

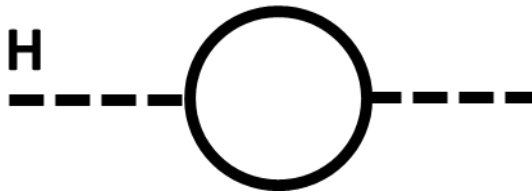
# Natural Supersymmetry in Warped Space

**Yuichiro Nakai (Harvard U.)**

**B. Heidenreich and YN, arXiv:1407.5095[hep-ph]**

# Questions in Higgs Mechanism

## Naturalness



Higgs potential  
is ad hoc.

Quantum correction destabilizes the EW scale.

## Dynamics of EWSB



Why ?

## Yukawa hierarchies

$$m_u \sim 3 \text{ MeV} \ll m_t \sim 173 \text{ GeV}$$

Yukawa couplings have to be hierarchical.

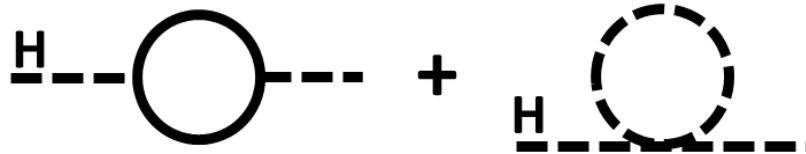
We need  
BSM physics !

# Supersymmetry

SUSY is a nice framework to address several questions !

## Naturalness

Quadratic divergence  
is cancelled.



$\gamma, Z, \dots$  (spin 1)  $\longleftrightarrow$  **gaugino** (spin 1/2)

quark, ...  $\longleftrightarrow$  **squark, ...**

Higgs  $\longleftrightarrow$  **Higgsino**

## Radiative EWSB



SUSY breaking can drive EWSB radiatively.

# Problems of SUSY

## Little hierarchy problem

$$m_{1\text{st}, 2\text{nd}} > 5 \times 10^4 \text{ TeV} \rightarrow \text{Fine-tuning !}$$

M. Bona, the UK Flavour Workshop  
(2013)

## Light Higgs

125 GeV Higgs is heavy for SUSY.

## Yukawa hierarchies

SUSY itself does not address  
Yukawa hierarchies.

## Stop mass bound

$$m_{\tilde{t}} > 650 \text{ GeV}$$

Tuning

To solve these questions, ...

# SUSY + Compositeness

is good.

Composite Higgs

AdS/CFT



# Warped Natural SUSY

Gherghetta, Pomarol (2003) , Sundrum (2009) , Larsen, Nomura, Roberts (2012) ...

**Combine SUSY and Randall-Sundrum !**

**Quadratic divergence is cutoff  
at the IR scale ( $\sim 10$  TeV ).**

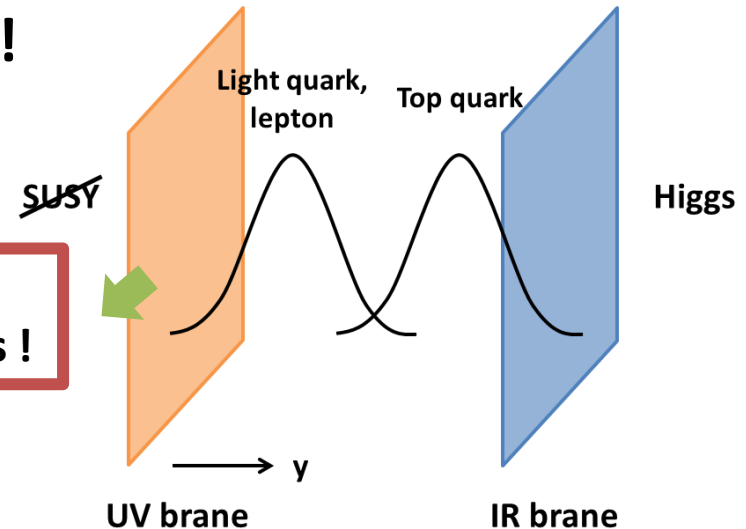


*For naturalness, we only need ...*

**Light stops, gauginos and Higgsinos.**

**The other squarks and sleptons are heavy.**

**Yukawa  
hierarchies !**



**Little hierarchy problem  
is solved !**

**This pattern of SUSY breaking is naturally realized.**

## Accidental SUSY

Even with large ~~SUSY~~ on UV brane, IR brane SUSY is maintained.

### SUSY Yang-Mills coupled to SCFT

Sundrum (2009)

$$\mathcal{L}(\mu) = -\frac{1}{4}F_{\mu\nu}^a{}^2 + \bar{\lambda}D.\sigma\lambda + \mathcal{L}_{SCFT} + \underline{g}A_\mu^a J_a^\mu + \underline{\tilde{g}}\lambda_a \Psi_J^a - \frac{1}{2}\underline{g_D^2}D^a D^a$$

$$\frac{d1/g^2}{d\ln\mu} = \frac{d1/\tilde{g}^2}{d\ln\mu} = \frac{d1/g_D^2}{d\ln\mu} \equiv -b_{CFT} \sim \frac{\mathcal{O}(N_{CFT})}{16\pi^2} \quad \Rightarrow \quad \text{IR free !}$$

$$\begin{aligned} \frac{\Delta\tilde{g}^2}{g^2} &\equiv \frac{\tilde{g}^2 - g^2}{g^2} \approx \frac{g^2}{g_0^2} \frac{\Delta\tilde{g}_0^2}{g_0^2} \\ \frac{\Delta g_D^2}{g^2} &\equiv \frac{g_D^2 - g^2}{g^2} \approx \frac{g^2}{g_0^2} \frac{\Delta g_{D0}^2}{g_0^2} \end{aligned}$$

**Focusing effect  
in the IR**

# Warped Natural SUSY

Gherghetta, Pomarol (2003) , Sundrum (2009) , Larsen, Nomura, Roberts (2012) ...

