

# Entanglement Entropy in Shock Wave Collisions

Based on work with Daniel Grumiller, Stefan Stricker 1506.02658 (JHEP);  
& Wilke van der Schee (15XX.XXXXX)

Christian Ecker

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DOKTORATSKOLLEG PI

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*Particles and Interactions*

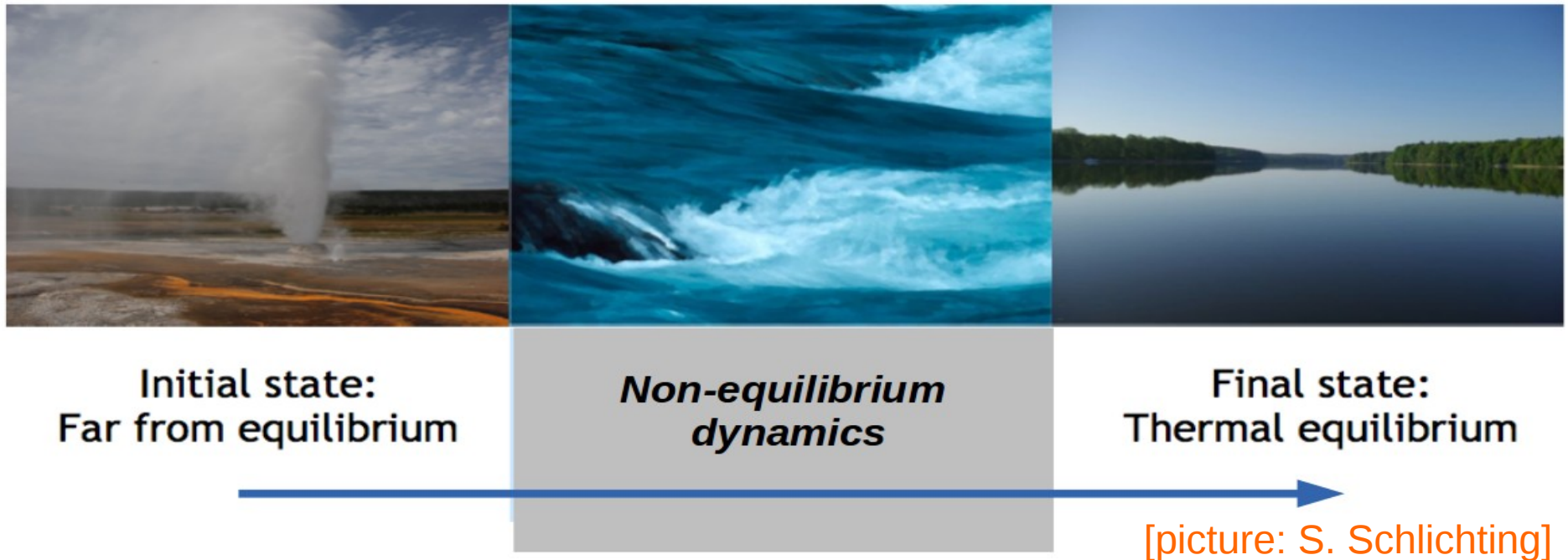


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# Motivation

## Central question:

How does a **strongly coupled** quantum system **evolve** from a **far-from equilibrium** state to **equilibrium**?



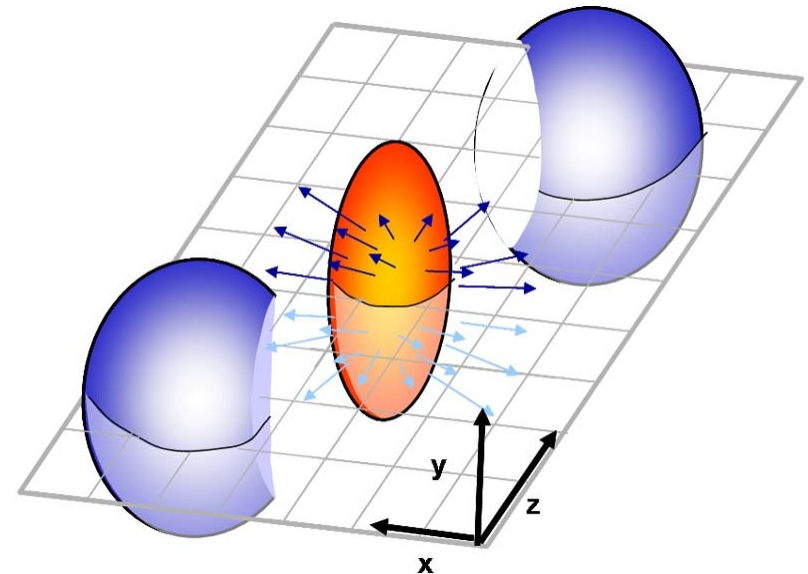
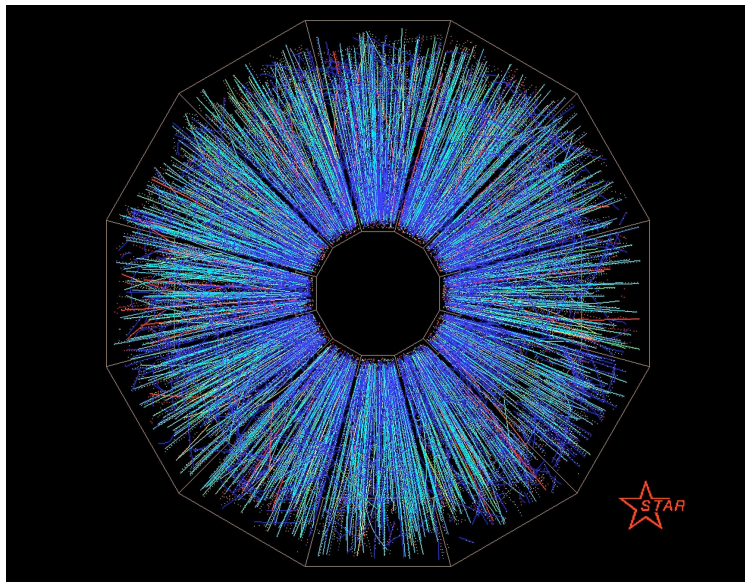
## Challenges:

- Due to **strong coupling** **perturbative** methods are **not applicable**.
- **Far-from equilibrium** we are **outside the regime** of **linear response theory**.

# Quark-gluon plasma in heavy ion collisions

[see talk by K. Landsteiner]

Quark-gluon plasma (QGP) is a **deconfined phase of quarks and gluons**. It is produced in **heavy ion collision (HIC)** experiments at **RHIC** and **LHC**.



## Why AdS/CFT?

- The QGP produced in HIC's behaves like a **strongly coupled liquid** rather than a **weakly coupled gas**.
- Due to strong coupling we **can't use perturbation theory**.
- To study real time dynamics we **can't use lattice QCD**.

# AdS/CFT correspondence

[see talks by K. Landsteiner,  
D.R. Fernandez]

AdS/CFT correspondence: [Maldacena 97]

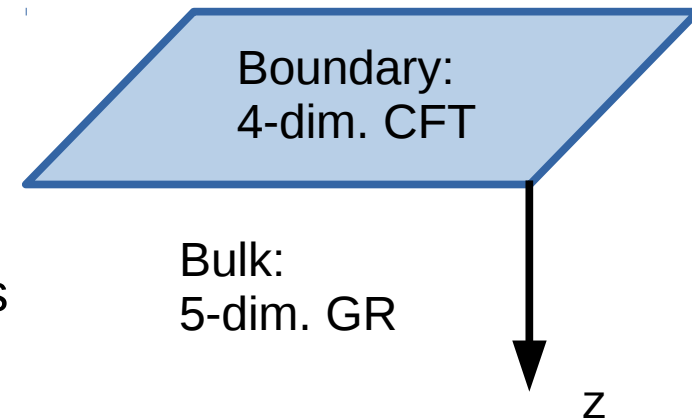
**Type IIB string theory** on  $\text{AdS}_5 \times S^5$  is equivalent to  $\mathcal{N}=4$  super symmetric  $\text{SU}(N_c)$  **Yang-Mills theory** in 4D.

Supergravity limit:

**Strongly coupled large  $N_c$   $\mathcal{N}=4$   $\text{SU}(N_c)$  SYM theory** is equivalent to **classical (super)gravity** on  $\text{AdS}_5$ .

Strategy:

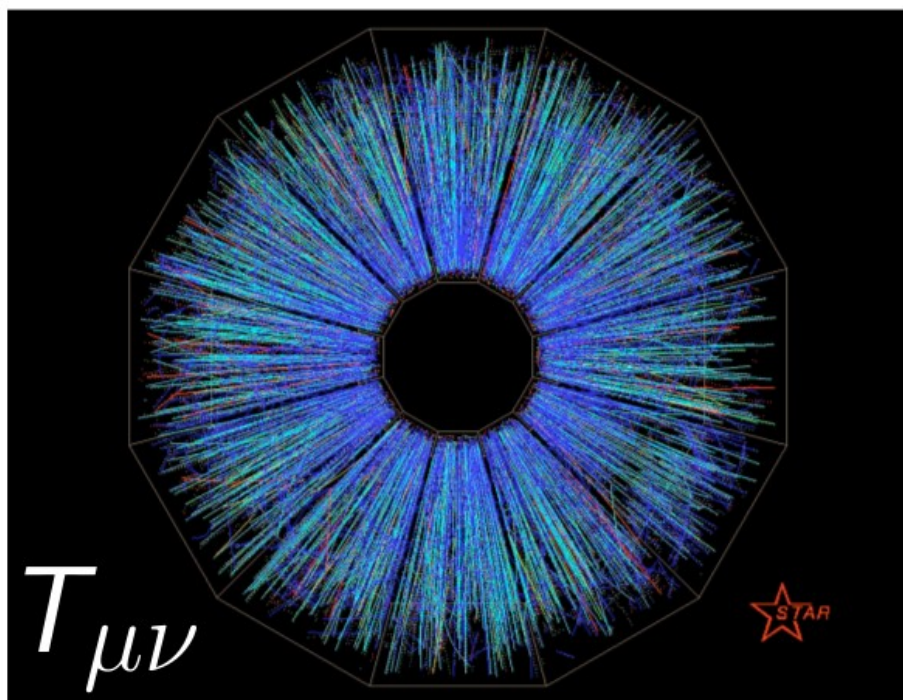
- Use  $\mathcal{N}=4$  SYM as **toymodel** for **QCD**.
- Build a **gravity model** dual to HICs, like colliding **gravitational shock waves**.
- Switch on the computer and solve the 5-dim. gravity problem **numerically**.
- Use the **holographic dictionary** to compute **observables in the 4 dim. field theory** from the gravity result. [Skenderis 01]





# Holographic thermalization

Thermalization = Black hole formation



- **Holographic dictionary** translates thermodynamic **properties (S,T,M) of black holes** to the corresponding **properties of the gauge theory**.
- Computing **black hole formation on AdS** in general requires methods from **numerical relativity**.
- The **observable** we use to **study thermalization** is **entanglement entropy**.

