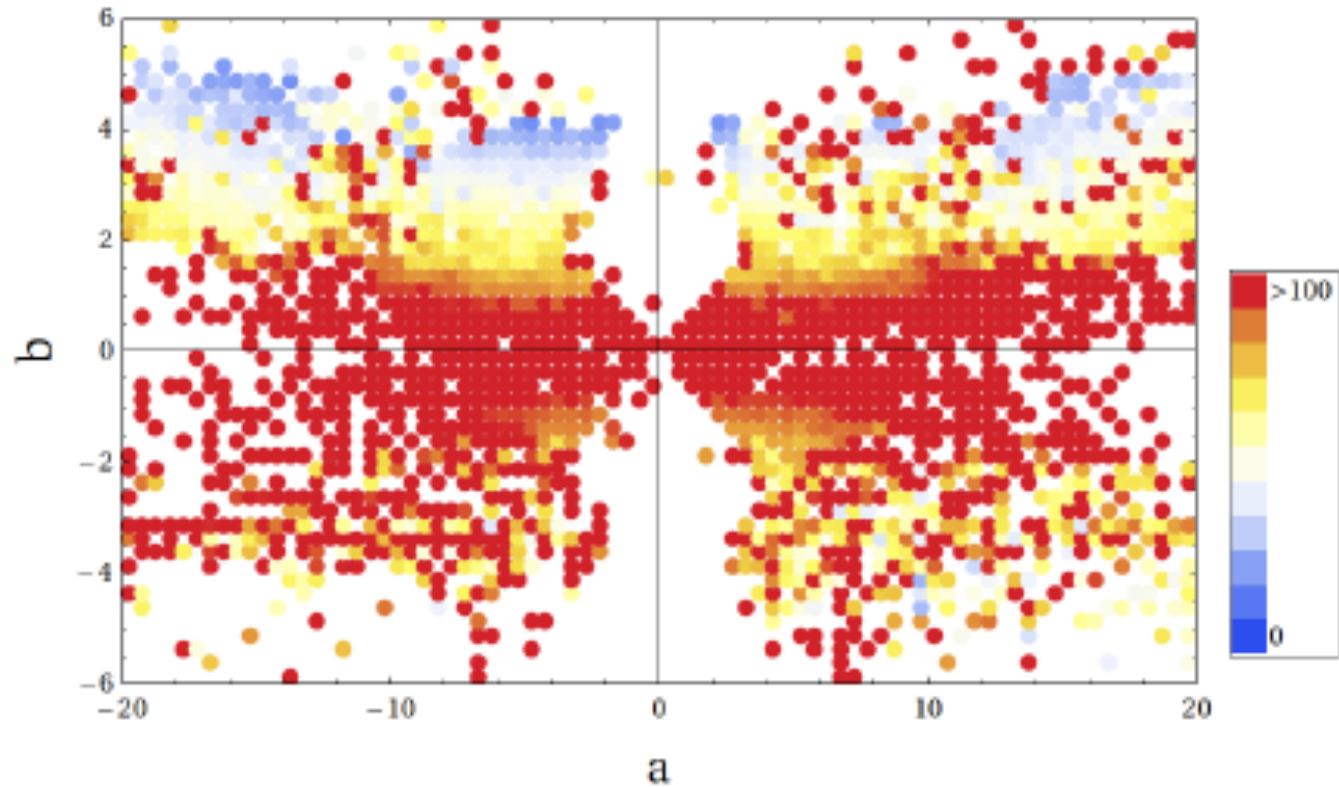
DM searches ✓

Δ



(uniform scan)