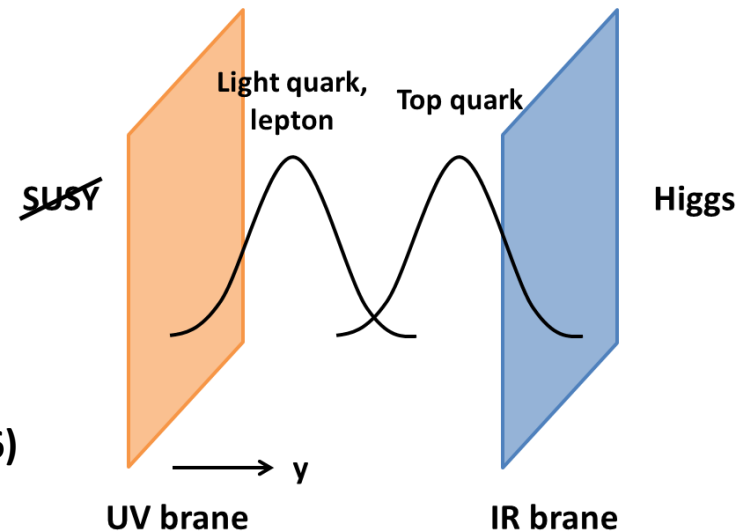
**EW breaking is driven by top/stop loop.**

**125 GeV Higgs is naturally obtained in  $\lambda$  SUSY.**

Barbieri, Hall, Nomura, Rychkov (2006)

$$W = \lambda S H_u H_d \quad \lambda > 0.7$$

( without encountering Landau pole )



**Let's pursue a fully realistic model !**



# SUSY RS Model

$$ds^2 = e^{-2k|y|} \eta_{\mu\nu} dx^\mu dx^\nu + dy^2 \quad (0 \leq |y| \leq \pi R)$$

Randall, Sundrum (1999)

$$S^1/Z_2$$

Extended SUSY in the bulk  $\rightarrow$   *$N = 1$  SUSY on the branes*

**Compactification scale :**  $k' \equiv k e^{-k\pi R} = \mathcal{O}(10) \text{ TeV}$      $kR \sim 10$

## SM gauge fields in the bulk

*Wavefunction profile of the zero mode is flat.*

$$A_\mu(x, y) \simeq \frac{1}{\sqrt{2\pi R}} A_\mu^{(0)}(x)$$



# SUSY RS Model

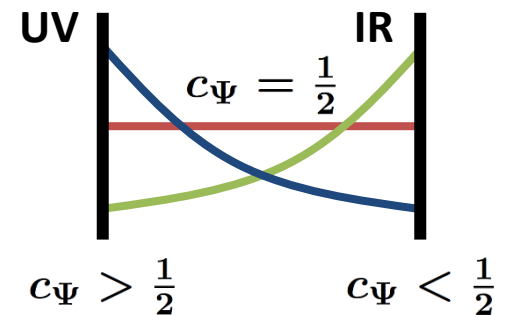
## Matter (hyper-)multiplets in the bulk

$$S_{\Psi} = \int d^5x \left\{ e^{-2k|y|} \int d^4\theta (\Psi^\dagger \Psi + \Psi^c \Psi^{c\dagger}) \right. \\ \left. + e^{-3k|y|} \int d^2\theta \Psi^c \left[ \partial_y - \left( \frac{3}{2} - c_{\Psi} \right) k\epsilon(y) \right] \Psi + \text{h.c.} \right\}$$

*Bulk mass parameter*

## *Wavefunction profile of the zero mode*

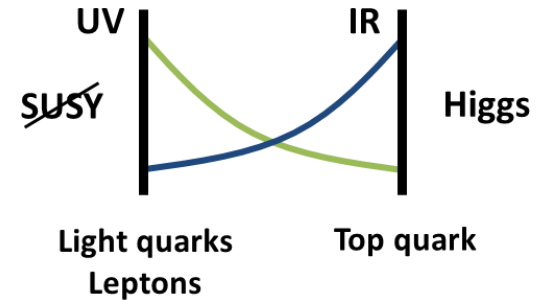
$$\Psi(x, y) \simeq \frac{e^{-(c_{\Psi} - \frac{3}{2})k|y|}}{\sqrt{\frac{1}{(c_{\Psi} - \frac{1}{2})k} \left( 1 - e^{-2\pi k R(c_{\Psi} - \frac{1}{2})} \right)}} \Psi^{(0)}(x)$$



# Yukawa hierarchy

## Yukawa coupling on IR brane

$$S_{\text{Yukawa}} = \int d^5x \delta(y - \pi R) e^{-3\pi k R} \left\{ \int d^2\theta \left( \tilde{y}_u^{ij} H_u Q_i \bar{u}_j + \tilde{y}_d^{ij} H_d Q_i \bar{d}_j + \tilde{y}_\nu^{ij} H_u L_i \bar{\nu}_j + \tilde{y}_e^{ij} H_d L_i \bar{e}_j \right) + \text{h.c.} \right\}$$



