

# Planck 2015 Cosmology results

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Planck Collaboration

8<sup>th</sup> June 2015

# Planck in 2015

- Parameters
  - Likelihood
  - Lensing
  - Dark Energy/Modified Gravity
- Inflation
  - Non-gaussianity
  - BKP
- (Geometry and Topology)
- (Isotropy and Statistics)

# Where are we now with Planck?

- A refined analysis
  - Same “cross-spectrum” methodology for the  $C_l$  likelihood as in 2013
  - But less conservative now
    - More aggressive use of sky
    - Can present results using data from the full mission
- Redundancy
  - Multiple surveys per detector
    - Good for probing systematics
  - Multiple frequencies
    - Good for mitigating foregrounds

# Main refinements since 2013

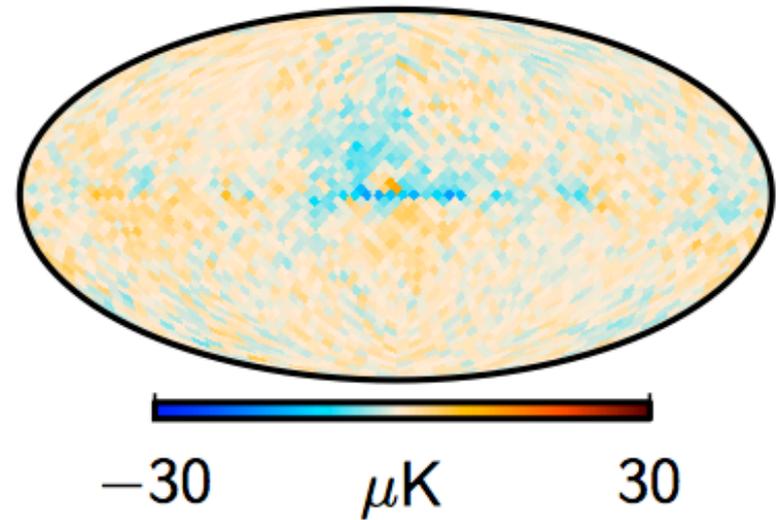
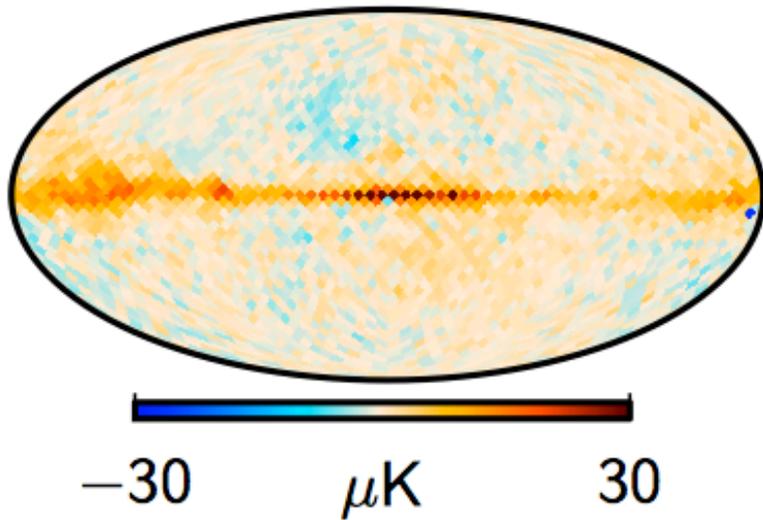
- Improved understanding of beams
  - Helped resolve calibration differences with WMAP
- Correction for the nonlinearity of the analogue-to-digital converters in the detectors
  - Removed much of the apparent gain variation seen in the 2013 data
  - Allowed us to calibrate properly off the orbital dipole

# Planck Polarization

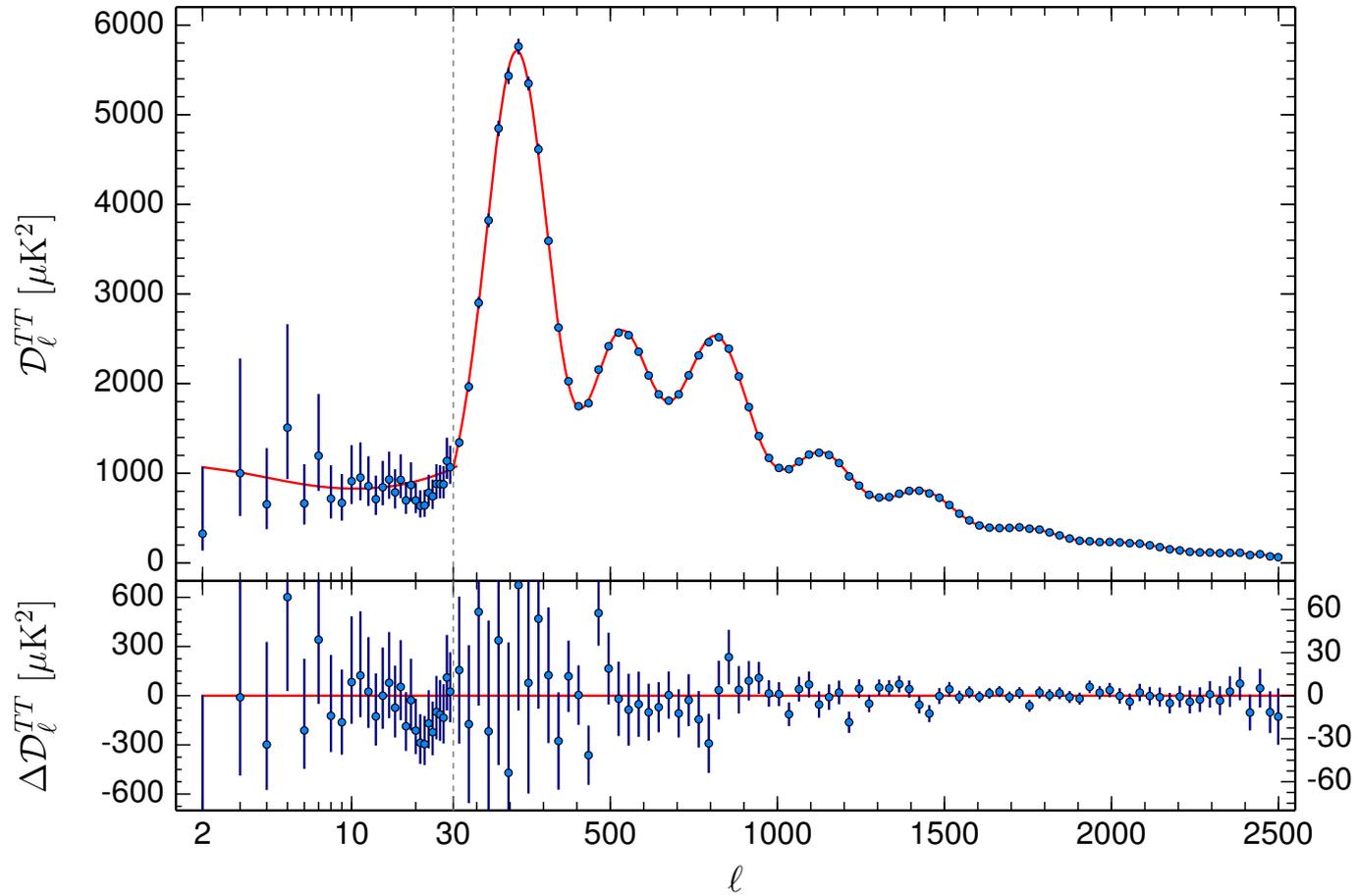
- What has P added?
  - Tighter error bars
  - Improvements in isocurvature constraints
  - Redundancy
    - Unresolved foregrounds in P are much less significant than in T
    - TE results are highly consistent with TT ones
- What complications has P bought?
  - T->P leakage (not fully quantified for 2015)