Fine tuning v/s gaugino mass ratios



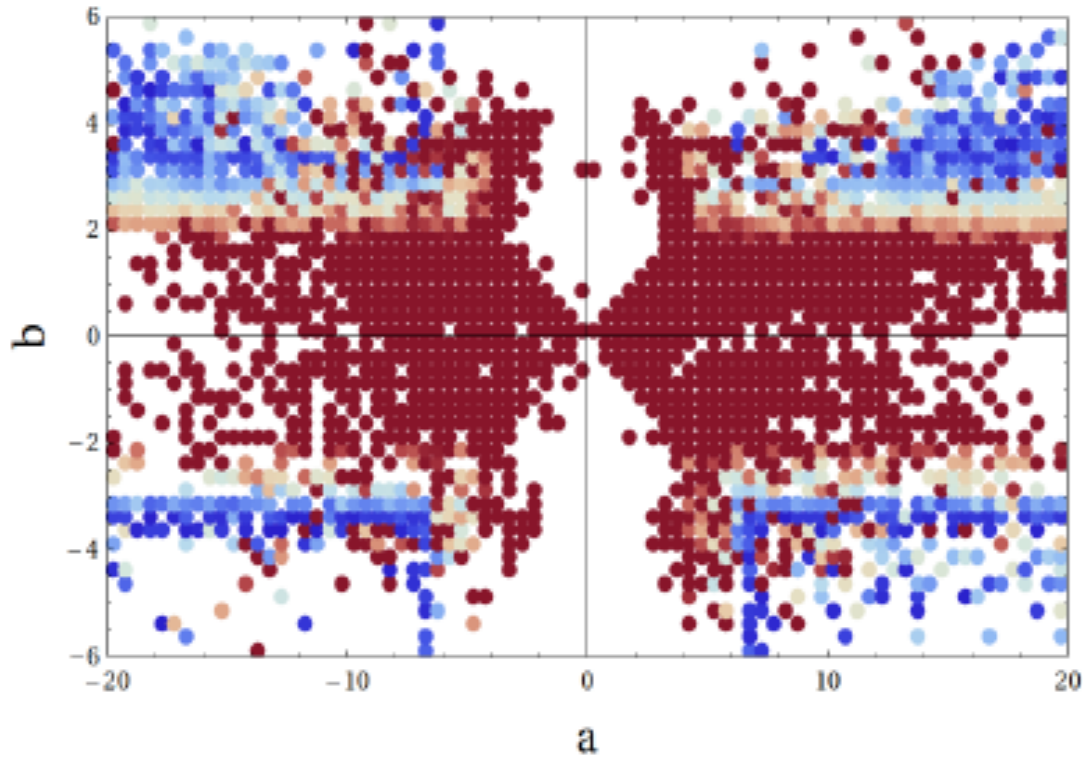
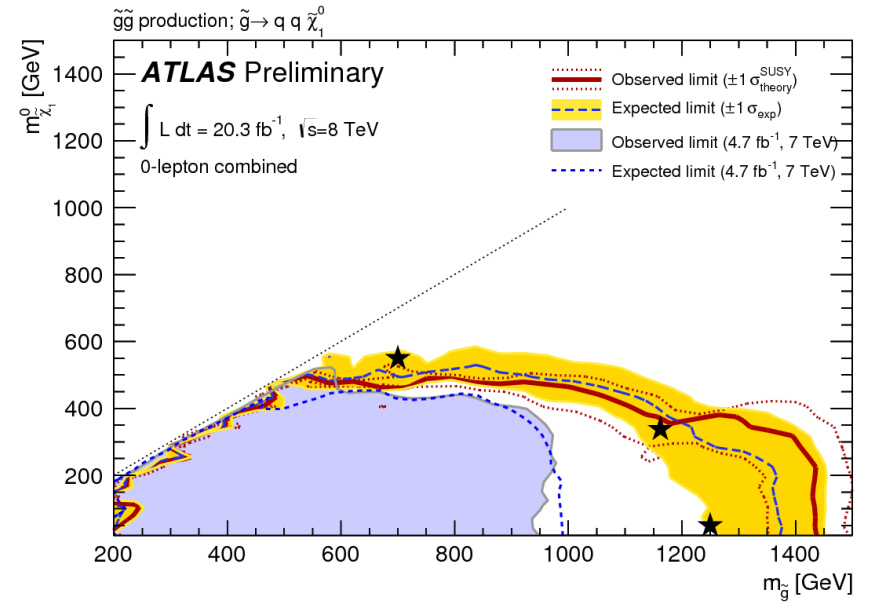
F.T.

