Hands-on Session

 γ -ray emission with HERMES [https://arxiv.org/abs/2105.13165]

ISAPP School $MAD^{(\gamma)}$ 2021 Gamma rays to shed light on Dark Matter 29, June 2021

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TZ RESEARCH FOR





Simulating the Galactic Multi-messenger Emissions with HERMES

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ABSTRACT

Context. The study of non-thermal processes such as synchrotron emission, inverse Compton scattering, bremsstrahlung and pion production is crucial to understand the properties of the Galactic cosmic-ray population, to shed light on their origin and confinement mechanisms, and to assess the significance of exotic signals possibly associated to new physics.

Aims. We present a public code called HERMES aimed at generating sky maps associated to a variety of multi-messenger and multi-wavelength radiative processes, spanning from the radio domain all the way up to high-energy gamma-ray and neutrino production. *Methods.* We describe the physical processes under consideration, the code concept and structure, and the user interface, with particular focus on the python-based interactive mode. We especially present the modular and flexible design that allows to easily further extend the numerical package according to the user's needs.

Results. In order to demonstrate the capabilities of the code, we describe in detail a comprehensive set of sky maps and spectra associated to all physical processes included in the code. We comment in particular on the radio, gamma-ray, and neutrino maps, and mention the possibility to study signals stemming from dark matter annihilation.

Conclusions. **HERMES** can be successfully applied to constrain the properties of the Galactic cosmic-ray population, improve our understanding of the diffuse Galactic radio, gamma–ray, and neutrino emission, and search for signals associated to particle dark matter annihilation or decay.

γ -ray emitting processes

Resulting energy bands [https://heasarc.gsfc.nasa.gov/docs/heasarc/headates/spectrum.html]

	CR involved	Target	Secondary ID	Secondary E		
Synchrotron		B -field		radio band		
Brems	e^{\pm}	ISM gas		broadband		
ICS		ISRF		high γ -rays		
$\pi^0 \to \gamma \gamma$	n He nuclei	ISM and h		high γ -rays		
$\pi^{\pm} ightarrow \mu_{\pm} {}^{`} \overline{ u}{}^{'}{}_{\mu}$	<i>p</i> , 11c, 11delet	decav	1/	high-energy ν		
$\mu^{\pm} ightarrow e^{\pm i} \overline{ u}_{\mu}^{\nu} \overline{ u}_{e}^{\nu}$	sec μ	liceuy		broadband		

 $E \lesssim \mathcal{O}(100) \,\mathrm{GHz}$

E > MeV

E > GeV





Gas longitudinal profile



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Introduction to γ -rays with HERMES

HERMES work-flow



Ring model

[Strong&Mattox, A&A, 308, L21 (1996)] \rightarrow based on HI and CO emissivity maps



The observable quantity

 γ -ray intensity $I_{\gamma} \equiv I_{\gamma}(l, b, E_{\gamma})$

Emissivity at the source

Source

$$\epsilon_{\pi^{0}}(E_{\gamma},\mathbf{r}) = 4\pi n_{\mathrm{H}}(\mathbf{r}) \int dE \left[\frac{d\Phi_{\mathrm{p}}}{dE}(E,\mathbf{r}) \left(\frac{d\sigma_{\mathrm{p-p}}}{dE_{\gamma}} + f_{\mathrm{He}} \frac{d\sigma_{\mathrm{He-p}}}{dE_{\gamma}} \right) + \frac{d\Phi_{\mathrm{He}}}{dE}(E,\mathbf{r}) \left(\frac{d\sigma_{\mathrm{p-He}}}{dE_{\gamma}} + f_{\mathrm{He}} \frac{d\sigma_{\mathrm{He-He}}}{dE_{\gamma}} \right) \right]$$