# And at low- $l$

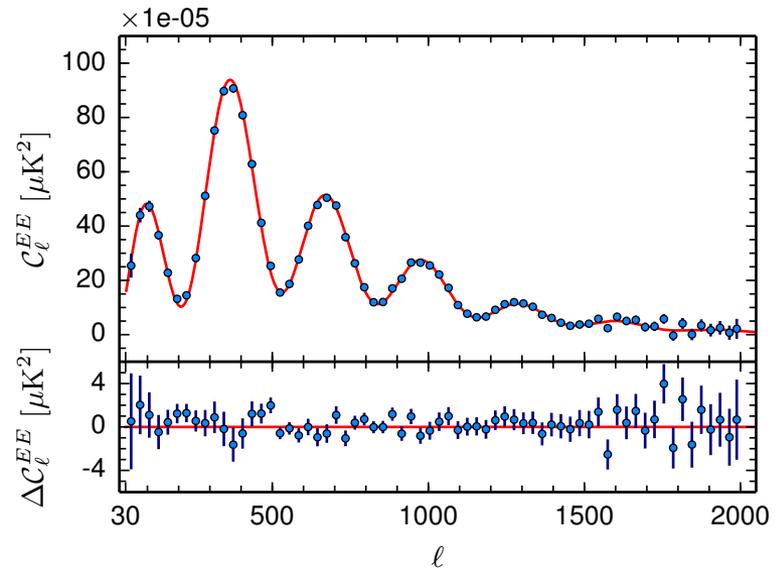
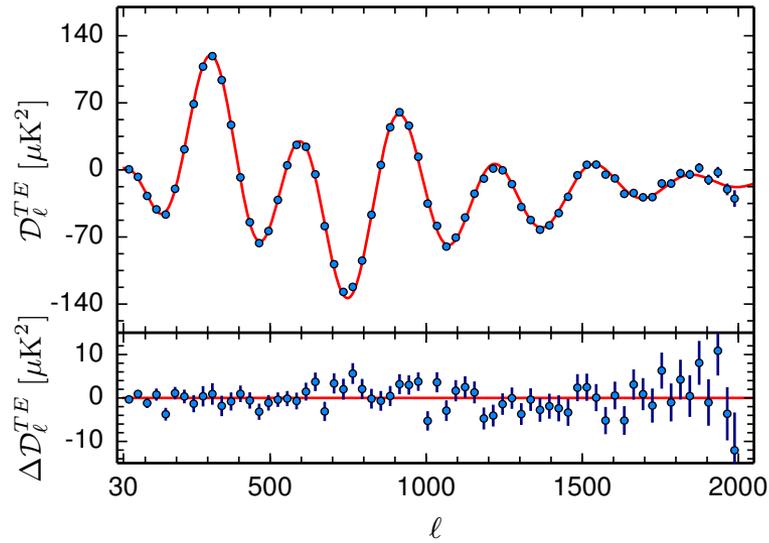
- We now use our own cleaned 70 GHz-based likelihood at low- $l$



