

# Model independent approach to galaxy clustering

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**TECHNION**  
Israel Institute  
of Technology

# Standard perturbation theory

Dark matter: non relativistic, non interacting, subject to gravitational interaction

$$\frac{\partial \delta(\mathbf{x}, \tau)}{\partial \tau} + \nabla_{\mathbf{x}} \left[ (1 + \delta(\mathbf{x}, \tau)) \mathbf{v}(\mathbf{x}, \tau) \right] = 0 \quad \text{Continuity equation}$$

$$\frac{\partial \mathbf{v}(\mathbf{x}, \tau)}{\partial \tau} + \mathcal{H}(\tau) \mathbf{v}(\mathbf{x}, \tau) + (\mathbf{v} \cdot \nabla_{\mathbf{x}}) \mathbf{v}(\mathbf{x}, \tau) + \nabla_{\mathbf{x}} \phi(\mathbf{x}, \tau) = \frac{1}{\rho} \nabla_{\mathbf{x}} \tau$$

Euler equation

$$\nabla_{\mathbf{x}}^2 \phi(\mathbf{x}, \tau) = \frac{3}{2} \Omega_m(\tau) \mathcal{H}^2(\tau) \delta(\mathbf{x}, \tau) \quad \text{Poisson equation}$$

# Standard perturbation theory

## Observational effects

From dark matter to galaxies: bias, redshift space, small scale effects

$$\delta(\mathbf{k}, \tau) \longrightarrow \delta_g^{\text{rs}}(\mathbf{k}, \tau) = \mathcal{F}[\delta, \partial_i \partial_j \phi, \dots](\mathbf{k}, \tau)$$

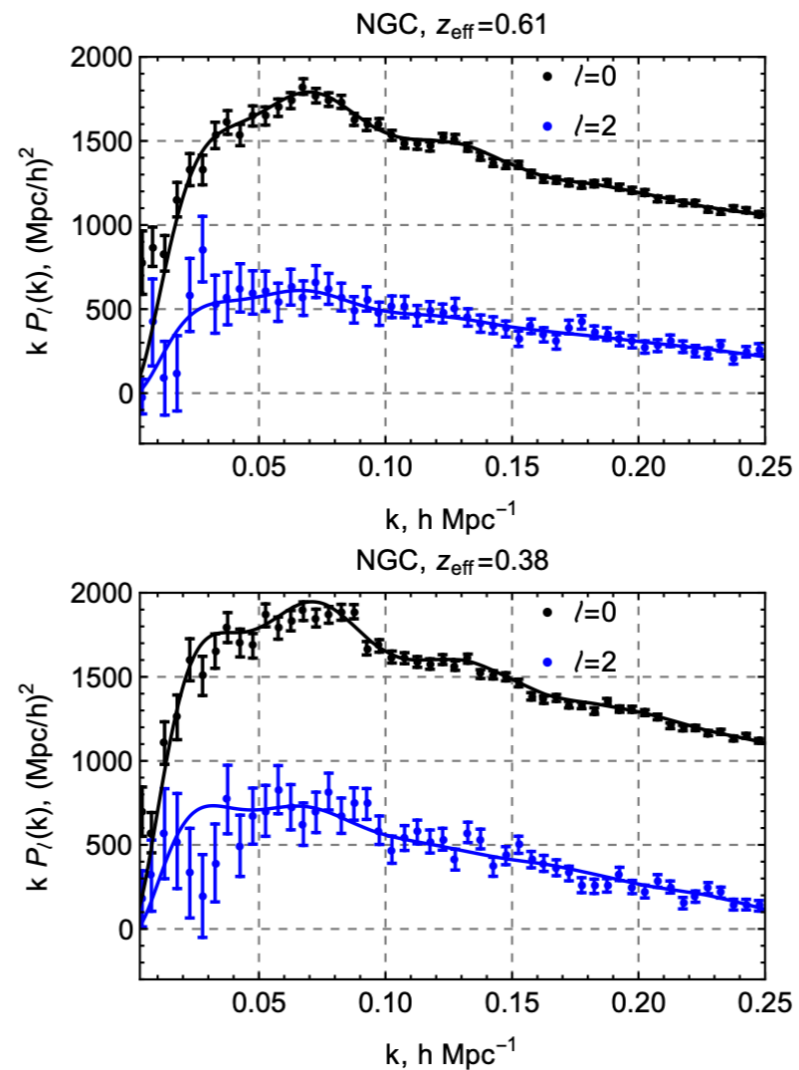
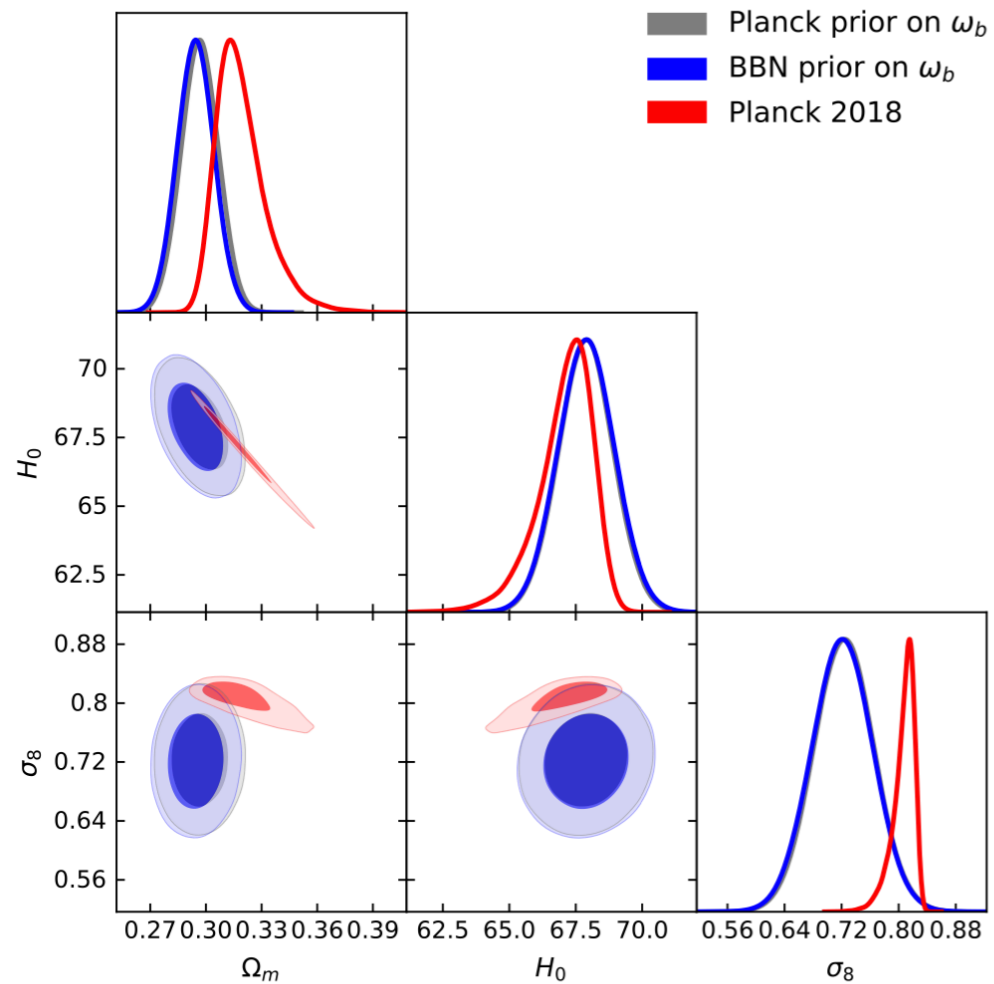
Perturbative solution

$$\delta_g^{\text{rs},(n)}(\mathbf{k}) = \int \frac{d^3 q_1}{(2\pi)^3} \cdots \int \frac{d^3 q_n}{(2\pi)^3} \delta_D(\mathbf{k} - \mathbf{q}_{1\dots n}) Z_n(\mathbf{q}_1, \dots, \mathbf{q}_n) \delta^{(1)}(\mathbf{q}_1) \cdots \delta^{(1)}(\mathbf{q}_n)$$

$$\delta_g^{\text{rs}}(\mathbf{k}) = \delta_g^{\text{rs,PT}}(\mathbf{k}) + \delta_g^{\text{rs,SN}}(\mathbf{k}) + \delta_g^{\text{rs,CT}}(\mathbf{k})$$

# Standard perturbation theory

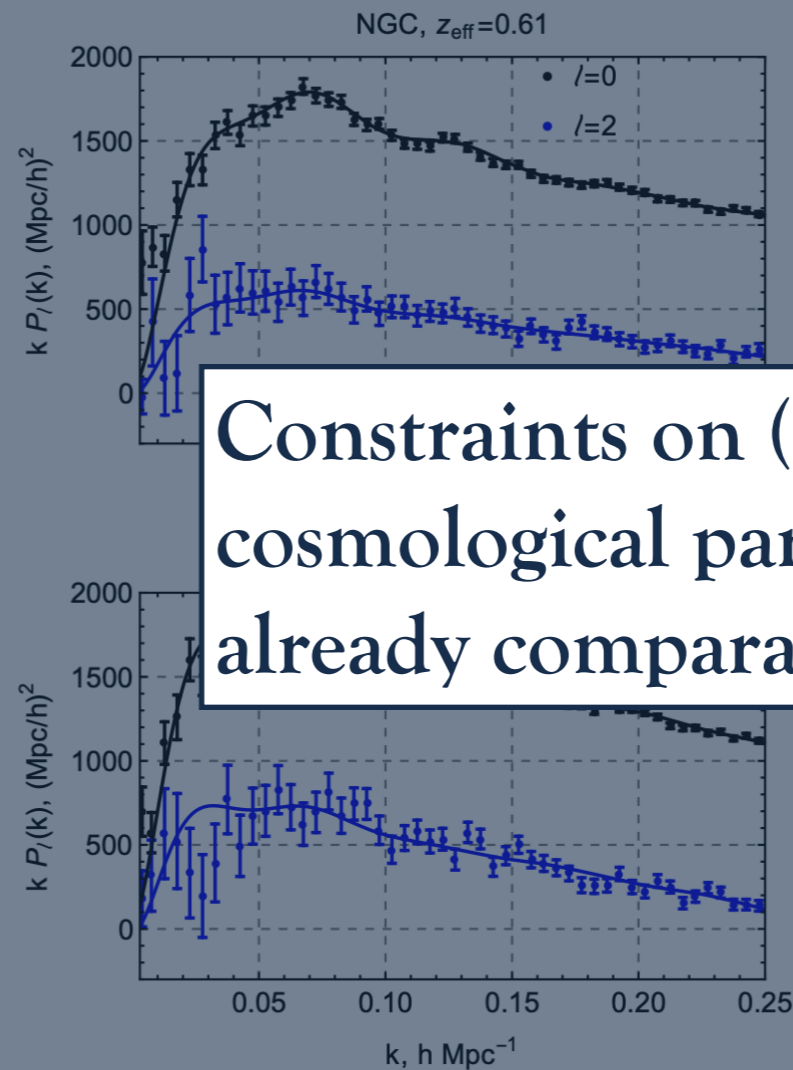
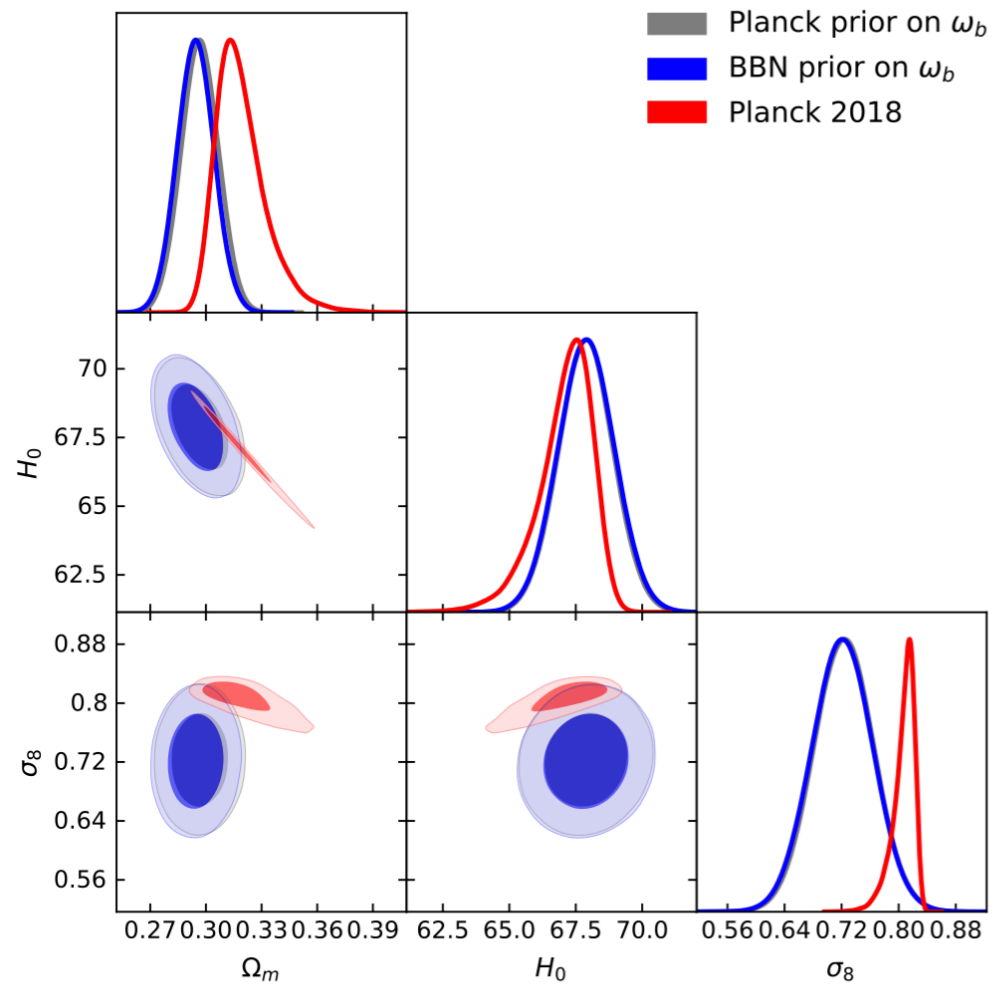
## Results