➔  $y_u^{ij} = \tilde{y}_u^{ij} k \zeta_{Q_i} \zeta_{\bar{u}_j}, \quad y_d^{ij} = \tilde{y}_d^{ij} k \zeta_{Q_i} \zeta_{\bar{d}_j}, \quad y_\nu^{ij} = \tilde{y}_\nu^{ij} k \zeta_{L_i} \zeta_{\bar{\nu}_j}, \quad y_e^{ij} = \tilde{y}_e^{ij} k \zeta_{L_i} \zeta_{\bar{e}_j}$

$$\zeta_\Psi \simeq \begin{cases} \sqrt{c_\Psi - \frac{1}{2}} e^{-(c_\Psi - \frac{1}{2})\pi k R} & (c_\Psi \gg 1/2) \\ \frac{1}{\sqrt{2\pi k R}} & (c_\Psi \sim 1/2) \\ \sqrt{\frac{1}{2} - c_\Psi} & (c_\Psi \ll 1/2) \end{cases}$$

➔ **Light quarks , Leptons**

➔ **Top quark**

# Proton Decay

*Even if we impose R-parity as usual, ...*

$$W_{\text{IR}} \sim \frac{1}{\Lambda_{\text{IR}}} QQQ L \times (\text{wavefunction factors})$$

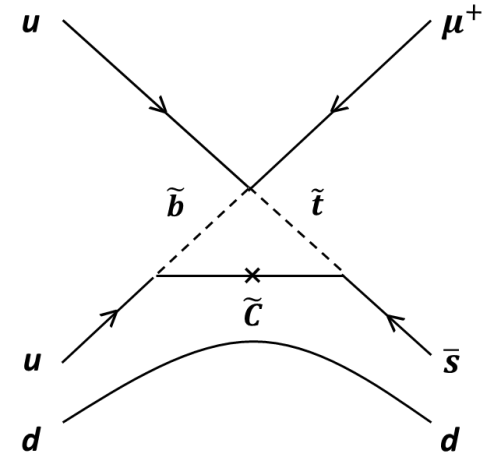
**O(10) TeV !**

**Rapid proton decay ...**

**$Z_3$  lepton number symmetry**

$$L \rightarrow e^{2\pi i/3} L, \quad \bar{\nu} \rightarrow e^{-2\pi i/3} \bar{\nu}, \quad \bar{e} \rightarrow e^{-2\pi i/3} \bar{e}$$

*Anomaly free* ➡ **Three generations !**



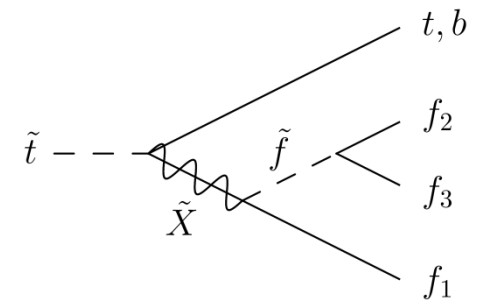
**RPV is natural  
in SUSY RS !**

# R-parity Violation

LSPs can decay promptly and evade searches based on missing transverse energy !

BNV couplings

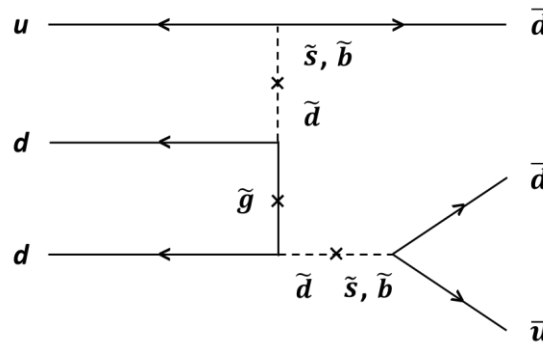
$$W_{\text{BNV}} = \frac{1}{2} \lambda''_{ijk} \bar{u}_i \bar{d}_j \bar{d}_k$$



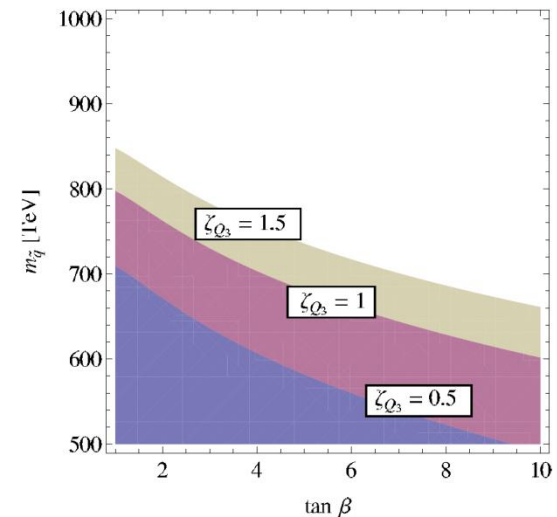
Stop and sbottom

$$m_{\tilde{b}, \tilde{t}} \gtrsim 100 \text{ GeV}$$

Constraints from  $\Delta B = 2$  processes are satisfied !



