

First order form of the CJS action [Julia & Silva, JHEP 2000]

$$S = \int_{M^{11}} \mathcal{L}_{11}[E^a, \psi^\alpha, \omega^{ab}, A_3, F_{a_1 a_2 a_3 a_4}]$$

$$\begin{aligned} \mathcal{L}_{11} &= \frac{1}{4} R^{ab} \wedge E_{ab}^{\wedge 9} - D\psi^\alpha \wedge \psi^\beta \wedge \bar{\Gamma}_{\alpha\beta}^{(8)} + \\ &+ \frac{1}{4} \psi^\alpha \wedge \psi^\beta \wedge (T^a + i/2 \psi \wedge \psi \Gamma^a) \wedge E_a \wedge \bar{\Gamma}_{\alpha\beta}^{(6)} + \\ &+ (dA_3 - a_4) \wedge (*F_4 + b_7) + \frac{1}{2} a_4 \wedge b_7 - \\ &- \frac{1}{2} F_4 \wedge *F_4 - \frac{1}{3} A_3 \wedge dA_3 \wedge dA_3 , \end{aligned}$$

$$\text{where } a_4 := \frac{1}{2} \psi^\alpha \wedge \psi^\beta \wedge \bar{\Gamma}_{\alpha\beta}^{(2)} :=$$

$$:= -\frac{1}{4} \psi^\alpha \wedge \psi^\beta \wedge E^a \wedge E^b \Gamma_{ab\alpha\beta} ,$$

$$b_7 := \frac{i}{2} \psi^\alpha \wedge \psi^\beta \wedge \bar{\Gamma}_{\alpha\beta}^{(5)} :=$$

$$:= \frac{i}{2 \cdot 5!} \psi^\alpha \wedge \psi^\beta \wedge E^{a_1} \wedge \dots \wedge E^{a_5} \Gamma_{a_1 \dots a_5 \alpha \beta} ,$$

and $F_4 = \frac{1}{4!} E^{a_4} \wedge \dots \wedge E^{a_1} F_{a_1 \dots a_4}$ with independent $F_{a_1 \dots a_4}(Z)$,

$$\begin{aligned} *F_4 &:= -\frac{1}{4!} E^{a_1 \dots a_4} F^{a_1 \dots a_4} \\ &\equiv \frac{1}{7!4!} E^{b_7} \wedge \dots \wedge E^{b_1} \varepsilon_{b_1 \dots b_7 a_1 \dots a_4} F^{a_1 \dots a_4} . \end{aligned}$$

We have also used the compact notation

$$\begin{aligned} \bar{\Gamma}_{\alpha\beta}^{(k)} &:= \frac{1}{k!} E^{a_k} \wedge \dots \wedge E^{a_1} \Gamma_{a_1 \dots a_k \alpha \beta} \quad \text{and} \\ E_{a_1 \dots a_k}^{\wedge(11-k)} &:= \frac{1}{k!} \varepsilon_{a_1 \dots a_k b_1 \dots b_{11-k}} E^{b_1} \wedge \dots \wedge E^{b_{11-k}} . \end{aligned}$$

Ex: to obtain eqs of motion from the first order action

$$\delta_A S = \int \mathcal{G}_8 \wedge \delta A_3, \quad \frac{\delta S}{\delta A_3} := \mathcal{G}_8 = d(*F_4 + b_7 - A_3 \wedge dA_3) = 0,$$

This can be written as $D_{[c_1} F_{c_2 \dots c_8]} - \frac{7!}{4!4!} F_{[c_1 \dots c_4} F_{c_5 \dots c_8]} = 0$, but includes the (auxiliary) independent field $F_{abcd} = F_{[abcd]}$. However,

$$\delta_F S = \int (dA_3 - a_4 - F_4) \wedge * \delta F_4,$$

so that $\delta S / \delta F^{a_1 \dots a_4} = 0$ can be written as $* \frac{\delta S}{\delta F_4} = (dA_3 - a_4 - F_4) = 0$ and identifies

$$F_4 := \frac{1}{4!} E^{a_4} \wedge \dots \wedge E^{a_1} F_{a_1 \dots a_4} = dA_3 - \frac{1}{2} \psi^\alpha \wedge \psi^\beta \wedge \bar{\Gamma}_{\alpha\beta}^{(2)}.$$