D'Amico et al. 1909.05271  
Ivanov et al. 1909.05277  
Colas et al. 1909.07951  
Chen et al. 2110.05530

# Standard perturbation theory

## Results



Constraints on (some) cosmological parameters already comparable with Planck

D'Amico et al. 1909.05271  
Ivanov et al. 1909.05277  
Colas et al. 1909.07951  
Chen et al. 2110.05530

# Role of symmetries in cosmology

# Role of symmetries in cosmology

## Equivalence principle



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graph TD; EP[Equivalence principle] --> MI[Model independent measurement using the EP in the BAO range]; EP --> CS[Constraints the perturbative kernels in a model independent way: Large Scale Structure Bootstrap!]
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Model independent measurement using the EP in the BAO range

- MM, Nishimichi T., Pietroni M., Phys. Rev. D100 (2019)
- MM, Nishimichi T., Pietroni M., JCAP 07 (2020)

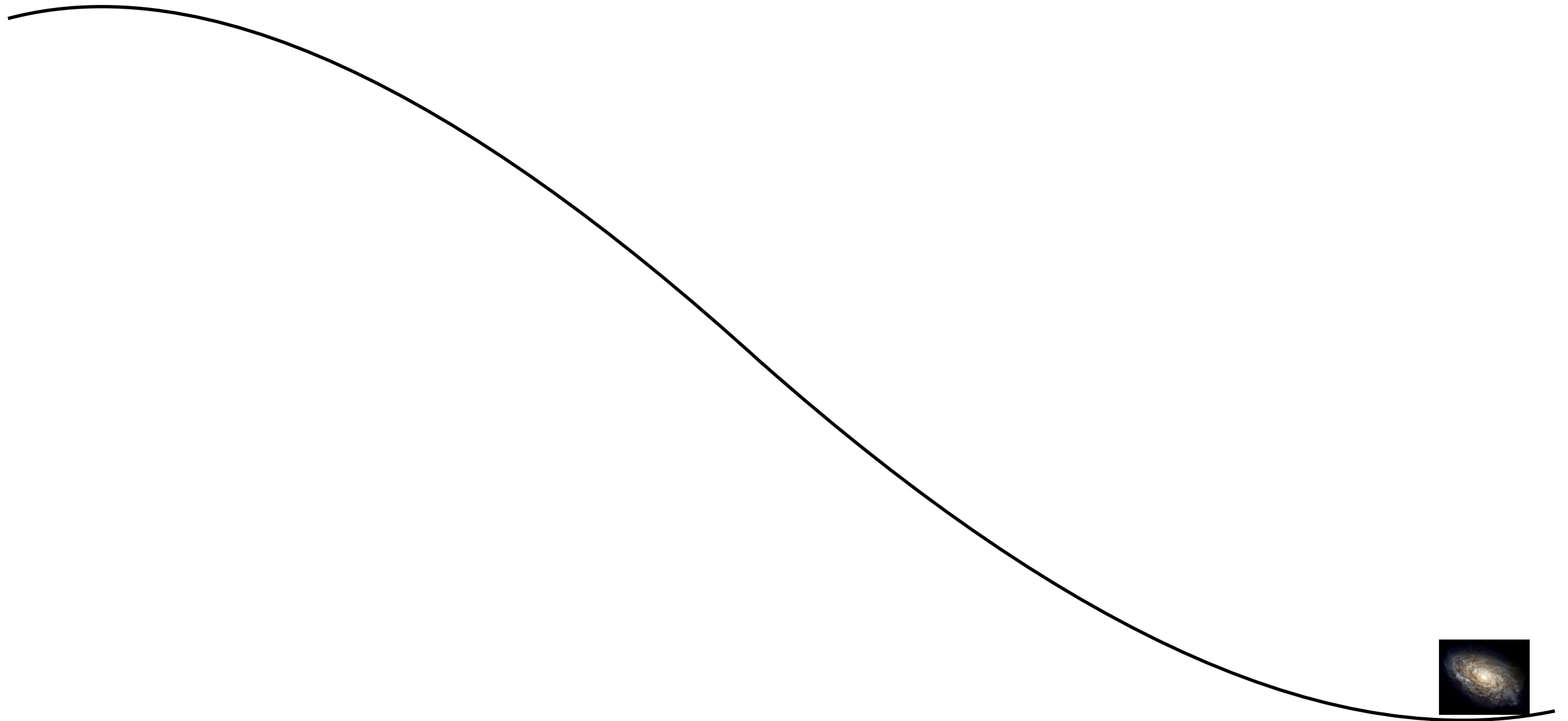
Constraints the perturbative kernels in a model independent way: **Large Scale Structure Bootstrap!**

- D'Amico G., MM, Pietroni M., Vernizzi F., JCAP 10 (2021)