# Entanglement entropy

**Divide** the system into **two parts** A,B.  
The total Hilbert space factorizes:

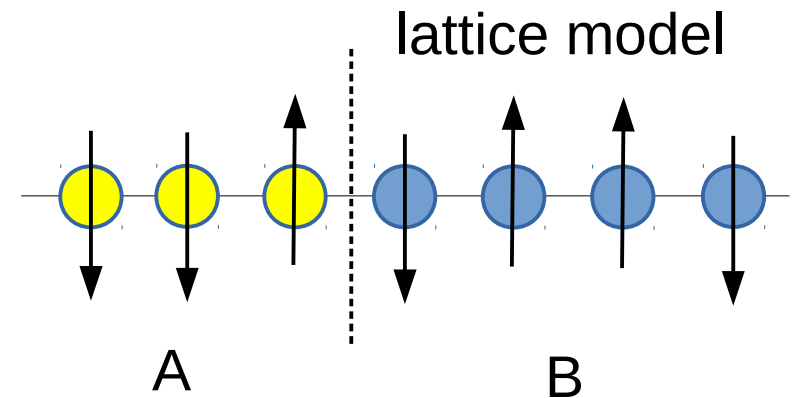
$$\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$$

The **reduced density matrix** of A is  
obtained by the trace over  $\mathcal{H}_B$

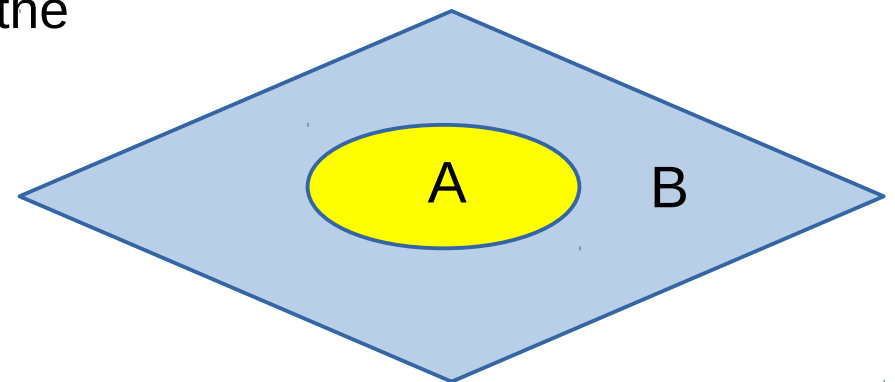
$$\rho_A = \text{Tr}_B \rho$$

**Entanglement entropy** is defined as the  
**von Neumann entropy** of  $\rho_A$ :

$$S_A = -\text{Tr}_A \rho_A \log \rho_A$$



quantum field theory



# Entanglement entropy in a two quantum bit system

Consider a quantum system of two spin 1/2 dof's.  
 Observer Alice has only access to one spin and Bob to the other spin.



A **product state (not entangled)** in a two spin 1/2 system:

$$S_A = 0$$

$$|\psi\rangle = \frac{1}{2} (|\uparrow_A\rangle + |\downarrow_A\rangle) \otimes (|\uparrow_B\rangle + |\downarrow_B\rangle)$$

Alice      Bob

A (maximally) **entangled state** in a two spin 1/2 system:

$$S_A = \log 2$$

$$|\psi\rangle = \frac{1}{\sqrt{2}} (|\uparrow_A\rangle \otimes |\downarrow_B\rangle - |\downarrow_A\rangle \otimes |\uparrow_B\rangle)$$

Alice      Bob

Entanglement entropy is a **measure** for how much a given quantum state is **entangled**.

# Entanglement entropy in quantum field theories

The Basic Method to compute entanglement entropy in quantum field theories is the **replica method**.

Involves path integrals over n-sheeted Riemann surfaces ~ it's **complicated!**

With the **replica method** one gets **analytic results** for **1+1 dim. CFTs**. [Holzhey-Larsen-Wilczek 94]

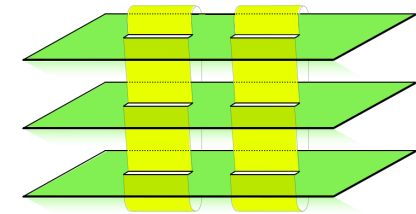
One finds **universal scaling** with interval size:

$$S_A = \frac{c}{3} \log \frac{L}{a} + \text{finite}$$

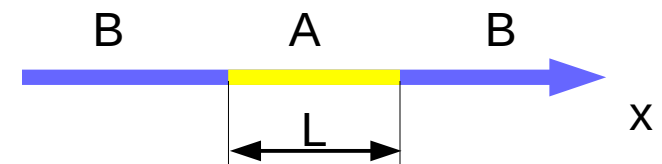
central charge of the CFT
UV cut off

**Message:** Computing entanglement entropy in interacting QFTs is complicated and analytically only possible in 1+1 dim. CFTs.

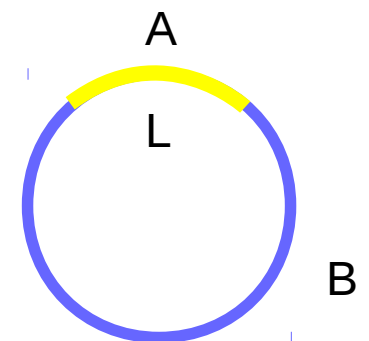
**AdS/CFT** provides a **simpler method** that works also in **higher dimensions**.