$n - \bar{n}$  oscillations



## U(1) D-term Problem

Strassler (2003)

Sundrum (2009)

**Heavy scalars**  $\rightarrow \mathcal{L}_{\text{FI}} \sim \int d^4\theta \frac{m_{\tilde{q}, \tilde{l}}^2}{16\pi^2} g_Y V_Y$  near UV brane

$\rightarrow$  **Large Higgs soft masses !**

The gauge field couples marginally to the CFT current  $J_Y^\mu$   
(  $\Delta = 3$  )

$\rightarrow$  Scaling dimension of  $D_Y$  :  $\Delta = 2$

$\rightarrow$  SCFT admits a relevant deformation :  $\Delta\mathcal{L} = M_D^2 D_Y$

**Conformal phase breaks down at  $M_D$**

## TeV Unification

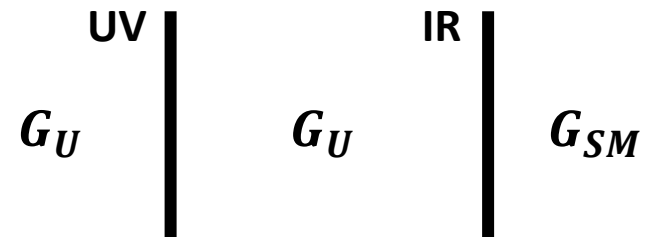
Extend the SM gauge group to forbid the relevant deformation

- Semi-simple group (  $SU(5)$  , ... )
- Left-right symmetry under which the  $U(1)$  D-term transforms nontrivially

$$SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$

The unified group is broken  
on IR brane by boundary conditions

cf. The Higgsless model of EWSB



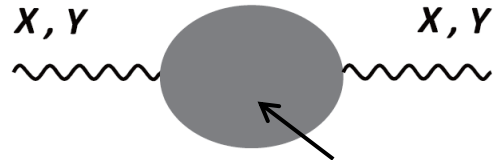
Extra gauge fields → Dirichlet boundary condition on IR brane

## ND Gauge Field

Extra gauge fields with ND boundary conditions satisfy

$$\frac{J_0(m/k)}{Y_0(m/k)} = \frac{J_1(m/k')}{Y_1(m/k')} \quad \longrightarrow \quad m_0 \simeq \sqrt{\frac{2}{\pi k R}} k'$$

$$\mathcal{L}_{SU(5)} \sim -\frac{1}{4} \left\{ \frac{1}{g_{UV}^2} + \frac{N_{\text{CFT}}}{16\pi^2} \log \left( \frac{M_{\text{pl}}}{\Lambda_{\text{IR}}} \right) \right\} \sum_{a \text{ (All)}} (F_{\mu\nu}^a)^2$$

$$+ \frac{1}{2} \frac{N_{\text{CFT}}}{16\pi^2} \Lambda_{\text{IR}}^2 \sum_{\alpha \text{ (Broken)}} (A_\mu^\alpha)^2$$


CFT particles

$$\longrightarrow \quad m_0^2 \sim \frac{\Lambda_{\text{IR}}^2}{\log(M_{\text{pl}}/\Lambda_{\text{IR}})} \quad \log(M_{\text{pl}}/\Lambda_{\text{IR}}) \simeq \pi k R$$

$k' = \mathcal{O}(10) \text{ TeV} \quad \longrightarrow \quad \text{ND gauge fields may be discovered at LHC !}$



## The SU(5) Model

Coupling unification can be realized  
by IR brane-localized kinetic term

$$\begin{array}{ccc} & \text{UV} & \text{IR} \\ \text{SU(5)} & \left| & \text{SU(5)} \right| G_{SM} \end{array}$$

$$S_{\text{IR}} = \int d^5x \delta(y - \pi R) \left\{ \frac{1}{4\tilde{g}_a^2} \int d^2\theta \text{Tr} W^{a\alpha} W^a_{\alpha} + \text{h.c.} \right\} \quad a = 3, 2, 1$$

Split multiplets for quarks and leptons :  $10_Q, 10_{\bar{u}, \bar{e}}$  &  $\bar{5}_{\bar{d}}, \bar{5}_L$

$Z_3$  lepton number symmetry  $\omega_3 \equiv e^{2\pi i/3}$

$$10_Q \rightarrow \omega_3 10_Q, \quad 10_{\bar{u}, \bar{e}} \rightarrow \omega_3^{-1} 10_{\bar{u}, \bar{e}}, \quad \bar{5}_L \rightarrow \omega_3 \bar{5}_L, \quad \bar{5}_{\bar{d}} \rightarrow \omega_3^{-1} \bar{5}_{\bar{d}}$$

Extra fields in split multiplets :  $Q', \bar{u}', \bar{d}', L', \bar{e}'$

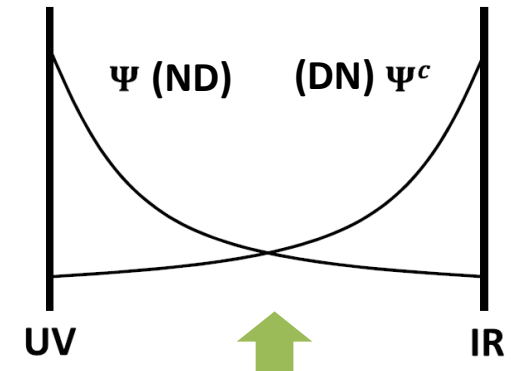
# Light Exotics

Obtain sizable masses of exotics by ND boundary conditions ?

Exotics satisfy  $\frac{J_{c-1/2}(m/k)}{Y_{c-1/2}(m/k)} = \frac{J_{c+1/2}(m/k')}{Y_{c+1/2}(m/k')}$   $\rightarrow$  No ...