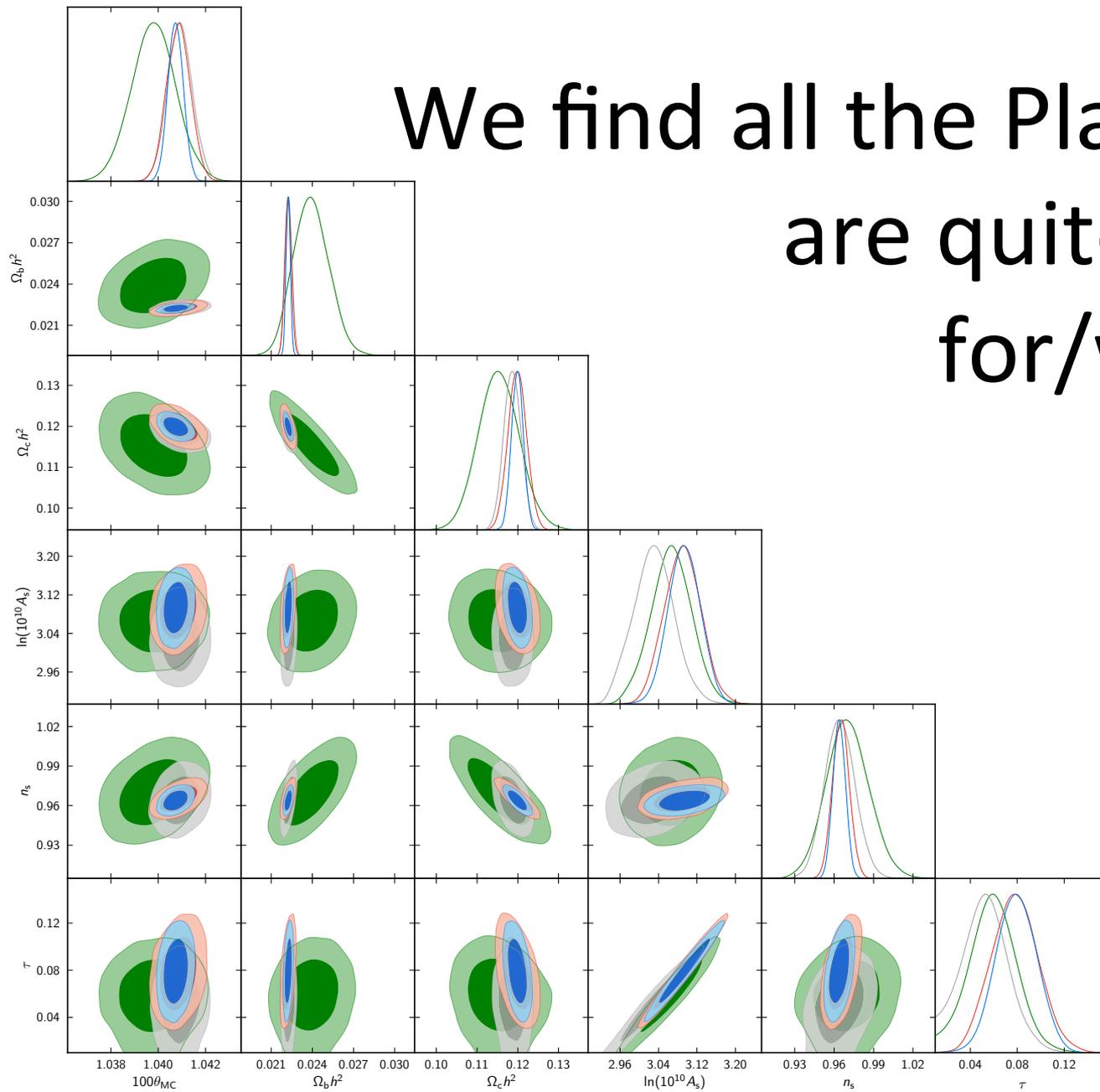
# Our new “combined” TT



# ...and now our TE & EE:

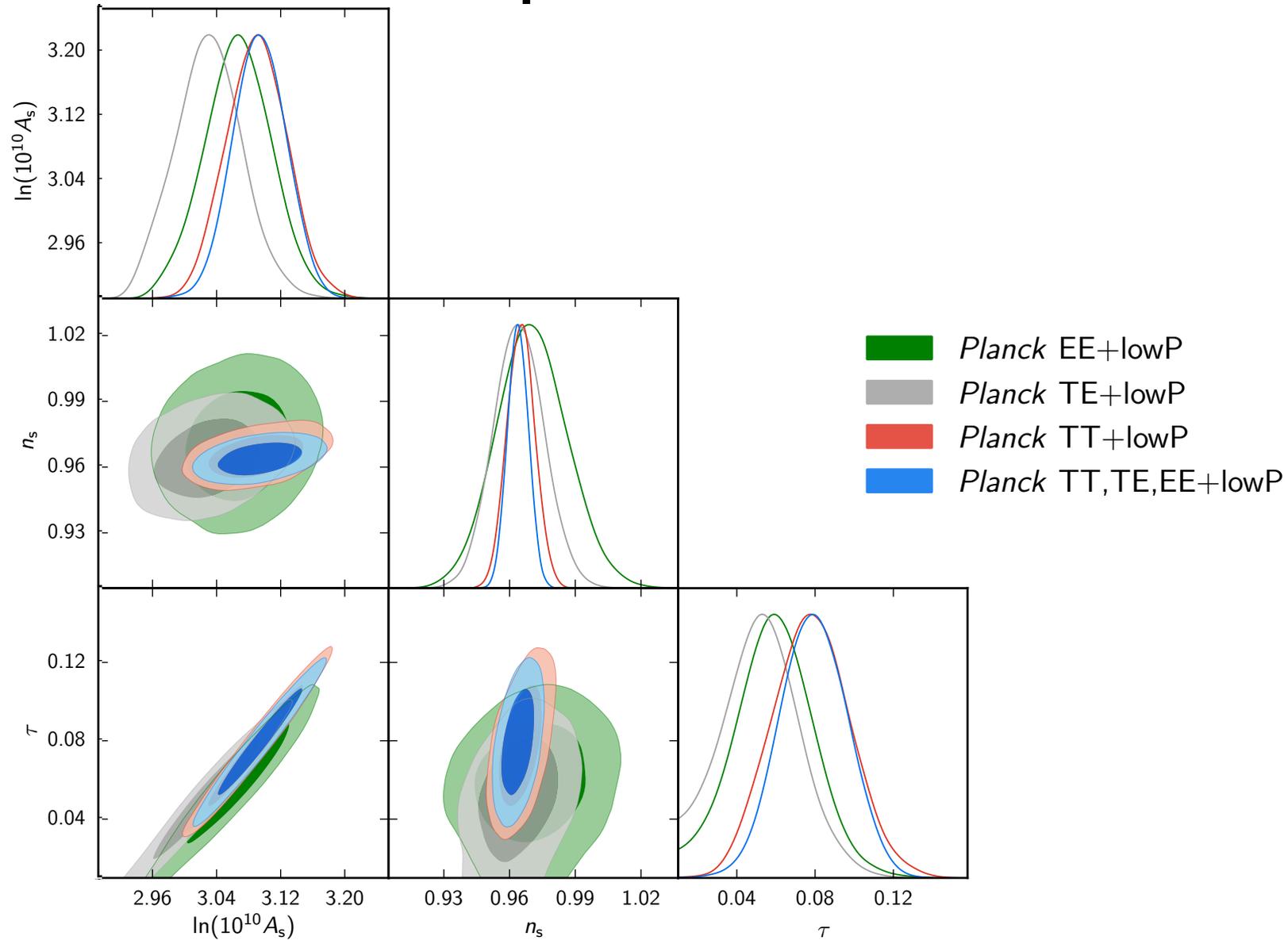


We find all the Planck spectra  
are quite consistent  
for/with  $\Lambda$ CDM!



