

Supergravity in spacetime

Local susy

$$e_{\mu}^{\alpha}, \psi_{\mu}^{\alpha}$$

$$\rightarrow e^{\alpha} = dx^{\mu} e_{\mu}^{\alpha}$$

← bosonic 1-forms

$$\psi^{\alpha} = dx^{\mu} \psi_{\mu}^{\alpha}$$

← fermionic 1-forms

Diff. forms (bosonic)

$$\Omega_q = \frac{1}{q!} dx^{\mu_1} \wedge \dots \wedge dx^{\mu_q} \Omega_{\mu_1 \dots \mu_q}(x)$$

Exterior derivative \equiv differential

$$d = dx^\mu \partial_\mu ; \quad d dx^\mu = 0, \quad dd \equiv 0$$

Acts from the right:

$$\begin{aligned} d\Omega_q &= \frac{1}{(q+1)!} dx^{\mu_1} \wedge \dots \wedge dx^{\mu_{q+1}} \partial_{\mu_1} \Omega_{\mu_2 \dots \mu_{q+1}}(x) \\ &\equiv \frac{1}{q!} dx^{\mu_1} \wedge \dots \wedge dx^{\mu_q} \wedge d\Omega_{\mu_1 \dots \mu_q}(x) \end{aligned}$$

$\llcorner \partial_{\mu_1} \Omega_{\mu_2 \dots \mu_{q+1}} \text{ with } (-)^{j-1}$
 cyclic perm. $(-)^{\text{parity}}$

Exterior product

$$\begin{aligned} \Omega_q \wedge \Omega_p &= (-)^{pq} \Omega_p \wedge \Omega_q, \\ dx^\mu \wedge dx^\nu &= -dx^\nu \wedge dx^\mu \end{aligned}$$

$$d(\Omega_q \wedge \Omega_p) = \Omega_q \wedge d\Omega_p + (-)^p d\Omega_q \wedge \Omega_p$$

Contraction with variation symbol (formal)

$$i_\delta(\Omega_q \wedge \Omega_p) = \Omega_q \wedge i_\delta \Omega_p + (-)^p i_\delta \Omega_q \wedge \Omega_p$$

(Generalized) Lie derivative formula

$$\delta \Omega_q = i_\delta d\Omega_q + d i_\delta \Omega_q$$

If we treat it literally:

$$i_{\delta} dx^{\mu} = \delta x^{\mu}$$

$$i_{\delta} \Omega_q = \frac{1}{(q-1)!} dx^{\mu_1} \wedge \dots \wedge dx^{\mu_{q-2}} \delta x^{\mu_{q-1}} \Omega_{\mu_1 \dots \mu_{q-1}}(x)$$

$$\delta \Omega_q = i_{\delta} d\Omega_q + d i_{\delta} \Omega_q$$

describes general coordinate or diffeomorphism transformation.

But generalized Lie derivative can be used in a wider context

$$\delta \Omega_q = i_{\delta} d\Omega_q + d i_{\delta} \Omega_q$$

e.g. if Ω_q is a q -form gauge field and the theory is inv under $\delta \Omega_q = d d_{q-1}$ (gauge symm.)

$i_{\delta} \Omega_q$ can be identified with parameter of gauge symm.

We will also use covariant Lie derivative formula for variation of covariant objects

$$\delta \Omega_q^A = i_{\delta} D \Omega_q^A + D i_{\delta} \Omega_q^A$$

$$\text{where } D \Omega_q^A = d \Omega_q^A - \Omega_q^B \wedge A_B^A$$

\nearrow connection
(1-form gauge field)

in the cases when $i_{\delta} A_B^A$ can be identified with parameter of (manifest) gauge symm. of the theory

(2) SUGRA IN spacetime & diff. forms.