For  $c \gg 1/2$

$$m \simeq 2\sqrt{c + \frac{1}{2}\zeta} k'$$



Exponentially small overlap !!

Light exotics always appear ...

$$M_{Q'_1} \sim 2\zeta_{\bar{u}_1} k' \ll M_Z$$

$\rightarrow$  The SU(5) model is excluded ...

## Split Couplings without Exotics

A way to avoid light exotics in split multiplets

$$\left[ \begin{array}{l} \Psi_A = (A, B') \\ \Psi_B = (A', B) \end{array} \right.$$

Introduce a new multiplet on UV brane :

$$\bar{\Psi}_{UV} = (\bar{A}_{UV}, \bar{B}_{UV}) \quad \text{A mass term : } M_{UV} \bar{\Psi}_{UV} (s_\theta \Psi_A - c_\theta \Psi_B)$$

➡ A light multiplet :  $\hat{\Psi} = (\hat{A}, \hat{B}) = c_\theta \Psi_A + s_\theta \Psi_B$

Yukawa couplings on IR brane :  $\mathcal{L}_{IR} = A\mathcal{O}_A + B\mathcal{O}_B + \dots$

➡  $\mathcal{L}_{\text{eff}} = y_A \hat{A} \mathcal{O}_A + y_B \hat{B} \mathcal{O}_B + \dots$   $\left[ \begin{array}{l} y_A \sim c_\theta \zeta_A \\ y_B \sim s_\theta \zeta_B \end{array} \right.$


## The Left-Right Model

The symmetry is broken on IR brane :  $SU(2)_R \times U(1)_{B-L} \rightarrow U(1)_Y$

$W', Z'$  may be discovered at LHC via  $W' \rightarrow \ell \nu$  ,  $Z' \rightarrow \ell^+ \ell^-$

Split multiplets :

$$\mathcal{Q}_Q = (Q, U'') \quad \mathcal{Q}_{\bar{u}} = (Q', U_{\bar{u}}) \quad \mathcal{Q}_{\bar{d}} = (Q'', U_{\bar{d}})$$


 $(\bar{u}, \bar{d}')$

Introduce new multiplets on UV brane to avoid light exotics :  $\bar{\mathcal{Q}}_{1,2}$

Yukawa couplings on IR brane :  $W_{\text{Yukawa}} = Q \bar{u} H_u + Q \bar{d} H_d$

## Summary

To pursue a fully realistic SUSY RS model ...

Proton decay problem



Lepton number symmetry



RPV

Light stop

U(1) D-term problem



TeV unification !

ND gauge fields may be  
discovered at LHC !

Viable pattern of RPV is naturally derived !

Thank you.

# Backup

# Yukawa hierarchy

## Wavefunction factors

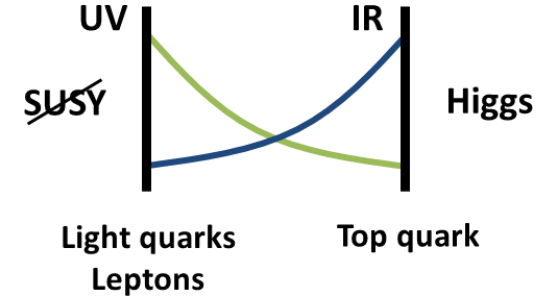
$$m_{u_i} \simeq \zeta_{Q_i} \zeta_{\bar{u}_i} v \sin \beta, \quad m_{d_i} \simeq \zeta_{Q_i} \zeta_{\bar{d}_i} v \cos \beta$$

$$|(V_{\text{CKM}})_{ij}| \simeq \frac{\zeta_{Q_j}}{\zeta_{Q_i}} \quad \text{for } j \leq i$$

$$|(V_{\text{CKM}})_{21}| \simeq \lambda, \quad |(V_{\text{CKM}})_{32}| \simeq \lambda^2, \quad |(V_{\text{CKM}})_{31}| \simeq \lambda^3 \quad \lambda \sim 0.2$$

$$\Rightarrow \left[ \begin{array}{lll} \zeta_{Q_1} \simeq \lambda^3 \zeta_{Q_3}, & \zeta_{Q_2} \simeq \lambda^2 \zeta_{Q_3}, & \\ \zeta_{\bar{u}_1} \simeq \frac{m_u}{\lambda^3 \zeta_{Q_3} v \sin \beta}, & \zeta_{\bar{u}_2} \simeq \frac{m_c}{\lambda^2 \zeta_{Q_3} v \sin \beta}, & \zeta_{\bar{u}_3} \simeq \frac{m_t}{\zeta_{Q_3} v \sin \beta}, \\ \zeta_{\bar{d}_1} \simeq \frac{m_d}{\lambda^3 \zeta_{Q_3} v \cos \beta}, & \zeta_{\bar{d}_2} \simeq \frac{m_s}{\lambda^2 \zeta_{Q_3} v \cos \beta}, & \zeta_{\bar{d}_3} \simeq \frac{m_b}{\zeta_{Q_3} v \cos \beta} \end{array} \right.$$

# R-parity violation



## RPV coupling on IR brane

$$S_{\text{RPV, IR}} = \int d^5x \delta(y - \pi R) e^{-3k\pi R} \left( \int d^2\theta \frac{1}{2} \tilde{\lambda}^{ijk} \bar{u}_i \bar{d}_j \bar{d}_k + \text{h.c.} \right)$$

$$\rightarrow W_{\text{RPV, IR}}^{4D} = \frac{1}{2} \lambda^{ijk} \bar{u}_i \bar{d}_j \bar{d}_k \quad \lambda^{ijk} = \tilde{\lambda}^{ijk} k^{3/2} \zeta_{\bar{u}_i} \zeta_{\bar{d}_j} \zeta_{\bar{d}_k}$$