- *Planck* EE+lowP
- *Planck* TE+lowP
- *Planck* TT+lowP
- *Planck* TT,TE,EE+lowP

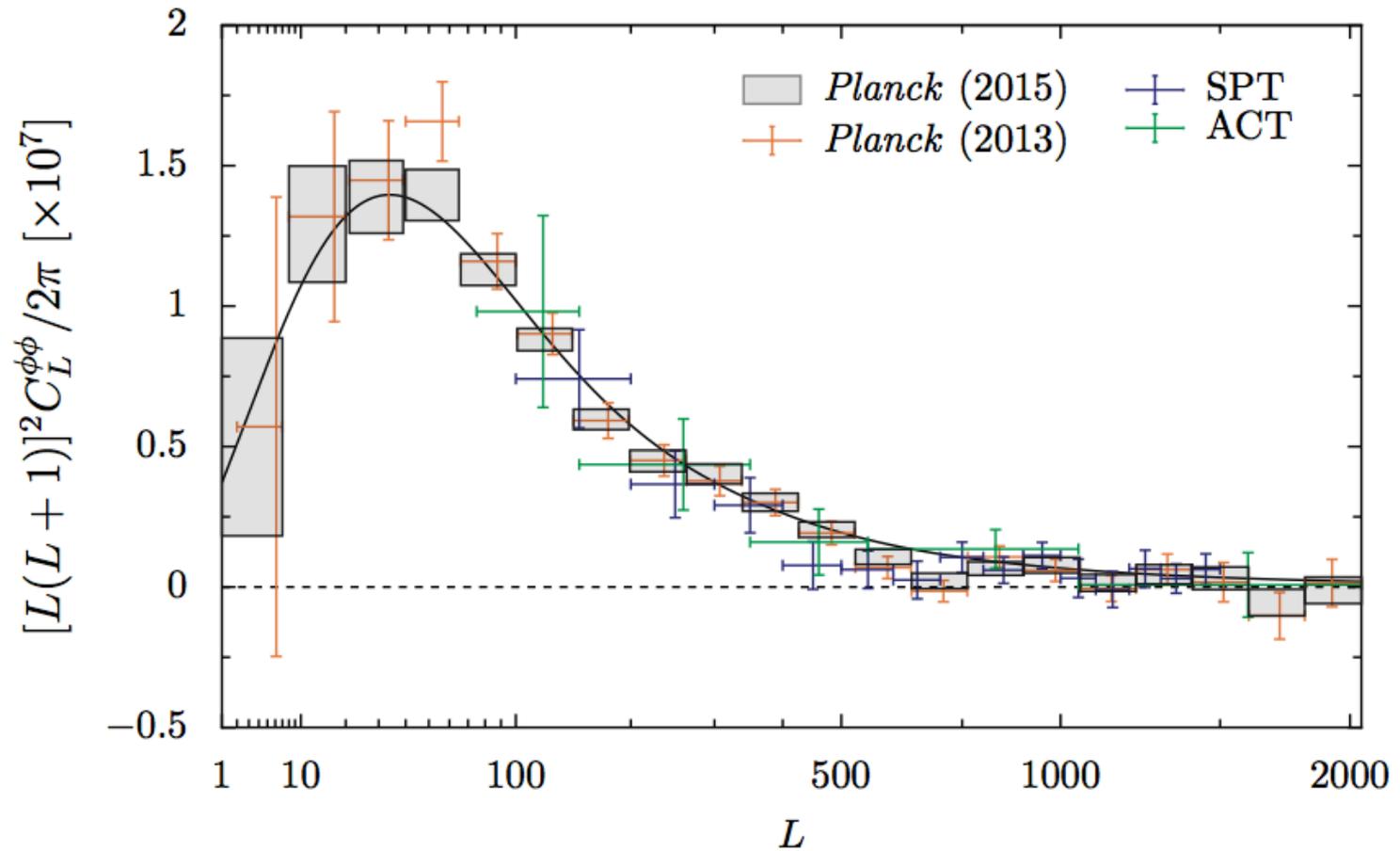
# An Example Zoom:



# Planck Lensing

- Mass inhomogeneities deflect the CMB photons on their way to us
- Can calculate the power spectrum of the deflection field and compare it to theory
  - i.e. a four-point function of the CMB
  - Mostly sensitive to matter nearby so can be thought of as an internal “low-redshift” dataset
  - Helps break degeneracies
- New for 2015 is the inclusion of polarization maps in the inference

# 2015 Lensing potential spectrum



# Extensions

- $\Omega_k$ 
  - TT:  $-0.052 \pm 0.05$  (95%)
  - TTTEEE:  $-0.040 \pm 0.04$  (95%)
  - TTTEEE+lensing:  $-0.004 \pm 0.015$  (95%)
- $dn_s / d \ln k$ 
  - TT:  $-0.008 \pm 0.016$  (95%)
  - TTTEEE:  $-0.006 \pm 0.014$  (95%)
  - TTTEEE+lensing:  $-0.002 \pm 0.013$  (95%)