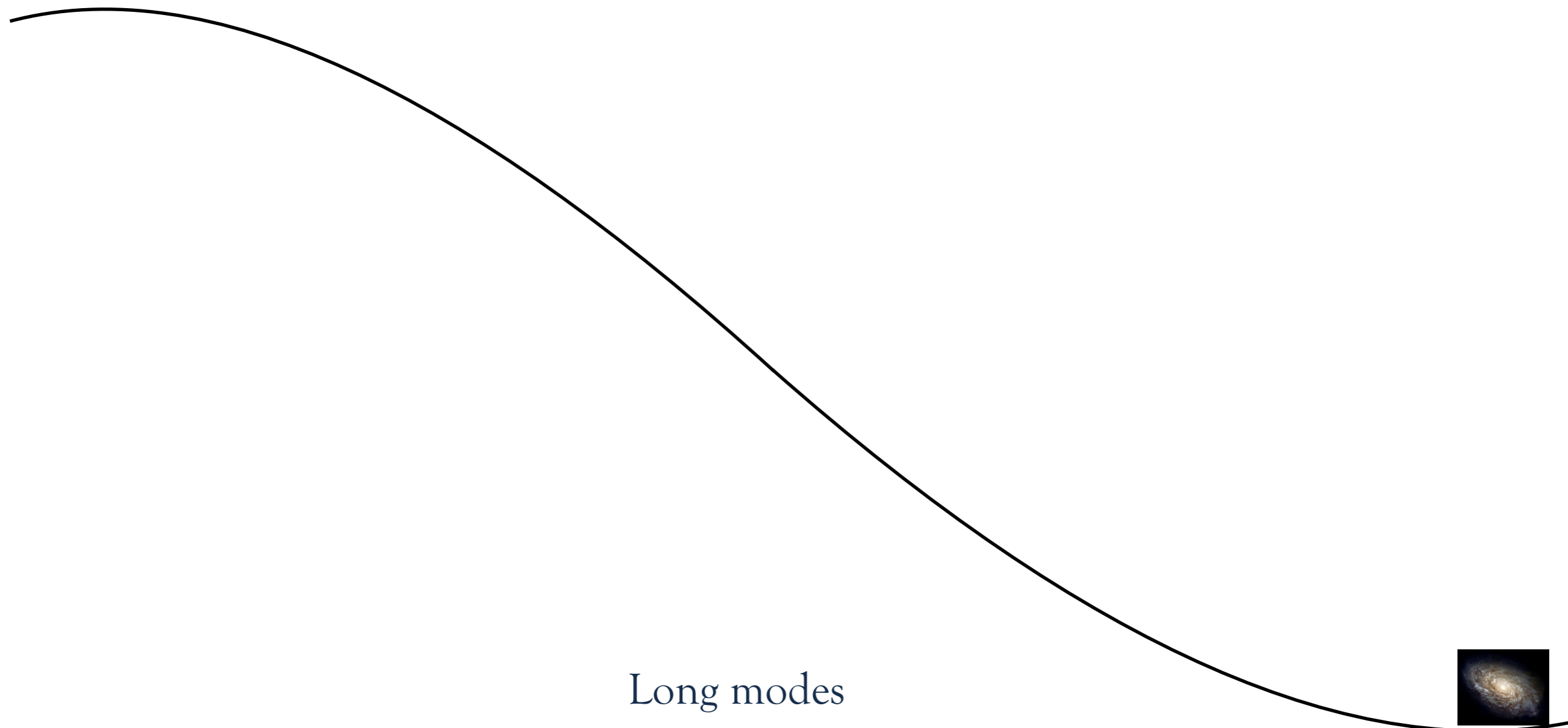
# Equivalence principle



# Equivalence principle



# Equivalence principle

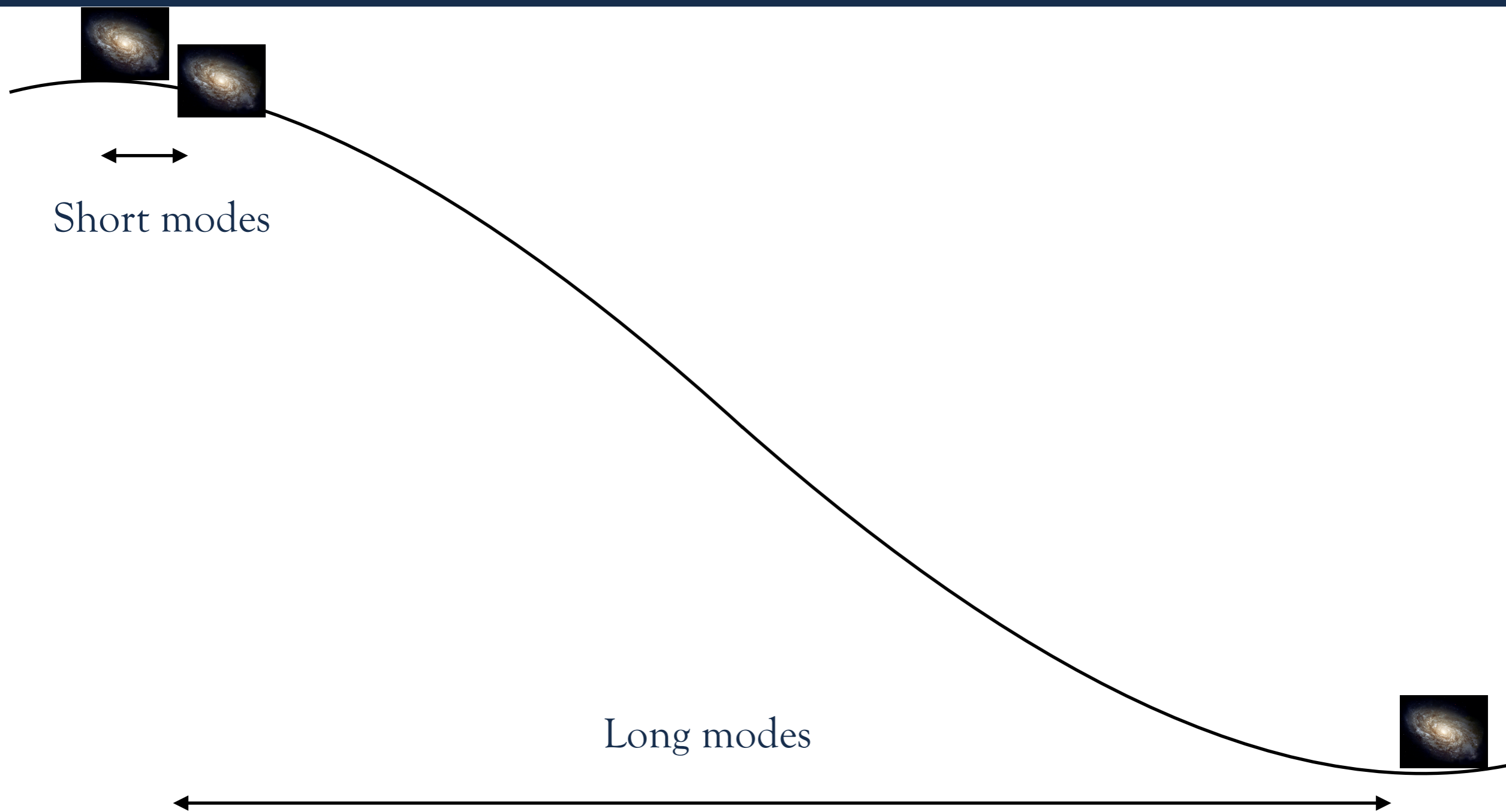


SPT

Equivalence principle

Bootstrap

# Equivalence principle

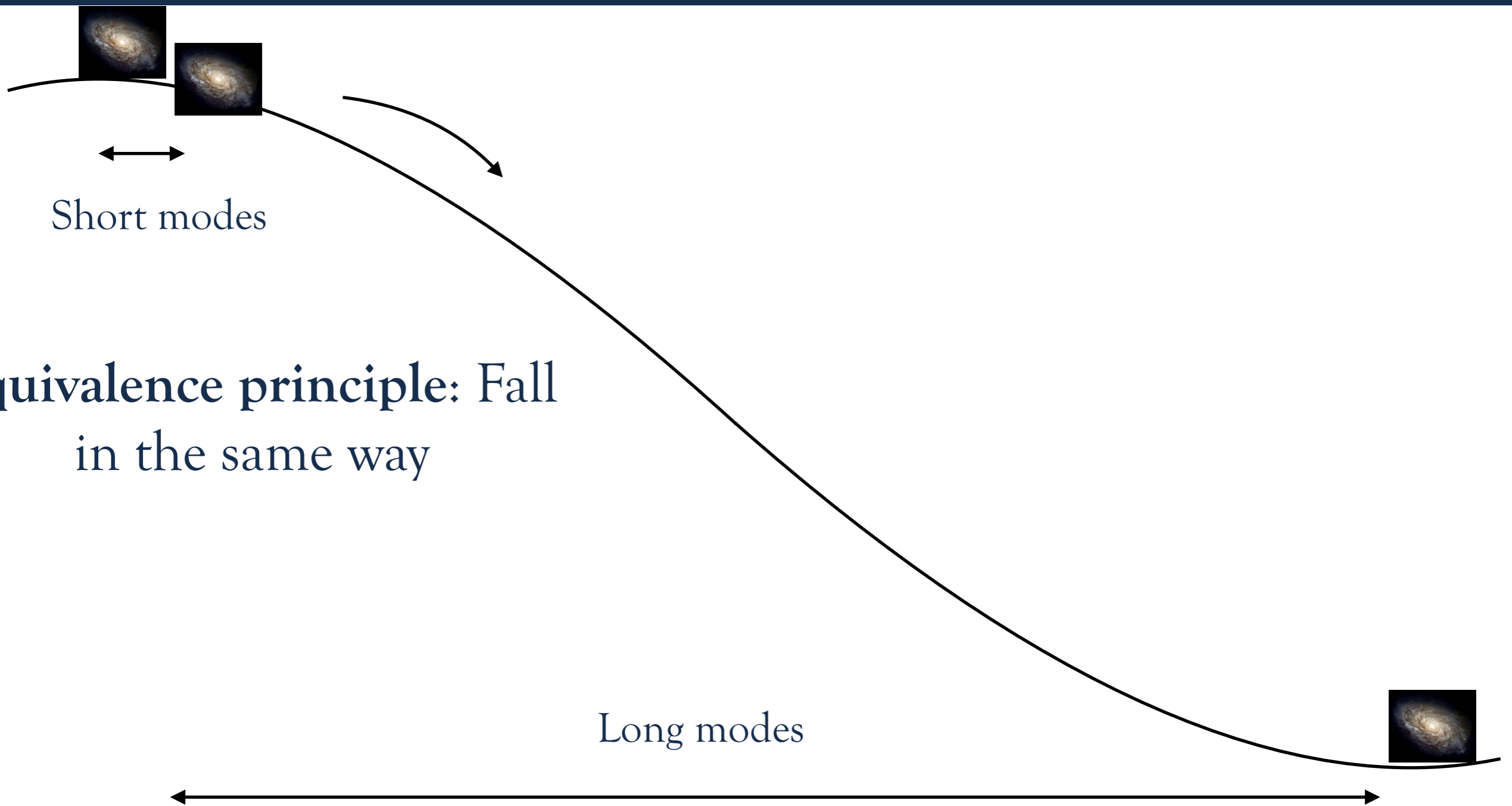


SPT

Equivalence principle

Bootstrap

# Equivalence principle



# Equivalence principle

## Consistency relations

Exact equalities among correlation functions of different order  
(Kehagias A., Riotto A., Nucl.Phys. 2013,  
Peloso M., Pietroni M., JCAP 2013,  
Creminelli P., Noreña J., Simonović M., Vernizzi F., JCAP 2013)

$$\langle \delta(\mathbf{x}_1, \tau_1) \dots \delta(\mathbf{x}_n, \tau_n) | \Phi_L \rangle = \langle \delta(\tilde{\mathbf{x}}_1, \tilde{\tau}_1) \dots \delta(\tilde{\mathbf{x}}_n, \tilde{\tau}_n) \rangle .$$

Effect of long modes on linear scales

→ **Equivalence Principle** (or Galilean Invariance)

$$\langle \delta_m(\mathbf{q}, \tau) \delta_g(\mathbf{k}_1, \tau_1) \dots \delta_g(\mathbf{k}_n, \tau_n) \rangle'_{\mathbf{q} \rightarrow 0} = -P_L(q, \tau) \sum_{i=1}^n \frac{D_+(\tau_i)}{D(\tau)} \frac{\mathbf{q} \cdot \mathbf{k}_i}{q^2} \langle \delta_g(\mathbf{k}_1, \tau_1) \dots \delta_g(\mathbf{k}_n, \tau_n) \rangle'$$

# LSS Bootstrap

- D'Amico G., MM, Pietroni M., Vernizzi F., JCAP 10 (2021)

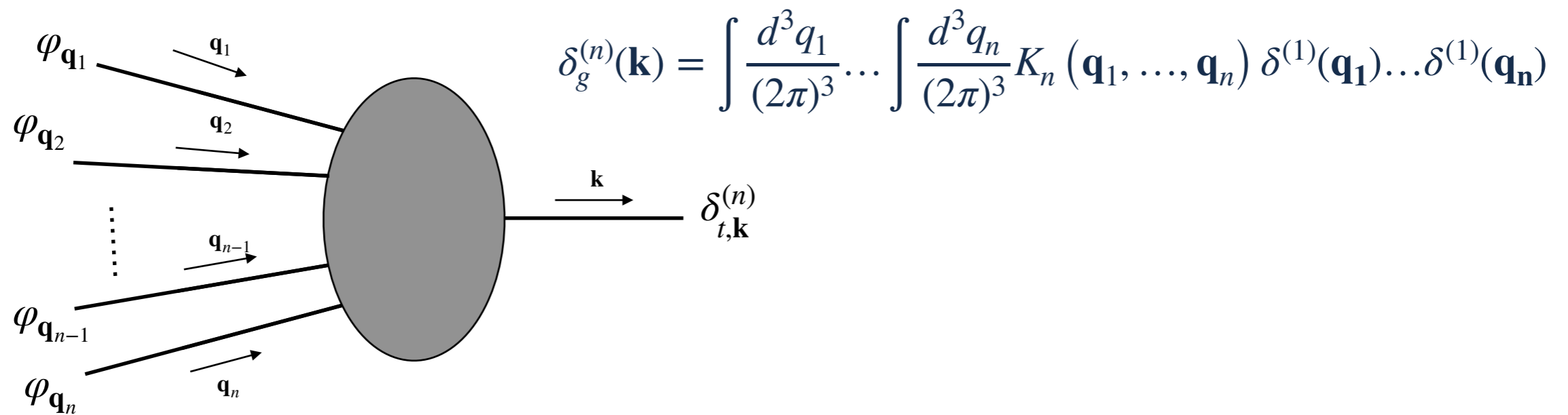
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- D'Amico G., MM, Pietroni M., Vernizzi F., JCAP 10 (2021)

$$\delta_g^{(n)}(\mathbf{k}) = \int \frac{d^3 q_1}{(2\pi)^3} \cdots \int \frac{d^3 q_n}{(2\pi)^3} K_n(\mathbf{q}_1, \dots, \mathbf{q}_n) \delta^{(1)}(\mathbf{q}_1) \cdots \delta^{(1)}(\mathbf{q}_n)$$

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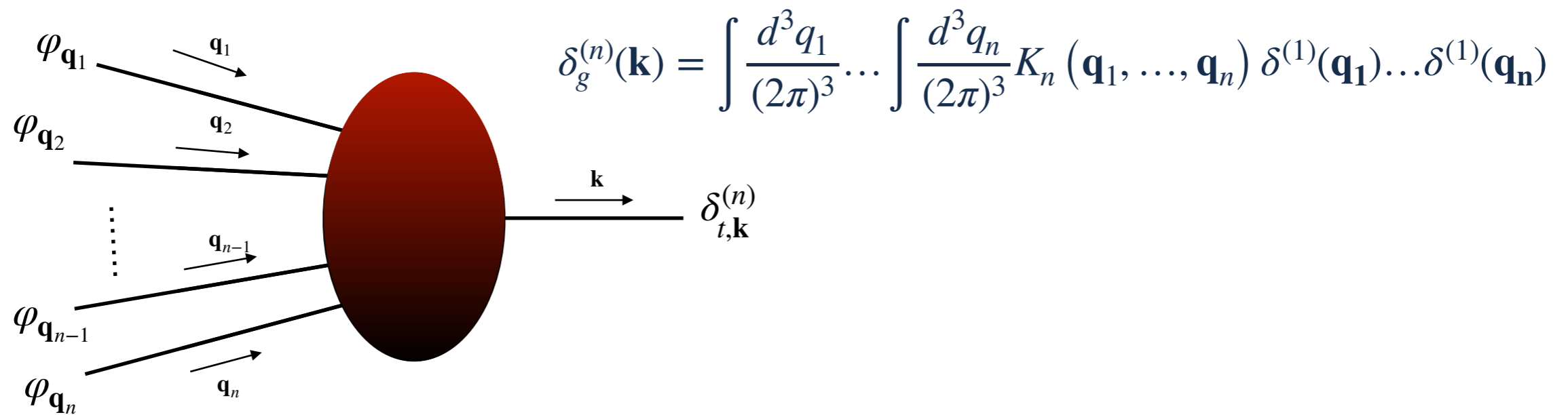
- D'Amico G., MM, Pietroni M., Vernizzi F., JCAP 10 (2021)





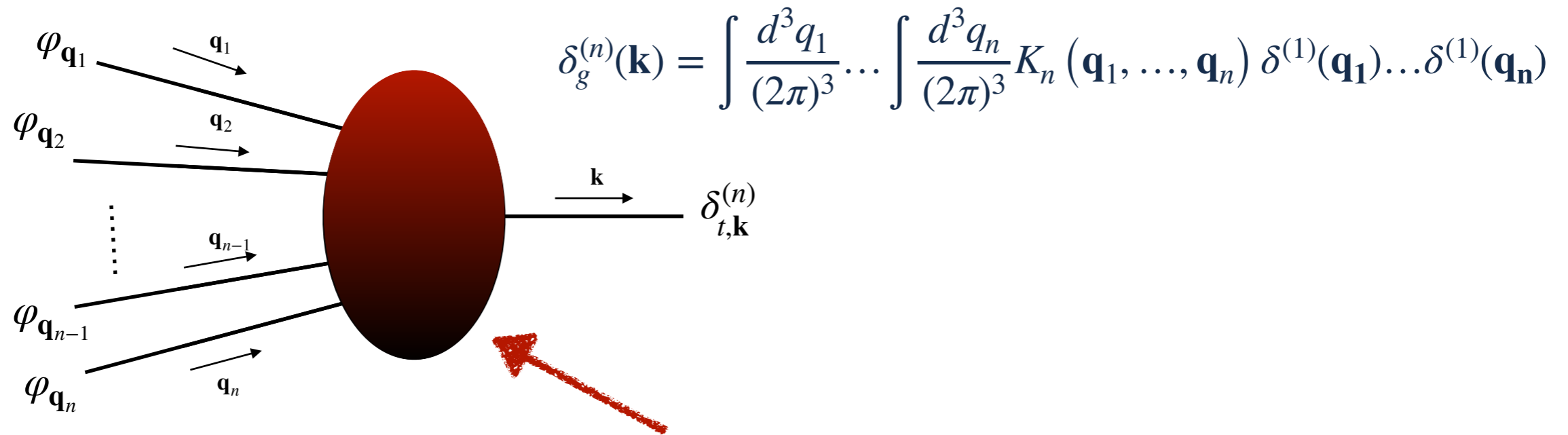
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## SYMMETRIES

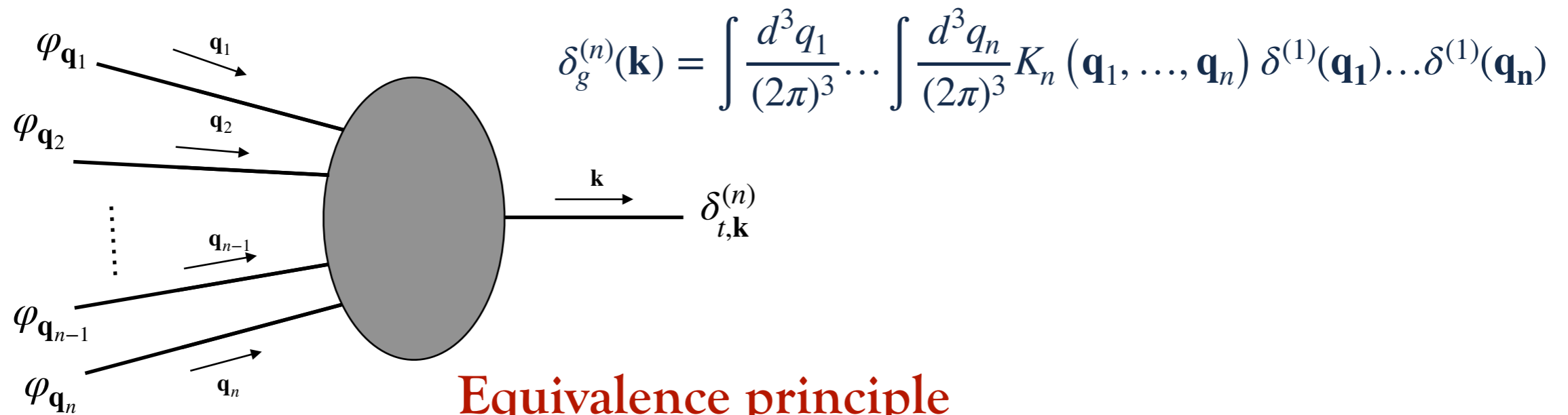
Homogeneity and isotropy

Mass and momentum conservation (only for dark matter)