3-sheeted Riemann surface



1+1 dim. CFTs

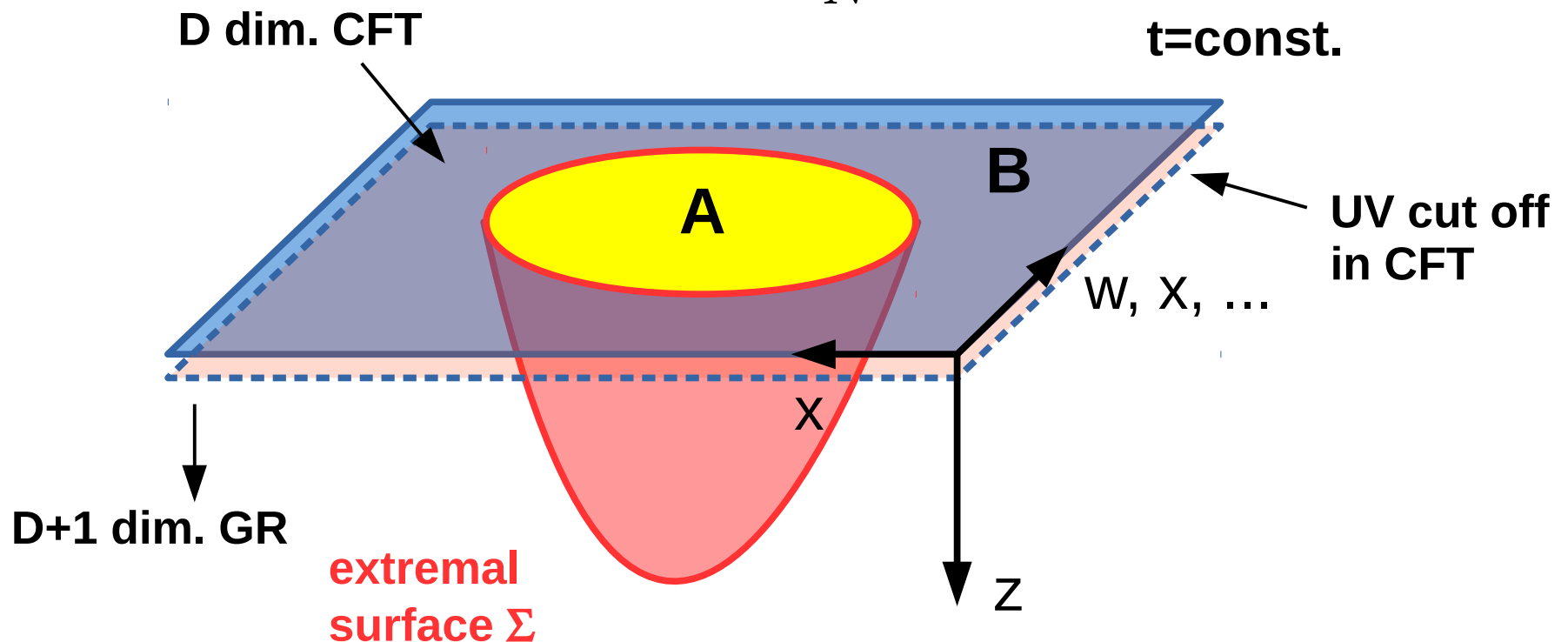




# Holographic entanglement entropy

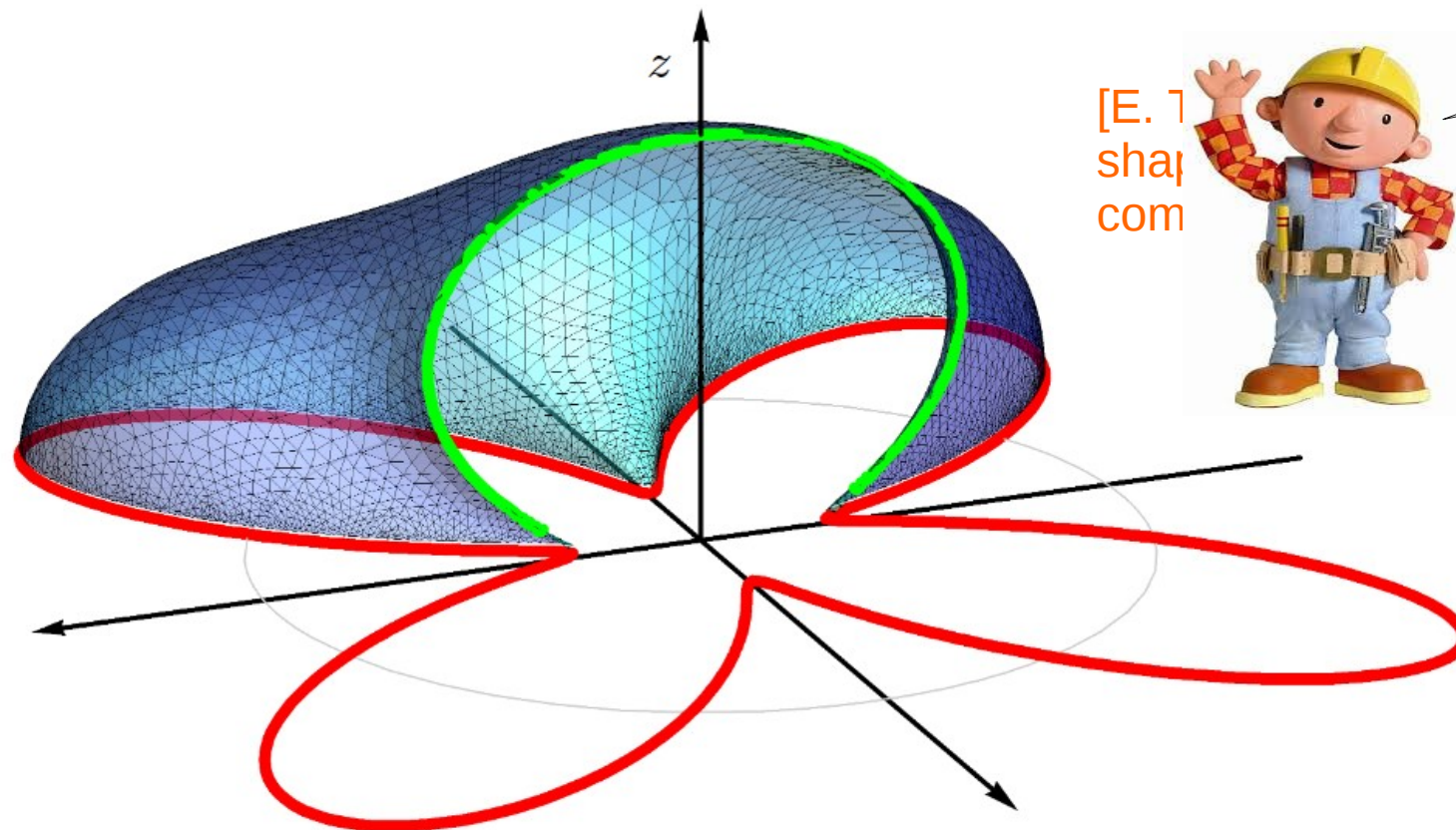
Within **AdS/CFT** entanglement entropy can be computed from the **area of minimal (extremal) surfaces** in the gravity theory.

$$S_A = \frac{\text{Area}(\Sigma)}{4G_N} \quad [\text{Ryu-Takayanagi 06, Hubeny-Rangamani-Takayanagi 07}]$$



# Holographic entanglement entropy

- In practice computing extremal co-dim. 2 hyper-surfaces is numerically involved.
- Can we somehow simplify our lives?



[E. 1  
shaj  
com

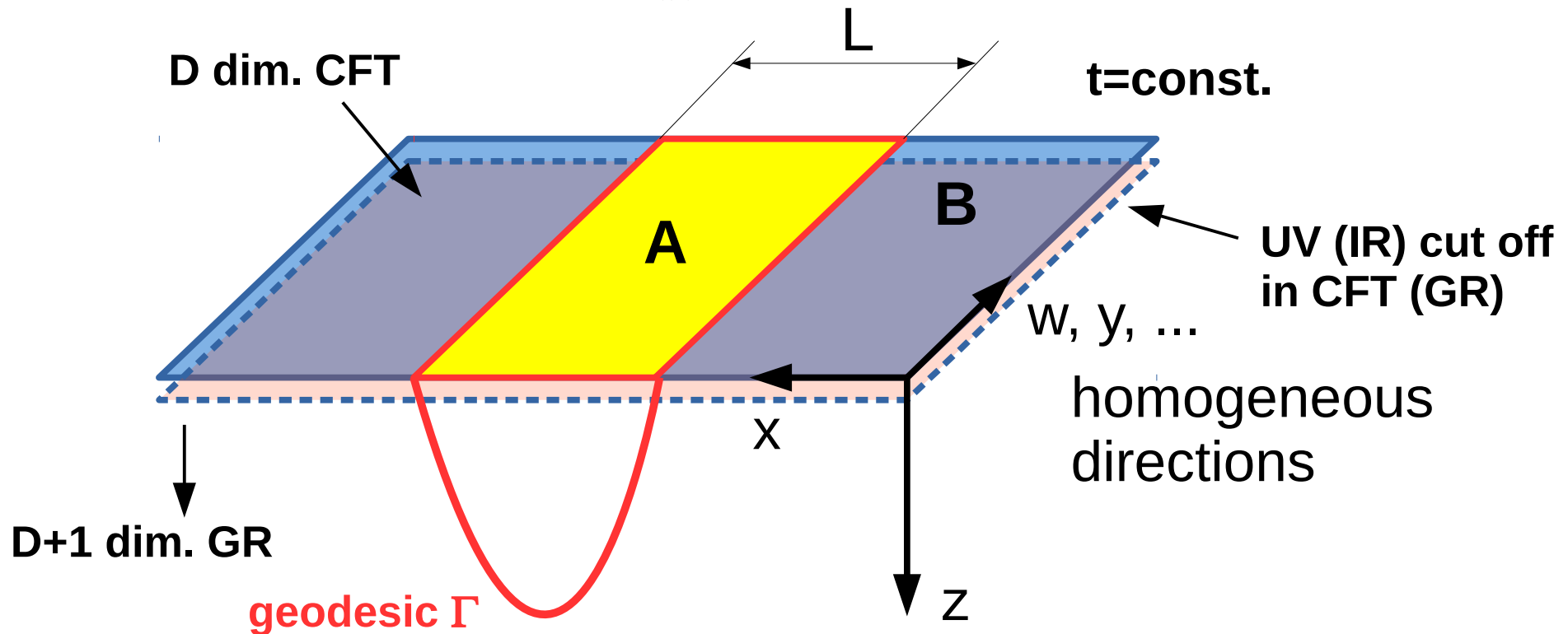


Yes we can!  
nal surface for a star  
region (red) in AdS4  
rface Evolver]

# Entanglement entropy from geodesics

Consider a **stripe region of infinite extent** in homogeneous directions of the geometry. The **entanglement entropy** is prop. to the **geodesics length** in an **auxiliary spacetime**.

$$S_A = \text{const.} \frac{\text{Length}(\Gamma)}{4G_N} \quad \tilde{g}_{\mu\nu} = \Omega(z, t, x)^2 g_{\mu\nu}$$

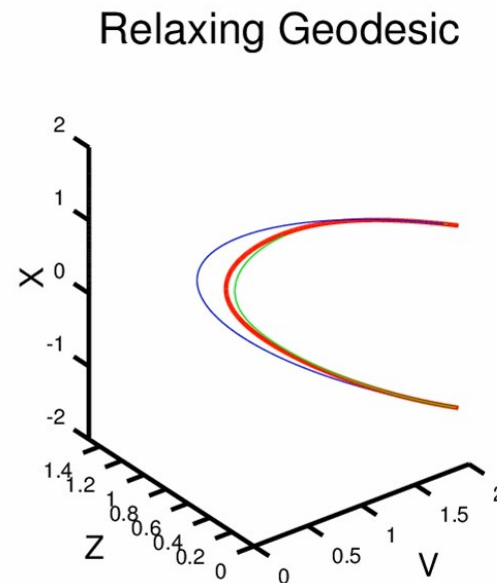
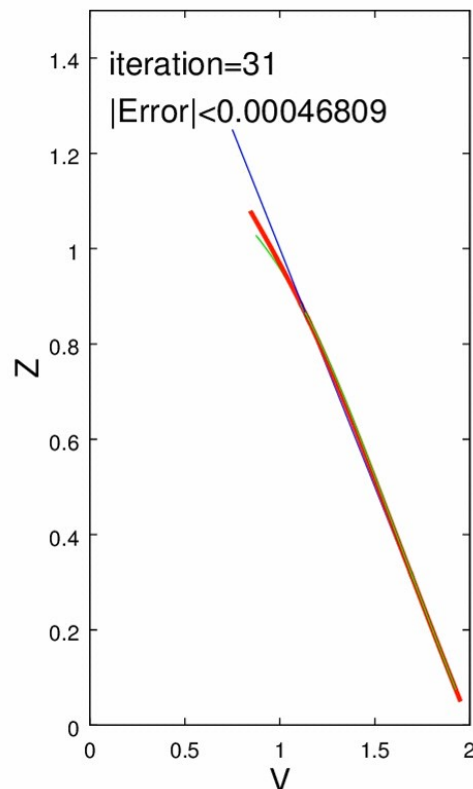


# Numerics: relax, don't shoot!

Geodesic equation as two point boundary value problem.

$$\ddot{X}^\mu(\tau) + \Gamma_{\alpha\beta}^\mu \dot{X}^\alpha(\tau) \dot{X}^\beta(\tau) = 0$$

$$BCs: (V(\pm 1), Z(\pm 1), X(\pm 1)) = (t_0, 0, L/2)$$



- There are two **standard numerical methods** for solving two point boundary value problems.