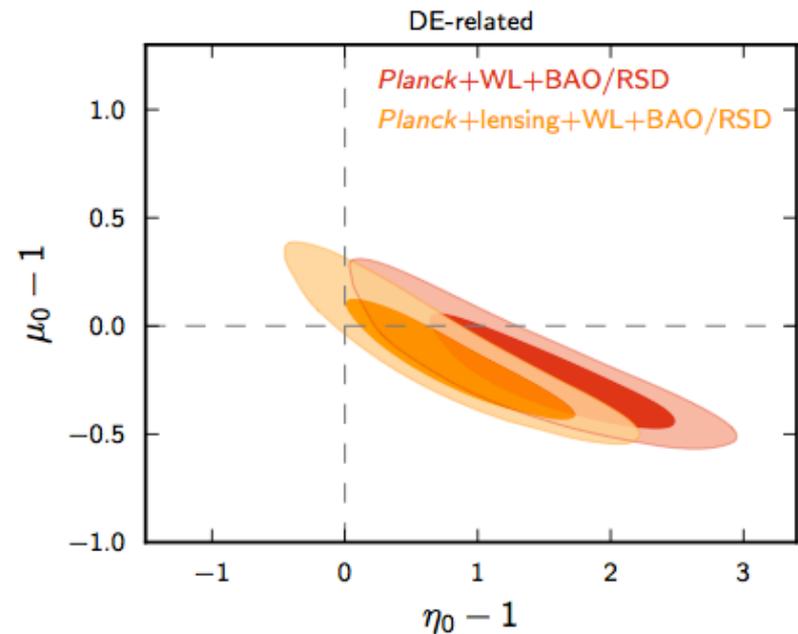
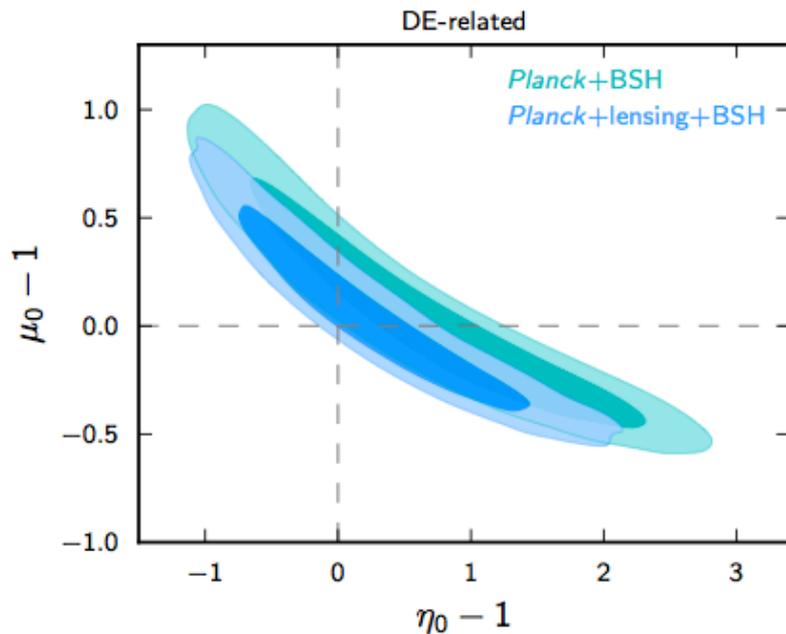
# So, the parameters bottom line:

- Power spectra are highly consistent with the flat  $\Lambda$ CDM model
  - Minor tension with curvature with Planck alone
- The varying analyses themselves are highly consistent within flat  $\Lambda$ CDM

# Dark Energy/Modified Gravity

- Meaningful constraints typically need additional low-redshift data over that needed for our  $\Lambda$ CDM analyses
  - Any modifications that only affect the background are *not* favoured over  $\Lambda$ CDM
  - Modifications affecting the perturbations also can be found that are favoured at 2-3  $\sigma$  depending on the choice of dataset
    - Go away when Planck lensing is used...

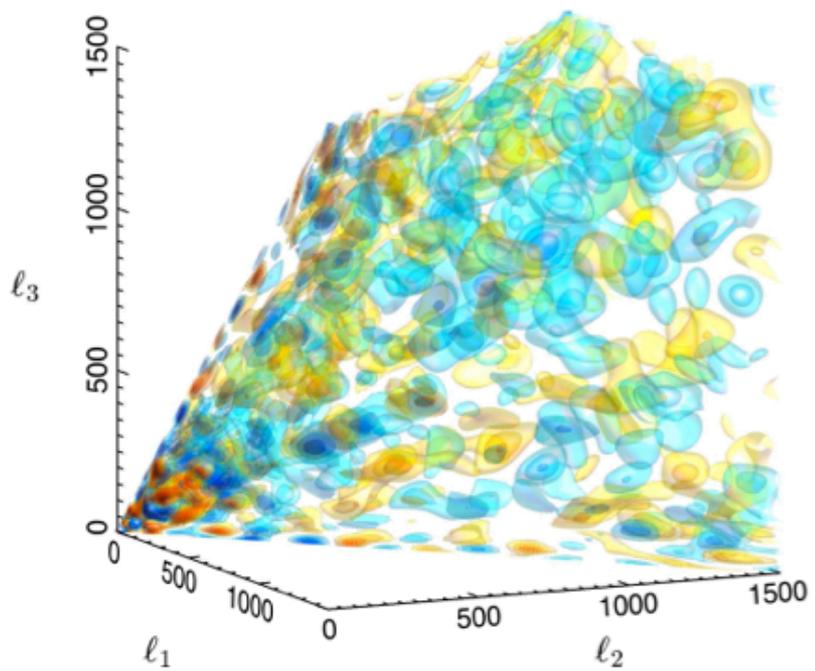
E.g. affecting the Poisson equation and  
no anisotropic stress condition...