**Equivalence principle**

# LSS Bootstrap

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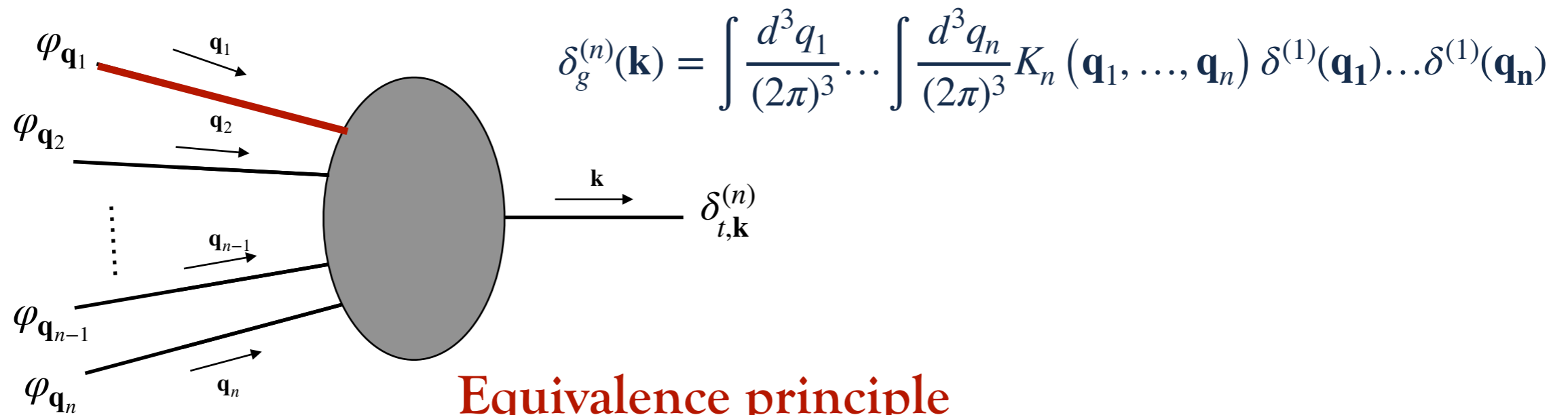


**Equivalence principle**

Leading Order: single momentum  
going  $\rightarrow 0$

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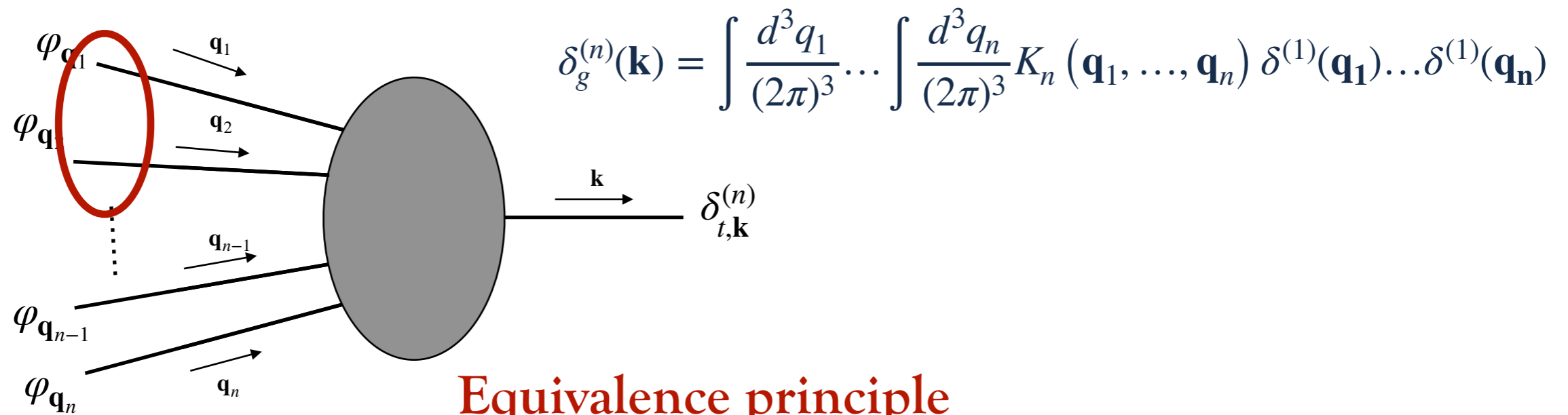


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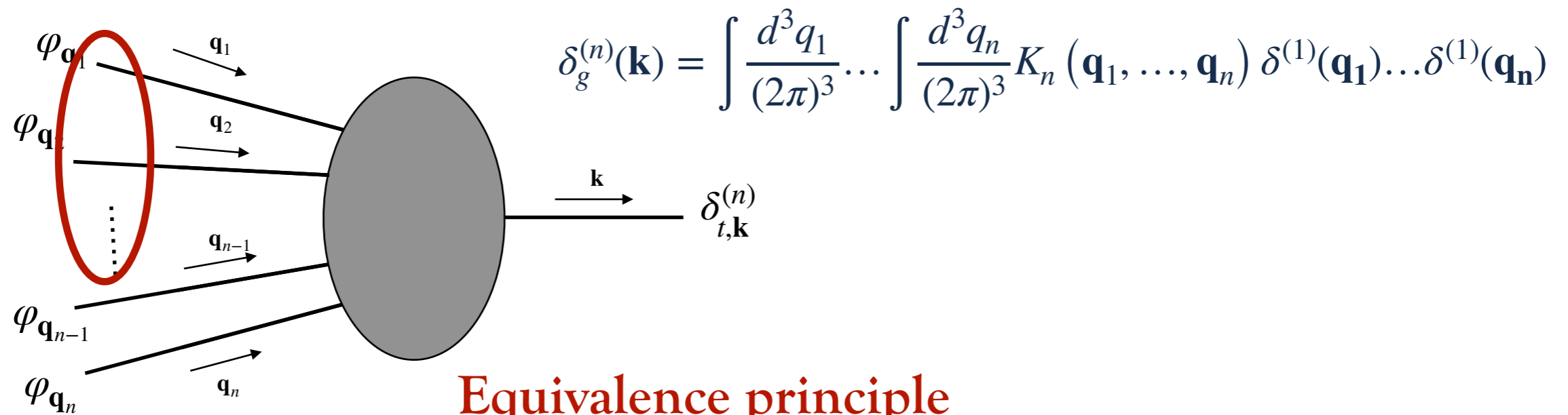
## Equivalence principle

Leading Order: single momentum going  $\rightarrow 0$

Next-to-Leading Order: sum of two momenta going  $\rightarrow 0$

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Leading Order: single momentum going  $\rightarrow 0$

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Next-to-Next-to-Leading Order: sum of three momenta going  $\rightarrow 0$

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$N^{l-1}$ -to-Leading Order: sum of  $l - 1$   
momenta going  $\rightarrow 0$



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$N^{l-1}$ -to-Leading Order: sum of  $l - 1$   
momenta going  $\rightarrow 0$

$$\lim_{\mathbf{q}_1, \dots, \mathbf{q}_m \rightarrow 0} K_n(\mathbf{q}_1, \dots, \mathbf{q}_m, \mathbf{q}_{m+1}, \dots, \mathbf{q}_n) = \frac{\mathbf{q}_1 \cdot \mathbf{Q}_{n,m}}{q_1^2} \dots \frac{\mathbf{q}_m \cdot \mathbf{Q}_{n,m}}{q_m^2} K_{n-m}(\mathbf{q}_{m+1}, \dots, \mathbf{q}_n) + \mathcal{O}\left(\left(\frac{1}{q}\right)^{m-1}\right)$$

# LSS Bootstrap

- D'Amico G., MM, Pietroni M., Vernizzi F., JCAP 10 (2021)

## Equivalence principle

Leading Order: single momentum going  $\rightarrow 0$

Next-to-Leading Order: sum of two momenta going  $\rightarrow 0$

N.<sup>*l*-1</sup>-to-Leading Order: sum of  $l - 1$  momenta going  $\rightarrow 0$

$$\lim_{\mathbf{q}_1 + \mathbf{q}_2 \rightarrow 0} K_n(\mathbf{q}_1, \mathbf{q}_2, \mathbf{q}_3, \dots, \mathbf{q}_n) \supset \frac{\mathbf{k} \cdot \mathbf{q}_{12}}{q_{12}^2} K_{n-2}(\mathbf{q}_3, \dots, \mathbf{q}_n) \int^\eta d\eta' f_+(\eta') \frac{D_+(\eta')^2}{D_+(\eta)^2} G_2(\mathbf{q}_1, \mathbf{q}_2; \eta')$$

# LSS Bootstrap

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## Equivalence principle

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$N^{l-1}$ -to-Leading Order: sum of  $l$  momenta going  $\rightarrow 0$

$$\lim_{Q_{l,0} \rightarrow 0} K_n(\mathbf{q}_1, \dots, \mathbf{q}_l, \mathbf{q}_{l+1}, \dots, \mathbf{q}_n) \supset \frac{\mathbf{k} \cdot \mathbf{Q}_{l,0}}{Q_{l,0}^2} \int^\eta d\eta' f_+(\eta') \left( \frac{D_+(\eta')}{D_+(\eta)} \right)^l G_l(\mathbf{q}_1, \dots, \mathbf{q}_l; \eta') K_{n-l}(\mathbf{q}_{l+1}, \dots, \mathbf{q}_n; \eta)$$

# LSS Bootstrap

Kernel at second order

$$K_1(\mathbf{q}_1) = c_0$$

$$K_2(\mathbf{q}_1, \mathbf{q}_2) = c_1 + c_\beta \beta(\mathbf{q}_1, \mathbf{q}_2) + c_\gamma \gamma(\mathbf{q}_1, \mathbf{q}_2)$$

$$\beta(\mathbf{q}_1, \mathbf{q}_2) = \frac{|\mathbf{q}_1 + \mathbf{q}_2|^2 \mathbf{q}_1 \cdot \mathbf{q}_2}{2q_1^2 q_2^2}$$

$$\gamma(\mathbf{q}_1, \mathbf{q}_2) = 1 - \left( \frac{\mathbf{q}_1 \cdot \mathbf{q}_2}{q_1 q_2} \right)^2$$

$$\alpha_a(\mathbf{q}_1, \mathbf{q}_2) = \frac{\mathbf{q}_1 \cdot \mathbf{q}_2}{q_1^2} - \frac{\mathbf{q}_1 \cdot \mathbf{q}_2}{q_2^2}$$

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$$\{c_0, c_1, c_\beta, c_\gamma\}$$

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Leading Order

# LSS Bootstrap

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$$\{c_0, c_1, \cancel{c_\beta}, c_\gamma\}$$

Only 3 parameters left!  
(tracers)

$$\beta(\mathbf{q}_1, \mathbf{q}_2) = \frac{|\mathbf{q}_1 + \mathbf{q}_2|^2 \mathbf{q}_1 \cdot \mathbf{q}_2}{2q_1^2 q_2^2}$$

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Leading Order

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Leading Order

Mass+momentum  
conservation (matter)



# LSS Bootstrap

Kernel at second order

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$$\{\cancel{c_0}, \cancel{c_1}, \cancel{c_\beta}, c_\gamma\}$$

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Leading Order

Mass+momentum  
conservation (matter)

Only 1 parameter left!  
(matter)

# LSS Bootstrap

Kernel at third order

$$\begin{aligned} K_3(\mathbf{q}_1, \mathbf{q}_2, \mathbf{q}_3) = & c_2 + c_{\gamma 1} \gamma(\mathbf{q}_1, \mathbf{q}_2) + c_{\gamma 2} \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\beta 1} \beta(\mathbf{q}_1, \mathbf{q}_2) + c_{\beta 2} \beta(\mathbf{q}_{12}, \mathbf{q}_3) \\ & + c_{\gamma \gamma} \gamma(\mathbf{q}_1, \mathbf{q}_2) \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\beta \beta} \beta(\mathbf{q}_1, \mathbf{q}_2) \beta(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\gamma \beta} \gamma(\mathbf{q}_1, \mathbf{q}_2) \beta(\mathbf{q}_{12}, \mathbf{q}_3) \\ & + c_{\beta \gamma} \beta(\mathbf{q}_1, \mathbf{q}_2) \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + \left( c_\alpha + c_{\gamma \alpha} \gamma(\mathbf{q}_1, \mathbf{q}_2) + c_{\beta \alpha} \beta(\mathbf{q}_1, \mathbf{q}_2) \right) \alpha_a(\mathbf{q}_{12}, \mathbf{q}_3) \end{aligned}$$

# LSS Bootstrap

Kernel at third order

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# LSS Bootstrap

Kernel at third order

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Leading Order

# LSS Bootstrap

Kernel at third order

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Leading Order

Next-to-Leading Order

# LSS Bootstrap

Kernel at third order

$$\begin{aligned} K_3(\mathbf{q}_1, \mathbf{q}_2, \mathbf{q}_3) = & c_2 + c_{\gamma 1} \gamma(\mathbf{q}_1, \mathbf{q}_2) + c_{\gamma 2} \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\beta 1} \beta(\mathbf{q}_1, \mathbf{q}_2) + c_{\beta 2} \beta(\mathbf{q}_{12}, \mathbf{q}_3) \\ & + c_{\gamma\gamma} \gamma(\mathbf{q}_1, \mathbf{q}_2) \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\beta\beta} \beta(\mathbf{q}_1, \mathbf{q}_2) \beta(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\gamma\beta} \gamma(\mathbf{q}_1, \mathbf{q}_2) \beta(\mathbf{q}_{12}, \mathbf{q}_3) \\ & + c_{\beta\gamma} \beta(\mathbf{q}_1, \mathbf{q}_2) \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + \left( c_\alpha + c_{\gamma\alpha} \gamma(\mathbf{q}_1, \mathbf{q}_2) + c_{\beta\alpha} \beta(\mathbf{q}_1, \mathbf{q}_2) \right) \alpha_a(\mathbf{q}_{12}, \mathbf{q}_3) \\ & \{ c_2, c_{\gamma 1}, \cancel{c_{\gamma 2}}, \cancel{c_{\beta 1}}, \cancel{c_{\beta 2}}, c_{\gamma\gamma}, \cancel{c_{\beta\beta}}, \cancel{c_{\gamma\beta}}, \cancel{c_{\beta\gamma}}, \cancel{c_\alpha}, c_{\gamma\alpha}, \cancel{c_{\beta\alpha}} \} \end{aligned}$$

**Leading Order**

**Next-to-Leading Order**

Only 4 parameters left!  
(tracers)