> 100

0

$$M_3 = m_{1/2}, M_2 = b.m_{1/2}, M_1 = a.m_{1/2}$$

Compressed spectrum



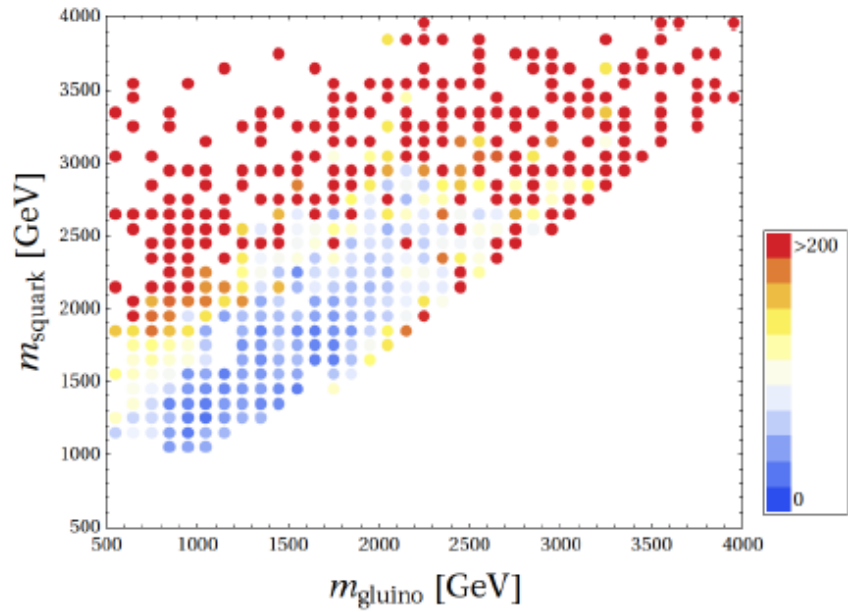
$$\frac{(M_{\tilde{g}} - M_{LSP}^{\text{neutralino}})}{\text{GeV}}$$

