Validating galaxy clustering models with *Fixed & Paired* and *Matched*-ICs simulations: application to Primordial Non-Gaussianities

arXiv:2204.11103

[check also references therein]

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In coll. w. Adrian Gutierrez Adame

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Primordial Non-Gaussianities (PNG)

- One of the few observables of Inflation
- Local PNG
 - Definition: $\Phi(\mathbf{x}) = \Phi_G(\mathbf{x}) + f_{\mathrm{NL}}^{\mathrm{loc}} (\Phi_G(\mathbf{x})^2 \langle \Phi_G^2 \rangle)$
 - If detected, ($O(f_{NL}) = 1$), a smoking gun for Multi-field inflation

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 - If detected, ($O(f_{NL}) = 1$), a smoking gun for Multi-field inflation
 - Creates a scale-dependent bias on halos

$$b(k) = b_g + 2\delta_c (b_g - p) \cdot f_{\rm NL} \cdot \frac{1}{\alpha(k, z)} \qquad \alpha(k, z) = \frac{2D(z)}{3\Omega_m(z)} \frac{c^2}{H_0^2} \frac{g(z_{\rm rad})}{g(0)} k^2 T(k) \propto k^2$$

Dalal+08, Slosar+08, Barreira+20, PNG-UNITsim in prep.



- Normal (Gaussian) Initial Conditions ($\Phi_i(k) \curvearrowleft \mathcal{N}_{\mathbb{C}}(\Phi_i(k); \mu = 0, \sigma^2 = P(k)/\alpha^2)$)
 - Modulus (Rayleigh) $|\Phi(k)|_i \curvearrowleft \mathcal{P}_{\text{Rayleigh}}(|\Phi(k)|)$
 - Phase (flat PDF)

$$\varphi_i \curvearrowleft \frac{1}{2\pi} \Theta_{[0,2\pi]}(\varphi)$$

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$$\delta(k,z) = \alpha(k,z) \Phi(k)$$

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Then we use 2LPT to generate ICs

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Goliat-PNG simulations

- 41 different seeds
- For $f_{\rm NL} = 0$ & $f_{\rm NL} = 100$
- And ICs:
 - Normal
 - Normal-Inverted
 - Fixed
 - Fixed-Inverted
- Halos with 10 parts. (AHF)
- At z=1
- L=1Gpc/h

Goliat-PNG simulations: halo power spectrum



• Similar reduction between

 $f_{\rm NL} = 0$ and $f_{\rm NL} = 100$



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• Ratio w.r.t. Normal Gaussian error

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- Pairing:

$$\sigma_{\text{Paired}} = \sqrt{2} \cdot \text{std} \left(\frac{1}{2} \left(P_{\varphi_i}(k) + P_{\varphi_i + \pi}(k) \right) \right)$$

 $P_{\text{Paired},i}(k) = \frac{1}{2} (P_{\varphi_i}(k) + P_{\varphi_i + \pi}(k))$



Halo bias

 Linear halo bias (valid at large scales)

$$\delta_{\text{halo}} = b \cdot \delta$$

 $P_{\text{halo}}(k) = b^2 \cdot P(k)$



Error-bar: simulations Lines/bands: theory

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$$\begin{split} b(k) &= b_g + 2\delta_c (b_g - p) \cdot f_{\rm NL} \cdot \frac{1}{\alpha(k, z)} \\ \alpha(k, z) &= \frac{2D(z)}{3\Omega_m(z)} \frac{c^2}{H_0^2} \frac{g(z_{\rm rad})}{g(0)} k^2 T(k) \quad \propto k^2 \end{split}$$



k [h⁻¹Mpc]
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 $k [h^{-1}Mpc]$

10-2

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 10^{-1}



• Normal sims.

$\langle \hat{f}_{\rm NL}^{100} \rangle \pm {\rm std}(\hat{f}_{\rm NL}^{100})$ 111.4 ± 60.5



$$p = 1$$

$$P_{\text{halo}}(k) = b^2 P_{\text{lin}}(k)$$
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 $\sigma(\Delta \hat{f}_{\rm NL})^2 = \sigma(\hat{f}_{\rm NL}^{100})^2 + \sigma(\hat{f}_{\rm NL}^0)^2 - 2 \cdot \rho \cdot \sigma(\hat{f}_{\rm NL}^{100})\sigma(\hat{f}_{\rm NL}^0)$



 ρ = 0.97



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ho = 0.97, 0.76



















PNG-UNITsim

In collab. with A.G. Adame, G. Yepes, V. Gonzalez-Perez, M. Pellejero, J. Garcia-Bellido et al.

- Same properties as UNITsim:
 - L=1Gpc/h, N=4096³, $m_p=1.2 \times 10^9 M_{sun}/h$
 - Fixed
 - Gadget, Rockstar, MergerTrees
- Re-run :
 - UNITsim1 f_{NL}=100 to z=0
 - UNITsim2 f_{NI} =-20 to z=1
- Also new FastPMs:
 - N=2048³, L=1 Gpc/h
 - f_{NI}=100 & f_{NI}=0
 - x 100 realisations
 - Fixed



Barcelona

Currently, under

Could be made

immediately

available

investigation

equivalent to an uncertainty of σ(f_{NL})~5

We also plan to run SAMs to know the *p* for different tracers See Knebe, Lopez-Cano, SA et al. 2022 MNRAS 510 5391

Take-home messages

- *Fixing* for local PNG:
 - Gives unbiased halo P(k)
 - Greatly reduces the σ(P(k))
- Fixing reduces $\sigma(f_{NL})$ by a factor ~5
- *Matching*-ICs (and explicitly using their correlation) reduces $\sigma(f_{NL})$ by a factor
 - ~ 3 For *Normal* simulations
 - ~2 for *Fixed-Paired* simulations
- One couple of *Fixed-Paired-Matched* L = 1 Gpc/h simulations can give us $\sigma(f_{NL}) = 5$
 - Comparable with Euclid/DESI
 - Motivation for PNG-UNITsim (V=2 x1 [Gpc/h]³, V_{eff}~140 [Gpc/h]³)

Additional Slides

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- Ratio w.r.t. Normal Gaussian error $\sigma_{\rm th}(k)^2 = \frac{4\pi^2}{V \, k^2 \, \Delta k} \left(P(k) + \frac{1}{n} \right)^2$
- Pairing: $P_{\text{Paired},i}(k) = \frac{1}{2}(P_{\varphi_i}(k) + P_{\varphi_i + \pi}(k))$ $\sigma_{\text{Paired}} = \sqrt{2} \cdot \text{std}\left(\frac{1}{2}(P_{\varphi_i}(k) + P_{\varphi_i + \pi}(k))\right)$
- Fit for Fixed (& Paired)

$$\sigma_{\rm r}(k) = \sqrt{\frac{4\pi^2}{V k^2 \Delta k}} \left(P(k) \cdot \left[R_{\rm cv} - (1 - R_{\rm cv}) \cdot \frac{2}{\pi} \arctan\left(\frac{k}{k_{\rm soft}}\right) \right] + \frac{1}{n} \cdot f_{\rm sn} \right)$$



Reference

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Validating galaxy clustering models with Fixed & Paired and Matched-ICs simulations: application to Primordial Non-Gaussianities

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