# LSS Bootstrap

Kernel at third order

$$\begin{aligned} K_3(\mathbf{q}_1, \mathbf{q}_2, \mathbf{q}_3) = & c_2 + c_{\gamma 1} \gamma(\mathbf{q}_1, \mathbf{q}_2) + c_{\gamma 2} \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\beta 1} \beta(\mathbf{q}_1, \mathbf{q}_2) + c_{\beta 2} \beta(\mathbf{q}_{12}, \mathbf{q}_3) \\ & + c_{\gamma\gamma} \gamma(\mathbf{q}_1, \mathbf{q}_2) \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\beta\beta} \beta(\mathbf{q}_1, \mathbf{q}_2) \beta(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\gamma\beta} \gamma(\mathbf{q}_1, \mathbf{q}_2) \beta(\mathbf{q}_{12}, \mathbf{q}_3) \\ & + c_{\beta\gamma} \beta(\mathbf{q}_1, \mathbf{q}_2) \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + \left( c_\alpha + c_{\gamma\alpha} \gamma(\mathbf{q}_1, \mathbf{q}_2) + c_{\beta\alpha} \beta(\mathbf{q}_1, \mathbf{q}_2) \right) \alpha_a(\mathbf{q}_{12}, \mathbf{q}_3) \end{aligned}$$

$$\{ \cancel{c_2}, \cancel{c_{\gamma 1}}, \cancel{c_{\gamma 2}}, \cancel{c_{\beta 1}}, \cancel{c_{\beta 2}}, c_{\gamma\gamma}, \cancel{c_{\beta\beta}}, \cancel{c_{\gamma\beta}}, \cancel{c_{\beta\gamma}}, \cancel{c_\alpha}, c_{\gamma\alpha}, \cancel{c_{\beta\alpha}} \}$$

**Leading Order**

**Next-to-Leading Order**

**Mass+mom. conservation**

Only 4 parameters left!  
(tracers)

# LSS Bootstrap

Kernel at third order

$$\begin{aligned} K_3(\mathbf{q}_1, \mathbf{q}_2, \mathbf{q}_3) = & c_2 + c_{\gamma 1} \gamma(\mathbf{q}_1, \mathbf{q}_2) + c_{\gamma 2} \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\beta 1} \beta(\mathbf{q}_1, \mathbf{q}_2) + c_{\beta 2} \beta(\mathbf{q}_{12}, \mathbf{q}_3) \\ & + c_{\gamma\gamma} \gamma(\mathbf{q}_1, \mathbf{q}_2) \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\beta\beta} \beta(\mathbf{q}_1, \mathbf{q}_2) \beta(\mathbf{q}_{12}, \mathbf{q}_3) + c_{\gamma\beta} \gamma(\mathbf{q}_1, \mathbf{q}_2) \beta(\mathbf{q}_{12}, \mathbf{q}_3) \\ & + c_{\beta\gamma} \beta(\mathbf{q}_1, \mathbf{q}_2) \gamma(\mathbf{q}_{12}, \mathbf{q}_3) + \left( c_\alpha + c_{\gamma\alpha} \gamma(\mathbf{q}_1, \mathbf{q}_2) + c_{\beta\alpha} \beta(\mathbf{q}_1, \mathbf{q}_2) \right) \alpha_a(\mathbf{q}_{12}, \mathbf{q}_3) \end{aligned}$$

$$\{ \cancel{c_2}, \cancel{c_{\gamma 1}}, \cancel{c_{\gamma 2}}, \cancel{c_{\beta 1}}, \cancel{c_{\beta 2}}, c_{\gamma\gamma}, \cancel{c_{\beta\beta}}, \cancel{c_{\gamma\beta}}, \cancel{c_{\beta\gamma}}, \cancel{c_\alpha}, c_{\gamma\alpha}, \cancel{c_{\beta\alpha}} \}$$

**Leading Order**

Only 4 parameters left!  
(tracers)

**Next-to-Leading Order**

Only 2 parameters left!  
(matter)

**Mass+mom. conservation**



# LSS Bootstrap

## General time dependence

Bootstrap expansion:  
valid for every scale-  
independent model

$$-\frac{k^2}{\mathcal{H}^2}\phi(\mathbf{k}, \eta) = \frac{3}{2}\Omega_m(\eta)\mu(\eta)\delta(\mathbf{k}, \eta) + S(\mathbf{k}, \eta)$$

$\Lambda$ CDM (exact time dep., Donath Y., Senatore L., 2020)

nDGP (Dvali et al. 2000, Schmidt F. 2009)

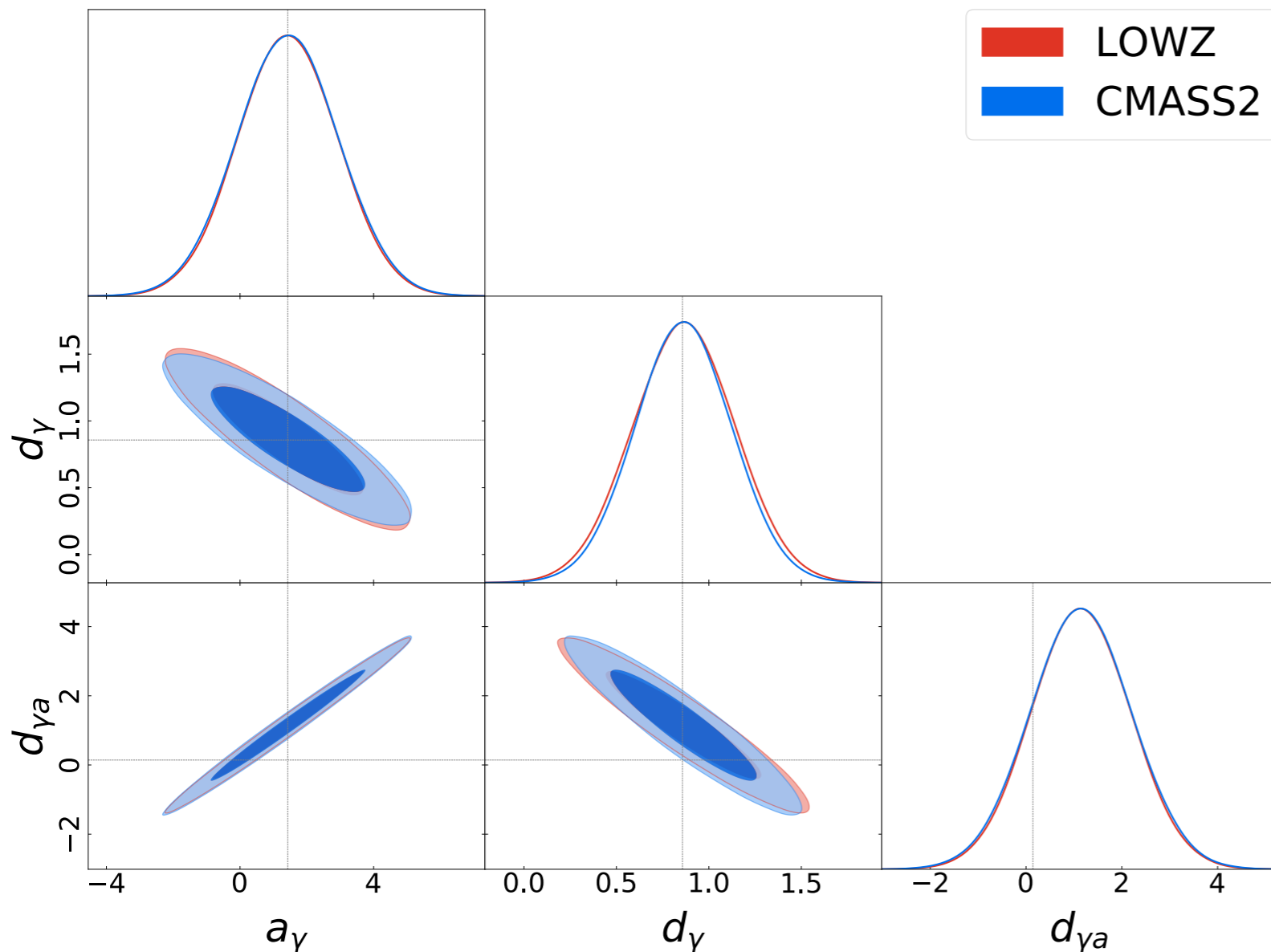
k-mouflage (Babichev et al. 2009, Brax and Valaegias 2014)

JBD (Brans C., Dicke R., 1961)

D'Amico G., MM, Piga L., Pietroni M., Vernizzi F., Wright B.,  
2209.XXXX

# LSS Bootstrap

## Matter and velocity kernels



In the 1-loop PS the only time dependent functions are

$$a_\gamma, d_\gamma, d_{\gamma a} \equiv d_{\gamma\alpha} - d_{\gamma\gamma}/2$$

Constrain (possible) deviations from  $\Lambda$ CDM with a model-independent approach

# LSS Bootstrap & PNG

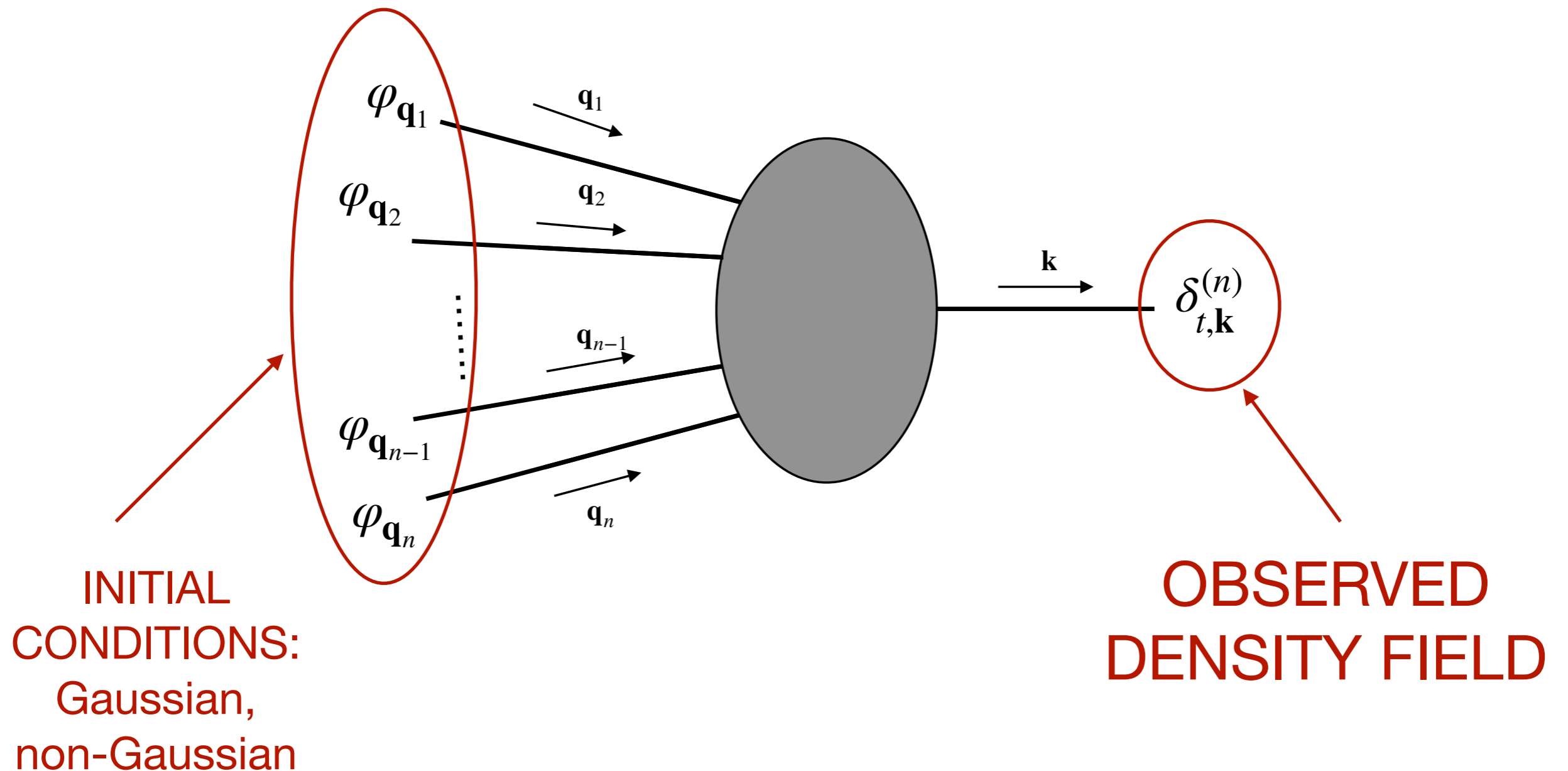
## Consistency relations

$$\lim_{k \rightarrow 0} \frac{B_\delta(k, q, |\mathbf{k} + \mathbf{q}|; \tau, \tau', \tau'')}{P_\delta(q; \tau', \tau'')P_\delta(k; \tau, \tau)} = -\frac{\mathbf{q} \cdot \mathbf{k}}{k^2} \frac{D_+(\tau') - D_+(\tau'')}{D_+(\tau)} + \frac{6f_{NL}\Omega_{m,0}H_0^2}{k^2 T(k)} \frac{D_+(\tau_0)}{D_+(\tau)} + \mathcal{O}(k^0, f_{NL}^2)$$

Peloso M., Pietroni M., *JCAP* 05 (2013) 031

Goldstein S., et al., 2209.06228

# LSS Bootstrap & PNG



# Conclusions

- Importance of Equivalence Principle in Cosmology
- Perturbative EP: LSS bootstrap
- Easily generalized to NG IC
- First step for a novel approach to Galaxy Clustering

Thanks for your attention

Backup slides

# In this thesis

## Equivalence principle

```
graph TD; EP[Equivalence principle] --> BAO[Model independent measurement using the EP in the BAO range]; EP --> Bootstrap[Constraints the perturbative kernels in a model independent way: Large Scale Structure Bootstrap!];
```

Model independent measurement using the EP in the BAO range

- MM, Nishimichi T., Pietroni M., Phys. Rev. D100 (2019)
- MM, Nishimichi T., Pietroni M., JCAP 07 (2020)

Constraints the perturbative kernels in a model independent way: **Large Scale Structure Bootstrap!**