Equations for the spin connection

$$\frac{\delta S_{11}}{\delta \omega^{ab}} = \frac{1}{4} E_{abc}^{\wedge 8} \wedge (T^c + i\psi^\alpha \wedge \psi^\beta \Gamma_{\alpha\beta}^c) = 0 \quad \Rightarrow$$

$$\Rightarrow T^a := DE^a = -i\psi^\alpha \wedge \psi^\beta \Gamma_{\alpha\beta}^a.$$

Gravitino equations have a compact form

$$\frac{\delta S_{11}}{\delta \psi^\alpha} = 0 \quad \Rightarrow \quad \Psi_{10\beta} := \hat{D}\psi^\alpha \wedge \bar{\Gamma}_{\alpha\beta}^{(8)} = 0,$$

in terms of "connection of generalized holonomy" $w_\beta^\alpha := \omega_\beta^\alpha + t_{1\beta}^\alpha$,

$$\hat{D}\psi^\alpha := d\psi^\alpha - \psi^\beta \wedge w_\beta^\alpha \equiv d\psi^\alpha - \psi^\beta \wedge (\omega_\beta^\alpha + t_{1\beta}^\alpha),$$

$$t_{1\beta}^\alpha = \frac{i}{18} E^a \left(F_{abcd} \Gamma^{bcd}{}_\beta{}^\alpha + \frac{1}{8} F^{bcde} \Gamma_{abcde}{}_\beta{}^\alpha \right).$$

Finally $\delta e^a \Rightarrow$ the Einstein equation for $D = 11$ supergravity

$$M_{10a} := R^{bc} \wedge E_{abc}^{\wedge 8} + 2(i_a F_4 \wedge *F_4 + F_4 \wedge i_a(*F_4)) + \mathcal{O}(\psi^{\wedge 2}) = 0$$

$$\Leftrightarrow R_{acb}{}^c = -\frac{1}{3} F_{ac_1 c_2 c_3} F_b{}^{c_1 c_2 c_3} + \frac{1}{36} \eta_{ab} F_{c_1 c_2 c_3 c_4} F^{c_1 c_2 c_3 c_4} + \mathcal{O}(\psi^{\wedge 2}).$$

SUSY transformations leaving the action invariant can be read of the supergravity constraints

$$T^a = -iE^\alpha \wedge E^\beta \Gamma_{\alpha\beta}^a ,$$

$$T^\alpha = -\frac{i}{18} E^a \wedge E^\beta \left(F_{ac_1c_2c_3} \Gamma^{c_1c_2c_3}{}_\beta^\alpha + \frac{1}{8} F^{c_1c_2c_3c_4} \Gamma_{ac_1c_2c_3c_4}{}^\alpha{}_\beta \right) + \frac{1}{2} E^a \wedge E^b T_{ba}{}^\alpha(Z) ,$$

$$\mathcal{F}_4 := dA_3 = \frac{1}{2} E^\alpha \wedge E^\beta \wedge \bar{\Gamma}_{\alpha\beta}^{(2)} + \frac{1}{4!} E^{c_4} \wedge \dots \wedge E^{c_1} F_{c_1\dots c_4}(Z) ,$$

$$\mathcal{F}_7 := dA_6 + A_3 \wedge dA_3 = \frac{i}{2} E^\alpha \wedge E^\beta \wedge \bar{\Gamma}_{\alpha\beta}^{(5)} + \frac{1}{7!} E^{c_7} \wedge \dots \wedge E^{c_1} F_{c_1\dots c_7}(Z) .$$

Namely

$$\delta e^a = -i\psi^\alpha \Gamma_{\alpha\beta}^a \epsilon^\beta , \quad \delta A_3 = \psi^\alpha \wedge \bar{\Gamma}_{\alpha\beta}^{(2)} \epsilon^\beta ,$$

$$\delta \psi^\alpha = \mathcal{D}\epsilon^\alpha := D\epsilon^\alpha - \epsilon^\beta t_{1\beta}{}^\alpha ,$$

where

$$t_{1\beta}{}^\alpha = -\frac{i}{18} E^a \left(F_{ac_1c_2c_3} \Gamma^{c_1c_2c_3}{}_\beta^\alpha + \frac{1}{8} F^{c_1c_2c_3c_4} \Gamma_{ac_1c_2c_3c_4}{}^\alpha{}_\beta \right) ,$$

plus suitable transformations of the spin connection.