mostly minus $\eta^{ab} = (+, -, \dots, -)$

2.1. GRAVITY AND DIFF. FORMS

$e^a_\mu, \psi^a \leftrightarrow dx^\mu e^a_\mu = e^a$
 $dx^\mu \psi^a = \psi^a$

$$S_{EH} = \frac{1}{2\kappa^2} \int \epsilon_{abcd} R^{ab} \wedge e^c \wedge e^d = \frac{1}{2\kappa^2} \int d^4x \sqrt{|g|} R(e, \omega)$$

Christoffel and spin connection
 $0 = \hat{\nabla}_\mu e^a_\nu = \partial_\mu e^a_\nu + \Gamma^\rho_{\mu\nu} e^a_\rho$

$\wedge \equiv$ EXTERIOR PRODUCT

CURVATURE

$R^{ab} := \frac{1}{2} dx^\nu \wedge dx^\mu R_{\mu\nu}{}^{ab} = \frac{1}{2} e^d \wedge e^c R_{cd}{}^{ab} = d\omega^{ab} - \omega^a{}_c \wedge \omega^c{}_b$

TORSION

$T^a := \frac{1}{2} dx^\nu \wedge dx^\mu T_{\mu\nu}{}^a = \frac{1}{2} e^c \wedge e^b T_{bc}{}^a = D e^a = d e^a - e^b \wedge \omega_b{}^a$

BIS: $DR^{ab} = 0$

$dx^\mu \wedge dx^\nu \wedge dx^\rho \wedge dx^\sigma = d^4x \in^{\mu\nu\rho\sigma}$

$\epsilon^{0123} = 1 = -\epsilon_{0123}$

$S_{EH} = \frac{1}{2\kappa^2} \int_{M^4} \epsilon_{abcd} R^{ab} \wedge e^c \wedge e^d$; FORMAL EXTERIOR DERIVATIVES
 $d \int_{M^4} \epsilon_{abcd} R^{ab} \wedge e^c \wedge e^d = \int_{M^4} \epsilon_{abcd} R^{ab} \wedge e^c \wedge T^d$

2.3. Variation of the Einstein action and equivalent forms of Einstein eq.

in tensor notation, but starting from $R^{ab} = d\omega^{ab} - \omega^a{}_c \wedge \omega^c{}_b \Rightarrow DR^{ab} = 0$

Lie derivative prescriptions:

- calculate d
 - use $i_\xi d = \delta$
- Notice: $i_\xi e^a = 0$
 $i_\xi \omega^{ab} = 0$
 for mathematical convenience (gauge symm. \rightarrow exact)

Covariant Lie derivative prescriptions:

- calculate D
- use $i_\xi D = \delta$, $i_\xi R^a{}_b = \delta^a{}_b$

$\Rightarrow \delta R^{ab} = D \delta \omega^{ab}$

$\Rightarrow \delta R_{\mu\nu}{}^{ab} = 2 D_\mu \delta \omega_{\nu}{}^{ab} = 2 (\delta \omega_{\nu}{}^{ab})_{;\mu} + 2 \delta \omega_{\nu}{}^{a}{}_{\mu} \omega^{\mu b}$
 Lorentz covariant derivative

$\delta \omega_p{}^a = i_\xi \omega_p{}^a + D i_\xi \omega_p{}^a$

$\Rightarrow \delta e^a = i_\xi T^a + D i_\xi e^a$
 target space copy of diff. gauge symm.
 formal expression for variation of E^a

Using identities (to prove:)

(Ex. 1:) $e D_a \xi^a = \partial_\mu (e \xi^\mu) - e T_{ba}{}^b \xi^a$

(Ex. 2:) $D_\mu (e e_a{}^\mu) = e T_{ba}{}^b$

$T_{\mu\nu}{}^c = D_\mu e_\nu{}^c$

we can show (Ex. 3:)

(Ex. 3) $\delta(\epsilon R) = -2 (R_a{}^b - \frac{1}{2} \delta_a{}^b R) e_b{}^\mu \delta e_\mu{}^a - (T_{ab}{}^c e_c{}^\mu + T_{ca}{}^c e_b{}^\mu) \delta e_\mu{}^b + 2 (e e_a{}^\mu \delta e_\mu{}^a)$
 $\hookrightarrow G_a{}^b =$ Einstein tensor