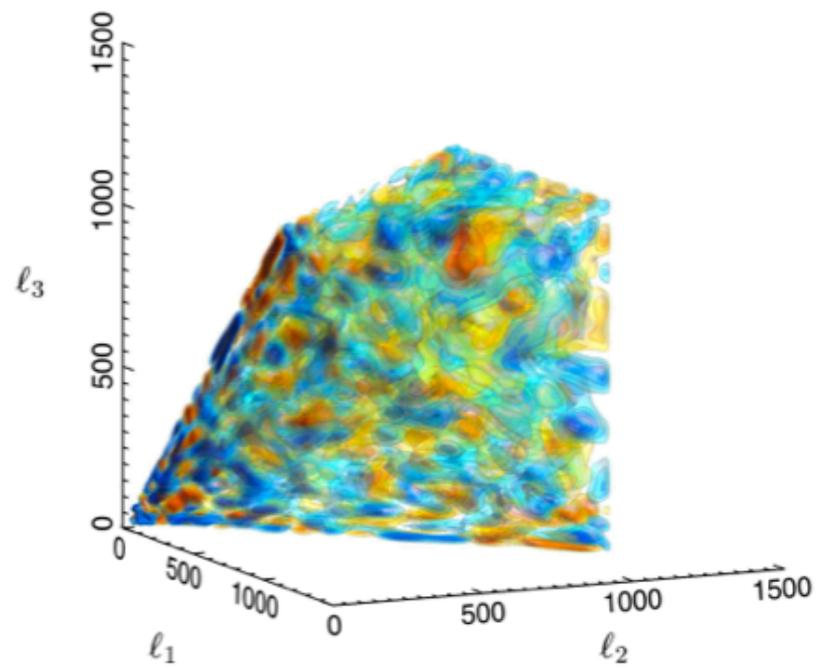
# Non-gaussianity

- Lensing gives a specific, non-gaussian, signature
- One can search more generally for non-gaussianities in the CMB
  - Various estimators, now in 2015 also including E
- No convincing detections made
  - Again consistent with  $\Lambda$ CDM

TTT



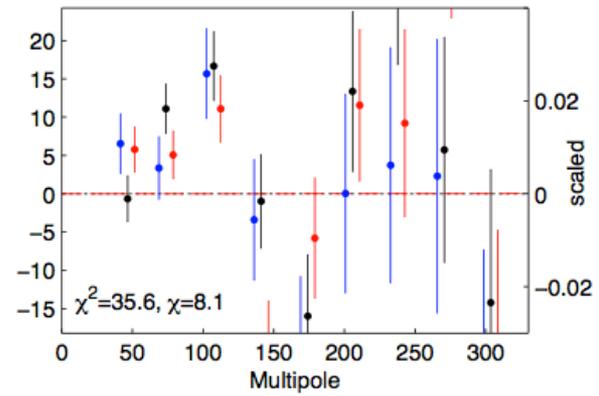
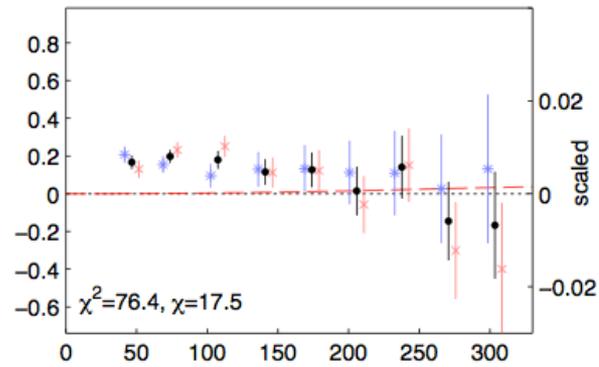
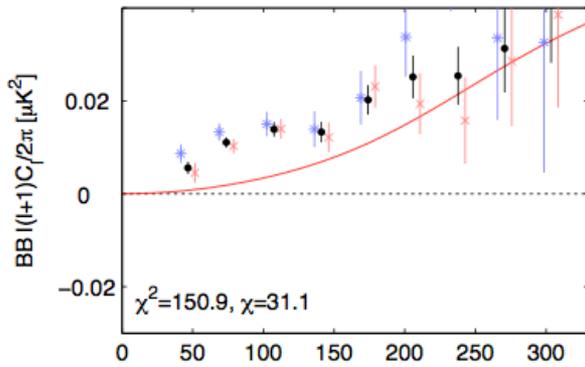
EEE