**Coupling size is proportional to wavefunction factors !**

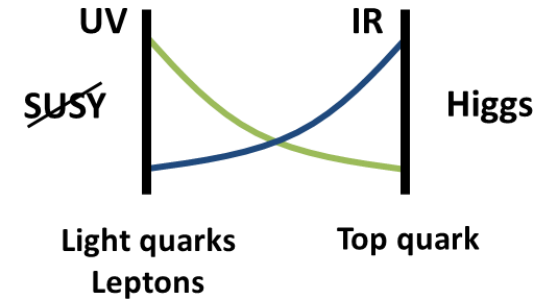
	$sb$	$bd$	$ds$
$u$	$8 \times 10^{-6}$	$2 \times 10^{-6}$	$1 \times 10^{-6}$
$c$	$7 \times 10^{-4}$	$2 \times 10^{-4}$	$1 \times 10^{-4}$
$t$	$3 \times 10^{-3}$	$1 \times 10^{-3}$	$6 \times 10^{-4}$

$$\tan \beta = 3 \text{ and } \zeta_{Q_3} = 1$$



# R-parity violation

## RPV coupling on UV brane



$$S_{\text{RPV, UV}} = \int d^5x \delta(y) \left( \int d^2\theta \frac{1}{2} \tilde{\lambda}'^{ijk} \bar{u}_i \bar{d}_j \bar{d}_k + \text{h.c.} \right)$$

$$\Rightarrow W_{\text{RPV, UV}}^{4D} = \frac{1}{2} \lambda'^{ijk} \bar{u}_i \bar{d}_j \bar{d}_k \quad \lambda'^{ijk} = \tilde{\lambda}'^{ijk} k^{3/2} \eta_{\bar{u}_i} \eta_{\bar{d}_j} \eta_{\bar{d}_k}$$

$$\eta_\Psi \simeq \begin{cases} \sqrt{c_\Psi - \frac{1}{2}} \simeq \sqrt{\frac{1}{\pi k R} \log \zeta_\Psi^{-1}} & (c_\Psi \gg 1/2) \\ \frac{1}{\sqrt{2\pi k R}} \simeq \zeta_\Psi & (c_\Psi \sim 1/2) \\ \sqrt{\frac{1}{2} - c_\Psi} e^{-(\frac{1}{2} - c_\Psi)\pi k R} \simeq \zeta_\Psi e^{-\pi k R \zeta_\Psi^2} & (c_\Psi \ll 1/2) \end{cases}$$

	$sb$	$bd$	$ds$
$u$	0.04	0.05	0.05
$c$	0.02	0.03	0.03
$t$	$5 \times 10^{-16}$	$6 \times 10^{-16}$	$6 \times 10^{-16}$