\Rightarrow Einstein eq. $G_{ab} = R_{ab} - \frac{1}{2} \eta_{ab} R = 0$, "metricity" $T_{ab}{}^c = 0$

2.2. (2.4) Rarita-Schwinger action in curved spacetime

$\sigma_{\alpha\beta}^a = \epsilon_{\alpha\beta} \epsilon_{ij} \tilde{\sigma}^{ab}$ - Relativistic Pauli matrices
 $\tilde{\sigma}^{ab} = \gamma^a \gamma^b - \frac{1}{2} \epsilon^{abcd} \tilde{\sigma}_{cd}$
 $\tilde{\sigma}^{ab} = \gamma^a \gamma^b - \frac{1}{2} \epsilon^{abcd} \tilde{\sigma}_{cd}$
 $\epsilon^{\alpha\beta\gamma} = \epsilon^{\alpha\beta\gamma}$

$R^{ab} = \frac{1}{2} R^{\alpha\beta} (\tilde{\sigma}^{ab})_{\alpha\beta} = \frac{1}{2} R^{\alpha\beta} (\tilde{\sigma}^{ab})_{\alpha\beta}$

$R_a^P = \frac{1}{4} R^{ab} (\tilde{\sigma}^{ab})_a^P$; $R_a^P = \frac{1}{4} R^{ab} (\tilde{\sigma}^{ab})_a^P$
 $\hookrightarrow d\omega_a^P - \omega_a^r \omega_r^P$ $\hookrightarrow d\tilde{\omega}_a^P - \tilde{\omega}_a^r \tilde{\omega}_r^P$

$\sigma_{\alpha\beta}^a = e^a_\mu \sigma_{\alpha\beta}^{\mu}$

$D\psi^\alpha = d\psi^\alpha - \psi^\beta \omega_\beta^\alpha$; $\omega_\beta^\alpha = \omega^\mu (\tilde{\sigma}^{\mu\alpha})_\beta = \pm (\tilde{\sigma}^{\mu\alpha})_\beta^*$
 $\sigma_{ab} \tilde{\sigma}_c + \tilde{\sigma}_c \sigma_{ab} = -2i \epsilon_{abcd}$

$\psi^\alpha \chi^\beta \psi^\gamma = + \psi^\beta \chi^\alpha \psi^\gamma$

$\int_{RS} = -4 \int d^4x \sigma_{\alpha\beta}^a \psi^\alpha \bar{\psi}^\beta + 4 \int d^4x \psi^\alpha \sigma_{\alpha\beta}^a D\bar{\psi}^\beta$
 $= +8 \int d^4x \epsilon^{\mu\nu\rho\sigma} e_\mu^a \psi_\nu \tilde{\sigma}_a^{\rho\sigma} \bar{\psi}$

Ex: 6

$d \int_{RS} = -4 \int d^4x (\partial_\mu \psi^\alpha \sigma_{\alpha\beta}^a \bar{\psi}^\beta - \psi^\alpha \partial_\mu \bar{\psi}^\beta \sigma_{\alpha\beta}^a) + 8 \int d^4x \psi^\alpha \partial_\mu \bar{\psi}^\beta \sigma_{\alpha\beta}^a + 4 \int d^4x (\partial_\mu \psi^\alpha \bar{\psi}^\beta + \psi^\alpha \partial_\mu \bar{\psi}^\beta)$
 $\hookrightarrow -4 \int d^4x (\tilde{\sigma}_a^{\mu\nu} R_{\mu\nu}^a + \sigma_{\mu\nu}^a R^{\mu\nu}) \psi^\alpha \bar{\psi}^\beta$
 $\frac{1}{4} R^{\mu\nu} \epsilon^{\alpha\beta\gamma\delta} (\tilde{\sigma}_{\mu\nu}^{\alpha\beta} + \tilde{\sigma}_{\mu\nu}^{\gamma\delta})$
 $= -2i \epsilon_{abcd} \tilde{\sigma}^a$