> 500

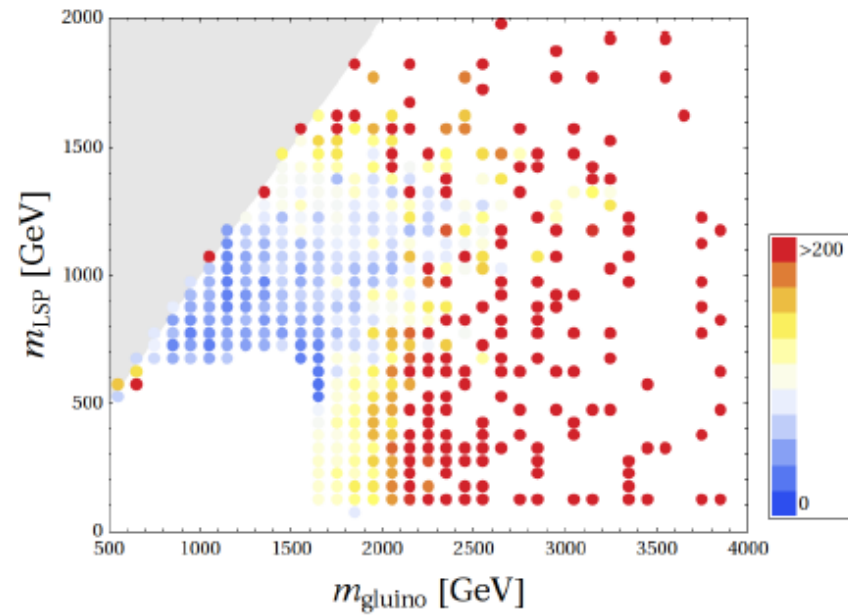
0

Masses v/s fine tuning

m_{squark}



m_{LSP}

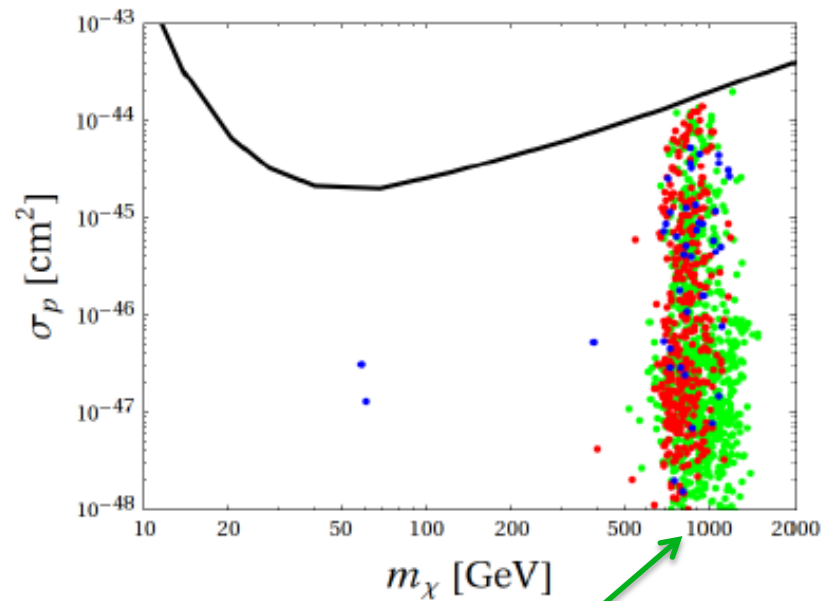


> 500

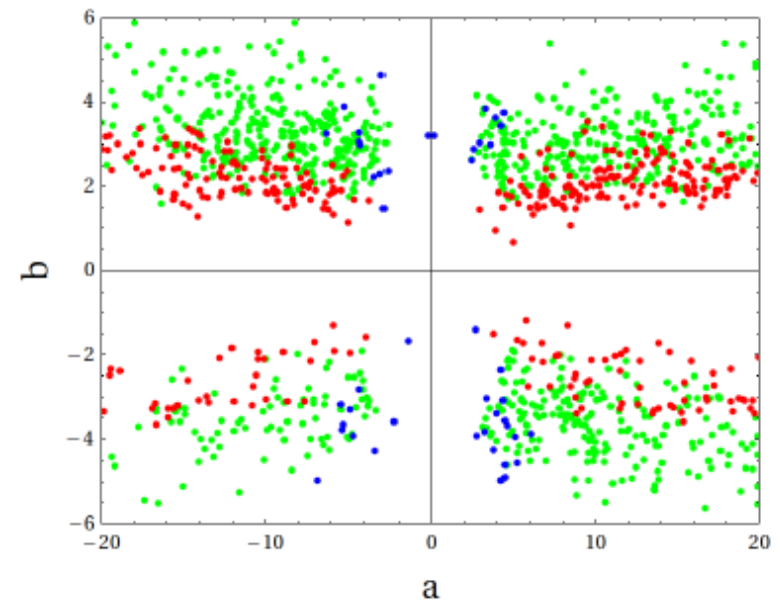
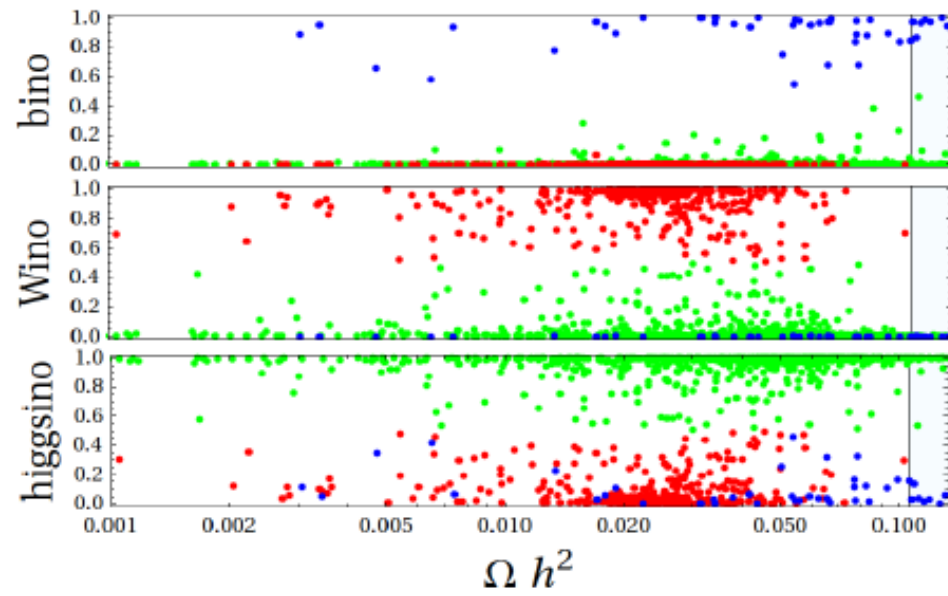
0

M_{gluino}

Dark matter



c.f. Roszkowski, Ruiz de Austria,
Trotta, Tsai, Varley



Summary

● GUTs \Rightarrow SUSY-GUTS (hierarchy problem)

