

Durham  
University



# SEARCHING FOR LIGHT PHYSICS AT THE LHC

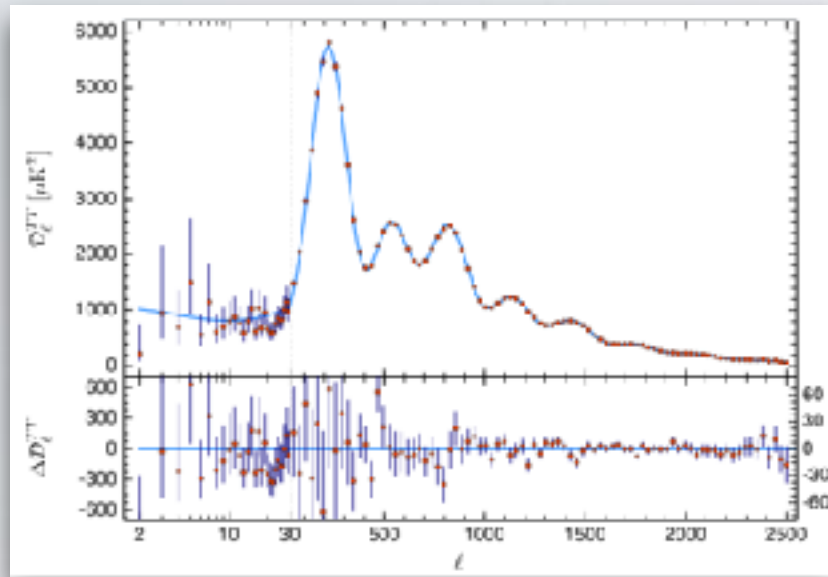
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Patrick Foldenauer

NuTs extended workshop – Jun 1, 2022

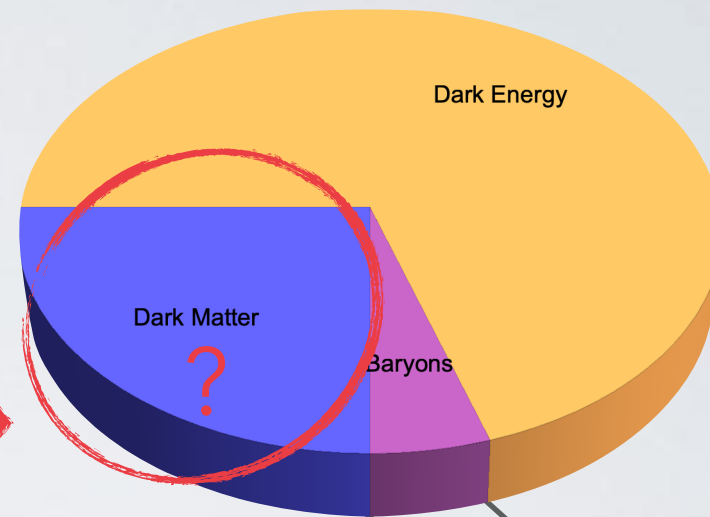
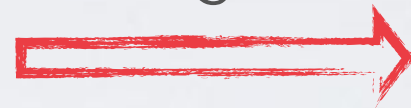
# WHERE TO LOOK BEYOND SM?

Two obvious targets:

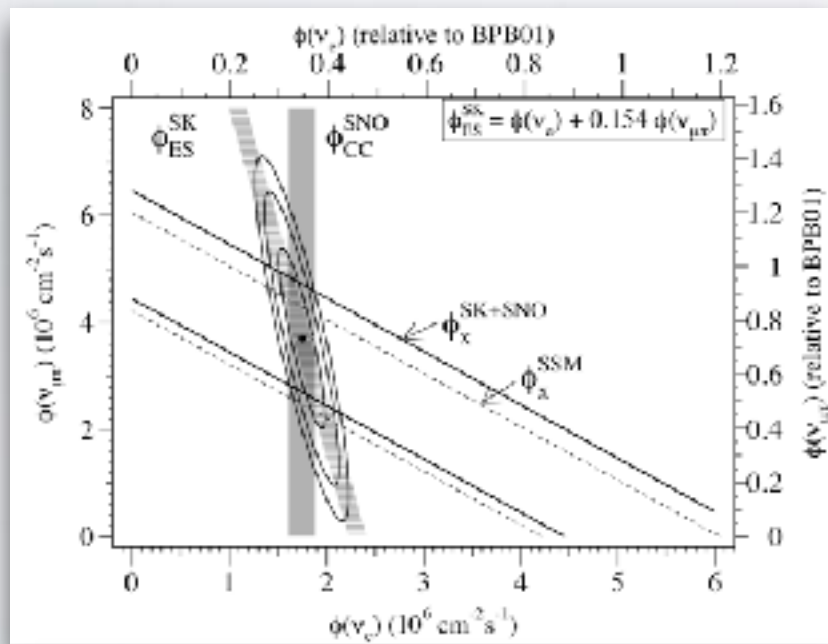


[Planck Collaboration; 1807.06209]

CMB informs us about energy budget

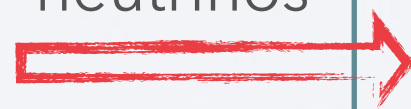


Part I



[SNO Collaboration PRL 87:071301]

Oscillations require massive neutrinos



	I	II	III	
mass	2.4 MeV	1.27 GeV	171.2 GeV	
charge	$2/3$	$2/3$	$2/3$	0
name	<b>u</b> Left up Right	<b>c</b> Left charm Right	<b>t</b> Left top Right	<b>g</b> gluon
	<b>d</b> Left down Right	<b>s</b> Left strange Right	<b>b</b> Left bottom Right	0 <b><math>\gamma</math></b> photon
Quarks	0 eV <b><math>\nu_e</math></b> electron neutrino	0 eV <b><math>\nu_{\mu}</math></b> muon neutrino	0 eV <b><math>\nu_{\tau}</math></b> tau neutrino	91.2 GeV <b>Z</b> weak force
	0.511 MeV <b>e</b> Left electron Right	105.7 MeV <b><math>\mu</math></b> Left muon Right	1.777 GeV <b><math>\tau</math></b> Left tau Right	80.4 GeV <b>W</b> $\pm$ weak force
Leptons				>114 GeV <b>H</b> Higgs boson spin 0
				Bosons (Forces) spin 1

[Gninenko et al., 1301.5516]

Part II

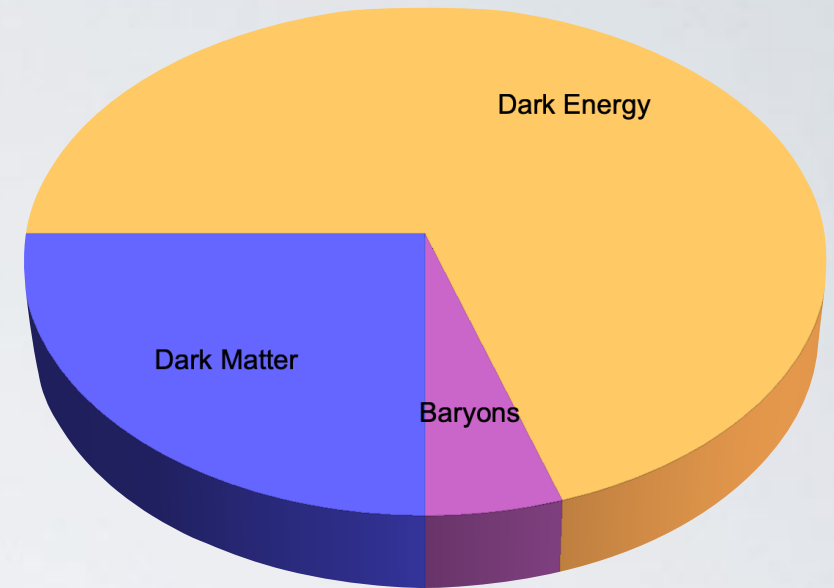
What can we learn about these @ **LHC**?

# ULTRALIGHT DM

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# QUICK REMINDER ABOUT DM

1. Stable, cold, (almost) collisionless, dissipationless substance
2. Interacts (only?) gravitationally
3. Makes up ~25 % of the energy density of the universe
4. Mass ???