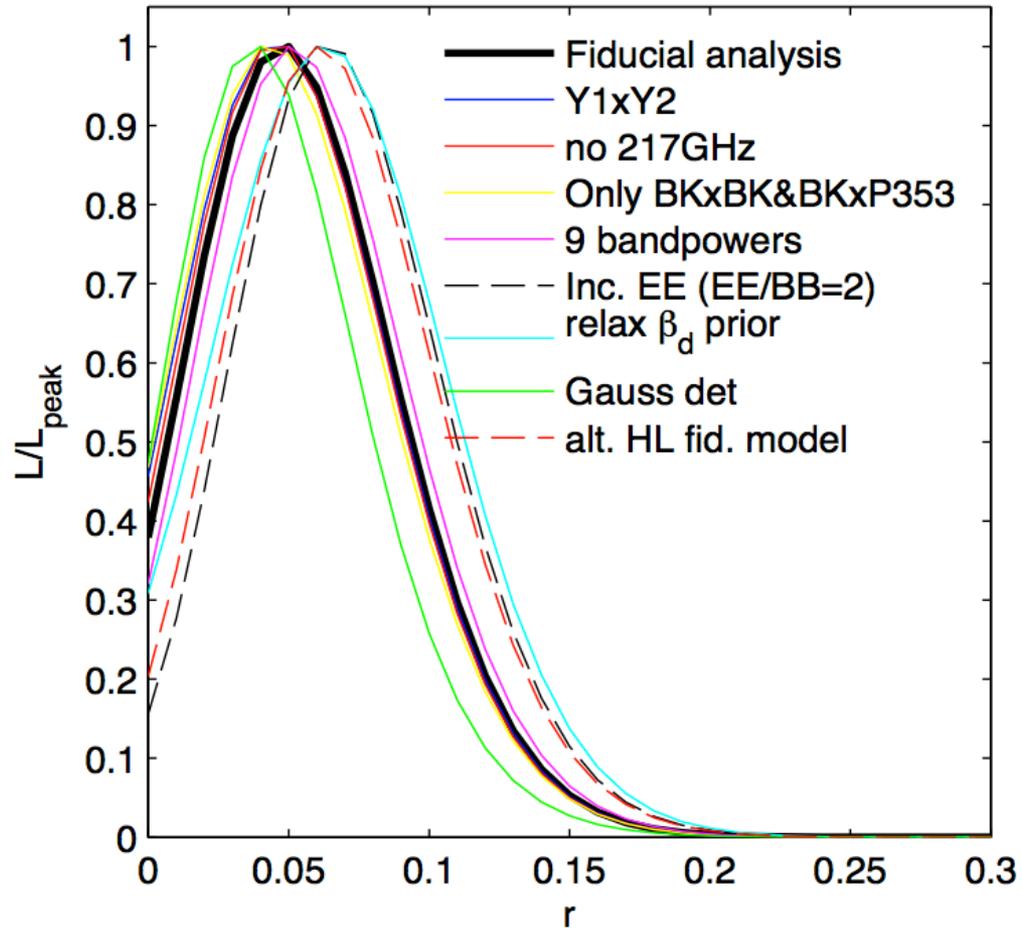
# BICEP2/Keck Array and Planck

- Planck added the 353 GHz polarization maps
  - Useful as a dust tracer
    - Dust has a different emission spectrum to the black-body CMB and is brighter at higher frequencies
  - Multiply by 0.04 to get it down to BK's 150 GHz
    - 353x353 spectrum is still very noisy even now
      - Suggests there is something there though
    - But the BKx353 spectrum is helpful
  - One then constructs a likelihood out of
    - BKxBK, BKx353 and 353x353

# The spectra



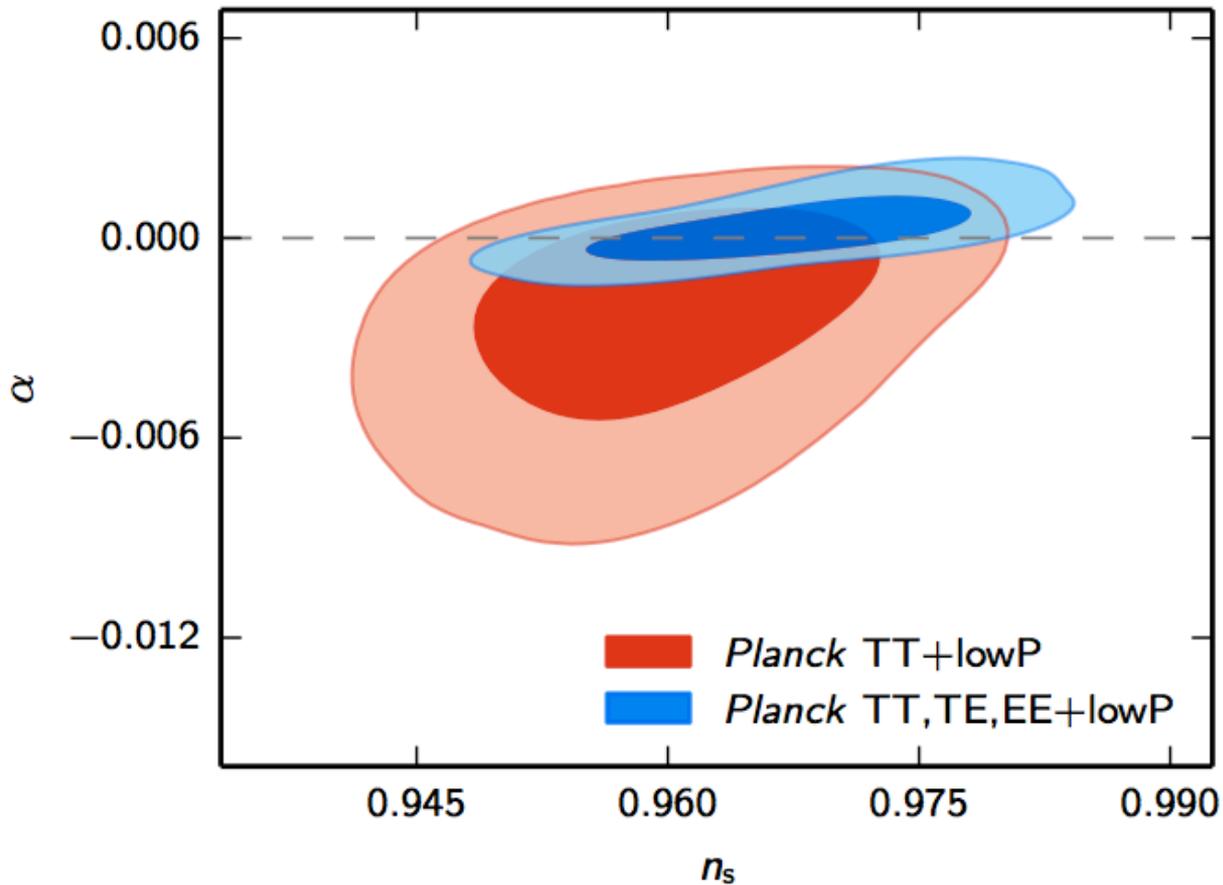
# Illustrative $r$ posteriors...