- D'Amico G., MM, Pietroni M., Vernizzi F., JCAP 10 (2021)



# CR and BAO

- MM, Nishimichi T., Pietroni M., Phys. Rev. D100 (2019)
- MM, Nishimichi T., Pietroni M., JCAP 07 (2020)

Squeezed limit of **Bispectrum** (real space)

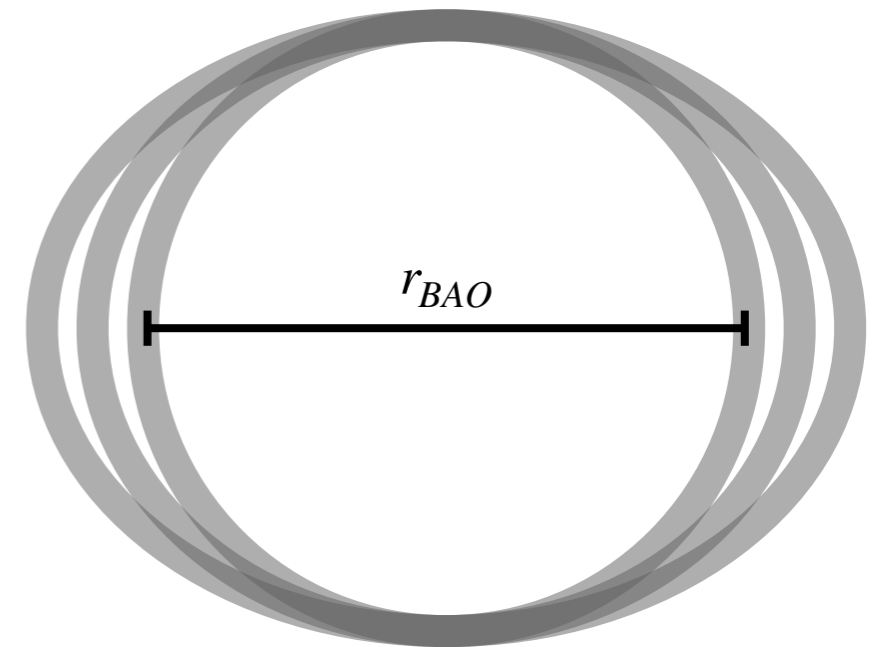
$$\lim_{q/k \rightarrow 0} B_{\alpha\beta\gamma}(q, k_+, k_-; t_\alpha, t_\beta, t_\gamma) \simeq \frac{\mathbf{k} \cdot \mathbf{q}}{q^2} P_{\alpha m}(q; t_\alpha) \left[ \frac{D(t_\beta)}{D(t_\alpha)} P_{\beta\gamma}(k_-; t_\beta, t_\gamma) - \frac{D(t_\gamma)}{D(t_\alpha)} P_{\beta\gamma}(k_+; t_\beta, t_\gamma) \right] + O\left(\left(\frac{q}{k}\right)^0\right)$$

Peloso M. and Pietroni M., JCAP (2013)

Kehagias A. and Riotto A., Nucl. Phys. (2013)

In configuration space  $\propto \mu^2 \frac{d \log \xi(r)}{d \log r}$

$$\mu = \hat{\mathbf{k}} \cdot \hat{\mathbf{q}}$$



# CR and BAO

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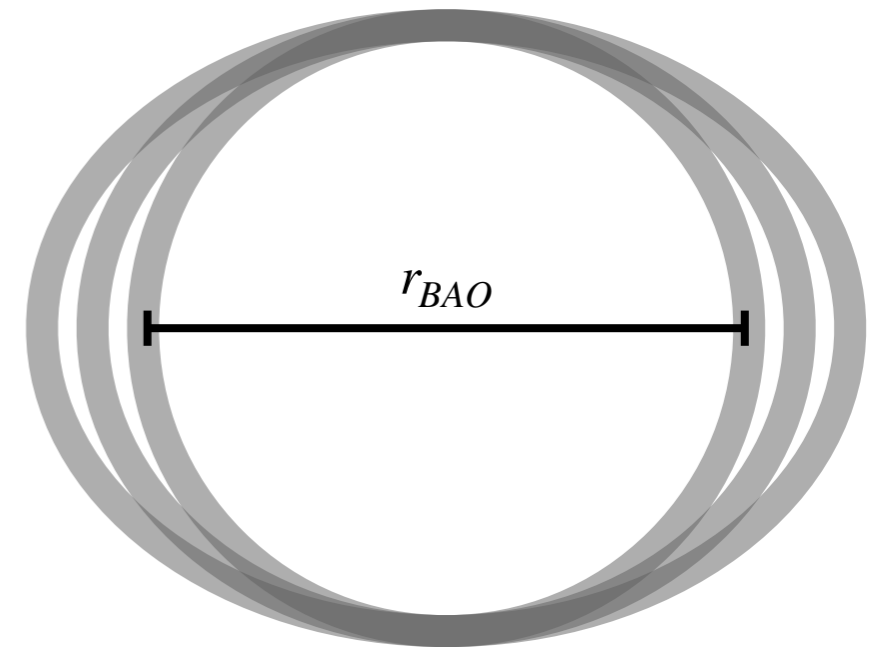
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Equal-time squeezed limit



# CR and BAO

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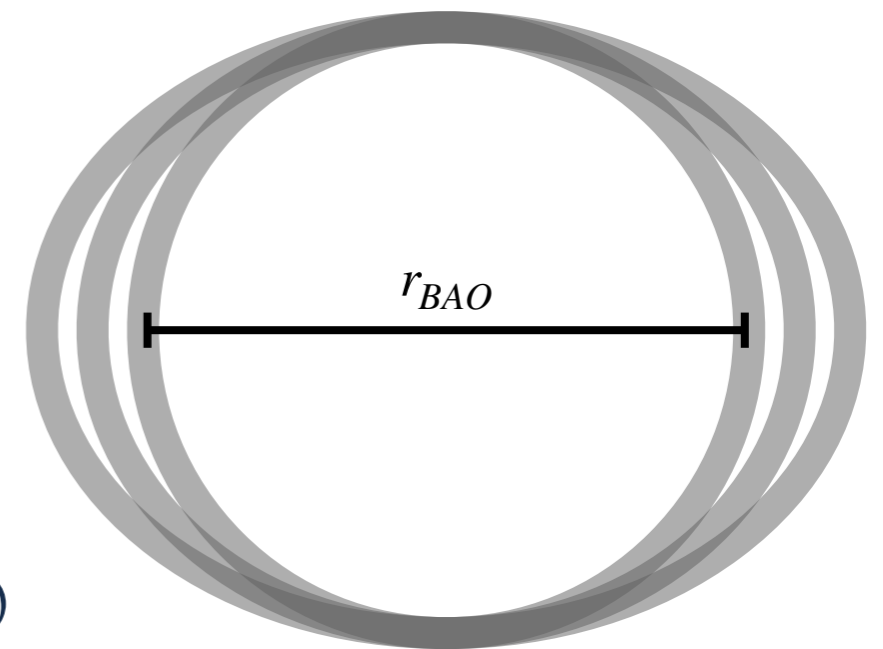
$$\mu = \hat{\mathbf{k}} \cdot \hat{\mathbf{q}}$$

## Equal-time squeezed limit

$$\lim_{q/k \rightarrow 0} \frac{B_{\alpha\alpha\alpha}(q, k_+, k_-)}{P_{\alpha\alpha}(q)P_{\alpha\alpha}(k)} = - \frac{\mu^2}{b_\alpha(q)} \frac{d \log P_{\alpha\alpha}(k)}{d \log k} + O\left(\left(\frac{q}{k}\right)^0\right)$$

Unchanged by nonlinearities!

Baldauf T. et al., Phys.Rev.D (2015)



# CR and BAO

- MM, Nishimichi T., Pietroni M., Phys. Rev. D100 (2019)
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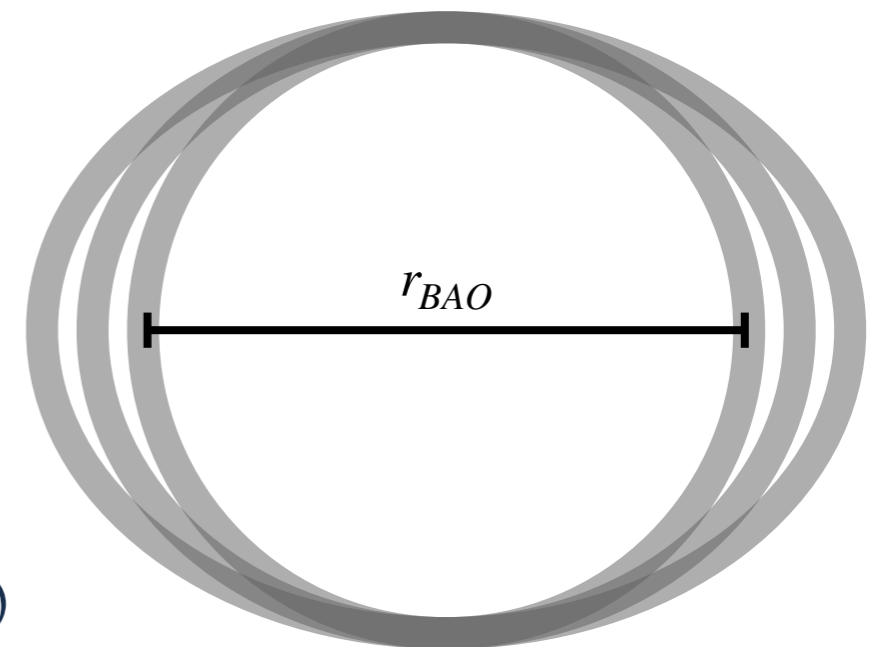
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In presence of a scale like the BAO, the oscillating part of the derivative is enhanced by a  $\sim k r_{BAO}$  factor, we can isolate it to verify CR and measure bias

# CR and BAO

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- MM, Nishimichi T., Pietroni M., JCAP 07 (2020)

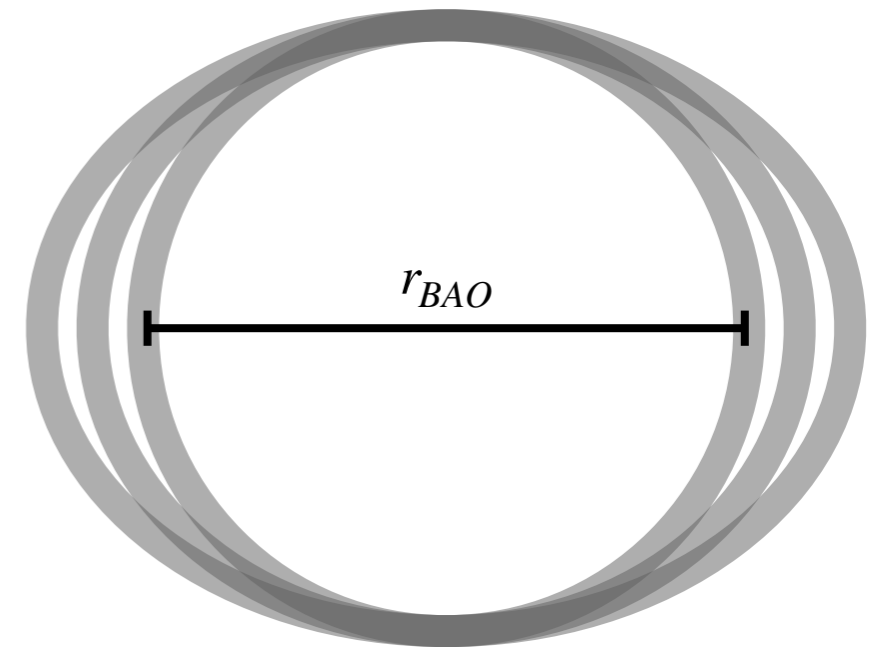
Squeezed limit of **Bispectrum** (real space)

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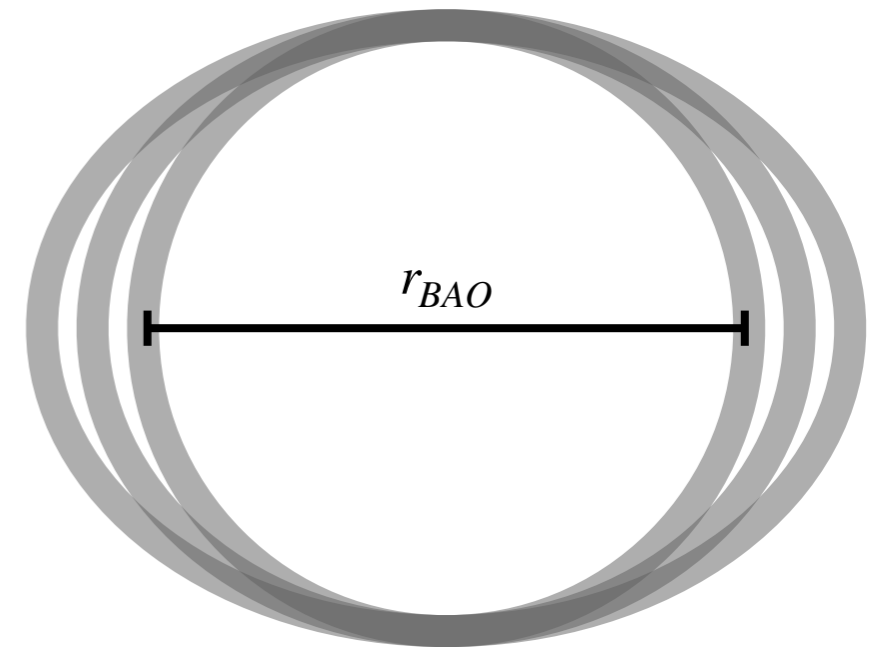
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Equal-time squeezed limit



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Peloso M. and Pietroni M., JCAP (2013)

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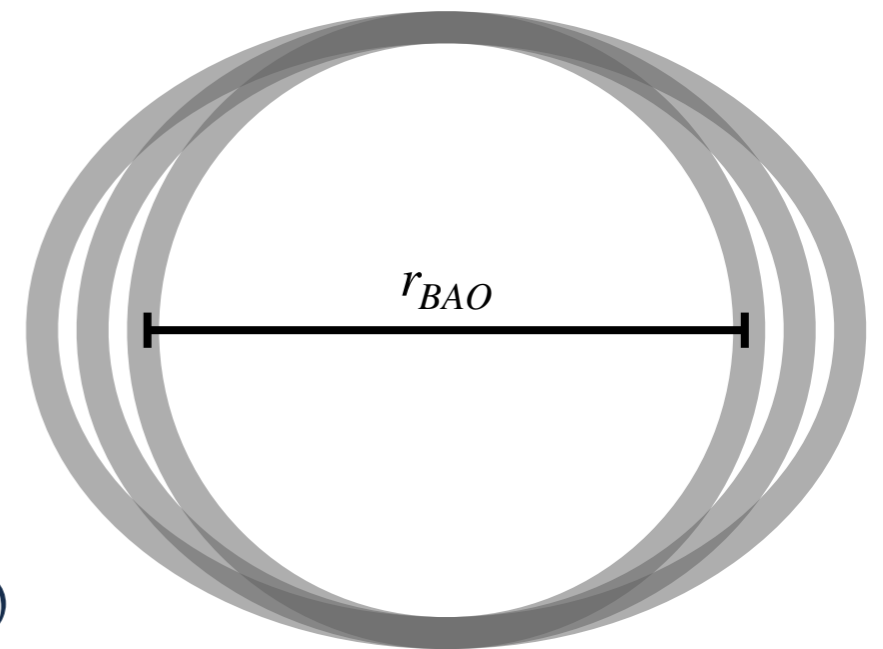
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Unchanged by nonlinearities!

