HIGHER DERIVATIVE TERMS IN M-THEORY REDUCTIONS

based on 1408.5136 and 1412.5073 with T.W. Grimm and M.Weissenbacher

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- Introduction
- M-theory Vacua with 3d N=2 Supersymmetry
- The 3d Effective Theory
- Conclusions

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INTRODUCTION

- Reductions of M-theory to three dimensions present many interesting features.
- In general the vacua include fluxes and warping, which require higher derivative corrections to the 11d action for global consistency.
- These reductions represent the M-theory duals of F-theory reductions to 4d and are important for deriving the F-theory effective action.
- Higher derivative corrections in M-theory can induce corrections to the 3d effective theory and the 4d F-theory lift.

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THE HIGHER DERIVATIVE VACUA

- The 11d action and field equations are given as an expansion in the 11d plank length
- So the solutions we will study exist as an expansion in $\alpha \sim (\kappa_{11})^{\frac{2}{3}} \sim l_p^3$

$$d\hat{s}^{2} = e^{\alpha^{2} Z} (e^{-2\alpha^{2} W^{(2)}} \eta_{\mu\nu} dx^{\mu} dx^{\nu} + 2e^{\alpha^{2} W^{(2)}} g_{m\bar{n}} dy^{m} dy^{\bar{n}})$$
$$g_{m\bar{n}} = g^{(0)}_{m\bar{n}} + \alpha^{2} g^{(2)}_{m\bar{n}}$$
$$\hat{G}_{m\bar{n}r\bar{s}} = \alpha G^{(1)}_{m\bar{n}r\bar{s}} \qquad \hat{G}_{\mu\nu\rho m} = \epsilon_{\mu\nu\rho} \partial_{m} e^{-3\alpha^{2} W^{(2)}}$$

• Where 3d N=2 supersymmetry at lowest order in α implies [Becker, Becker] $dJ^{(0)} = d\Omega^{(0)} = 0 \ G^{(1)} \wedge J^{(0)} = 0 \ G^{(1)} \in H^{2,2}(Y_4)$

$\mathcal{O}(\alpha^2)$ Constraints on the Vacua

 And at second order in α the solution is constrained by the warp factor equation
 d[†]dW⁽²⁾ + G⁽¹⁾ ∧ G⁽¹⁾ + X₈ = 0

And the internal space part of the metric field

equation [Becker, Becker] [Lu, Pope, Stelle, Townsend] $R_{m\bar{n}}^{(2)} = \partial_m^{(0)} \bar{\partial}_{\bar{n}}^{(0)} Z \qquad Z = *^{(0)} (J^{(0)} \wedge c_3^{(0)})$ $c_3^{(0)} = i \operatorname{Tr}(\mathcal{R}^{(0)} \wedge \mathcal{R}^{(0)} \wedge \mathcal{R}^{(0)})$

• This is solved by [Grimm, TP, Weissenbacher]

$$c_3^{(0)} = H c_3^{(0)} + i \partial^{(0)} \bar{\partial}^{(0)} F$$
$$g_{m\bar{n}}^{(2)} = \nabla_m^{(0)} \bar{\nabla}_{\bar{n}}^{(0)} *^{(0)} (F \wedge J^{(0)} \wedge J^{(0)})$$

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PERTURBATIONS FOR KÄHLER MODULI

• We make an ansatz for the perturbations that correspond to Kähler moduli where

 $\delta g_{m\bar{n}}^{(0)} = i\delta v^i \omega_{im\bar{n}}^{(0)} \qquad \delta \hat{G} = F^i \wedge \omega_i^{(0)}$ $F^i = dA^i \qquad d *^{(0)} \omega_i^{(0)} = d\omega_i^{(0)} = 0$

• The metric shift induces a shift of the higher derivative parts of the ansatz

$$\delta Z = i \delta v^i Z^{\bar{n}m} \omega_{im\bar{n}}^{(0)} + \dots$$

• And induces a shift of the warping

$$W^{(2)}| \to W^{(2)}| + \partial_i W^{(2)}|\delta v^i + \dots$$

(See Martucci's Talk)

REDUCING THE 11D ACTION

• Next we reduce the action on the background described

$$S = S^{(0)} + \alpha^2 S^{(2)}_{\hat{R}^4} + \alpha^2 S^{(2)}_{\hat{G}^2 \hat{R}^3} + \alpha^2 S^{(2)}_{(\hat{\nabla} \hat{G})^2 \hat{R}^2}$$