● Fine tuning sensitive to SUSY spectrum

...scalar and gaugino focus points

● $\Delta^{CMSSM} > 350$ ^x $\Delta^{(C)MSSM} > 60$ ^x
 $\Delta^{CGMSSM} > 60$ ^x $\Delta^{(C)GNMMS} > 20$ [✓]

c.f. $\Delta_{Low\ scale}^{CMSSM} = (10 - 30), \quad m_{\tilde{t}} = (1 - 5)TeV$

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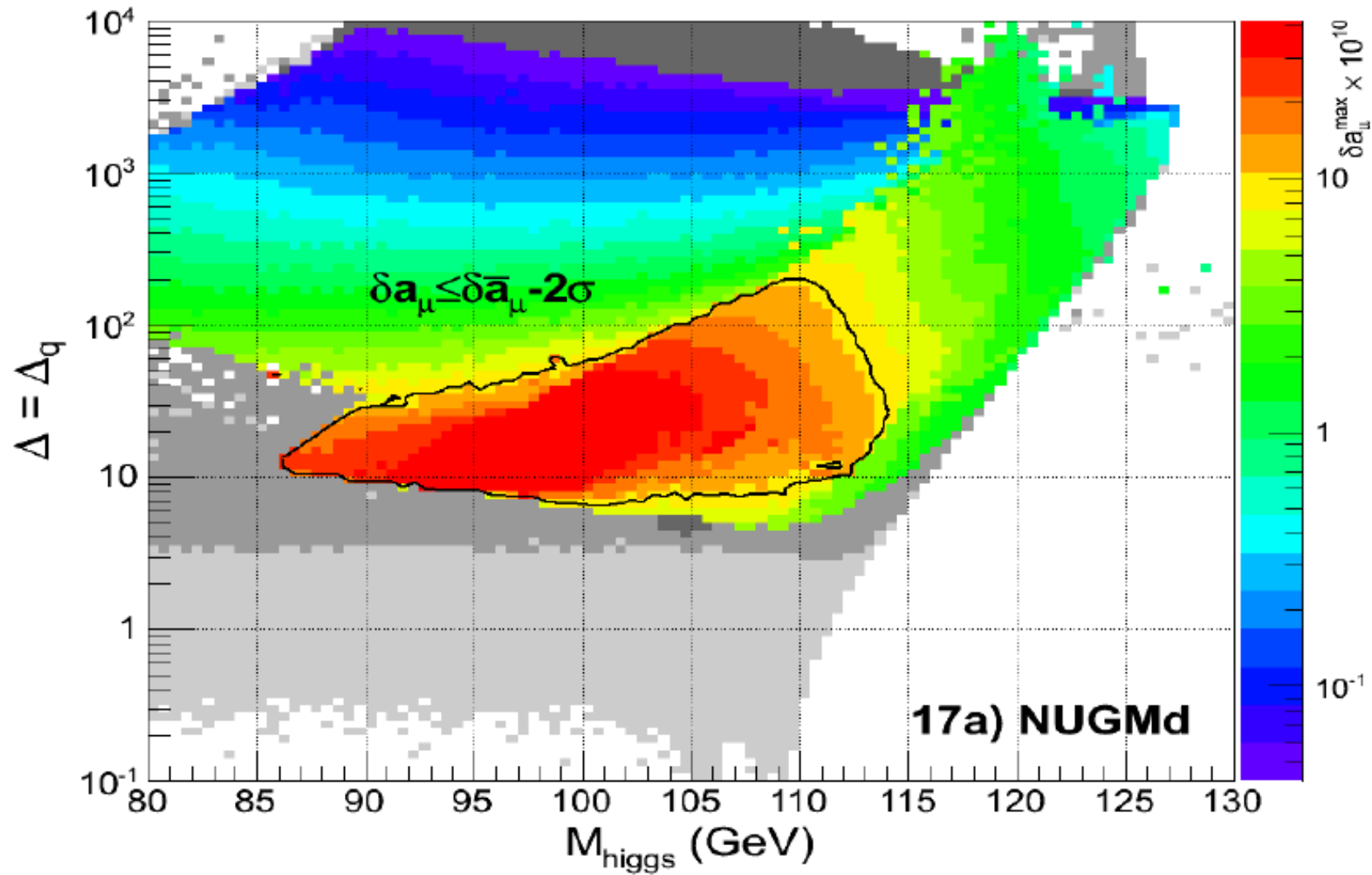
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c.f. $\Delta_{Low\ scale}^{CMSSM} = (10 - 30), \quad m_{\tilde{t}} = (1 - 5)TeV$

- Well motivated SUSY models remain to be tested

LHC14?