Killing spinor equations

Can purely bosonic solutions be supersymmetric? As

$$\delta\psi_\mu^\alpha = \mathcal{D}_\mu\epsilon^\alpha := D_\mu\epsilon^\alpha - \epsilon^\beta t_{\mu\beta}^\alpha$$

$$t_{1\beta}^\alpha = dx^\mu t_{\mu\beta}^\alpha = -\frac{i}{18}e^a \left(F_{a[3]}\Gamma_{\beta}^{[3]\alpha} + \frac{1}{8}F^{[4]}\Gamma_{a[4]\beta}^\alpha \right),$$

if $\psi_\mu^\alpha = 0$, then ($\delta e^a = 0$ and $\delta A_3 = 0$ are satisfied identically while)

$\delta\psi_\mu^\alpha = 0$ gives a nontrivial **Killing spinor equation**

$$\boxed{\mathcal{D}\epsilon^\alpha := D\epsilon^\alpha - \epsilon^\beta t_{1\beta}^\alpha = 0}. \quad (*)$$

- If the bosonic solution is such that (*) has no solution \Rightarrow that bosonic solution is not supersymmetric.
- If the bosonic solution is such that the solution of (*) is non-trivial, then it is supersymmetric or **BPS solution**
- If (*) is solved in terms of arbitrary 32 component constant spinor— the bosonic solution is completely SUSY (solution is 32-parametric); [Figueroa-O'Farrill & Papadopoulos 2002:] all such solutions of 11D SUGRA are locally isomorphic to
 - flat 11D space with $F_4 = 0$,
 - Freund-Rubin $AdS_4 \times S^7$ and $AdS_7 \times S^4$ solutions,
 - Hpp-wave solution
- If the solution of (*) has k arbitrary parameters one says that bosonic solution has k Killing spinors, or that it is $k/32$ SUSY or that it describes $k/32$ BPS state. The most interesting in the String/M-theory perspective are 1/2 BPS M2-brane and M5-brane solutions.

SUSY conditions and equations of motion

Why it is easier to SUSY solutions? Because for them the Killing spinor eq $\mathcal{D}\epsilon^\alpha := d\epsilon^\alpha - \epsilon^\beta w_\beta^\alpha = D\epsilon^\alpha - \epsilon^\beta t_{1\beta}^\alpha = 0$ plays the role of the 'Lax pair' (associated linear system) for SUGRA equations.

Its selfconsistency ('integrability') conditions is

$$0 = \mathcal{D}\mathcal{D}\epsilon^\alpha = -\epsilon^\beta \mathcal{R}_\beta^\alpha,$$

where
$$\begin{aligned} \mathcal{R}_\beta^\alpha &:= dw_\beta^\alpha - w_\beta^\gamma \wedge w_\gamma^\alpha \\ &= \frac{1}{4} R^{ab} (\Gamma_{ab})_\alpha^\beta + Dt_{1\alpha}^\beta - t_{1\alpha}^\gamma \wedge t_{1\gamma}^\beta \end{aligned}$$

is the so-called 'curvature of generalized holonomy', and

$$t_{1\beta}^\alpha = e^a t_{a\beta}^\alpha = \frac{i}{18} e^a \left(F_{ac_1c_2c_3} \Gamma^{c_1c_2c_3} + \frac{1}{8} F^{c_1c_2c_3c_4} \Gamma_{ac_1c_2c_3c_4} \right) \alpha_\beta.$$

For completely SUSY solutions $\epsilon^\beta \mathcal{R}_{ab\beta}^\alpha = 0 \quad \forall \epsilon^\beta$ and hence

$$32 \text{ susy} \Rightarrow \mathcal{R}_{ab\beta}^\alpha = 0.$$

\Rightarrow completely SUSY configurations of bosonic fields always solve the bosonic equations of motion of D=11 supergravity, which, when $\psi_\mu^\alpha = 0$, can be collected in (see hep-th/0501007=PLB2005)

$$\mathcal{R}_{ab\beta}^\gamma \Gamma_\gamma^b{}^\alpha = 0,$$

and when $\psi_\mu^\alpha \neq 0$ (see hep-th/0501007=PLB2005) in

$$*\mathcal{M}_\alpha^{a\beta} = \mathcal{R}_{bc\alpha}^\gamma \Gamma^{abc}{}_{\gamma\beta} - 4i((\hat{D}\psi)_{bc} \Gamma^{[abc]})_\beta (\psi_d \Gamma^d)_\alpha = 0 \text{ or equiv.}$$