[Niikura et al., Nat. Astr. 3 (2019) 6]

Galaxy formation

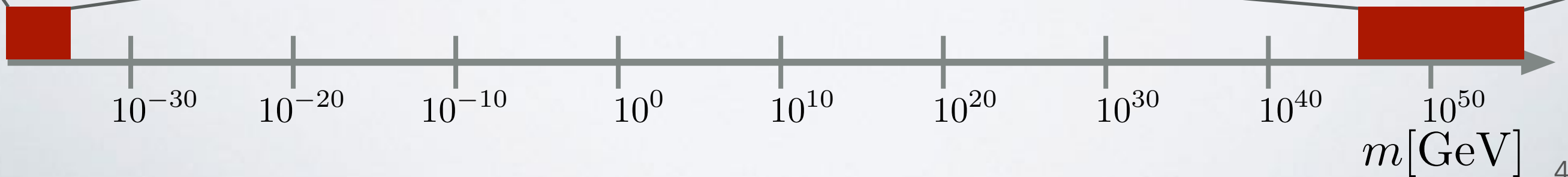
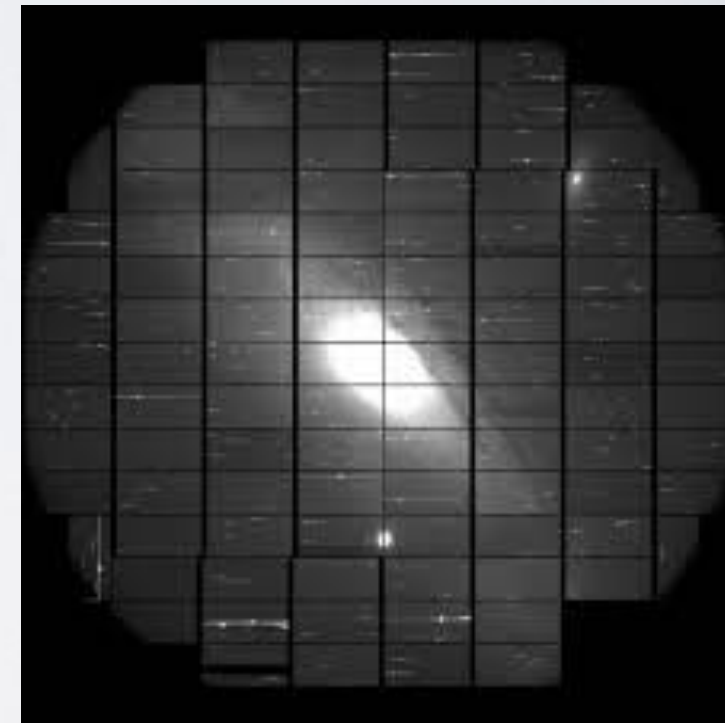
$$\lambda_{dB} = \frac{2\pi}{mv} \lesssim 100 \text{ kpc}$$

$$m \gtrsim 10^{-24} \text{ eV}$$

[Hlozek et al., PRD **91** (2015)]

microlensing searches of PBHs

$$m \lesssim 10^{46} \text{ GeV}$$





# THE FUZZY DM PARADIGM

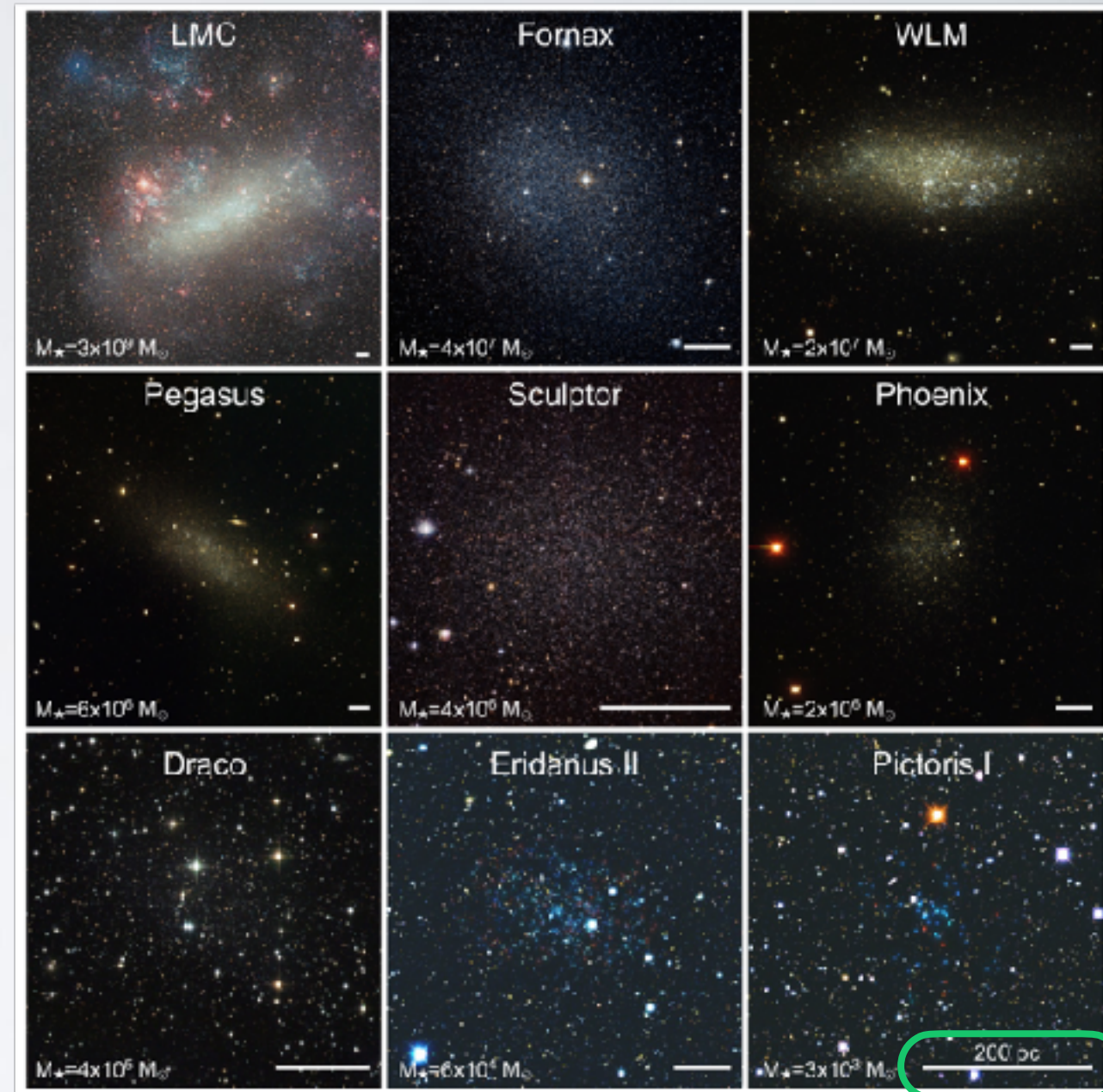
## Dwarf galaxies

- Standard CDM typically produces too much small scale structure
- Can be suppressed if DM de Broglie wavelength prohibits small scale structures:

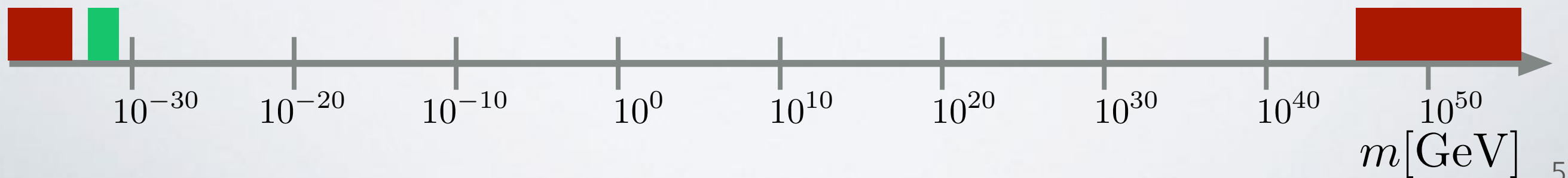
$$m_{\text{DM}} \approx 10^{-22} \text{ eV} \Rightarrow \lambda_{\text{dB}} \gtrsim 1 \text{ kpc}$$

[Hu, Barkana, Gruzinov, PRL 85 (2000)]

Better fit to small  
scale structure!



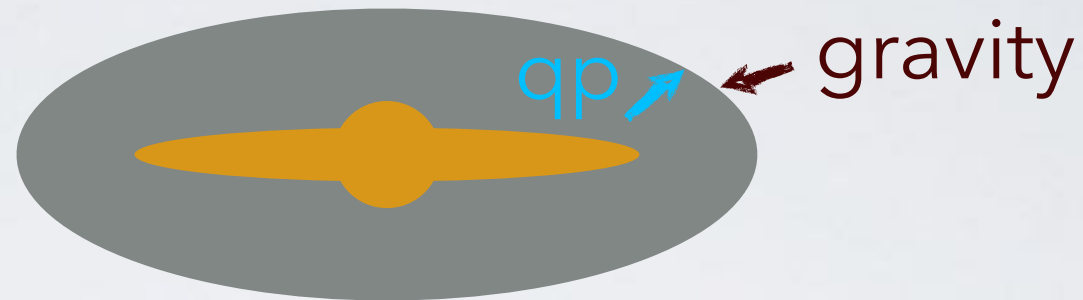
[Bullock et al., Ann.Rev.Astron.Astrophys. 55 (2017)]



# THE FUZZY DM PARADIGM

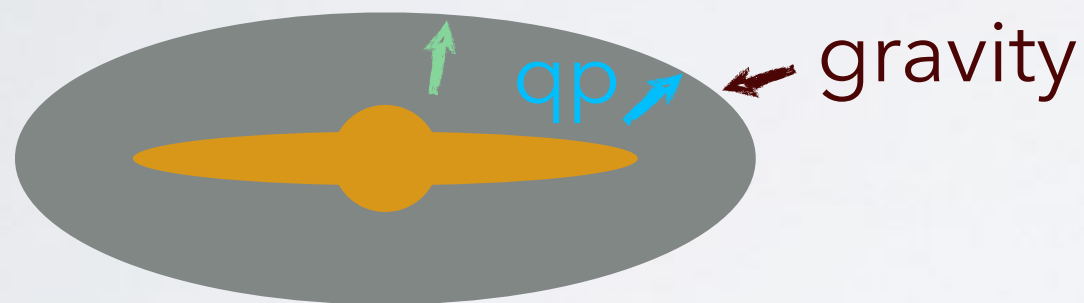
- Small scale is set by a balance of gravity and quantum pressure:

No self-interactions!