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# CR and BAO

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Squeezed limit of Bispectrum (real space)

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In configuration space  $\propto \int d^3x e^{i\mathbf{q}\cdot\mathbf{x}} \xi(r)$

Equal-time

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$$\lim_{q/k \rightarrow 0} \frac{B_{\alpha\alpha\alpha}(q, k_+, k_-)}{P_{\alpha\alpha}(q)P_{\alpha\alpha}(k)} = \frac{b_\alpha(q)}{d \log k} + O\left(\left(\frac{q}{k}\right)^0\right)$$

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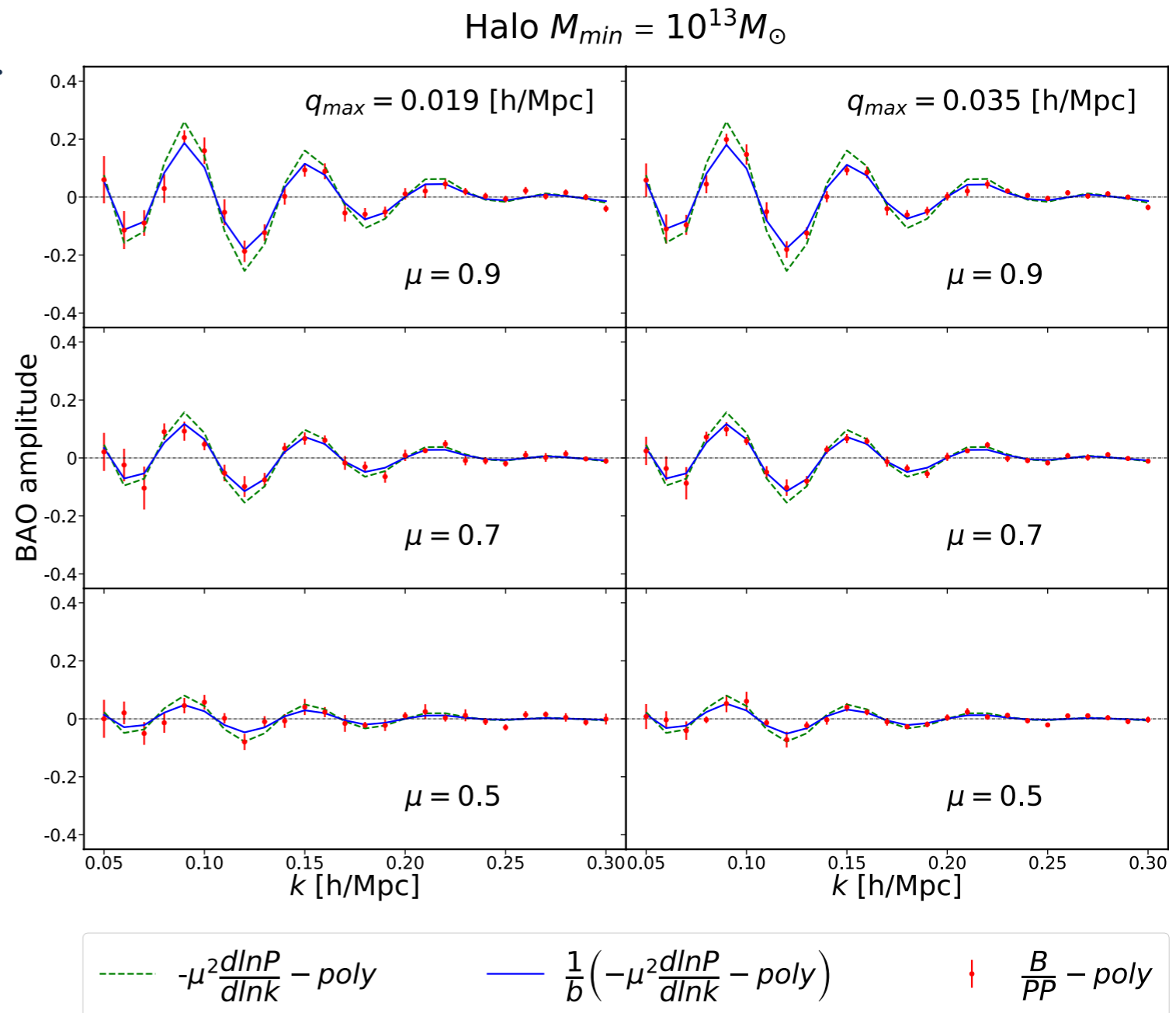
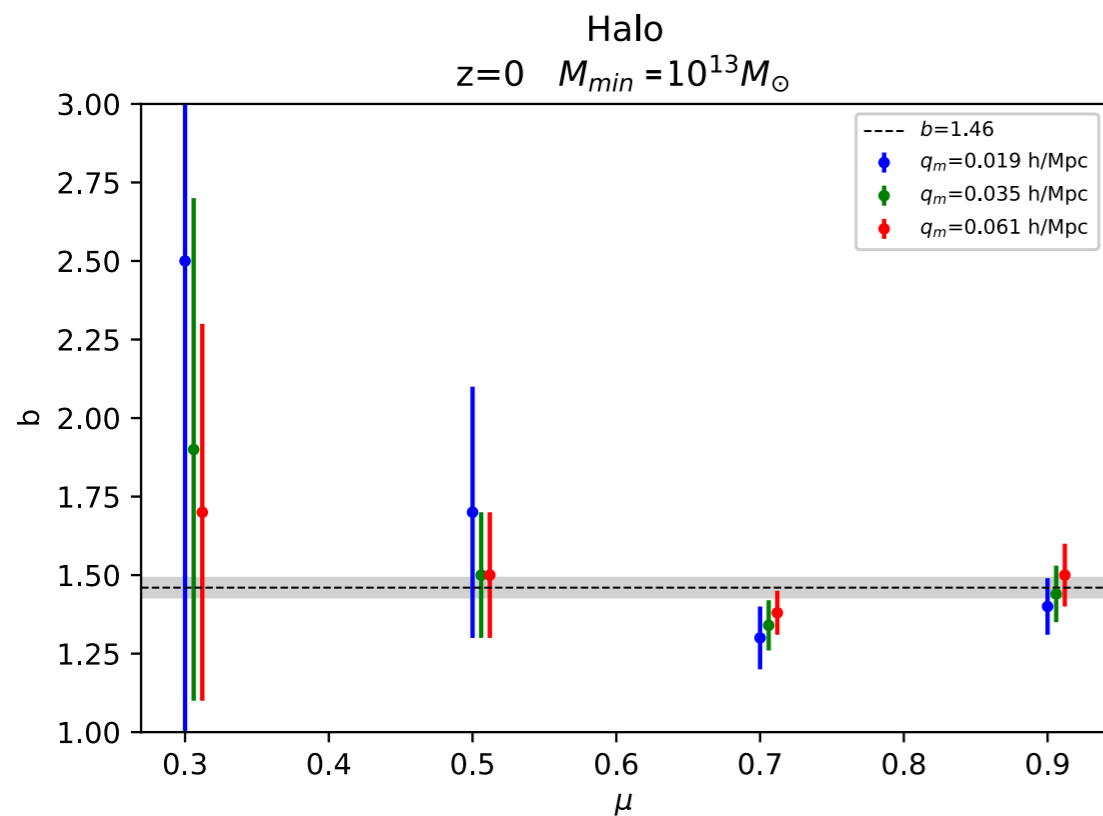


# CR and BAO

N-body simulations: real space w/ biased tracers, MM+ (2019)

$$\lim_{q/k \rightarrow 0} \frac{B_t(q, k_+, k_-)}{P_t(q)P_t(k)} = -\frac{\mu^2}{b_t} \frac{d \log P(k)}{d \log k} + \dots$$

Bias parameter  $b_t = \lim_{q \rightarrow 0} \frac{P_{tt}(q)}{P_{tm}(q)}$



# CR and BAO

N-body simulations: redshift space w/ biased tracers, MM+ (2020)

## Multipoles + Kaiser

$$\lim_{q/k \rightarrow 0} \frac{B_t^{(0)}(q, k)}{P_t^{(0)}(q)P_t^{(0)}(k)} = - \left[ \frac{1}{3b_t} + \frac{b_t - 1}{9b_t} \beta_t \frac{1 + \frac{3}{5}\beta_t}{1 + \frac{2}{3}\beta_t + \frac{1}{5}\beta_t^2} \right] \frac{\partial \log P_t^{(0)}(k)}{\partial \log k} + \dots$$

$$\lim_{q/k \rightarrow 0} \frac{B_t^{(l_k=2)}(q, k)}{P_t^{(0)}(q)P_t^{(0)}(k)} = - \frac{2\beta_t}{45b_t} \frac{2 + b_t(5 + 3\beta_t)}{1 + \frac{2}{3}\beta_t + \frac{1}{5}\beta_t^2} \frac{\partial \log P_t^{(0)}(k)}{\partial \log k} + \dots$$

$$\lim_{k \rightarrow 0} \frac{P^{(2)}(k)}{P_t^{(0)}(k)} = \frac{4\beta_t}{21} \frac{7 + 3\beta_t}{1 + \frac{2}{3}\beta_t + \frac{1}{5}\beta_t^2}$$

Angles

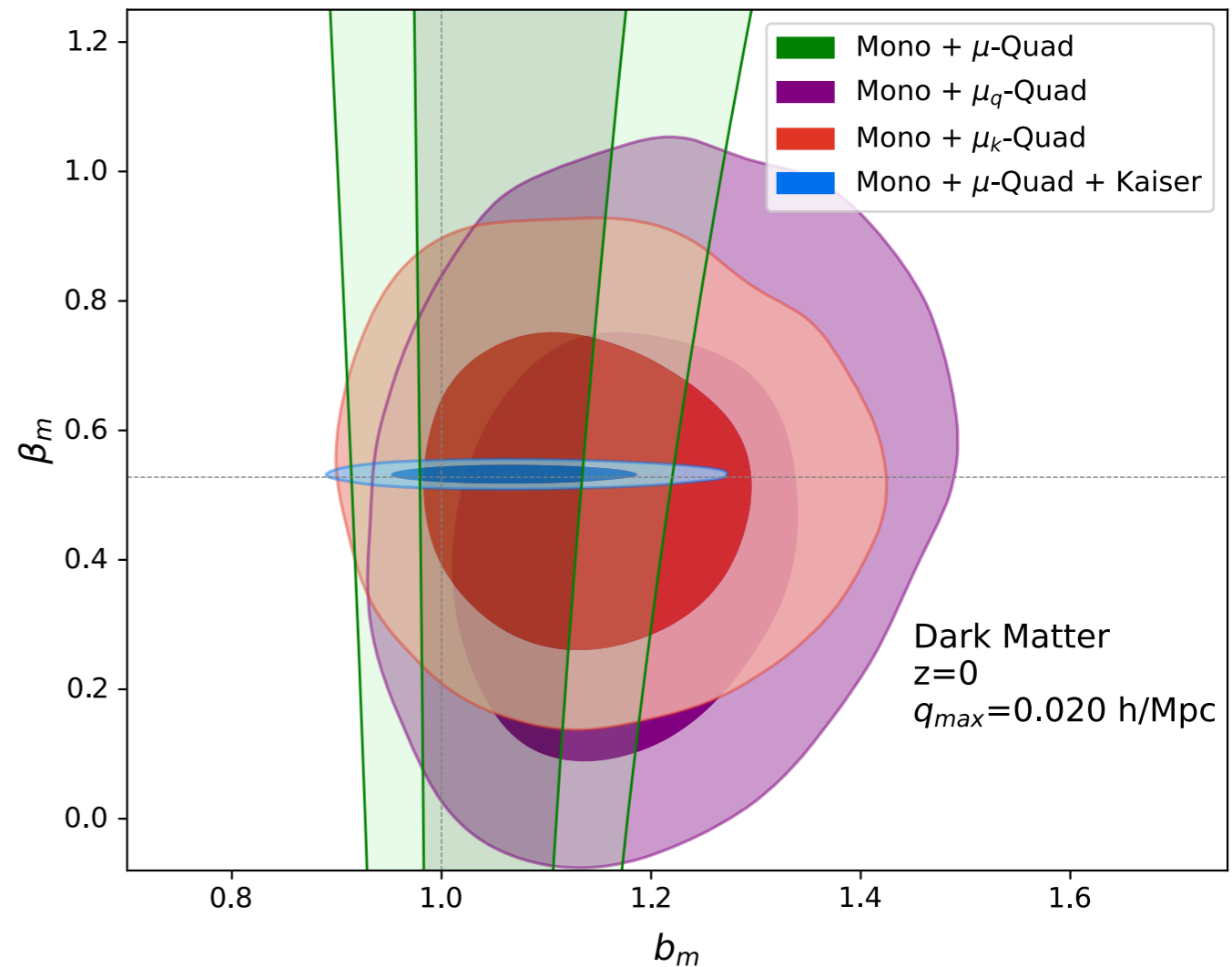
$$\mu \equiv \hat{\mathbf{k}} \cdot \hat{\mathbf{q}}$$