$$\tan \beta = 3 \text{ and } \zeta_{Q_3} = 1$$

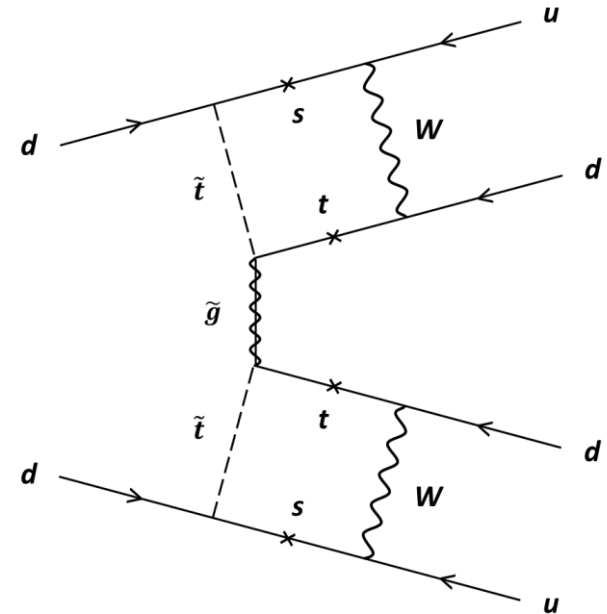
# R-parity violation

## $n - \bar{n}$ oscillations (RPV on IR brane)

Constraint :  $\tau_{n-\bar{n}} \geq 2.44 \times 10^8 \text{ s}$

*If the scalars of light quarks  
are very heavy, ...*

**The leading diagram must involve  
only light superpartners.**



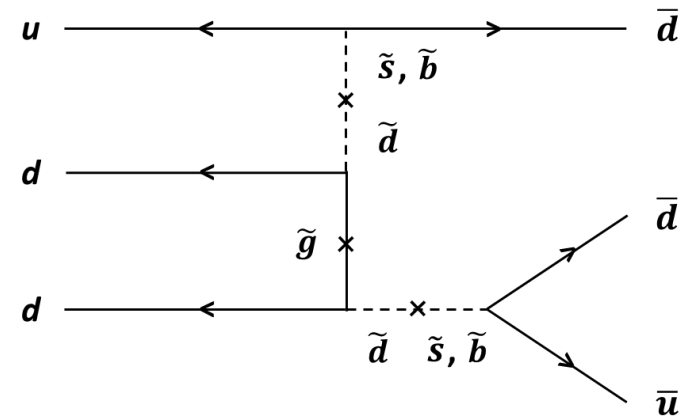
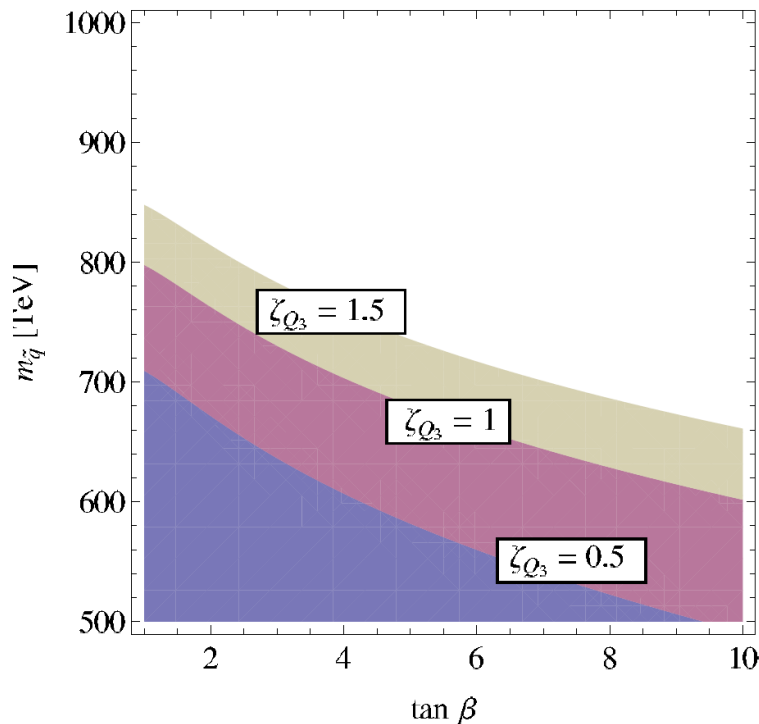
$$\tau_{n-\bar{n}} \sim (3 \times 10^{10} \text{ s}) \left( \frac{\lambda_{tds}}{6 \times 10^{-4}} \right)^{-2} \left( \frac{m_{\tilde{g}}}{1.2 \text{ TeV}} \right) \left( \frac{m_{\tilde{t}}}{300 \text{ GeV}} \right)^4$$

**The bound is easily satisfied !**

# R-parity violation

## $n - \bar{n}$ oscillations (RPV on UV brane)

Sizable coupling for light quarks.



$$\mathcal{M}_{n-\bar{n}} \sim 4\pi\alpha_3 \lambda_{uds}'^2 \tilde{\Lambda} \left( \frac{\tilde{\Lambda}}{m_{\tilde{g}}} \right) \left( \frac{\tilde{\Lambda}^2}{m_{\tilde{q}}^2} \right)^2$$

$$\tilde{\Lambda} \sim \Lambda_{\text{QCD}} \sim 250 \text{ MeV}$$



**The bound is weaker than the FCNC bound.**

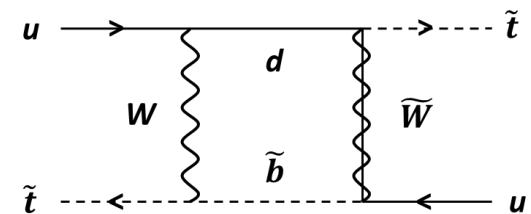
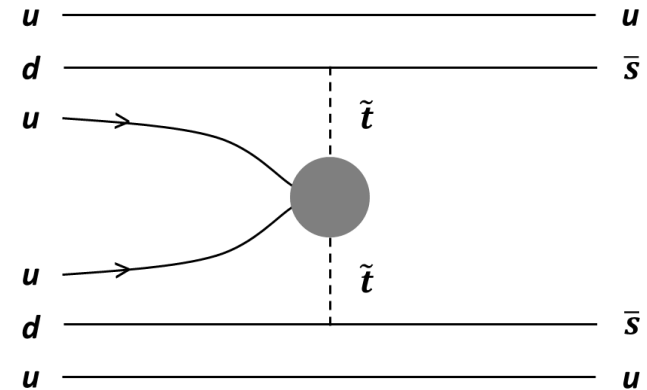
# R-parity violation

## Dinucleon decay (RPV on IR brane)

Constraint :  $\tau_{pp \rightarrow K^+ K^+} \geq 1.7 \times 10^{32} \text{ yrs}$

*If the scalars of light quarks  
are very heavy , ...*

**The leading diagram must involve  
only light superpartners.**



$$\tau_{pp \rightarrow K^+ K^+} \sim (4 \times 10^{39} \text{ yrs}) \left( \frac{\lambda_{tds}}{6 \times 10^{-4}} \right)^{-4} \left( \frac{m_{\tilde{W}}}{600 \text{ GeV}} \right)^2 \left( \frac{m_{\tilde{t}, \tilde{b}}}{300 \text{ GeV}} \right)^{12}$$

# R-parity violation

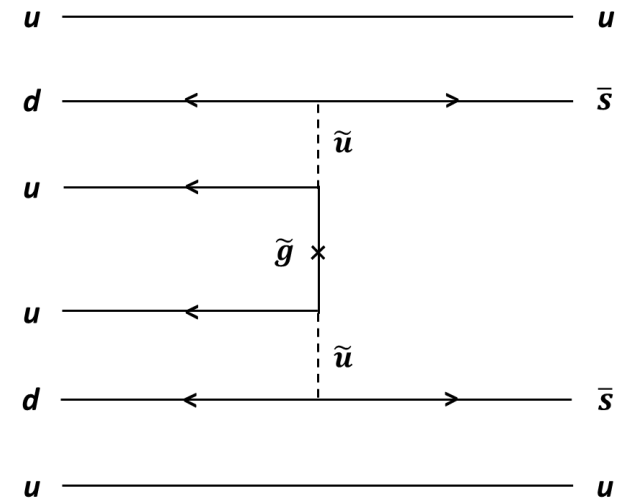
## Dinucleon decay (RPV on UV brane)

Scalars of light quarks are very heavy,  
but sizable coupling for light quarks.