- Self-interactions may drastically alter situation:

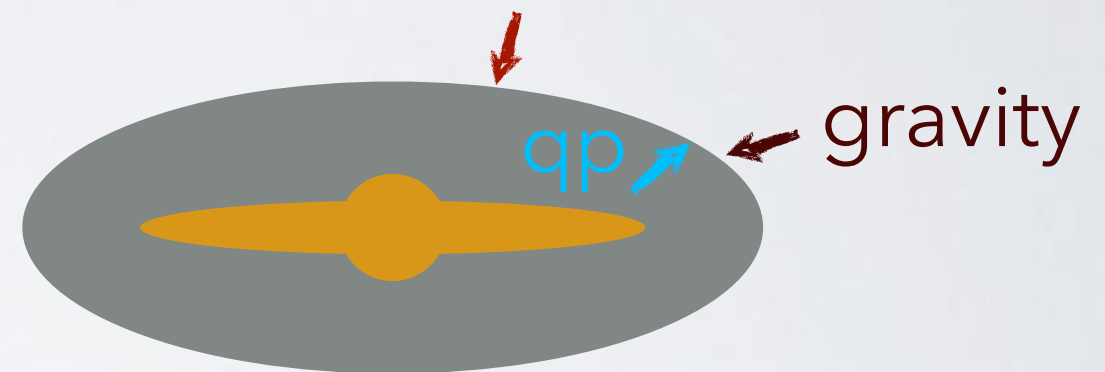
repulsive  $\lambda > 0$



Relaxed mass range: [Ferreira, 2005.03254]

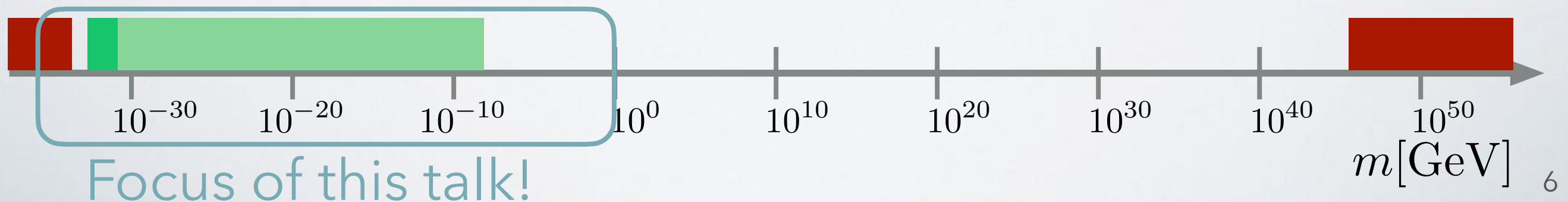
$$m_{\text{DM}} \approx 10^{-22} - 1 \text{ eV}$$

attractive  $\lambda < 0$



Instabilities!

[Guth et al. PRD **92**, 2015]



# QFT TOY MODELS FOR FUZZY DM

- Typical searches for Fuzzy DM (FDM) employ their properties of a **classical background field**.
- **QUESTION:** Can we write down QFT toy models for this kind of DM and learn something about the microscopic properties?
- **STRATEGY:** Explore complementary search strategies using the particle properties of QFT toy models:
  - i) FDM can be **scalar**  $s$  or **pseudo-scalar**  $a$
  - ii) **Coupled** to SM via **Higgs**  $H$  or **new heavy singlet mediator**  $\phi$

# 1. SCALAR DM — HIGGS PORTAL

- Most economic way to couple fuzzy DM to SM via Higgs Portal:

$$\mathcal{L} \supset \frac{1}{2} \partial_\mu s \partial^\mu s - \frac{1}{2} m_s^2 s^2 - \frac{1}{4!} \lambda_s s^4 - \frac{1}{2} \lambda_{hs} s^2 H^\dagger H$$

- DM is protected by a  $Z_2$  symmetry and has positively bounded potential  $\lambda_s > 0$

$\Rightarrow$  FDM can have wide mass range (but for no good reason) due to repulsive self-interactions

- In the FDM regime momenta are small and the occupation numbers are huge

$$n \lambda_{\text{dB}}^3 \approx 6.35 \cdot 10^5 \left( \frac{\text{eV}}{m} \right)^4$$

$\Rightarrow$  **can be treated as a classical wave**

How do we search for wave DM?



# VARIATION OF CONSTANTS

- Fundamental constants like  $m_f$ ,  $\alpha_{\text{em}}$  or  $m_V$  are described by SM operators

$$\mathcal{L}_{\text{SM}} \supset - \sum_f m_f \bar{f} f - \frac{F_{\mu\nu} F^{\mu\nu}}{4} + \sum_V \delta_V m_V^2 V_\mu V^\mu$$

- In the presence of ULDM these operators are modified, e.g. in the *Higgs portal*

$$\mathcal{L} \supset \underbrace{\frac{\lambda_{hs}}{2} \frac{m_f}{m_h^2} s^2}_{\delta m_f} \bar{f} f - \underbrace{\frac{\lambda_{hs} g_{h\gamma\gamma}}{2} \frac{1}{m_h^2} s^2}_{\delta \alpha_{\text{em}}} F_{\mu\nu} F^{\mu\nu} - \underbrace{\lambda_{hs} \delta_V \frac{m_V^2}{m_h^2} s^2}_{\delta m_V} V_\mu V^\mu$$