$d \int_{RS} = -8 \int d^4x \sigma_{\alpha\beta}^a \psi^\alpha D\bar{\psi}^\beta + 2i \epsilon_{abcd} R^{\mu\nu} \epsilon^{\alpha\beta\gamma\delta} \psi^\alpha \bar{\psi}^\beta + 4 \int d^4x (\psi^\alpha \partial_\mu \bar{\psi}^\beta + \psi^\alpha \partial_\mu \bar{\psi}^\beta)$

$S^{RS} = \frac{1}{k^2} \int \int_{RS} = -\frac{4}{k^2} \int d^4x (\partial_\mu \psi^\alpha \sigma_{\alpha\beta}^a \bar{\psi}^\beta - \psi^\alpha \partial_\mu \bar{\psi}^\beta \sigma_{\alpha\beta}^a)$
 $= -\frac{8}{k^2} \int d^4x \epsilon^{\mu\nu\rho\sigma} \psi_\mu \tilde{\sigma}_\nu^{\rho\sigma} \bar{\psi}$
 $\hookrightarrow e_\nu^a \sigma_{\alpha\beta}^a$

25) Variation of the free RS action in flat spacetime & pre-susy

$$d\mathcal{L}_4^{RS(0)} = -8 d\psi^\alpha \delta_{\alpha i}^{0\gamma} \bar{\psi}^i$$

$$\delta\mathcal{L}_4^{RS(0)} = i_8 d d\mathcal{L}_4^{RS(0)} + d i_8 \mathcal{L}_4^{RS(0)} = -8 (d\psi^\alpha \delta_{\alpha i}^{0\gamma}) \delta\bar{\psi}^i + 8 \delta\psi^\alpha (\delta_{\alpha i}^{0\gamma} d\bar{\psi}^i) + d(-4\delta\psi^\alpha \delta_{\alpha i}^{0\gamma} \bar{\psi}^i + 4\psi^\alpha \delta_{\alpha i}^{0\gamma} \delta\bar{\psi}^i)$$

Eqs: $(d\psi^\alpha \delta_{\alpha i}^{0\gamma})_i = 0$, $(\delta_{\alpha i}^{0\gamma} d\bar{\psi}^i)_\alpha = 0$

$$\begin{aligned} &\leftarrow \int dx^1 dx^2 dx^3 (\partial_\mu \psi^\alpha \delta_{\alpha i}^{0\gamma})_i \\ &\leftarrow \int dx^1 dx^2 dx^3 \epsilon^{\mu\nu\rho\sigma} (\partial_\nu \psi^\alpha \delta_{\alpha i}^{0\gamma})_i \end{aligned}$$

RS eqs. $\epsilon^{\mu\nu\rho\sigma} (\partial_\nu \psi^\alpha \delta_{\alpha i}^{0\gamma})_i = 0$

$$\delta_{\alpha i}^{0\gamma} = dx^\alpha \delta_{\alpha i}^{0\gamma} \Rightarrow d\delta_{\alpha i}^{0\gamma} = 0$$

$$d((d\psi^\alpha \delta_{\alpha i}^{0\gamma})_i) \equiv 0$$

Noether identity for the fermionic gauge symm. ("PRE-SUSY")
 $\delta\psi^\alpha = d\epsilon^\alpha$ or, eqv. $\delta\psi^\alpha_\mu = \partial_\mu \epsilon^\alpha$

IN CURVED SPACETIME:

$$d \rightarrow D = d + \omega$$

$$\delta_{\alpha i}^{0\gamma} = e^\alpha_\mu \delta_{\alpha i}^{0\gamma} \equiv dx^\mu e^\alpha_\mu \delta_{\alpha i}^{0\gamma}$$