The observable quantity

 γ -ray intensity $I_{\gamma} \equiv I_{\gamma}(l, b, E_{\gamma})$

Emissivity at the source

$$\epsilon_{\pi^{0}}(E_{\gamma},\mathbf{r}) = 4\pi n_{\mathrm{H}}(\mathbf{r}) \int dE \left[\frac{d\Phi_{\mathrm{p}}}{dE}(E,\mathbf{r}) \left(\frac{d\sigma_{\mathrm{p-p}}}{dE_{\gamma}} + f_{\mathrm{He}} \frac{d\sigma_{\mathrm{He-p}}}{dE_{\gamma}} \right) + \frac{d\Phi_{\mathrm{He}}}{dE}(E,\mathbf{r}) \left(\frac{d\sigma_{\mathrm{p-He}}}{dE_{\gamma}} + f_{\mathrm{He}} \frac{d\sigma_{\mathrm{He-He}}}{dE_{\gamma}} \right) \right]$$

Source

$$\left\langle \epsilon_{\pi^{0}}(E_{\gamma},\mathbf{r})\right\rangle^{i} = \frac{\int d\mathbf{r} \,\epsilon_{\pi^{0}}(E_{\gamma},\mathbf{r}) \,p_{\mathrm{HI}}(\mathbf{r}) \,\Theta^{i}(\mathbf{r})}{\int d\mathbf{r} \,p_{\mathrm{HI}}(\mathbf{r}) \,\Theta^{i}(\mathbf{r})}$$

The observable quantity

 γ -ray intensity $I_{\gamma} \equiv I_{\gamma}(l, b, E_{\gamma})$

Emissivity at the source

$$\epsilon_{\pi^{0}}(E_{\gamma},\mathbf{r}) = 4\pi n_{\mathrm{H}}(\mathbf{r}) \int dE \left[\frac{d\Phi_{\mathrm{p}}}{dE}(E,\mathbf{r}) \left(\frac{d\sigma_{\mathrm{p-p}}}{dE_{\gamma}} + f_{\mathrm{He}} \frac{d\sigma_{\mathrm{He-p}}}{dE_{\gamma}} \right) + \frac{d\Phi_{\mathrm{He}}}{dE}(E,\mathbf{r}) \left(\frac{d\sigma_{\mathrm{p-He}}}{dE_{\gamma}} + f_{\mathrm{He}} \frac{d\sigma_{\mathrm{He-He}}}{dE_{\gamma}} \right) \right]$$

Emissivity averaged over the ring *i*-th $\left\langle \epsilon_{\pi^{0}}(E_{\gamma},\mathbf{r})\right\rangle^{i} = \frac{\int d\mathbf{r} \,\epsilon_{\pi^{0}}(E_{\gamma},\mathbf{r}) \,p_{\mathrm{HI}}(\mathbf{r}) \,\Theta^{i}(\mathbf{r})}{\int d\mathbf{r} \,p_{\mathrm{HI}}(\mathbf{r}) \,\Theta^{i}(\mathbf{r})}$

Observed intensity

$$I_{\gamma}(l,b,E_{\gamma}) = \frac{1}{4\pi} \sum_{i} N_{\mathrm{H}}^{i}(l,b) \left\langle \epsilon_{\pi^{0}}(E_{\gamma},\mathbf{r}) \right\rangle^{i}$$



Source

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The outcome

π^0 -decay at given energy

Pion Decay HI - $E_{\gamma} = 10 \text{ GeV}$ Pion Decay H2 - $E_{\gamma} = 10 \text{ GeV}$ 75 75 50 50 25 25 b [deg] b [deg] 0 0 -25-25-50-50-75-75150 100 50 -50-100-150150 100 50 -50-100-1500 0 1 [deg] l [deg] -3.5-3.0-2.5-2.0-1.5 $-2.00 \ -1.75 \ -1.50 \ -1.25 \ -1.00 \ -0.75 \ -0.50 \ -0.25$ -1.0 $\log E^2$ Flux [GeV cm⁻² sr⁻¹ s⁻¹] $\log E^2$ Flux [GeV cm⁻² sr⁻¹ s⁻¹]

Support slides

Run HERMES Full-Sky

<u>https://github.com/ottaviofornieri/ISAPP-</u> school-2021_HandsOn-Diffuse_HERMES

> nside = 32 $t \sim 58 \min$

Resolution parameters

HEALPix convention

The number of pixels in a file is computed from its resolution index (Res).

nside = 2^{res} : The number of pixels per side.

npix = 12 nside^2 : The total number of pixels in the map.