$$\mathcal{M}_{10\alpha\beta} := \mathcal{R}_\beta^\gamma \wedge E_{abc}^{\wedge 8} \Gamma_{\gamma\alpha}^{abc} + i\hat{D}\psi^\delta \wedge \psi^\gamma \wedge E_{a_1\dots a_4}^{\wedge 7} \Gamma_{\delta\alpha}^{[a_1a_2a_3] \Gamma_{\beta\gamma}^{a_4]} = 0.$$

one can show that

$$\mathcal{M}_{10\alpha\beta} = -3i \left(-2i\Gamma_{\beta\alpha}^a M_{10a} + \mathcal{G}_8 \wedge \bar{\Gamma}_{\beta\alpha}^{(2)} \right)$$

$\nwarrow R^k{}_\alpha E_{abc}^{\wedge 8}$
Einstein eq. of 11D SUGRA

$\llcorner d(*F_7 - A_3 \wedge dA_3 + b_7) = 0$
 $\llcorner \frac{i}{2} \psi_\alpha \psi_\beta \Gamma_{\alpha\beta}^{(5)}$
 A_3 field eqs

where

$$M_{10a} := R^{bc} \wedge E_{abc}^{\wedge 8} + 2(i_a F_4 \wedge *F_4 + F_4 \wedge i_a(*F_4)) + \mathcal{O}(\psi^{\wedge 2}) = 0$$

$$\mathcal{G}_8 := d(*F_4 + b_7 - A_3 \wedge dA_3) = 0 ,$$

the Einstein equation and the A_3 gauge field equations of 11D SUGRA ($b_7 := \frac{i}{2} \psi^\alpha \wedge \psi^\beta \wedge \bar{\Gamma}_{\alpha\beta}^{(5)}$). In other words:

$$M_{10a} = \alpha \operatorname{tr}(\Gamma_a \mathcal{M}_{10}) ,$$

$$\mathcal{G}_8 \wedge E^a \wedge E^b = \alpha \operatorname{tr}(\Gamma^{ab} \mathcal{M}_{10}) .$$

$AdS_4 \times S^7$ Freund-Rubin solution

The $AdS_4 \times S^7$ metric

$$ds^2 = \left(\frac{r}{R}\right)^4 dx^a \eta_{ab} dx^b - \left(\frac{R}{r}\right)^2 \left(dr^2 + r^2 ds_{S^7}^2\right) .$$

Flux

$$F_{\tilde{a}\tilde{b}\tilde{c}\tilde{d}} = \frac{3}{R} \varepsilon_{\tilde{a}\tilde{b}\tilde{c}\tilde{d}} , \quad \text{others} = 0 , \quad \tilde{a}, \tilde{b}, \tilde{c}, \tilde{d} = 0, 1, 2, 3 .$$

The only nonvanishing component of the dual tensor (of the seven form flux) is

$$(F_7)^{\hat{I}_1 \dots \hat{I}_7} = (*F_4)^{\hat{I}_1 \dots \hat{I}_7} = -\frac{1}{8R} \varepsilon^{\hat{I}_1 \dots \hat{I}_7} .$$

***** A BIT MORE DETAILS: *****

Vielbein forms

$$e^a = \left(\frac{r}{R}\right)^2 dx^a , \quad e^r \equiv e^2 = \frac{R}{r} dr , \quad e^{\hat{I}} = R \Omega^{\hat{I}} ,$$

The natural vielbein and spin connection on the S^7 sphere is provided by the set of Cartan forms $\Omega^{\hat{I}}$ (such that $ds_{S^7} = \Omega^{\hat{I}} \otimes \Omega^{\hat{I}}$) and $\omega^{\hat{J}\hat{I}} = -\omega^{\hat{I}\hat{J}}$ which $\vec{n} := (n_J)$ which obey the Maurer-Cartan equations of the $SO(8)$ group

$$D\Omega^{\hat{I}} = d\Omega^{\hat{I}} - \Omega^{\hat{J}} \wedge \omega^{\hat{J}\hat{I}} = 0 , \quad d\omega^{\hat{J}\hat{I}} + \omega^{\hat{J}\hat{K}} \wedge \omega^{\hat{K}\hat{I}} = \Omega^{\hat{I}} \wedge \Omega^{\hat{J}} .$$