[see [Numerical Recipes](#)]



**Shooting:**

Very **sensitive to initialization** on **asymptotic AdS** spacetimes.



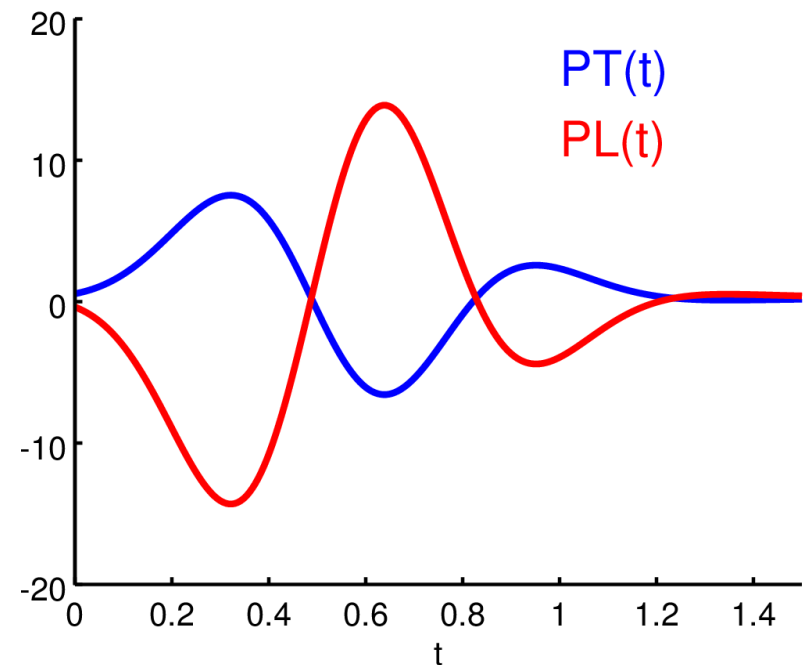
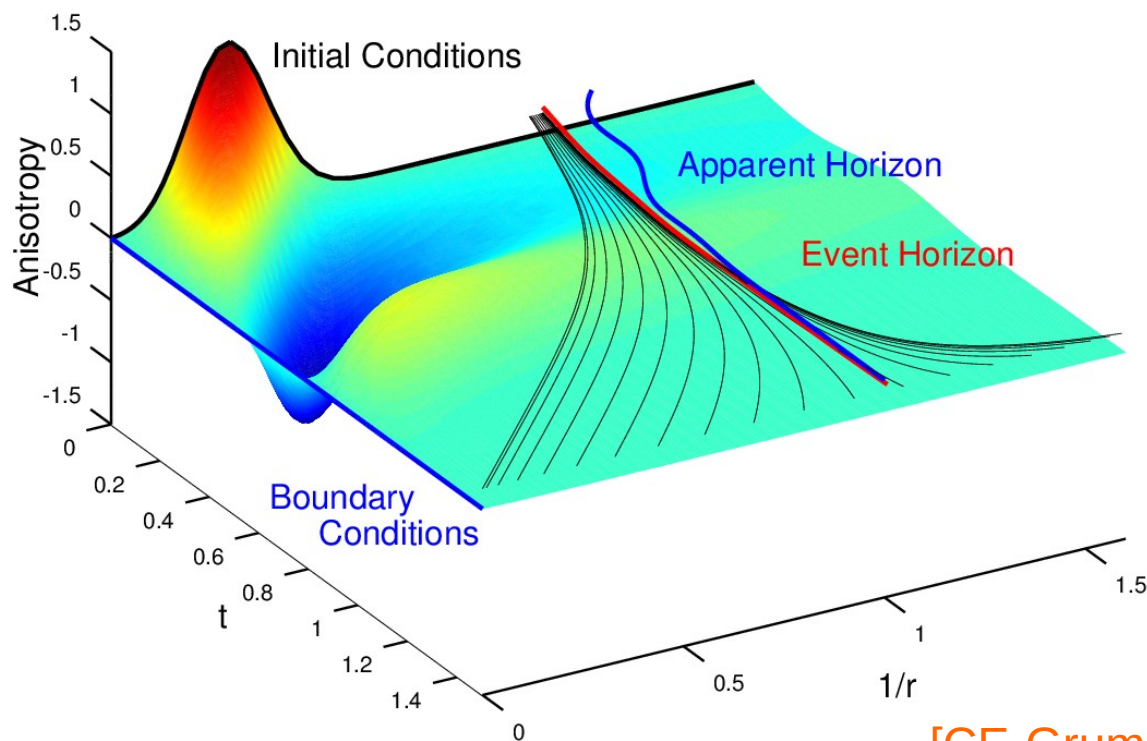
**Relaxation:**

**Converges very fast** if **good initial geodesic** is provided.

# Isotropization of homogeneous plasma

An initially highly anisotropic (N=4 SYM) plasma relaxates to its isotropic equilibrium state. [Chesler-Yaffe 09]

The dual gravity model describes the formation of a black brane in an anisotropic AdS<sub>5</sub> geometry.

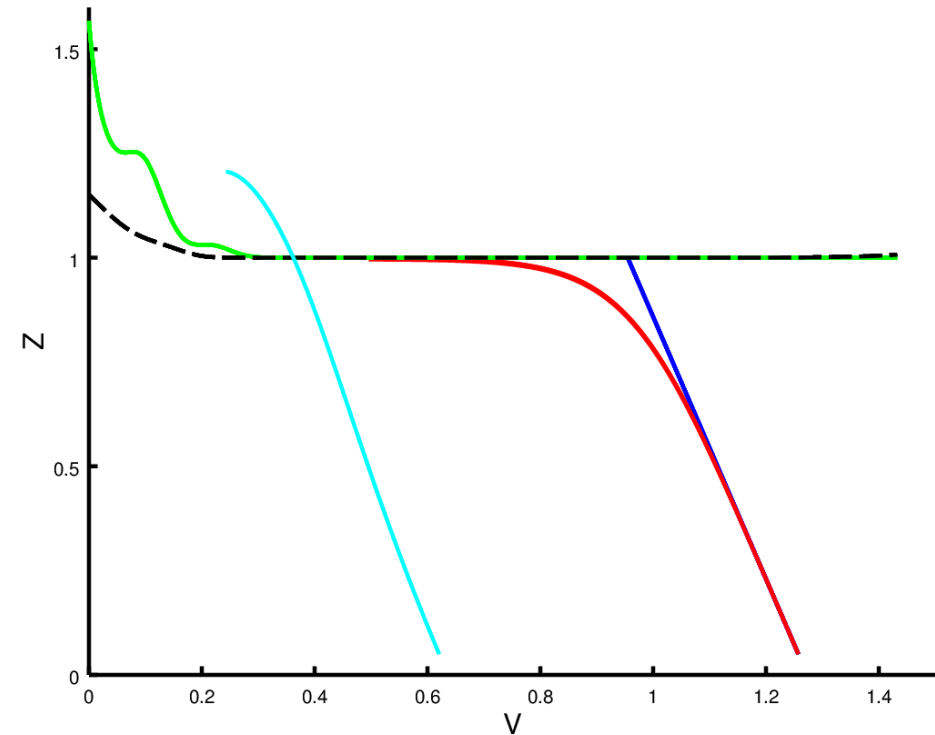
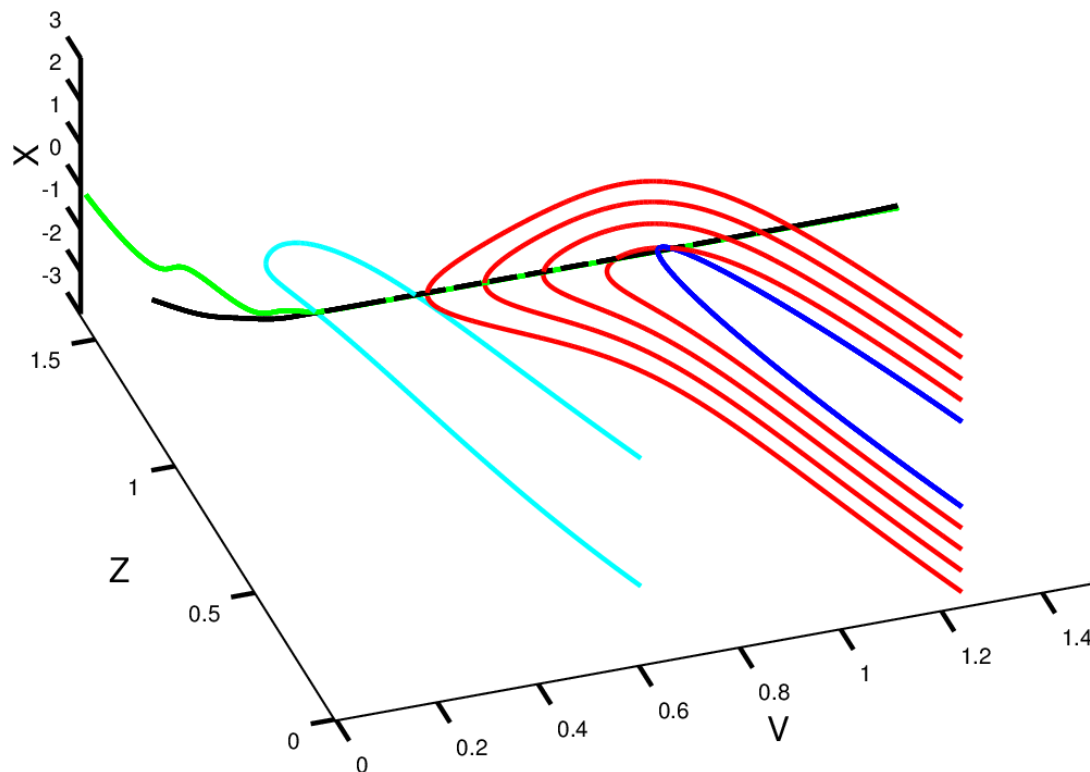


[CE-Grumiller-Stricker 15]



# Geodesics in anisotropic $\text{AdS}_5$ black brane background

- Far-from equilibrium geodesics can go beyond the horizon.
- Near equilibrium geodesics stay outside the horizon.