HEALPix Pixel Information								
Res	NSide	NPixels	Mean Spacing (deg)	Area (sterad)				
0	1	12	58.6323	1.0471976 X 10 ⁺⁰⁰				
1	2	48	29.3162	2.6179939 X 10 ⁻⁰¹				
2	4	192	14.6581	6.5449847 X 10 ⁻⁰²				
3	8	768	7.3290	1.6362462 X 10 ⁻⁰²				
4	16	3072	3.6645	4.0906154 X 10 ⁻⁰³				
5	32	12288	1.8323	1.0226539 X 10 ⁻⁰³				
6	64	49152	0.9161	2.5566346 X 10 ⁻⁰⁴				
7	128	196608	0.4581	6.3915866 X 10 ⁻⁰⁵				
8	256	786432	0.2290	1.5978967 X 10 ⁻⁰⁵				
9	512	3145728	0.1145	3.9947416 X 10 ⁻⁰⁶				
10	1024	12582912	0.0573	9.9868541 X 10 ⁻⁰⁷				

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Progress of a full-sky run with nside = 32

12 log-spaced energy points for each process

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	50	# Lepto	hermes::Integrator::initCacheTable: Number of Threads: 8							
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	61	skymapI	<pre>isrf_range_compute()</pre>							
	63	nameIC2	= SkyMapsOutputsFolder + 'BaseRunIC isrf nside' + str(nside)							
	64	skymapI	<pre>C_isrf_range.save(outputs.HEALPixFormat("!{}.fits.gz".format(nameIC2)))</pre>							
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In [*]: 1 # Path to the HERMES output

3 SkyMapsFolder = SkyMapsOutputsFolder + 'Sky_Maps/'

Progress of a full-sky run with nside = 32

12 log-spaced energy points for each process

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	64 skymapl	<pre>IC_isrf_range.save(outputs.HEALPixFormat("!{}.fits.gz".format(" IC_on_TSRE_done")</pre>	nameIC2)))			
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	68 69					
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In [*	*]: 1 # Path	to the HERMES output #				
	3 SkyMaps	Folder = SkyMapsOutputsFolder + 'Sky_Maps/'				

Progress of a full-sky run with nside = 32

12 log-spaced energy points for each process

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	pi0 H2 done		
	Brems on HI done		
	IC on CMB done		
	IC on ISRF done		
	Total time: 3481.3141323109157 seconds		
In [9]:	<pre>1 # Path to the HERMES output # 2 3 SkyMapsFolder = SkyMapsOutputsFolder + 'Sky Maps/'</pre>		

Cartesian projection of the Full Sky

Paper plot: π^0 on HI

 $E_{\gamma} = 10.0 \text{ GeV}$



 π^0 on $HI + H_2$

 $E_{\gamma} = 10.0 \text{ GeV}$



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Runs HERMES Cygnus-X region

<u>https://github.com/ottaviofornieri/ISAPP-</u> school-2021_HandsOn-Diffuse_HERMES

 $l \in [77^{\circ}, 82^{\circ}], b \in [-1^{\circ}, 4^{\circ}]$ nside = 512, 1024, 2048 $t \leq 30 \text{ min}$





Rectangular mask: cross-check $l \in [60^\circ, 75^\circ], b \in [-30^\circ, 15^\circ]$



Diffuse emission from DRAGON CR distribution $E_{\gamma} = 10 \,\text{GeV}$



Integral map for HI and H_2 10 GeV < E_{γ} < 10 TeV, $\Delta \phi \approx 0.057^{\circ}$