where the DM field is described by the **classical wave**

$$s^2 = s_0^2 \cos^2(m_s t) \rightarrow \frac{s_0^2}{2} (1 + \cos(2m_s t))$$

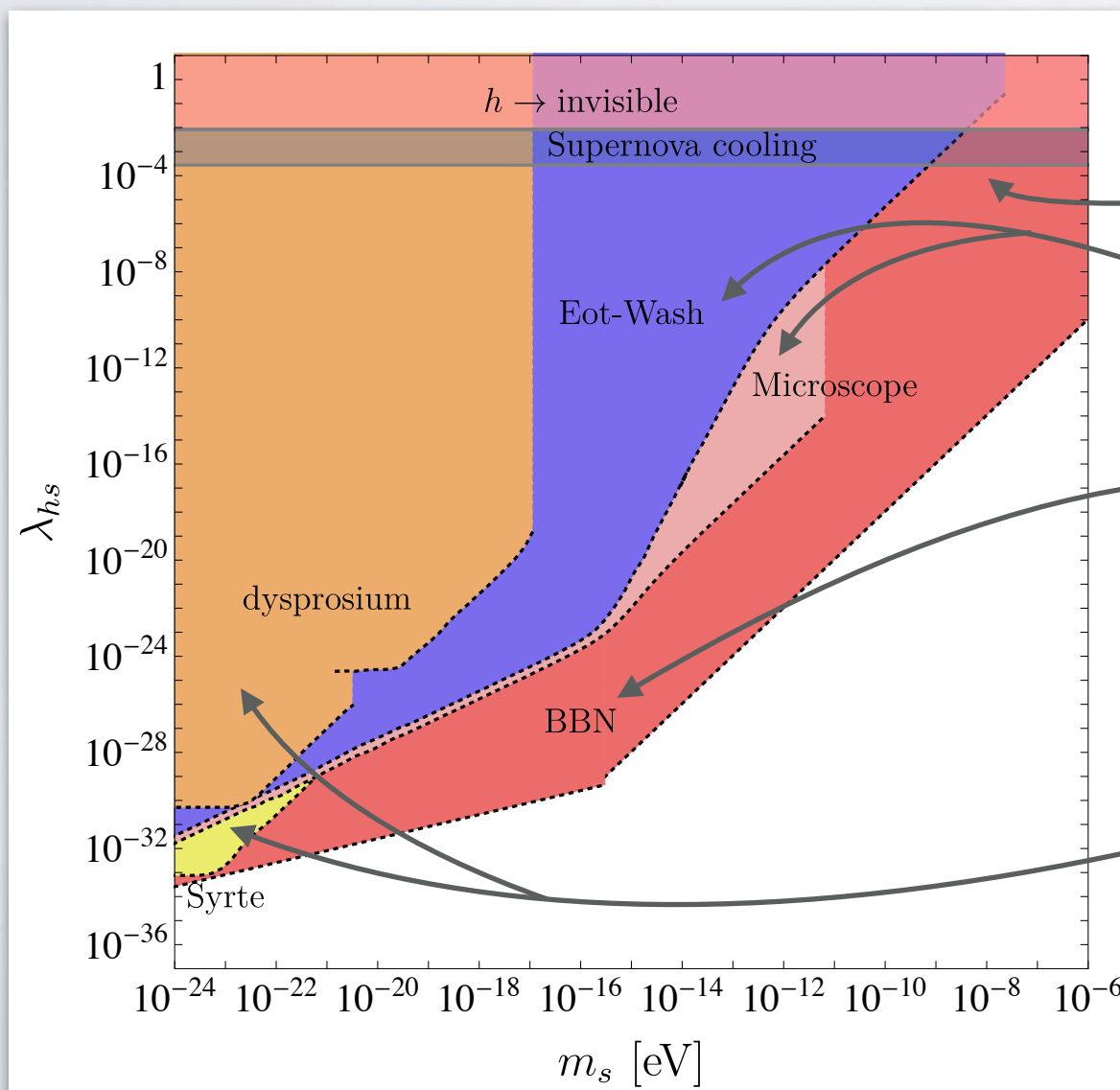
# 1. SCALAR DM — HIGGS PORTAL

- At low momenta Higgs portal mediates an effective DM-nucleon coupling

$$\mathcal{L} \supset -\frac{1}{2} \lambda_{hs} s^2 H^\dagger H \longrightarrow c_{sNN} s^2 \bar{N} N$$

where classically  $s^2 = s_0^2 \cos^2(m_s t) \rightarrow \frac{s_0^2}{2} (1 + \cos(2m_s t))$

$$c_{sNN} = \lambda_{hs} \frac{m_N}{m_h^2} \frac{2n_H}{3(11 - \frac{2}{3}n_L)}$$



Supernova

fifth force

primordial helium abundance

$$m_N - m_P \propto c_{sNN} s_0^2$$

oscillating energy levels

[Brax et al., PRD **97**, 2018]

[Hees et al., PRD **98**, 2018]

[Bauer, PF, Reimitz, Plehn, 2005.13551]



# 1. SCALAR DM — HIGGS PORTAL

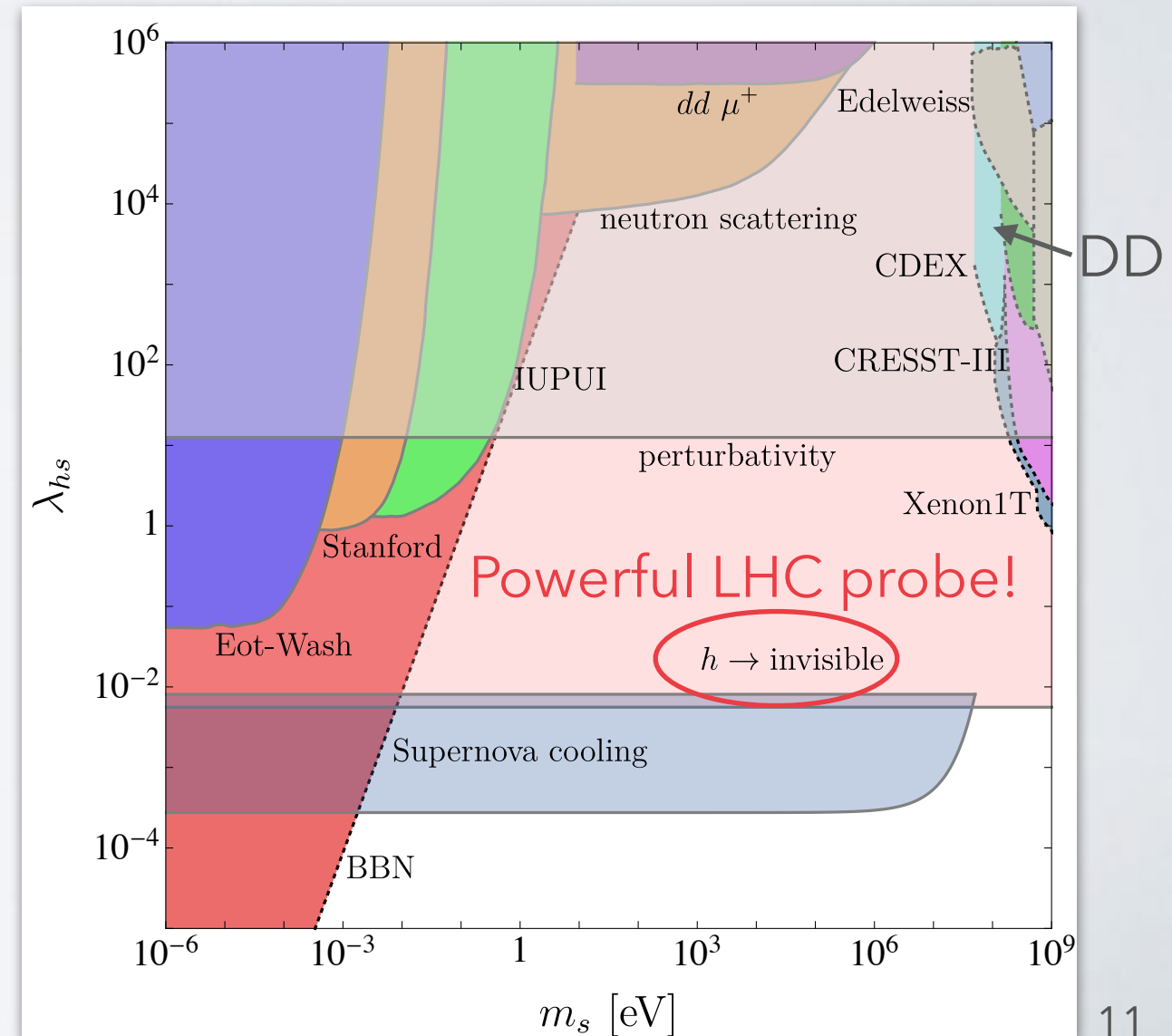
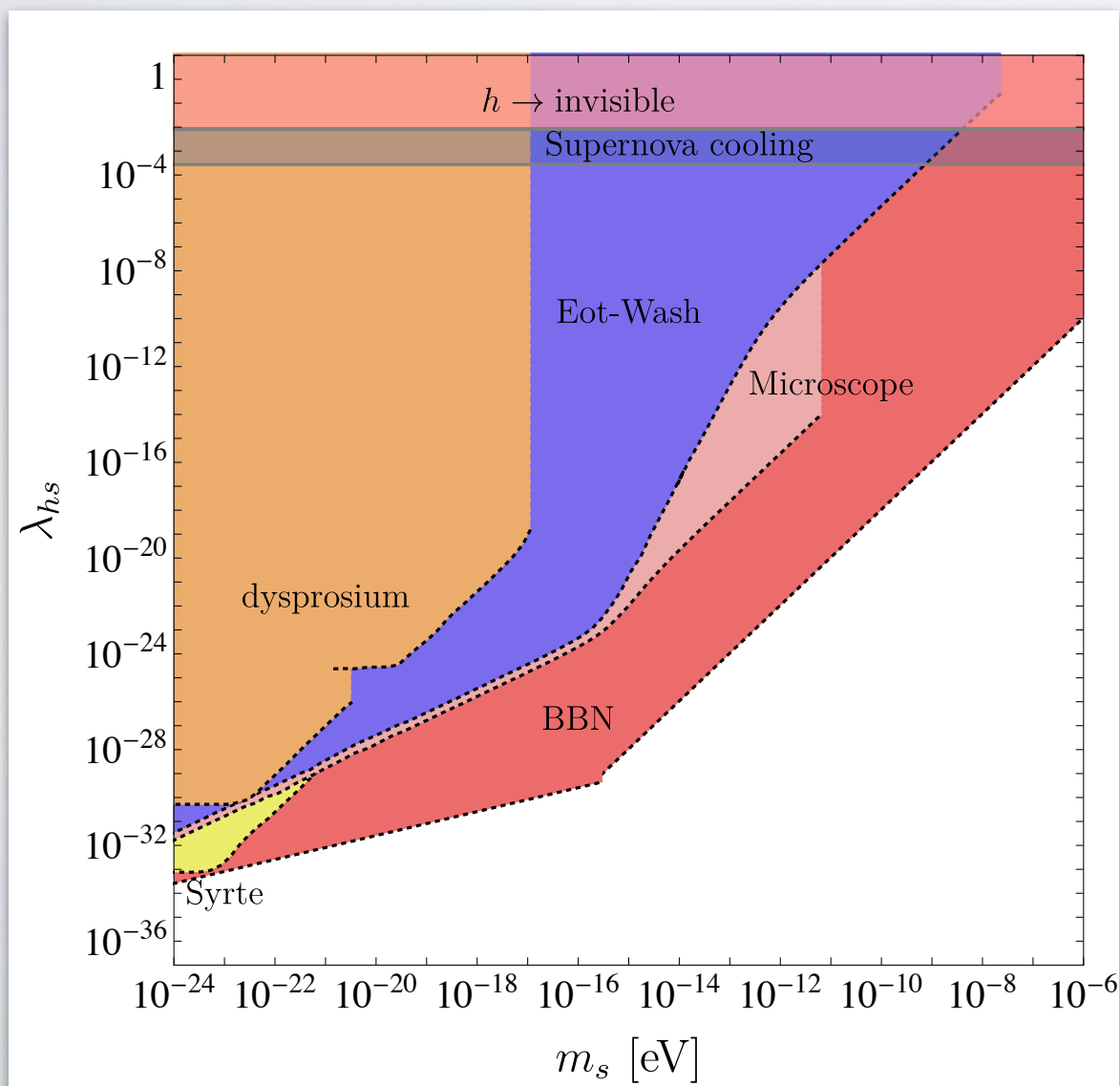
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[Bauer, PF, Reimitz, Plehn, 2005.13551]