Compressed spectra, TeV squarks and gluinos

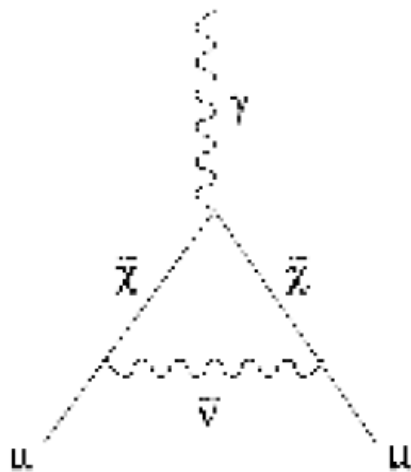


2-loop fine tuning in 75 case

Muon g-2

a_μ is a plausible location for a new physics signal!!

eg could be light SUSY (now tension with LHC)



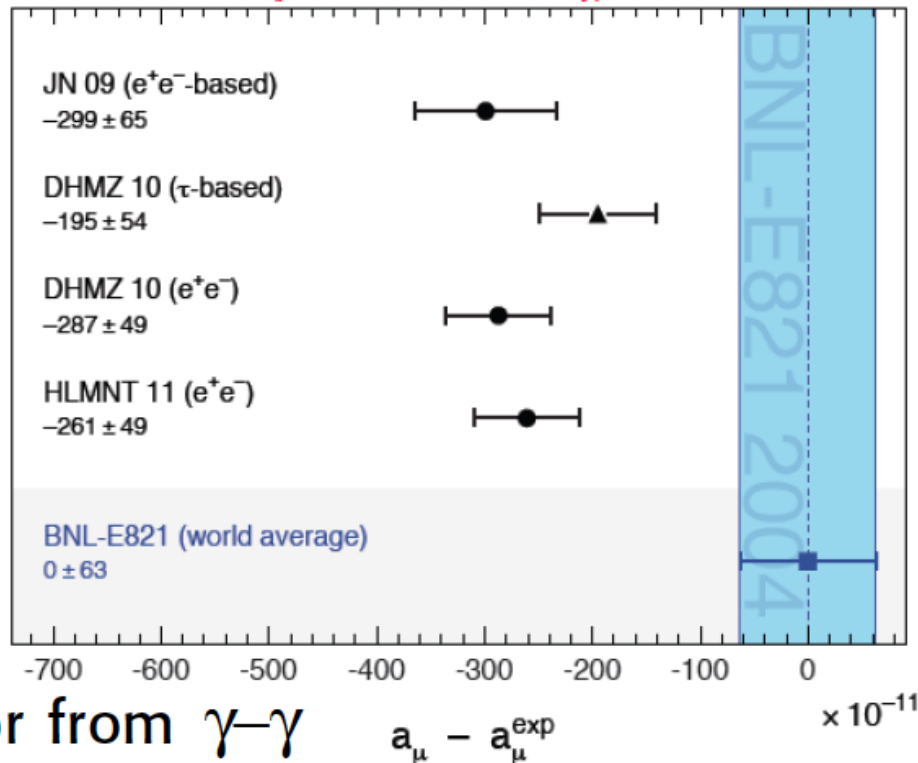
$$a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = (28.7 \pm 8.0) \times 10^{-10}$$

→ 3.6 "standard deviations" (e^+e^-)

→ 2.4 "standard deviations" (τ)

$$\delta a_\mu = 13 \cdot 10^{-10} \left(\frac{100 \text{ GeV}}{M_{\text{SUSY}}} \right)^2 \text{tg}\beta$$

Status: summer 2011 (published results shown only)