 $+ \mathcal{O}(\hat{G}^3 \alpha^2) + \mathcal{O}(\alpha^3)$ [Liu, Minasian]

- We derive the effective action to second order in 3d derivatives, α , δv^i and A^i . This is sufficient to determine the higher derivative contributions to 3d kinetic terms and mass terms.
- This results in an effective action given in terms of $Z_{m\bar{n}r\bar{s}} = (\epsilon_8^{(0)} \epsilon_8^{(0)} R^{(0)3})_{m\bar{n}r\bar{s}}$

 $Z_{m\bar{m}} = i Z_{m\bar{m}n}{}^n = (*^{(0)} c_3^{(0)})_{m\bar{m}} \quad Z_m{}^m = Z$

KINETIC TERMS IN THE EFFECTIVE ACTION

• Substituting into the 11d action and integrating we find

$$S_{\rm kin} = \frac{1}{2\kappa_{11}} \int_{\mathcal{M}_3} \left[R * 1 - (G_{ij}^T + \mathcal{V}_T^{-2} K_i^T K_j^T) Dv^i \wedge *Dv^j - \mathcal{V}_T^2 G_{ij}^T F^i \wedge *F^j \right]$$

• Where

$$G_{ij}^{T} \sim \frac{1}{\mathcal{V}_{0}} \int_{Y_{4}} \left[e^{3\alpha^{2}W^{(2)} + \alpha^{2}Z} \omega_{im\bar{n}}^{(0)} \omega_{j}^{(0)\bar{n}m} + \alpha^{2}Z_{m\bar{n}r\bar{s}} \omega_{j}^{(0)\bar{n}m} \omega_{i}^{(0)\bar{s}r} \right] *^{(0)} 1$$
$$K_{i}^{T} = \int_{Y_{4}} \left[e^{3\alpha^{2}W^{(2)} + \alpha^{2}Z} \omega_{im}^{(0)m} - Z_{m\bar{n}} \omega_{i}^{(0)\bar{n}m} \right] *^{(0)} 1$$
$$\mathcal{V}_{T} = \int_{Y_{4}} e^{3\alpha^{2}W^{(2)} + \alpha^{2}Z} *^{(0)} 1$$

SCALAR MASS TERMS

• The additional mass terms for the scalar fluctuations cancel

$$\int \hat{t}_8 \hat{t}_8 \hat{R}^4 \hat{*}1|_{\text{mass}} = -\int_{\mathcal{M}_3} \delta v^i \delta v^j *_3 1 \int_{Y_4} \nabla_r \nabla^r Z \omega_{im\bar{n}}^{(0)} \omega_j^{(0)\bar{n}m} *_8 1$$
$$\int \hat{R} \hat{*}1|_{\text{mass}} = \alpha^2 \int_{\mathcal{M}_3} \delta v^i \delta v^j *_3 1 \int_{Y_4} \nabla_r \nabla^r Z \omega_{im\bar{n}}^{(0)} \omega_j^{(0)\bar{n}m} *_8 1$$

• So that the only potential terms are

$$S_{\text{pot}} = \frac{\alpha^2}{4\kappa_{11}} \int_{\mathcal{M}_3} *_3 1 \int_{Y_4} (G^{(1)} \wedge *'G^{(1)} - G^{(1)} \wedge G^{(1)})$$

• These come with the associated Chern-Simons terms for the vectors

$$S_{CS} = \alpha \int_{\mathcal{M}_3} A^i \wedge F^j \int_{Y_4} G^{(1)} \wedge \omega_i \wedge \omega_j$$

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CONCLUSIONS

- Effective theories including higher derivative corrections for 3d moduli may be derived in the way we have described.
- This process involves corrections to the vacua, ansatz and action parameterized by the 11d Planck length.
- These computations required significant use of computer algebra packages such as xAct.
- The masses of the 3d fluctuations are not altered.
- All corrections may be written in terms of $Z_{m\bar{n}r\bar{s}} = (\epsilon_8^{(0)} \epsilon_8^{(0)} R^{(0)3})_{m\bar{n}r\bar{s}}$ and the warp factor.

THANK YOU