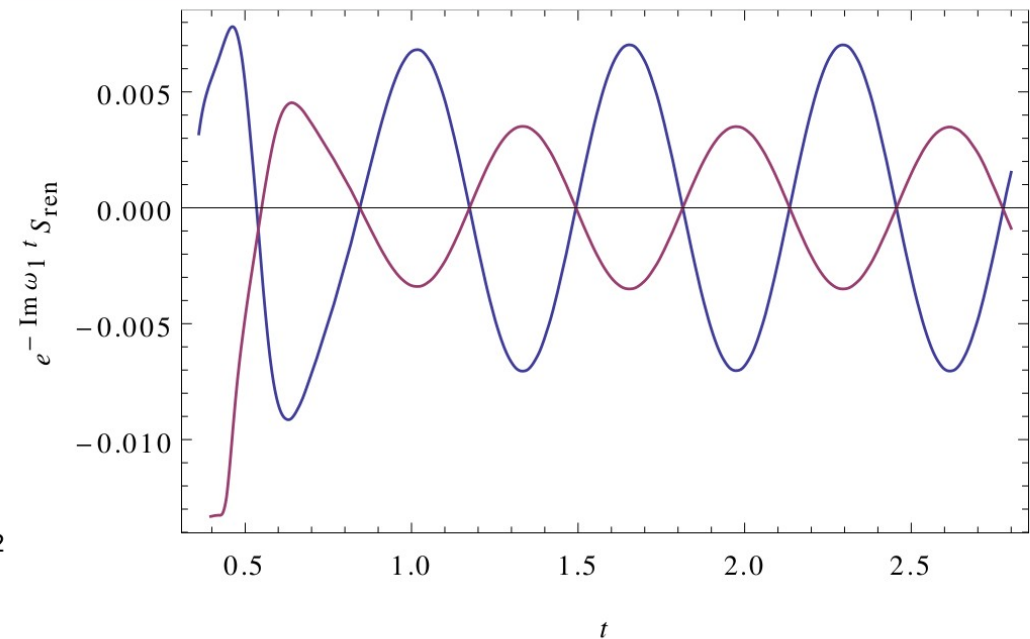
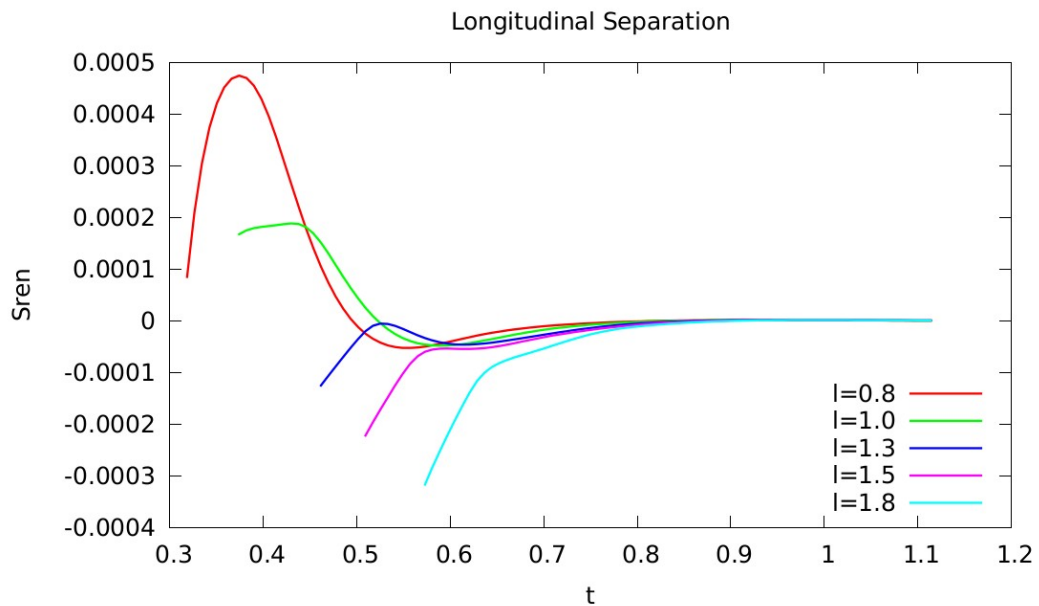
[CE-Grumiller-Stricker 15]

# Quasinormal ringdown of entanglement entropy

[see also talk by W. Sybesma]

The **late time dynamics** of EE is captured by a **single (complex) number**:

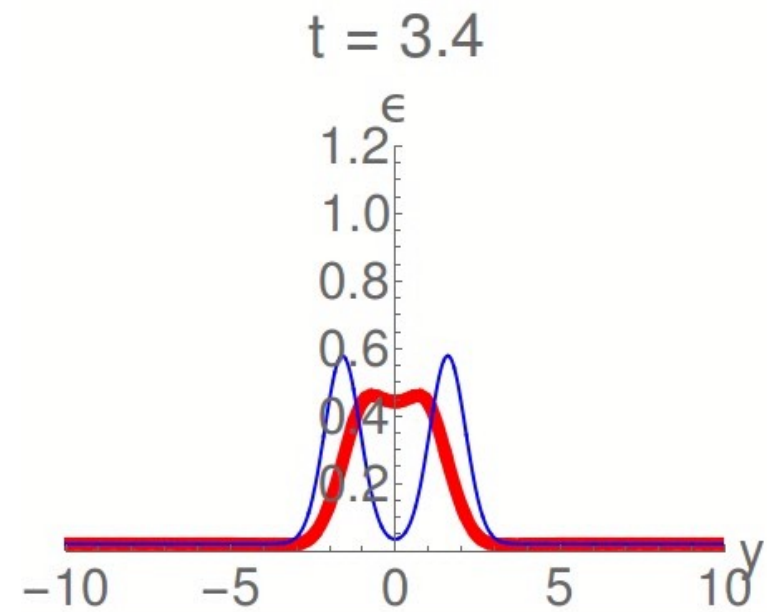
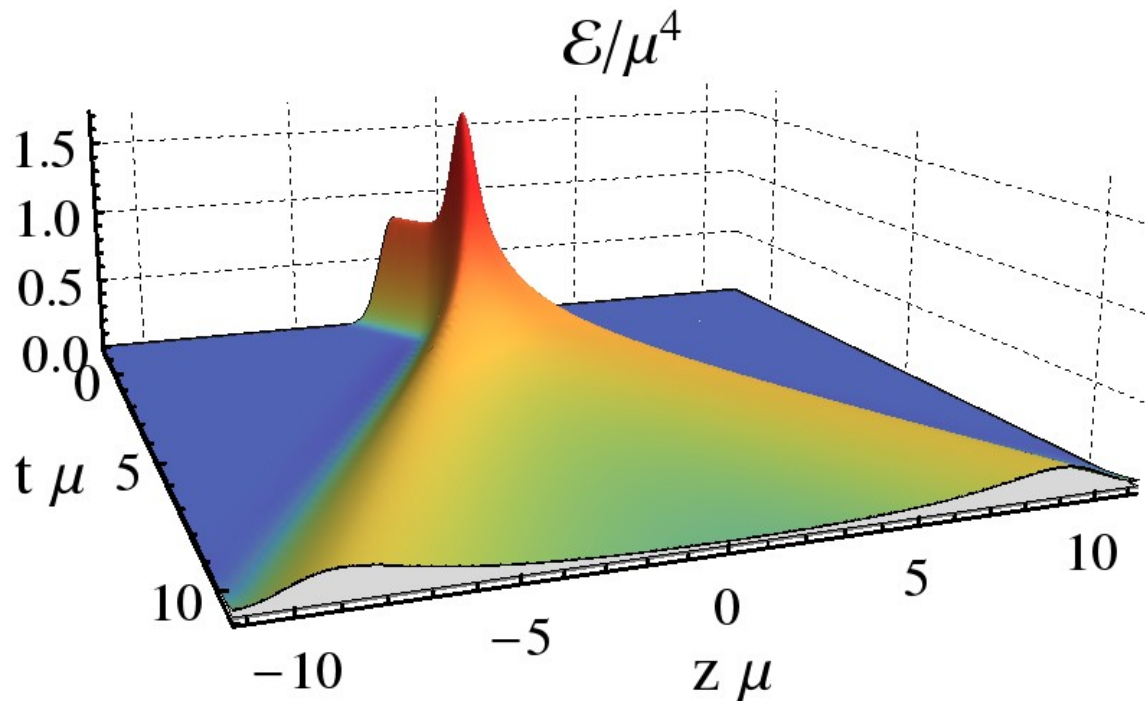
$$\frac{\omega_1}{\pi T} = \pm 3.1119452 - 2.746676i$$



[CE-Grumiller-Stricker 15]

# Holographic shock wave collisions

HIC is modeled by **two colliding sheets of energy** with **infinite extend in transverse direction** and **Gaussian profile in beam direction**. [Chesler-Yaffe 10]



Solving the numerical GR problem is tricky: [Chesler-Yaffe 13]

- **Characteristic formulation** decouples EE to **nested system of ODEs**
- ODEs solved by **spectral method** [see textbooks by Boyd, Trefethen]
- **Time evolution** with simple **Adams-Bashforth** algorithm