# 2. SCALAR DM — NEW MEDIATOR

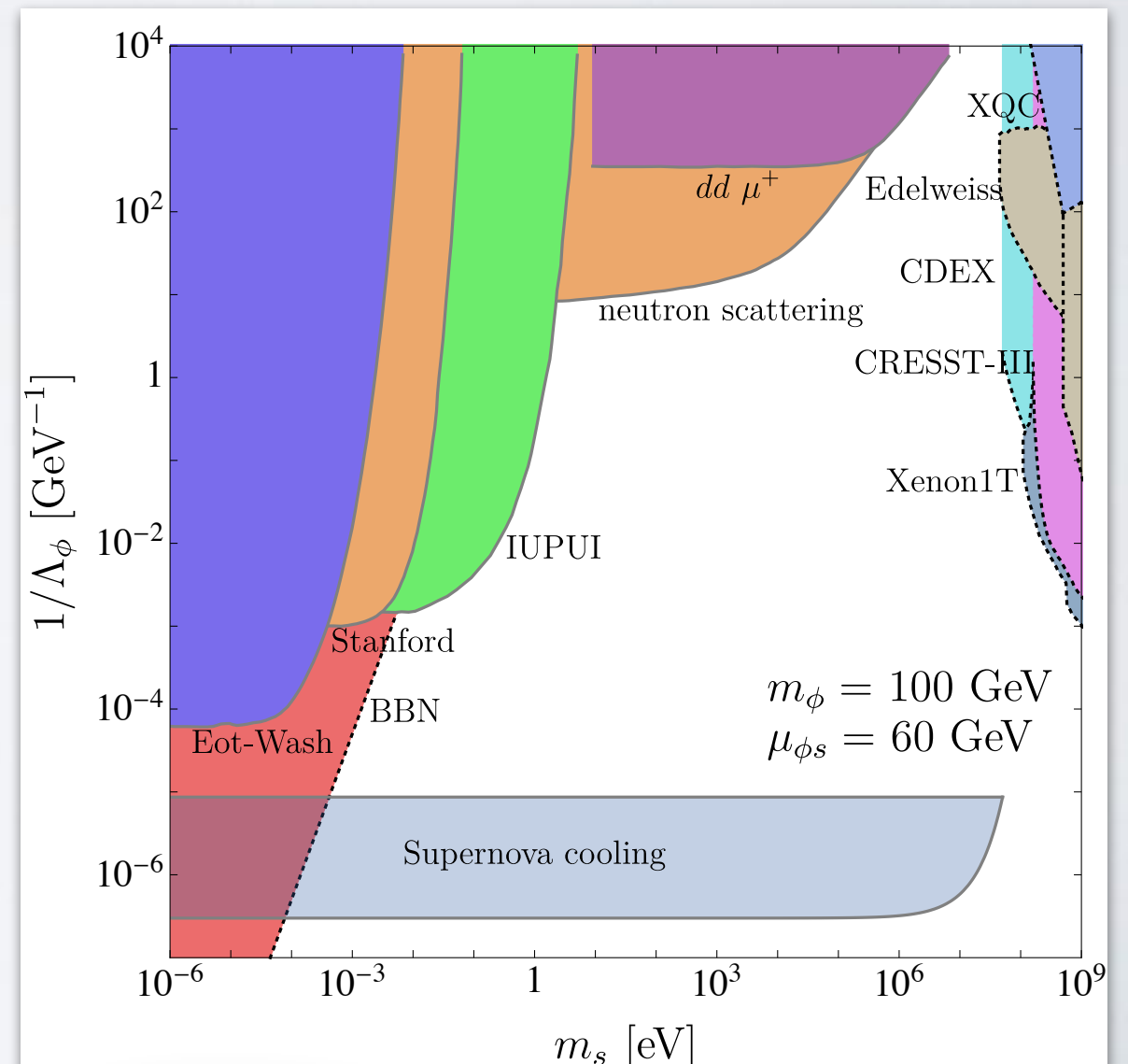
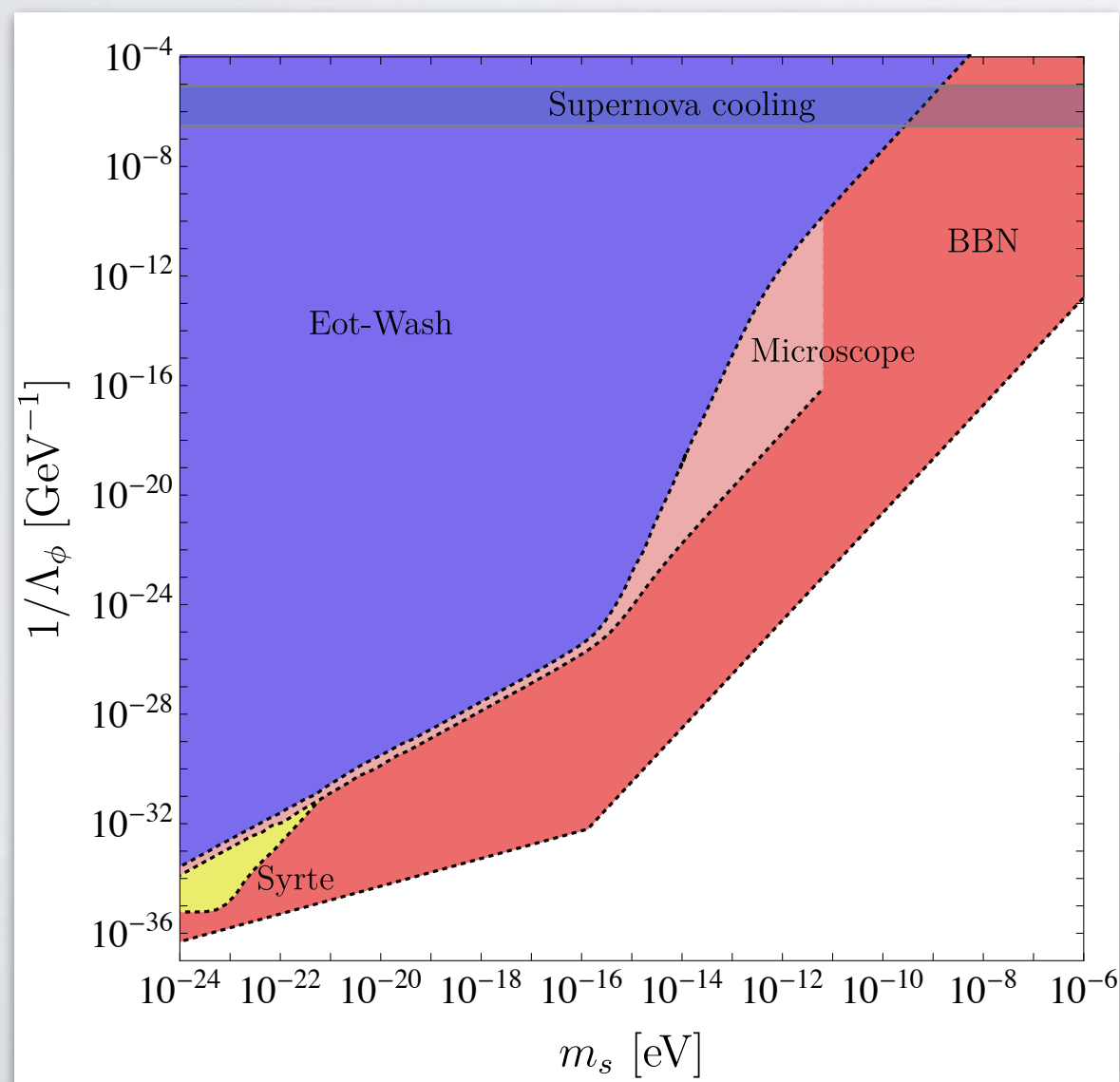
- Consider model with new weak scale mediator  $\phi$

$$\mathcal{L} \supset -\frac{1}{2}m_\phi^2\phi^2 - \frac{\mu_{\phi s}}{2}\phi s^2 - \frac{\alpha_S}{\Lambda_\phi}\phi \text{Tr}[G_{\mu\nu}G^{\mu\nu}] \longrightarrow c_{sNN} s^2 \bar{N}N$$