can be constructed from the so-called spherical harmonic variables

$$\vec{u}^{\hat{I}} := (u_{J^{\hat{I}}}) , \quad J = 1, \dots, 8, \quad \hat{I} = 1, \dots, 7$$

$$\Omega^{\hat{I}} = \vec{n} d\vec{u}^{\hat{I}} , \quad \omega^{\hat{I}\hat{J}} = -\vec{u}^{\hat{I}} d\vec{u}^{\hat{J}} ,$$

$$(\vec{n}, \vec{u}^{\hat{I}}) := (n_J, u_{J^{\hat{I}}}) \in SO(8) \leftrightarrow \begin{cases} \vec{n}\vec{n} = 1 , \\ \vec{n}\vec{u}^{\hat{J}} = 0 , \\ \vec{u}^{\hat{I}}\vec{u}^{\hat{J}} = \delta^{\hat{I}\hat{J}} \end{cases} .$$

1/2 supersymmetric solutions: M2-brane and M5-brane

M2-brane also known as 11D supermembrane

The M2-brane metric ($a, b = 0, 1, 2; I, J = 1, \dots, 8$)

$$\begin{aligned}
 ds^2 &= e^a \otimes e^b \eta_{ab} - e^I \otimes e^I = \\
 &= \left(1 + k/r^6\right)^{-2/3} dx^\mu \otimes dx^\nu \eta_{\mu\nu} - \left(1 + k/r^6\right)^{-1/3} dy^m dy^n \delta_{mn}, \\
 \mu, \nu &= 0, 1, 2, \quad m, n = 1, \dots, 8.
 \end{aligned}$$

$$A_{\mu\nu\rho} = \epsilon_{\mu\nu\rho} \left(1 + \frac{k}{r^6}\right)^{-2/3}, \quad \text{others} = 0, \quad (1)$$

This solution interpolates between $AdS_4 \times S^7$ and flat M^{11} [G. Gibbons & P. Townsend 1993] This is to say it tends to $AdS_4 \times S^7$ when $r \mapsto 0$ (so that $\frac{k}{r^6} \gg 1$) and to 11D Minkowski spacetime when $r \mapsto \infty$ (so that $\frac{k}{r^6} \ll 1$).

SUSY preserved by M2-brane ($\underline{\alpha}, \underline{\beta} = 1, \dots, 32$)

$$\epsilon^\alpha = (1 + \bar{\gamma})^\alpha_{\underline{\beta}} \kappa^{\underline{\beta}} \quad \Leftrightarrow \quad \epsilon^\alpha = \bar{\gamma}^\alpha_{\underline{\beta}} \epsilon^{\underline{\beta}} \quad (2)$$

is defined with the use of the projector $(1 + \bar{\gamma})$ where

$$\bar{\gamma} := \frac{i}{3!} \epsilon^{abc} \delta_a^a \delta_b^b \delta_c^c \Gamma_{abc}$$

has the properties $\bar{\gamma}^2 = I, tr(\bar{\gamma}) = 0$.

We will be back to the origin of this projector...

In a special Lorentz frame 11D $\bar{\gamma} = i\Gamma_1\Gamma_2\Gamma_3$ and, with a certain Γ matrix representation $\bar{\gamma} = \begin{pmatrix} I_{16 \times 16} & 0 \\ 0 & -I_{16 \times 16} \end{pmatrix}$ so that the solution the preserved susy is characterized by $\epsilon^\alpha = \begin{pmatrix} \epsilon^\alpha \\ 0 \end{pmatrix}$ with $\alpha = 1, \dots, 16$. Hence $1/2 = 16/32$ part of susy is preserved.

M5-brane solution (also known as 5-brane of M-theory)

M2-brane solution is similar to 'electric charge' solution in electrodynamics (although it is not a point-like, but rather extended object).