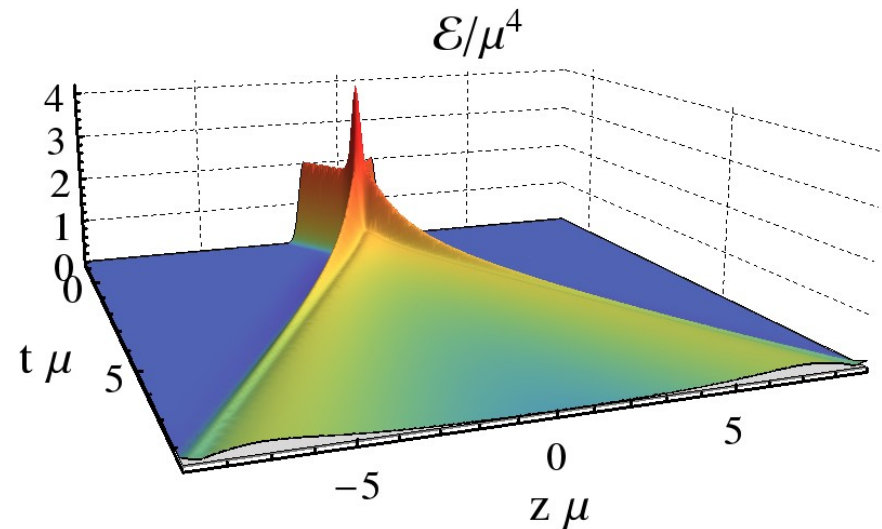
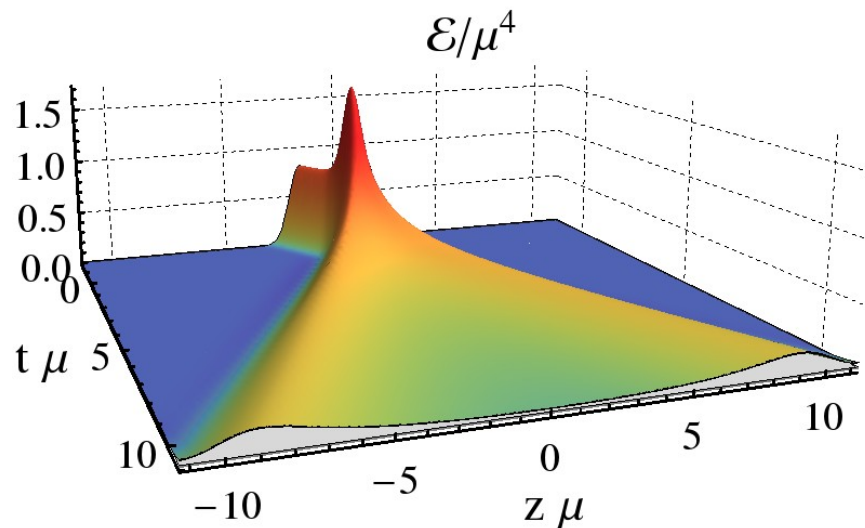
# Wide vs. narrow shocks

Two qualitatively different dynamical regimes

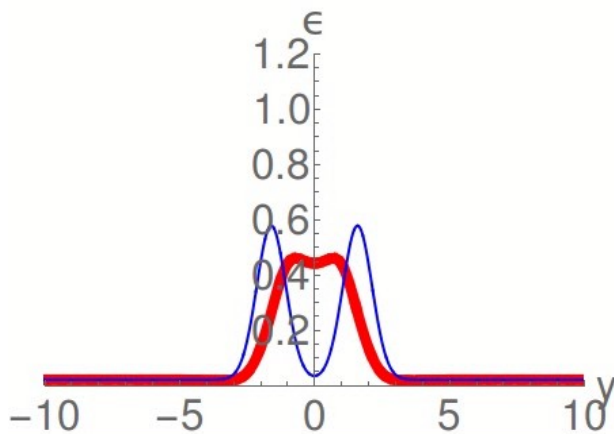
[Solana-Heller-Mateos-van der Schee 12]

- **Wide shocks (~RHIC): full stopping**

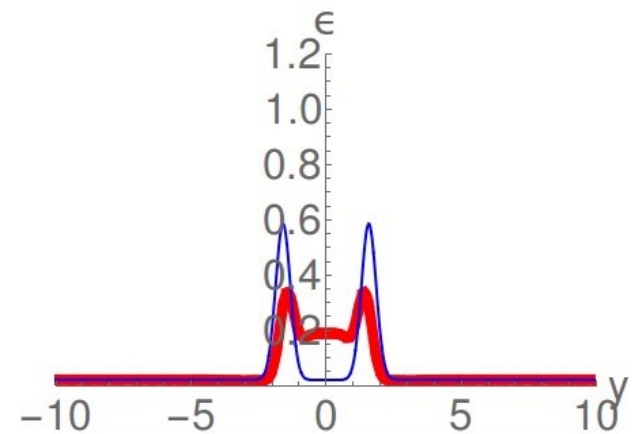
- **Narrow shocks (~LHC): transparency**