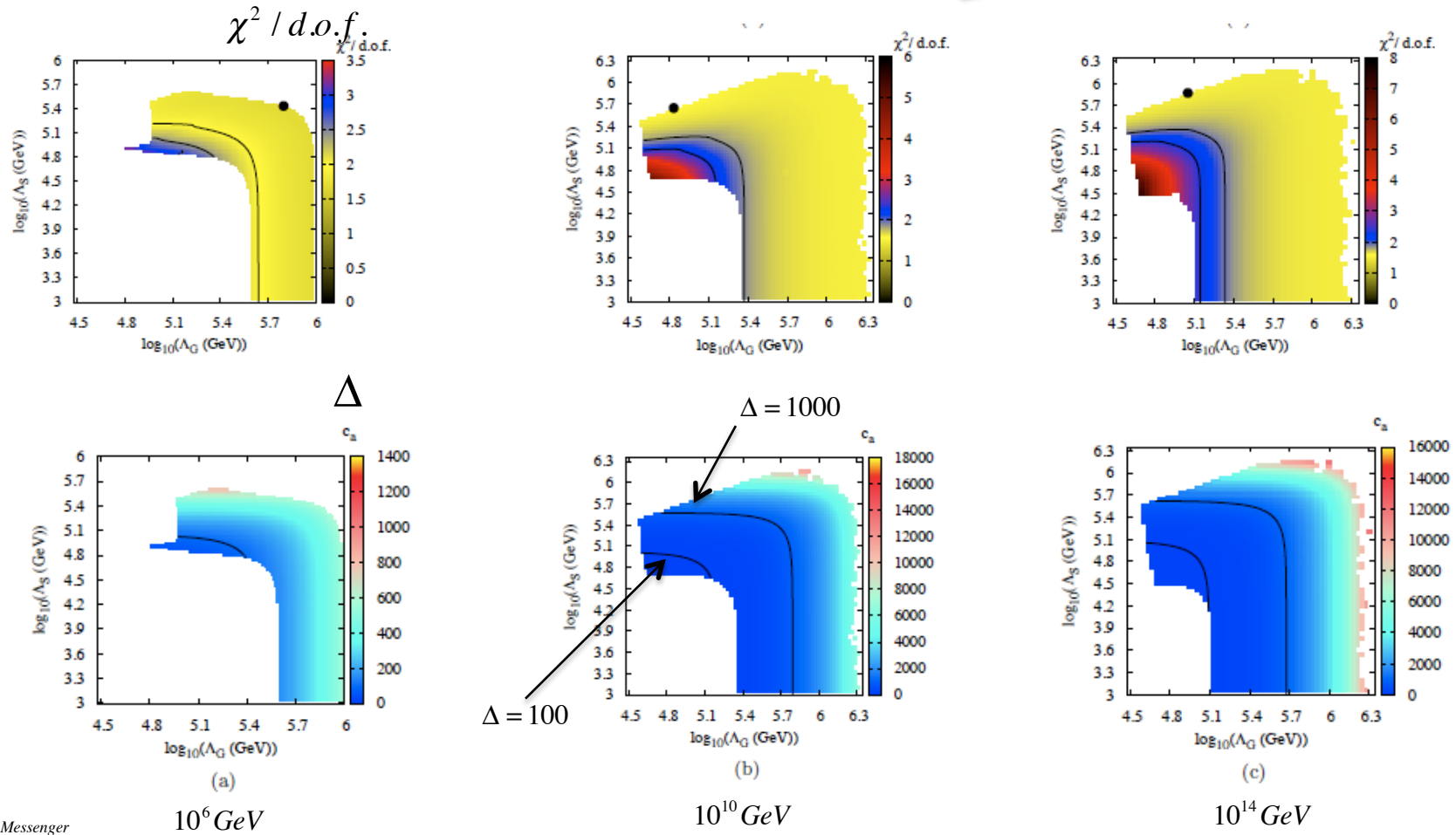


Error dominated by th error from γ - γ

$a_\mu - a_\mu^{\text{exp}} \times 10^{-11}$

Fine tuning in General Gauge Mediation

$B \rightarrow X_s \gamma, B \rightarrow \tau \mu, B \rightarrow \mu^+ \mu^-, B \rightarrow D \tau \mu,$
 $D_s \rightarrow \mu \nu, D_s \rightarrow \tau \nu, K \rightarrow \mu \nu / \pi \rightarrow \mu \nu, \Delta_{0-}$



$\Delta > 100$ no focus point

Abel, Dolan, Jaeckel, Khoze
 (Giusti, Romanino, Strumia)