For  $\tan \beta = 3$  and  $\zeta_{Q_3} = 1$

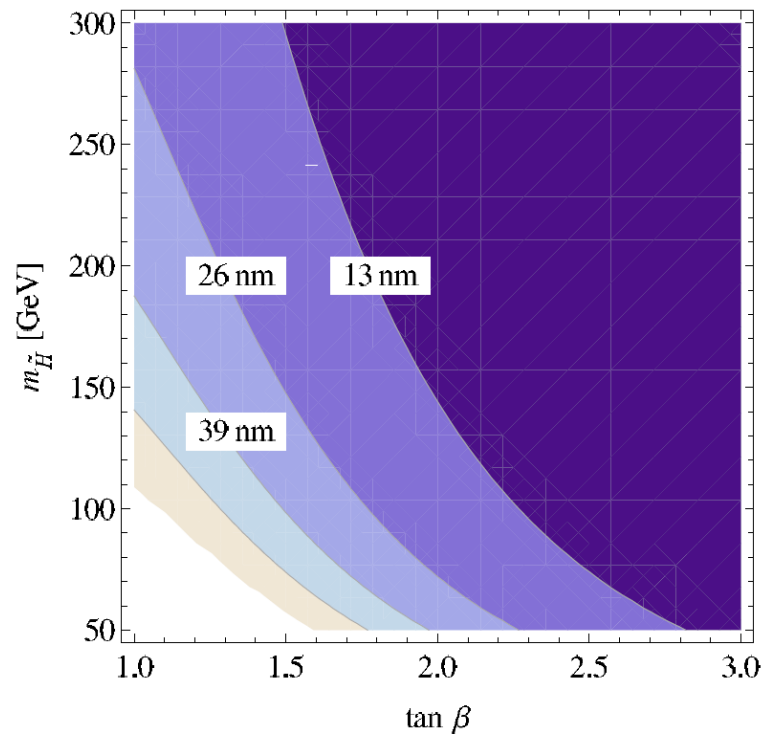
$$\tau_{pp \rightarrow K^+ K^+} \sim (5 \times 10^{35} \text{ yrs}) \left( \frac{\lambda'_{uds}}{0.05} \right)^{-4} \left( \frac{m_{\tilde{g}}}{1.2 \text{ TeV}} \right)^2 \left( \frac{m_{\tilde{q}}}{1000 \text{ TeV}} \right)^8$$

**Constraints from  $\Delta B = 2$  processes are satisfied !**

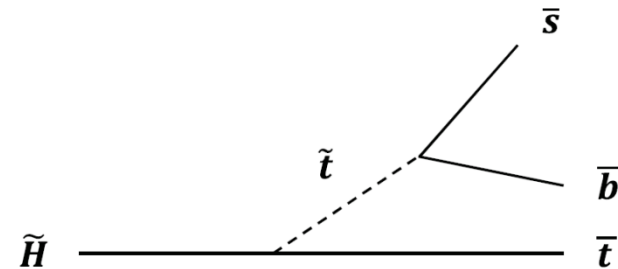


# R-parity violation

## LSP decay ( Constraint from displaced vertex )



### Higgsino LSP



$$\Gamma_{\tilde{H}} \sim \frac{m_{\tilde{H}}}{128\pi^3} |\lambda_{tsb}|^2$$

**If LSP is lighter than top quark,  
decay length is still short.**

# U(1) D-term problem

Strassler (2003)

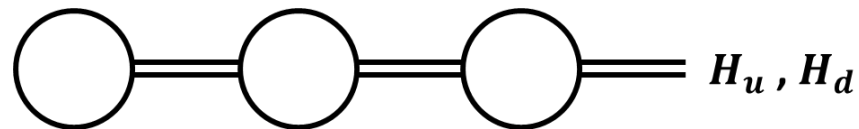
Sundrum (2009)

Heavy scalars  $\rightarrow \mathcal{L}_{\text{FI}} \sim \int d^4\theta \xi V_1$

$\rightarrow$  Large Higgs soft masses !

## Three-site model

	$U(1)_1$	$U(1)_2$	$U(1)_3$
$\Sigma_1$	1	-1	0
$\bar{\Sigma}_1$	-1	1	0
$\Sigma_2$	0	1	-1
$\bar{\Sigma}_2$	0	-1	1
$H_u$	0	0	1/2
$H_d$	0	0	-1/2



$$W \sim X_1 (\Sigma_1 \bar{\Sigma}_1 - v_1^2) + X_2 (\Sigma_2 \bar{\Sigma}_2 - v_2^2)$$

$$\rightarrow |\Sigma_1|^2 - |\bar{\Sigma}_1|^2 \sim |\Sigma_2|^2 - |\bar{\Sigma}_2|^2 \sim \frac{\xi}{g_Y}$$