$$\mu_k \equiv \hat{\mathbf{k}} \cdot \hat{\mathbf{z}}$$

$$\mu_q \equiv \hat{\mathbf{q}} \cdot \hat{\mathbf{z}}$$

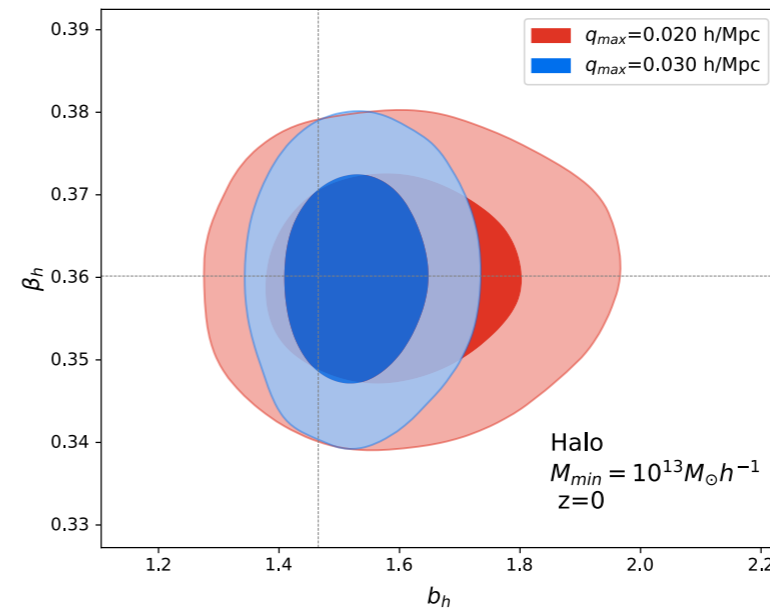
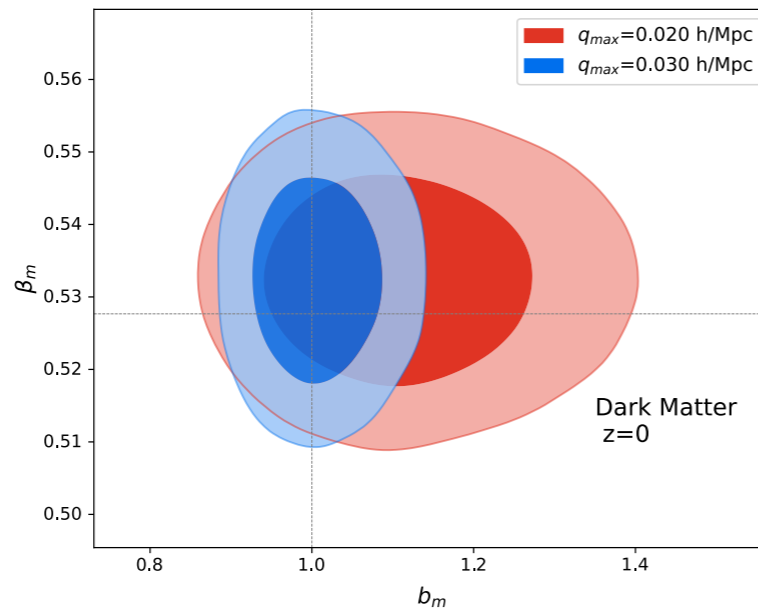
Growth rate  $f \equiv \frac{d \log D}{d \log a}$

$$\beta_t \equiv \frac{f}{b_t}$$



# CR and BAO

N-body simulations: redshift space w/ biased tracers, MM+ (2020)

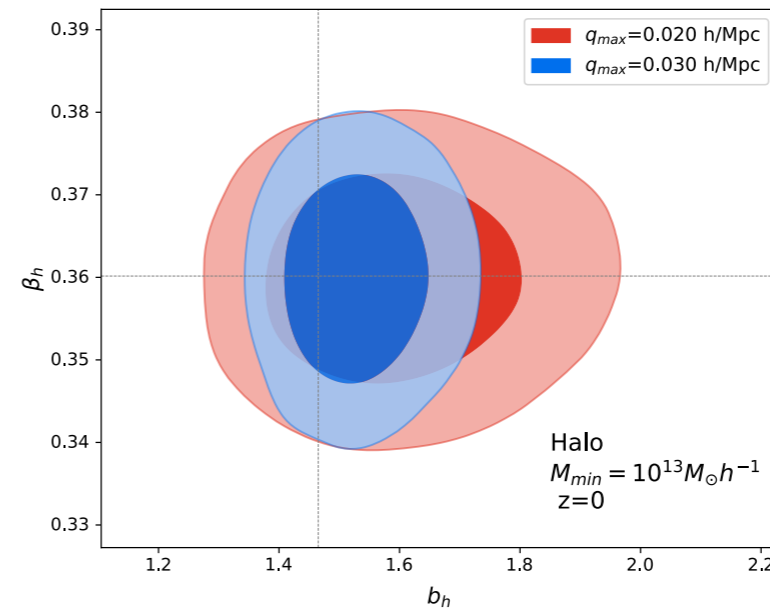
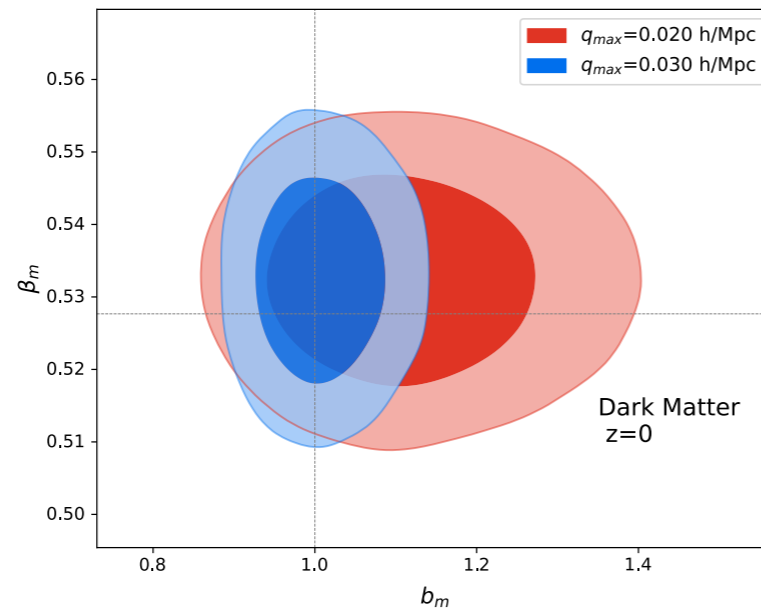


$M_{\min} = 10^{13} h^{-1} M_{\odot} \quad z = 0$				
$q_{\max}$ (h/Mpc)	$b_h$	$b_h^{fid}$	$f = \beta_h b_h$	$f^{fid}$
0.020	$1.58^{+0.15}_{-0.13}$	1.47	$0.57^{+0.06}_{-0.06}$	0.528
0.030	$1.53^{+0.08}_{-0.08}$	1.47	$0.55^{+0.04}_{-0.04}$	0.538
$M_{\min} = 10^{13} h^{-1} M_{\odot} \quad z = 1$				
$q_{\max}$ (h/Mpc)	$b_h$	$b_h^{fid}$	$f = \beta_h b_h$	$f^{fid}$
0.020	$2.85^{+0.39}_{-0.32}$	2.686	$0.93^{+0.14}_{-0.14}$	0.877
0.030	$2.58^{+0.17}_{-0.16}$	2.686	$0.82^{+0.07}_{-0.07}$	0.877
$M_{\min} = 10^{14} h^{-1} M_{\odot} \quad z = 0$				
$q_{\max}$ (h/Mpc)	$b_h$	$b_h^{fid}$	$f = \beta_h b_h$	$f^{fid}$
0.020	$3.40^{+0.83}_{-0.61}$	2.446	$0.73^{+0.18}_{-0.18}$	0.528
0.030	$2.29^{+0.21}_{-0.18}$	2.446	$0.49^{+0.06}_{-0.06}$	0.528

**Model independent measurement of  $f(z)$  and  $b_t$  at 10%!**

# CR and BAO

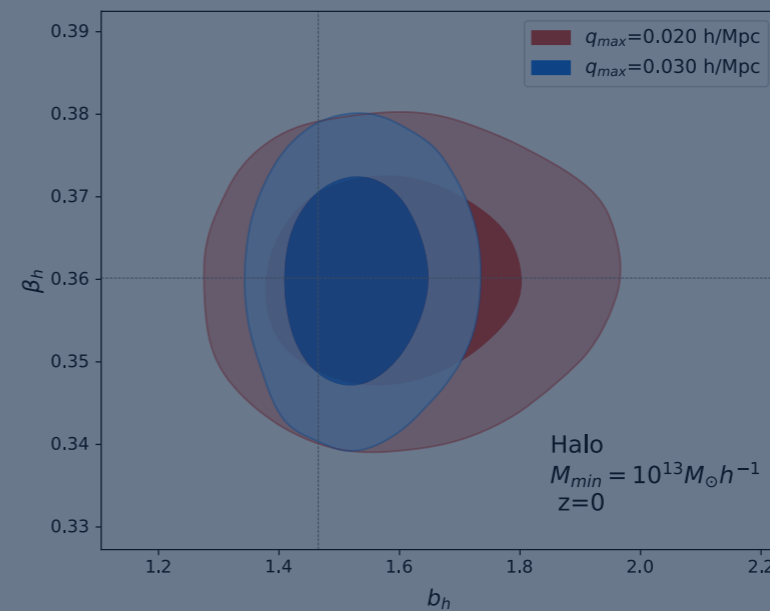
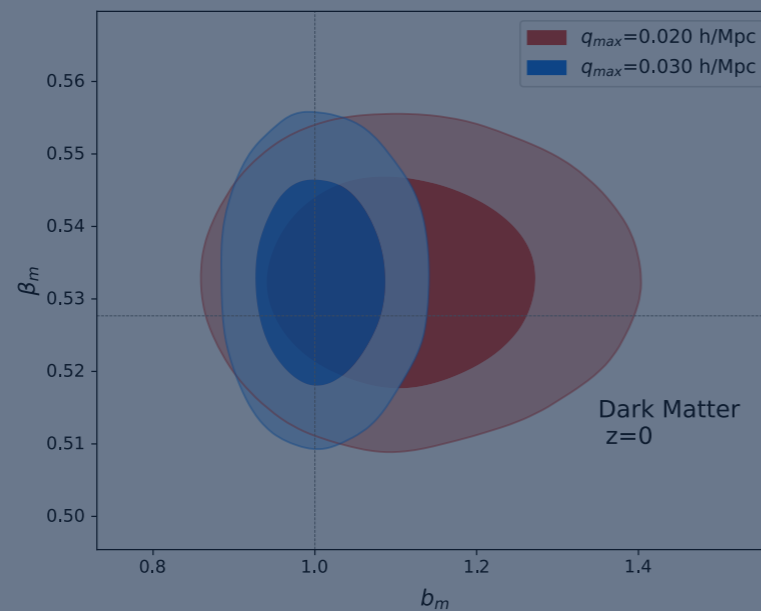
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$q_{\max}$ (h/Mpc)	$b_h$	$b_h^{fid}$	$f = \beta_h b_h$	$f^{fid}$
0.020	$1.58^{+0.15}_{-0.13}$	1.47	$0.57^{+0.06}_{-0.06}$	0.528
0.030	$1.53^{+0.08}_{-0.08}$	1.47	$0.55^{+0.04}_{-0.04}$	0.538
$M_{\min} = 10^{13} h^{-1} M_{\odot} \quad z = 1$				
$q_{\max}$ (h/Mpc)	$b_h$	$b_h^{fid}$	$f = \beta_h b_h$	$f^{fid}$
0.020	$2.85^{+0.39}_{-0.32}$	2.686	$0.93^{+0.14}_{-0.14}$	0.877
0.030	$2.58^{+0.17}_{-0.16}$	2.686	$0.82^{+0.07}_{-0.07}$	0.877
$M_{\min} = 10^{14} h^{-1} M_{\odot} \quad z = 0$				
$q_{\max}$ (h/Mpc)	$b_h$	$b_h^{fid}$	$f = \beta_h b_h$	$f^{fid}$
0.020	$3.40^{+0.83}_{-0.61}$	2.446	$0.73^{+0.18}_{-0.18}$	0.528
0.030	$2.29^{+0.21}_{-0.18}$	2.446	$0.49^{+0.06}_{-0.06}$	0.528

# CR and BAO

N-body simulations: redshift space w/ biased tracers, MM+ (2020)



$M_{\min} = 10^{13} h^{-1} M_{\odot} \quad z = 0$				
$q_{\max}$ (h/Mpc)	$b_h$	$b_h^{fid}$	$f = \beta_h b_h$	$f^{fid}$
0.020	$1.58^{+0.15}_{-0.13}$	1.47	$0.57^{+0.06}_{-0.06}$	0.528
0.030	$1.53^{+0.08}_{-0.08}$	1.47	$0.55^{+0.04}_{-0.04}$	0.538
$M_{\min} = 10^{13} h^{-1} M_{\odot} \quad z = 1$				
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**Model independent measurement of  $f(z)$  and  $b_t$  at 10%!**