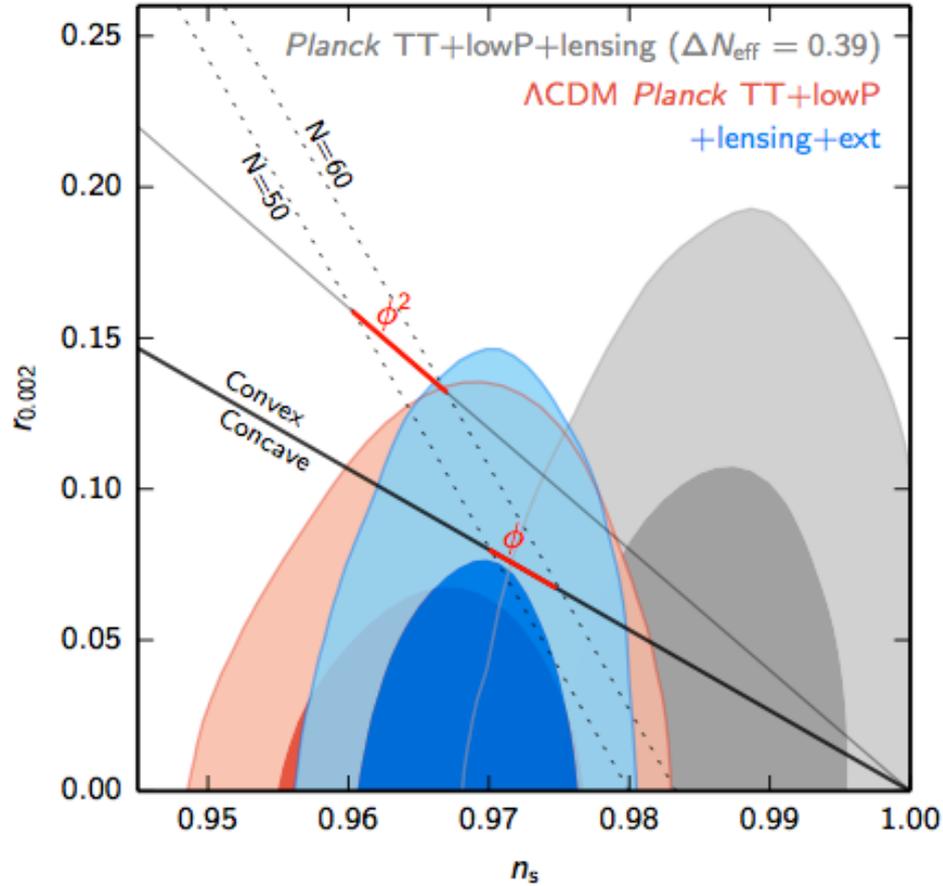
# Inflation

- Adiabaticity of the perturbations
- Implications of  $n_s$  &  $r$  constraints for inflation
- (Reconstruction of the primordial power spectrum)

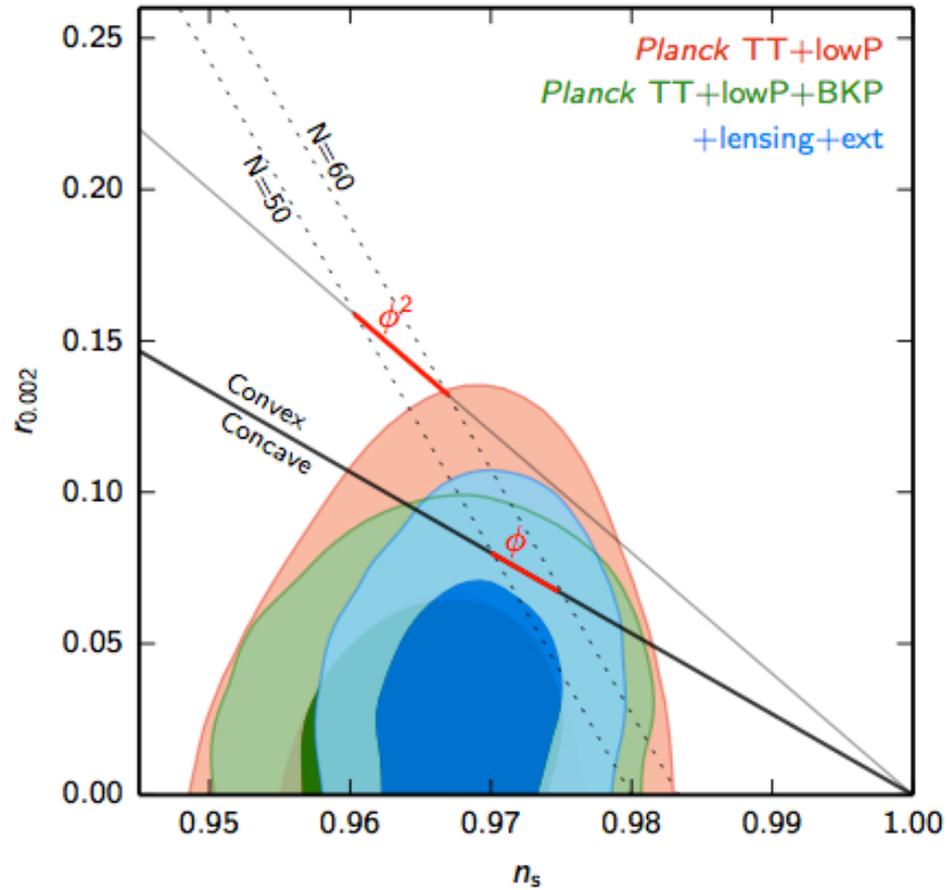
# Perturbations are consistent with being adiabatic...



# Inflation: Planck $n_s$ vs $r$



# Inflation: Planck+BKP $n_s$ vs $r$



# What is still to come from Planck?

- 2015
  - Likelihood paper
  - Likelihood
  - 100, 143 & 217 polarization maps
  - Low-ell paper
- 2016
  - Analysis with Improved Timeline Processing

**The scientific results that we present today are a product of the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada.**



Planck is a project of the European Space Agency, with instruments provided by two scientific Consortia funded by ESA member states (in particular the lead countries: France and Italy) with contributions from NASA (USA), and telescope reflectors provided in a collaboration between ESA and a scientific Consortium led and funded by Denmark.

# What does it mean?

- Flat  $\Lambda$ CDM works *really well*
- Inflation?
  - Results are highly consistent with “*single field with a flat potential*” inflation
    - How generic are such models in string theory?
    - Then, if such inflation is possible, is it *probable*?
  - BKP *doesn't rule out* a significant tensor contribution to the fluctuations

# What is still needed from theory?

- Well-motivated models that make subtle, distinctive and correlated predictions
  - A dip at  $l=25$  *and* {another effect} anybody?
- Alternatives to inflation
- Better understandings of measures on cosmological physics and histories

Thanks!