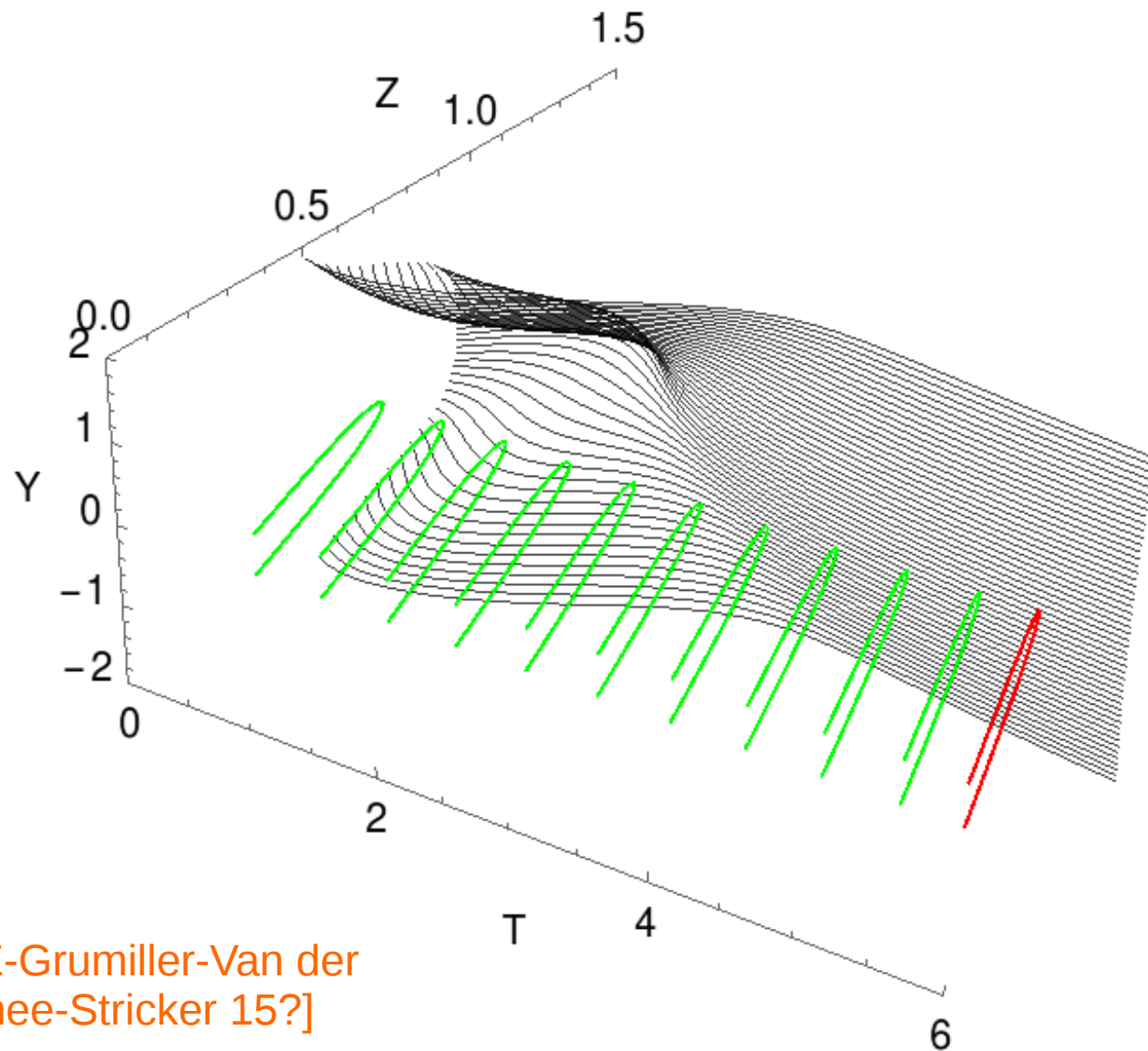
$t = 3.4$



$t = 3.4$



# Geodesics and apparent horizon



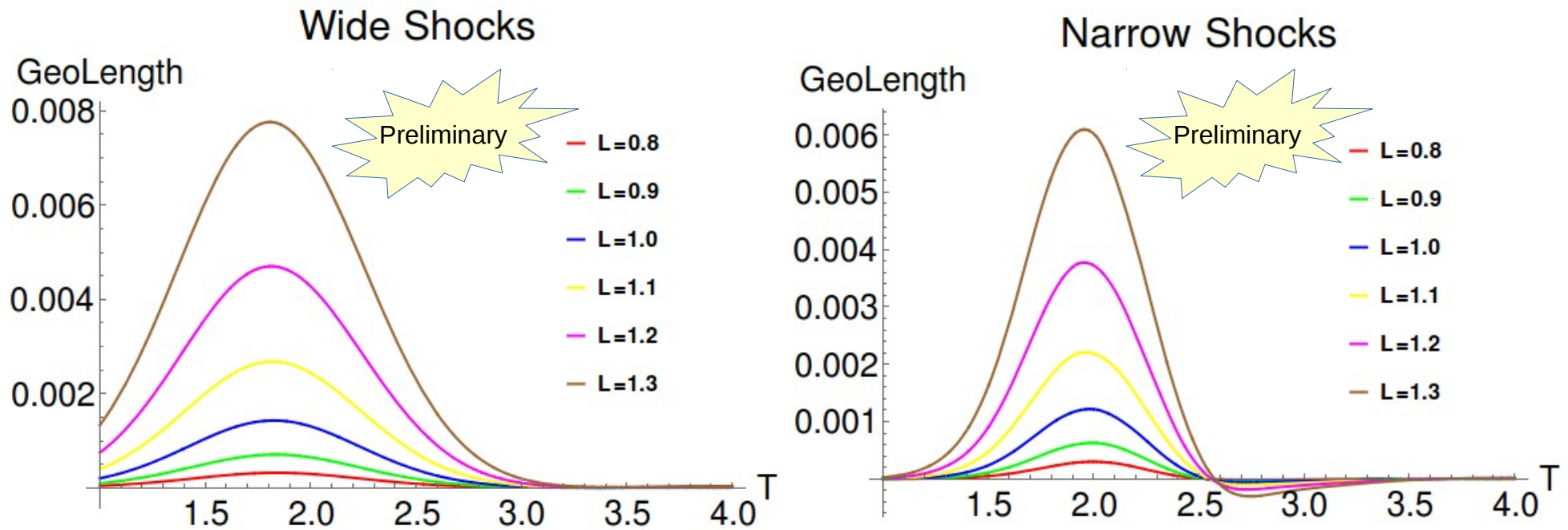
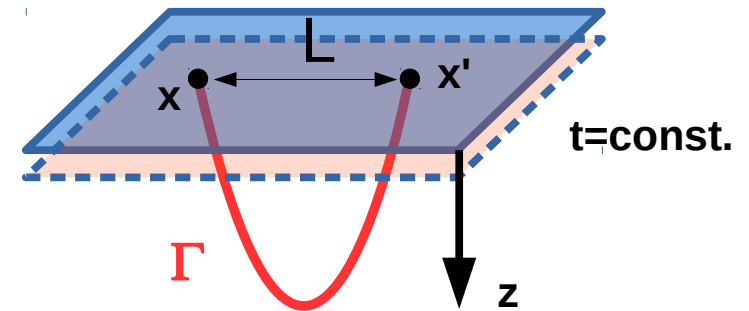
[CE-Grumiller-Van der  
Schee-Stricker 15?]



# Two point functions

Within AdS/CFT **two point functions** for operators  $\mathcal{O}(t,x)$  of large conformal weight  $\Delta$  can be computed from the **length of geodesics**. [Balasubramanian-Ross 00]

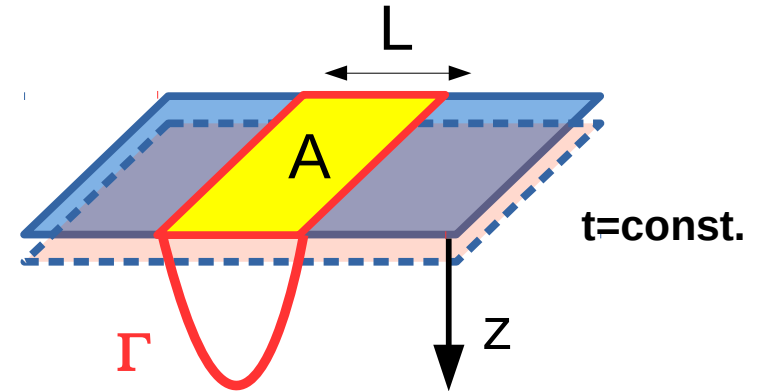
$$\langle \mathcal{O}(t, x) \mathcal{O}(t, x') \rangle \propto e^{-\Delta \text{Length}(\Gamma)}$$



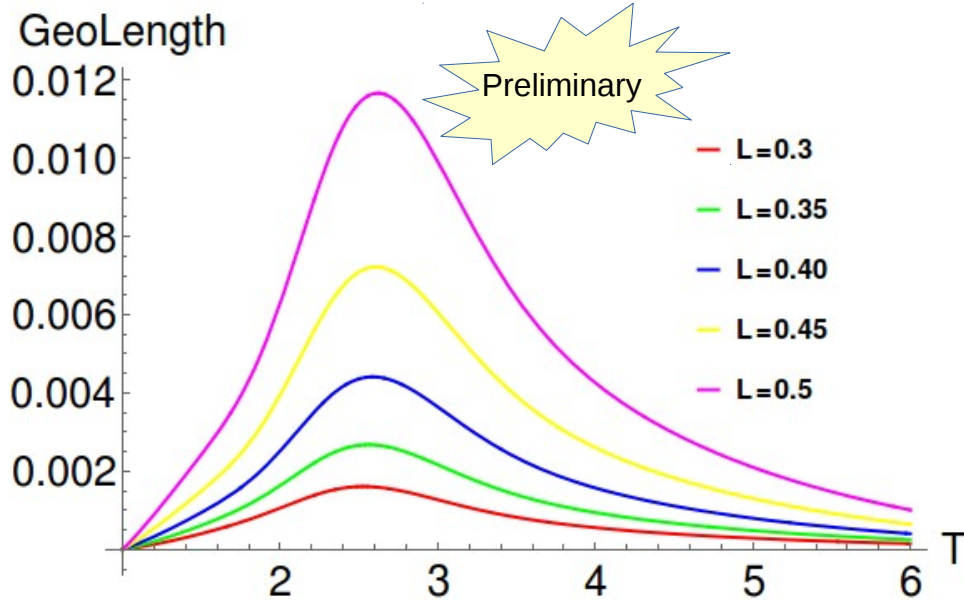
[CE-Grumiller-Van der Schee-Stricker 15?]

# Entanglement entropy

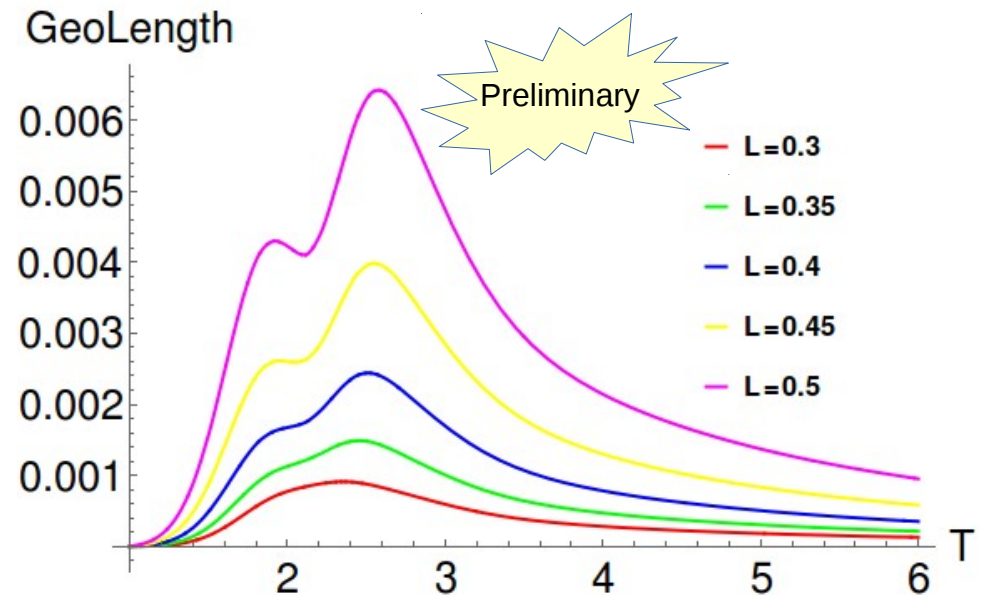
$$S_A = \text{const.} \frac{\text{Length}(\Gamma)}{4G_N}$$



Wide Shocks



Narrow Shocks

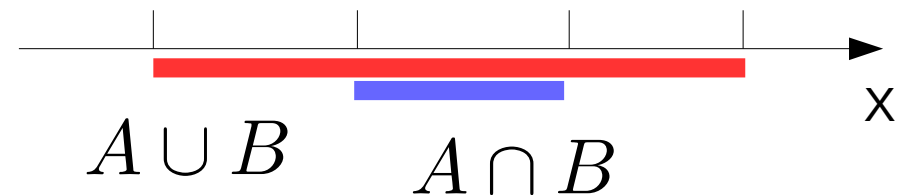
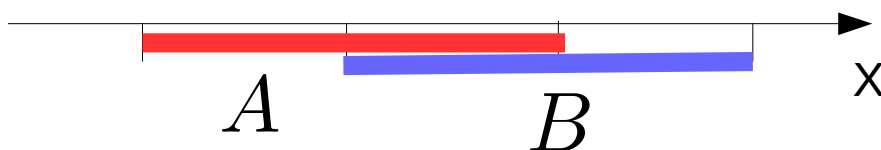


[CE-Grumiller-Van der Schee-Stricker 15?]

# Strong subadditivity

- A **fundamental property** of entanglement entropy is strong subadditivity.
- Hard to prove within QFT, very **intuitive** in the **dual gravity picture**.