- High mass window rather unconstrained!

$$c_{sNN} = \frac{\mu_{\phi s}}{\Lambda_\phi} \frac{m_N}{m_\phi^2} \frac{8\pi}{11 - \frac{2}{3}n_L}$$

[Bauer, PF, Reimitz, Plehn, 2005.13551]



# LIGHT DM — ALPS

- Maybe **best motivated** candidate for FDM is an **axion-like particle**. It has a reason to be very light!
- Axions are Nambu-Goldstone particles, protected by shift symmetry:

$$S = \frac{s + f}{\sqrt{2}} e^{ia/f} \quad e^{ia/f} \rightarrow e^{i(a+c)/f} = e^{ia/f} e^{ic/f}$$

- Mass is generated by small explicit breaking:

$$V(a) = \Lambda^4 \left[ 1 - \cos \left( \frac{a}{f} \right) \right] = \frac{\Lambda^4}{2f^2} a^2 + \dots$$

**Suppressed** by heavy axion scale  $f = \mathcal{O}(f_{\text{GUT}})$

# 3. ALP DM — HIGGS MEDIATOR

- Can couple the Goldstone mode  $a$  of complex scalar  $S$  to the Higgs via Dim-6 operator

$$\mathcal{L} = \frac{(\partial_\mu S)(\partial^\mu S)^\dagger}{\Lambda_{ha}^2} H^\dagger H \supset \frac{\partial_\mu a \partial^\mu a}{2\Lambda_{ha}^2} H^\dagger H \longrightarrow c_{aNN} \partial_\mu a \partial^\mu a \bar{N} N$$

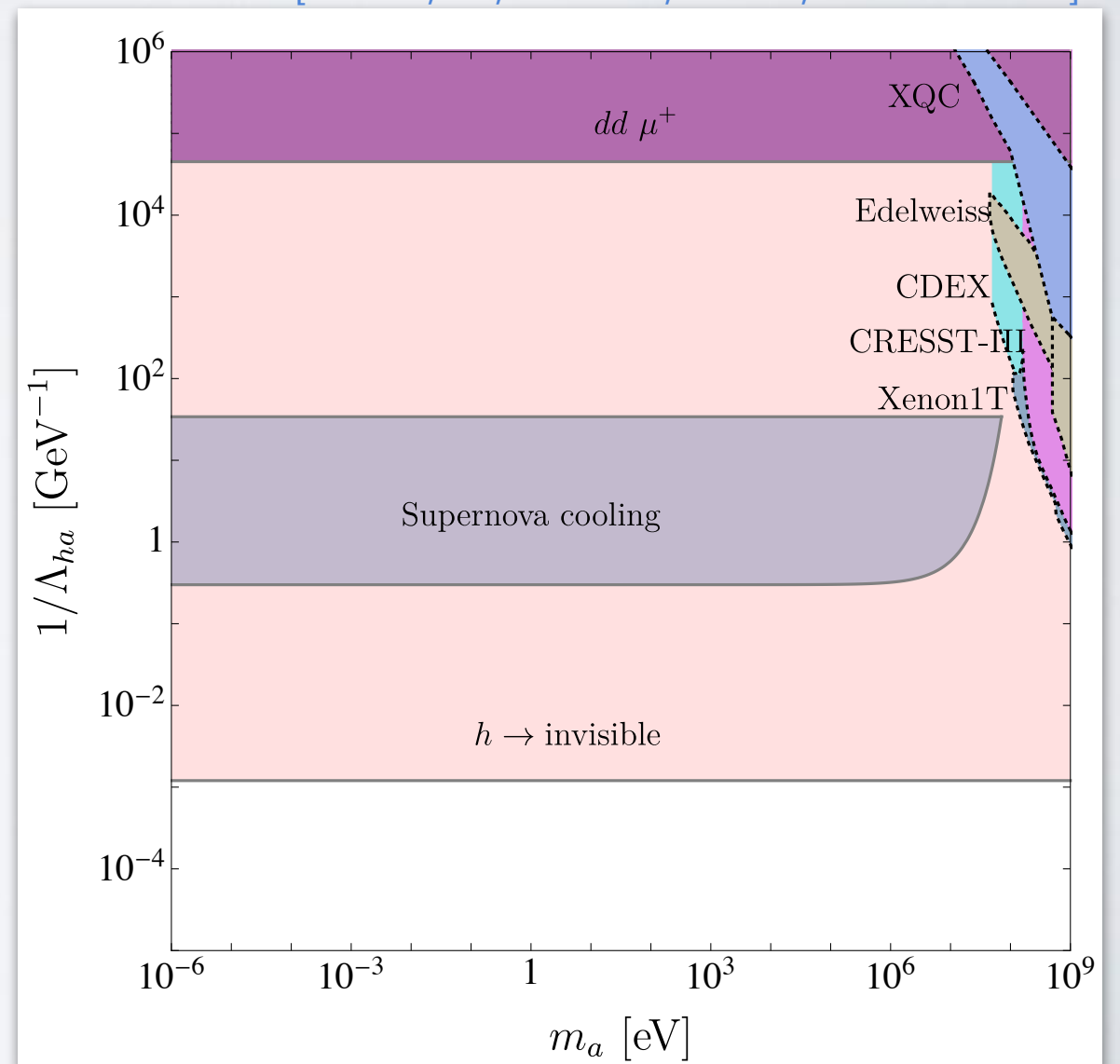
[Bauer, PF, Reimitz, Plehn, 2005.13551]

with 
$$c_{aNN} = \frac{1}{\Lambda_{ha}^2} \frac{m_N}{m_h^2} \frac{2n_H}{3(11 - \frac{2}{3}n_L)}$$

- Strong** model independent Higgs to invisible bound:

$$\Gamma(h \rightarrow aa) \approx \frac{v^2 m_h^3}{128\pi \Lambda_{ha}^4}$$

$$\Lambda_{ha} \gtrsim 832 \text{ GeV}$$



# 4. ALP DM — NEW MEDIATOR

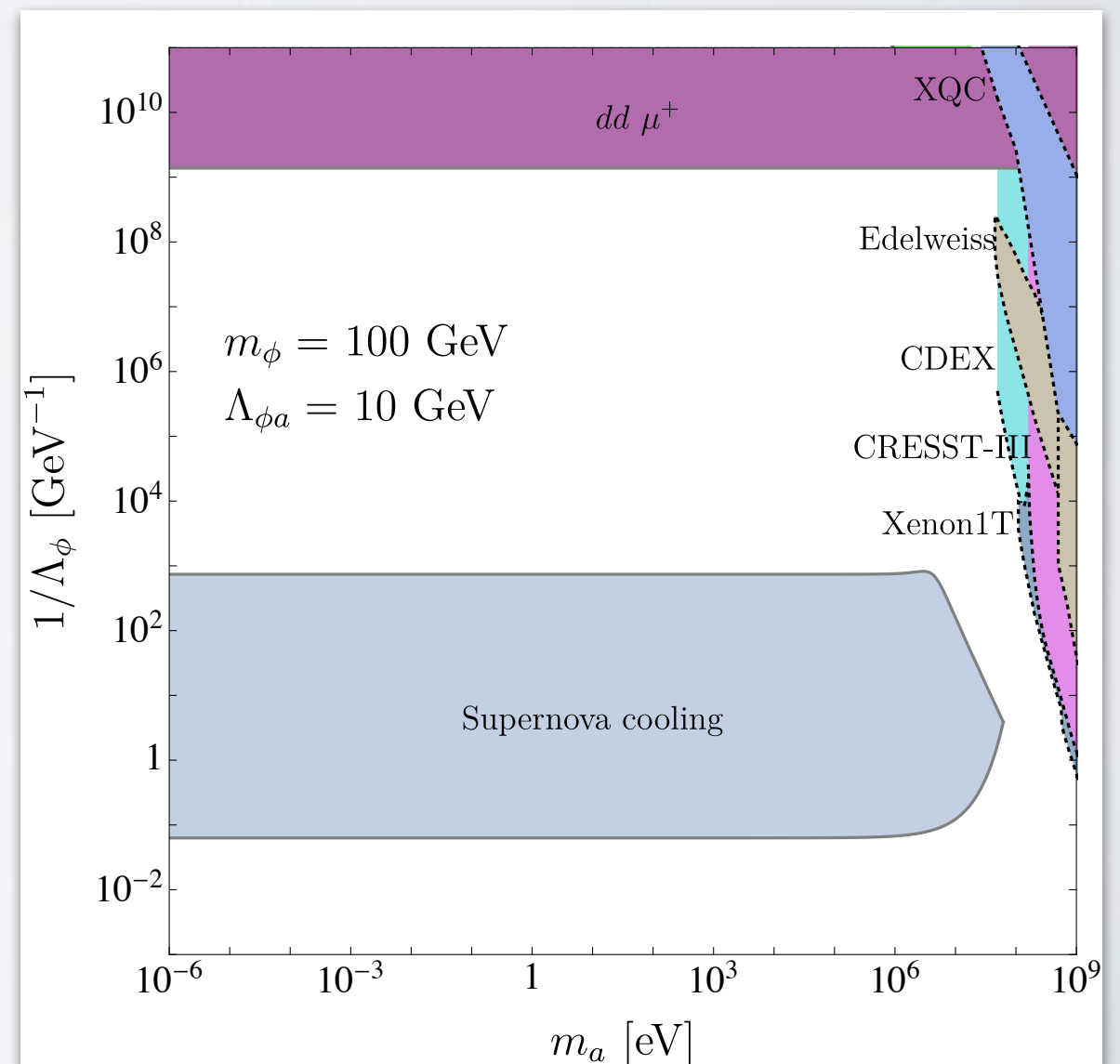
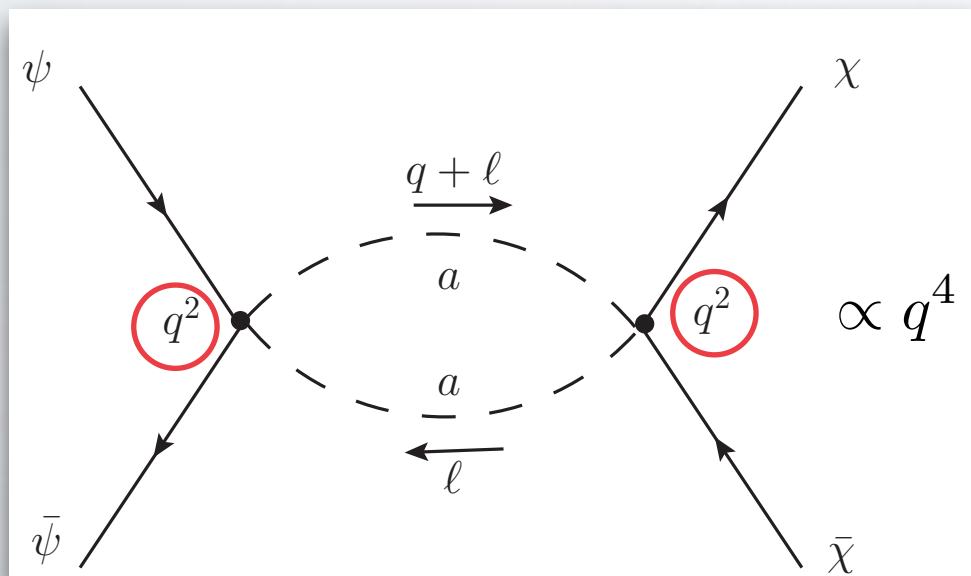
- Consider model with new **weak scale mediator**  $\phi$  and **ALP**  $a$ . Only shift-symmetric couplings allowed:

$$\mathcal{L} \supset -\frac{1}{2}m_\phi^2\phi^2 - \frac{\partial_\mu a \partial^\mu a}{2\Lambda_{\phi a}}\phi - \frac{\alpha_S}{\Lambda_\phi}\phi \text{Tr}[G_{\mu\nu}G^{\mu\nu}] \longrightarrow c_{aNN} \partial_\mu a \partial^\mu a \bar{N}N$$

with 
$$c_{aNN} = \frac{m_N}{\Lambda_{\phi a}\Lambda_\phi m_\phi^2} \frac{8\pi}{11 - \frac{2}{3}n_L}$$

[Bauer, PF, Reimitz, Plehn, 2005.13551]

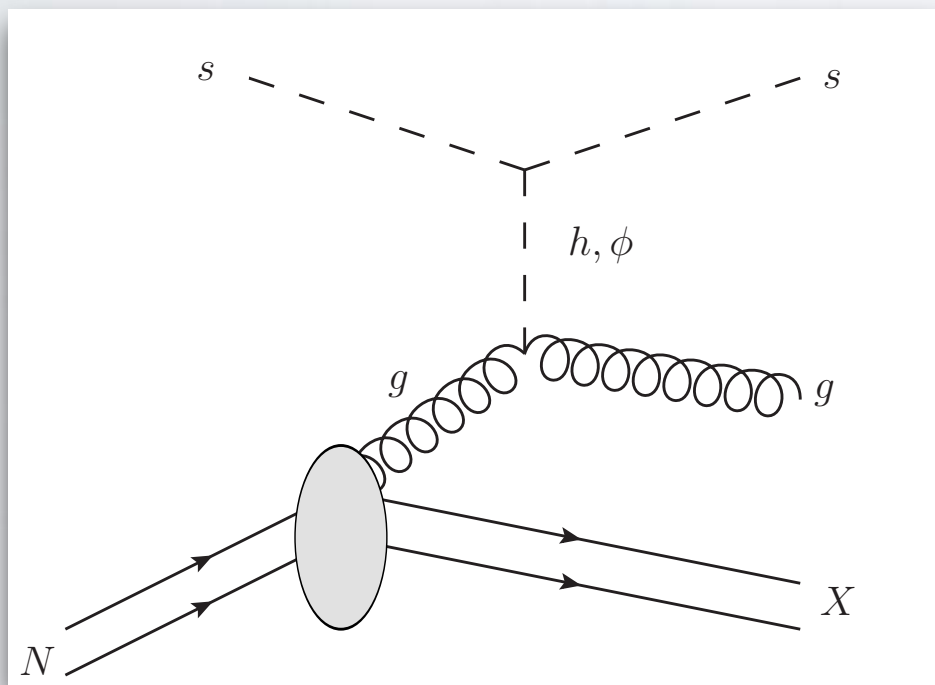
- Almost unconstrained at low masses (momenta) because of momentum suppression:



# NEW SEARCH STRATEGIES AT LHC

- Conventional direct and indirect DM search strategies hopeless due to low momenta of (U)LDM
- But production at LHC enhances momenta:

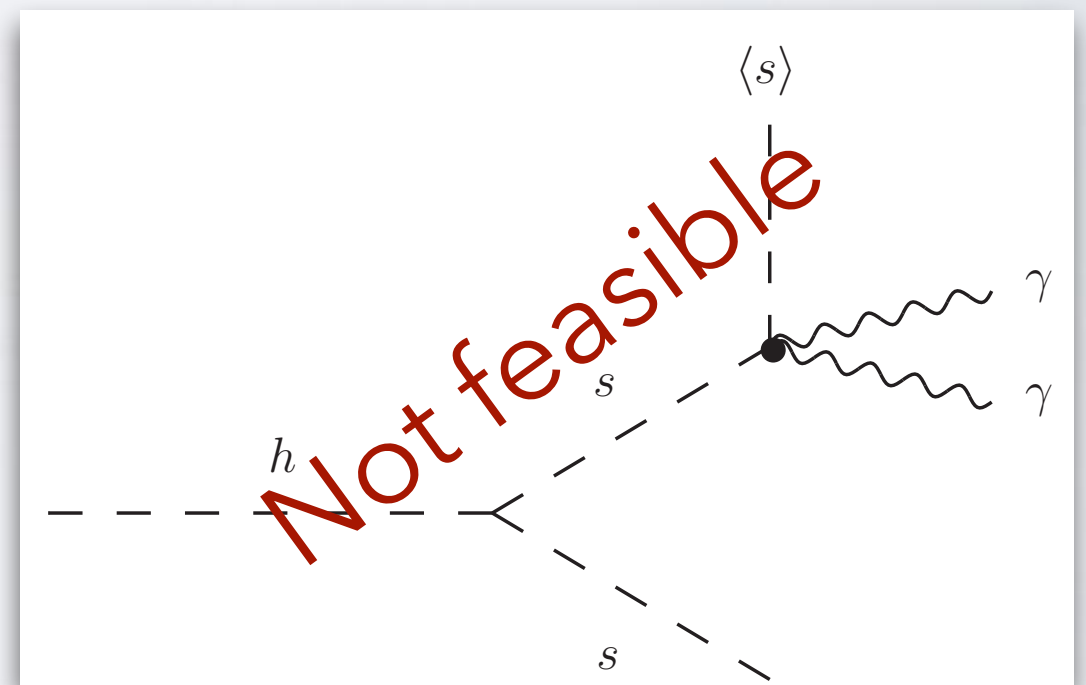
## Direct detection @ LHC



(Deep inelastic scattering)

[Bauer, PF, Reimitz, Plehn, 2005.13551]