There is also 1/2 BPS state (=1/2 susy solution of 11D SUGRA) similar to Dirac monopole (also not a point-like but rather extended object). The M2-brane metric ($a, b = 0, 1, \dots, 5; I, J = 1, \dots, 6$)

$$\begin{aligned} ds^2 &= e^a \otimes e^b \eta_{ab} - e^I \otimes e^J = \\ &= (1 + k/r^3)^{-1/3} dx^\mu \otimes dx^\nu \eta_{\mu\nu} - (1 + k/r^3)^{2/3} dy^m dy^n \delta_{mn} , \end{aligned}$$

$\mu, \nu = 0, 1, \dots, 5, m, n = 1, \dots, 6$. As for the Dirac monopole in $d=4$, the 3-form potential of 11D M5-brane solution is not well defined, so that the solution is characterized by the field strength

$$F_{mnpq} = 3k \epsilon_{mnpqr} \frac{y^r}{r^5} . \quad \text{others} = 0 , \quad (3)$$

This solution interpolates between $AdS_7 \times S^4$ and flat M^{11} [G. Gibbons & P. Townsend 1993] This is to say it tends to $AdS_7 \times S^4$ when $r \mapsto 0$ (so that $\frac{k}{r^6} \gg 1$) and to 11D Minkowski spacetime when $r \mapsto \infty$ (so that $\frac{k}{r^6} \ll 1$).

Mp-branes and most general 11D SUSY algebra

Most general susy algebra in 11D is [Van Holten & Van Proeyen, J.Phys.A15:3763,1982.]

$$\{Q_{\underline{\alpha}}, Q_{\underline{\beta}}\} = \mathcal{P}_{\underline{\alpha}\underline{\beta}} = i\Gamma_{\underline{\alpha}\underline{\beta}}^a P_a + \Gamma_{\underline{\alpha}\underline{\beta}}^{ab} Z_{ab} + i\Gamma_{\underline{\alpha}\underline{\beta}}^{a_1 \dots a_5} Z_{a_1 \dots a_5}$$

where $Z_{\underline{ab}} = -Z_{\underline{ba}} = Z_{[ab]}$ and $Z_{\underline{a_1 \dots a_5}} = Z_{[a_1 \dots a_5]}$ are so-called tensorial central charges.

They are called central because $[Z, P] = 0 = [Z, Q]$, $[Z, Z] = 0$. 'Central' are because they transform nontrivially under the $SO(1, 10)$.

These tensorial central charge can be associated with the existence of supersymmetric extended objects - M2-brane and M5-brane in the case of 11D. Schematically, in the presence of M2-brane

$$Z^{ab} = \alpha \int_{R^2} d\xi^1 d\xi^2 \epsilon^{0mn} \partial_m \hat{X}^a \partial_n \hat{X}^b, \quad \hat{X}^a = \hat{X}^a(\tau, \xi^1, \xi^2)$$

and in the presence of M5-brane ($\hat{X}^a = \hat{X}^a(\tau, \vec{\xi})$)

$$Z^{a_1 \dots a_5} = \alpha \int_{R^5} d\xi^1 d\xi^2 d\xi^3 d\xi^4 d\xi^5 \epsilon^{0m_1 \dots m_5} \partial_{m_1} \hat{X}^{a_1} \dots \partial_{m_5} \hat{X}^{a_5},$$

Preserved susy implies that $\boxed{\det(\mathcal{P}_{\underline{\alpha}\underline{\beta}}) = 0}$ (otherwise $\epsilon^\alpha \mathcal{P}_{\underline{\alpha}\underline{\beta}} = 0$ has only trivial solution).

Preserved 1/2 of susy: $rank(\mathcal{P}_{\underline{\alpha}\underline{\beta}}) = 16$ which implies that $C\mathcal{P}$ is a projector ($C\mathcal{P}C\mathcal{P} = \alpha C\mathcal{P}$).

Membrane (supermembrane or M2-brane) ground state:

$$P^a = \alpha \delta_0^{[a} T \quad Z^{ab} = \alpha T \delta_1^{[a} \delta_2^{b]}$$

$$\boxed{\{Q_{\underline{\alpha}}, Q_{\underline{\beta}}\} = iT \Gamma_{(\underline{\alpha}\underline{\gamma}}^0 (\delta + i\Gamma^0 \Gamma^1 \Gamma^2)^{\underline{\gamma}\underline{\beta}}) = iT (\Gamma^0 (I + \bar{\gamma}))_{\underline{\alpha}\underline{\beta}}}$$

Thus $\epsilon^\alpha Q_{\underline{\alpha}} |M2\rangle = 0$ iff $\boxed{(I + \bar{\gamma})^\alpha_{\underline{\beta}} \epsilon^{\underline{\beta}} = 0}$ (the same relation as follows from Killing spinor equation).