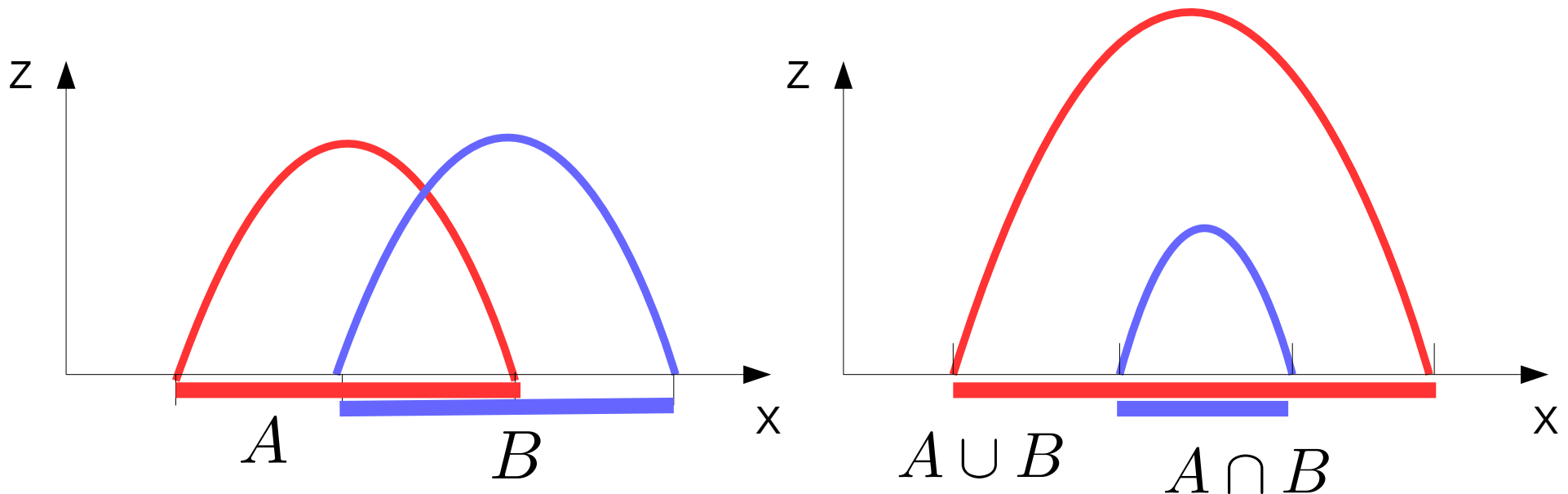
$$S_A + S_B \geq S_{A \cup B} + S_{A \cap B}$$



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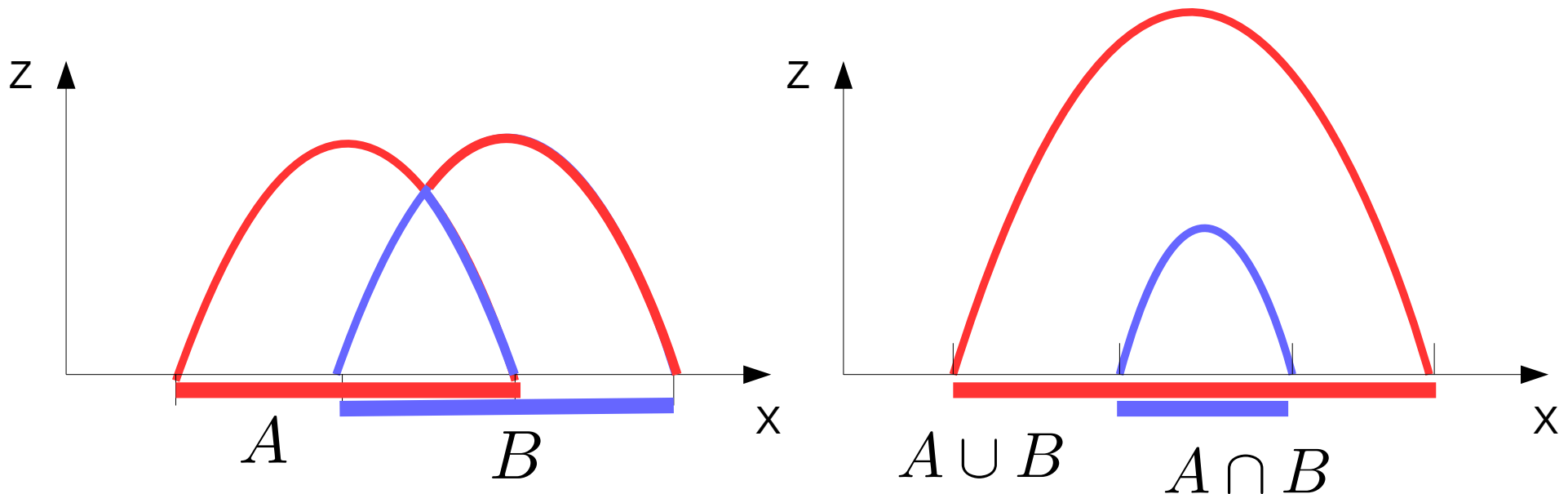
$$\boxed{S_A} + \boxed{S_B} \geq \boxed{S_{A \cup B}} + \boxed{S_{A \cap B}}$$



# Strong subadditivity

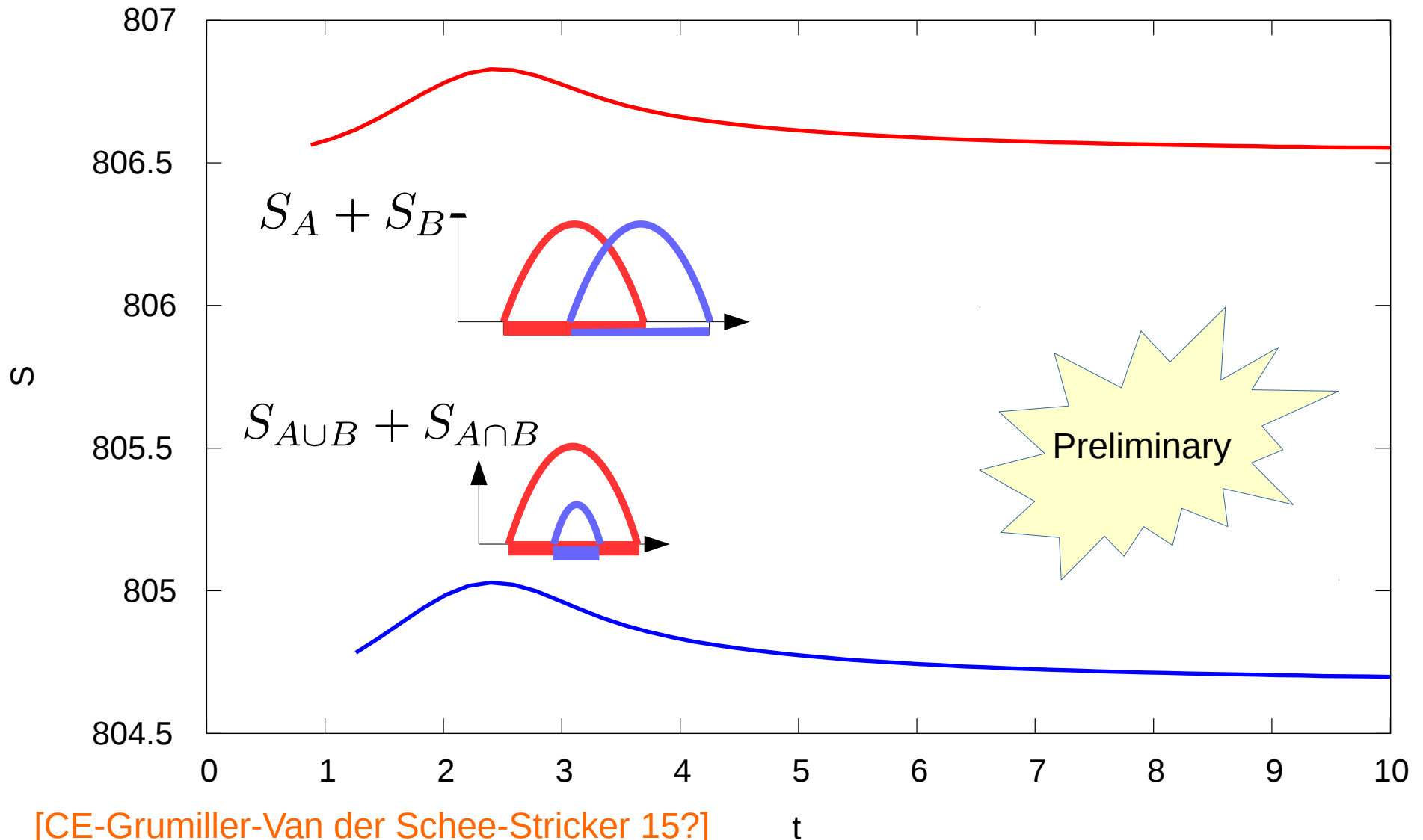
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$$\boxed{S_A} + \boxed{S_B} \geq \boxed{S_{A \cup B}} + \boxed{S_{A \cap B}}$$





# Numerical check of strong subadditivity



[CE-Grumiller-Van der Schee-Stricker 15?]

$t$

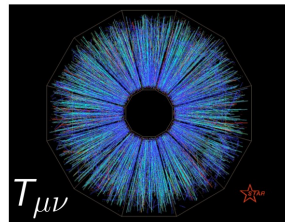
# Summary

- The **near equilibrium dynamics** of holographic entanglement shows **quasinormal mode** behaviour.
- In holographic **shock wave collisions** the **entanglement entropy** and the **two point function** may serve as **order parameter** for the **full stopping–transparency transition**.
- We **numerically checked** the **strong subadditivity** condition.

# Take home message

Within AdS/CFT complicated stuff in the CFT often has a very intuitive geometric interpretation on the AdS side.

**thermalization**

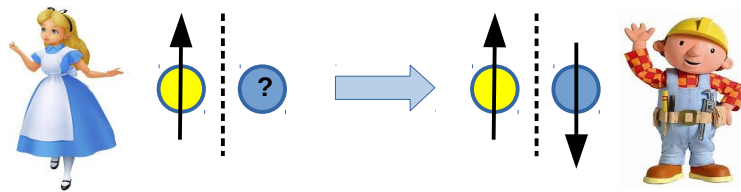


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**black hole formation**

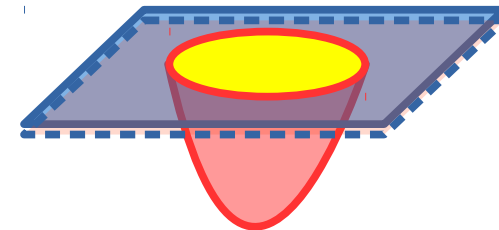


**entanglement entropy**



=

**area of extremal surface**



**two point function**

$$\langle \mathcal{O}(t, x) \mathcal{O}(t, x') \rangle$$

=

**length of geodesic**