## Indirect detection @ LHC



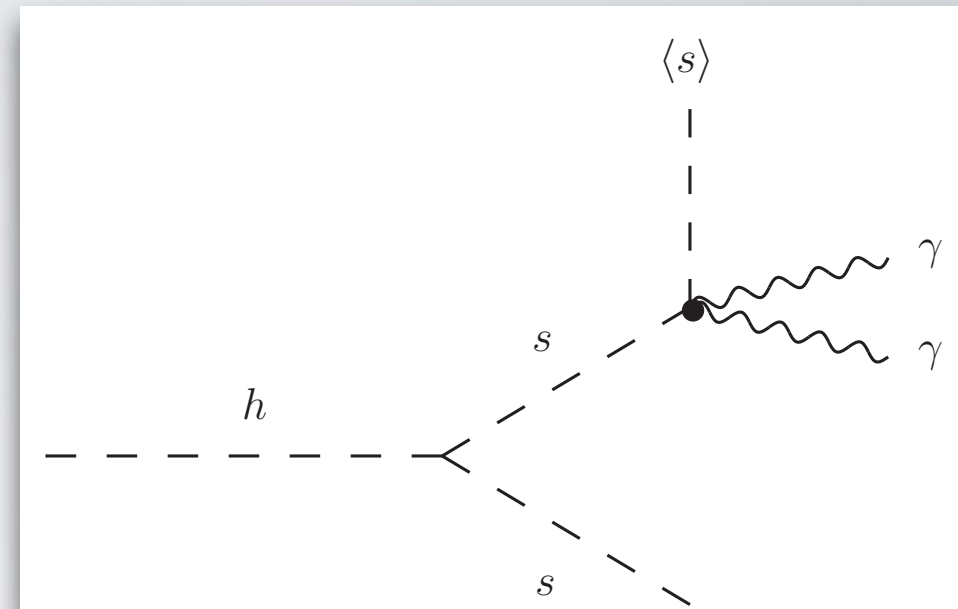
(Background annihilation)



# INDIRECT DETECTION @ LHC

- ULDM has huge occupation numbers. Can it annihilate with the halo background field if produced at LHC?

$$n_{\text{DM}} = \frac{\rho_{\text{DM}}}{m_s} \approx \frac{3 \times 10^{30}}{\text{cm}^3} \left( \frac{10^{-22} \text{ eV}}{m_s} \right)$$



- But cross section scales with mass  $\sigma_{\langle s \rangle s \rightarrow \gamma\gamma} \approx \frac{\lambda_{hs}^2 g_{h\gamma\gamma}^2}{4\pi} \frac{m_s}{m_h^3}$
- **Mean free path independent of mass and very large**

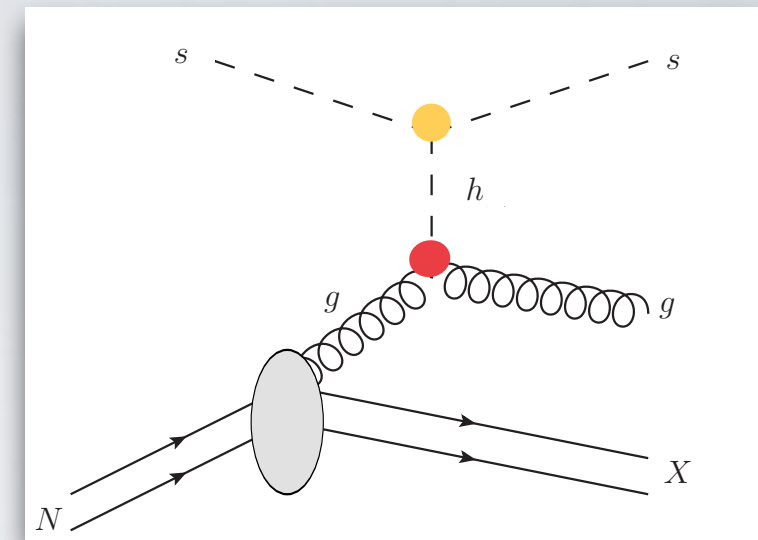
$$\lambda = \frac{1}{n_{\text{DM}} \sigma_{\langle s \rangle s \rightarrow \gamma\gamma}} = \frac{4\pi}{\lambda_{hs}^2 g_{h\gamma\gamma}^2} \frac{m_h^3}{\rho_{\text{DM}}} \gtrsim 10^{43} \text{ m} \quad \text{⚡}$$

- Larger cross section above electron threshold, but also lower densities!

$$\sigma_{\langle s \rangle s \rightarrow \bar{f}f} = \frac{\lambda_{hs}^2}{8\pi} \frac{m_f^2}{m_h^4} \left( 1 - \frac{4m_f^2}{m_s m_h} \right)$$

# DIRECT DETECTION

- Boosted DM can undergo DIS in detector material and produce jets.



1. E.g. Higgs Portal:

$$N_{\text{DIS}} = \mathcal{L}_{\text{HL}} \sigma_h \text{BR}_{h \rightarrow ss} P_{\text{DIS}}$$

with  $P_{\text{DIS}} = 1 - e^{-L_{\text{det}} n_X \sigma_X}$

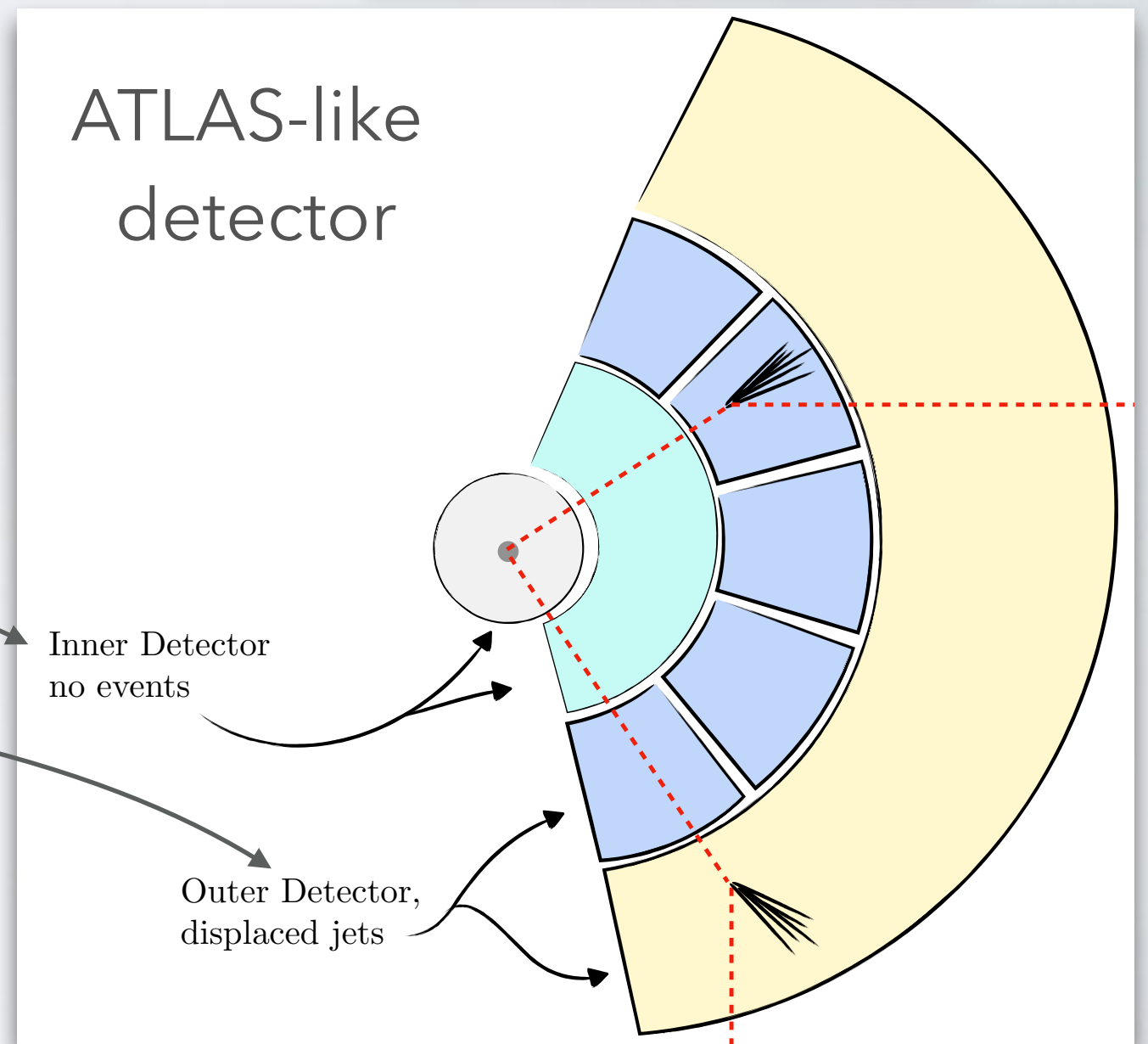
Distinguishable from LLPs by location of interaction:

$$n_{Pb} \gg n_{Xe}$$

But unfortunately for HP:

$$\frac{d^2 \hat{\sigma}_{\text{DIS}}}{dx dy} = \frac{\lambda_{hs}^2 g_{hgg}^2}{4\pi \hat{s}} \frac{Q^4}{(Q^2 + m_h^2)^2}$$

$$P_{\text{DIS}} = 1 - e^{-L_E n_{Pb} \sigma_{Pb}} e^{-L_H n_{Fe} \sigma_{Fe}} \approx 7.5 \cdot 10^{-21} \text{ ⚡}$$

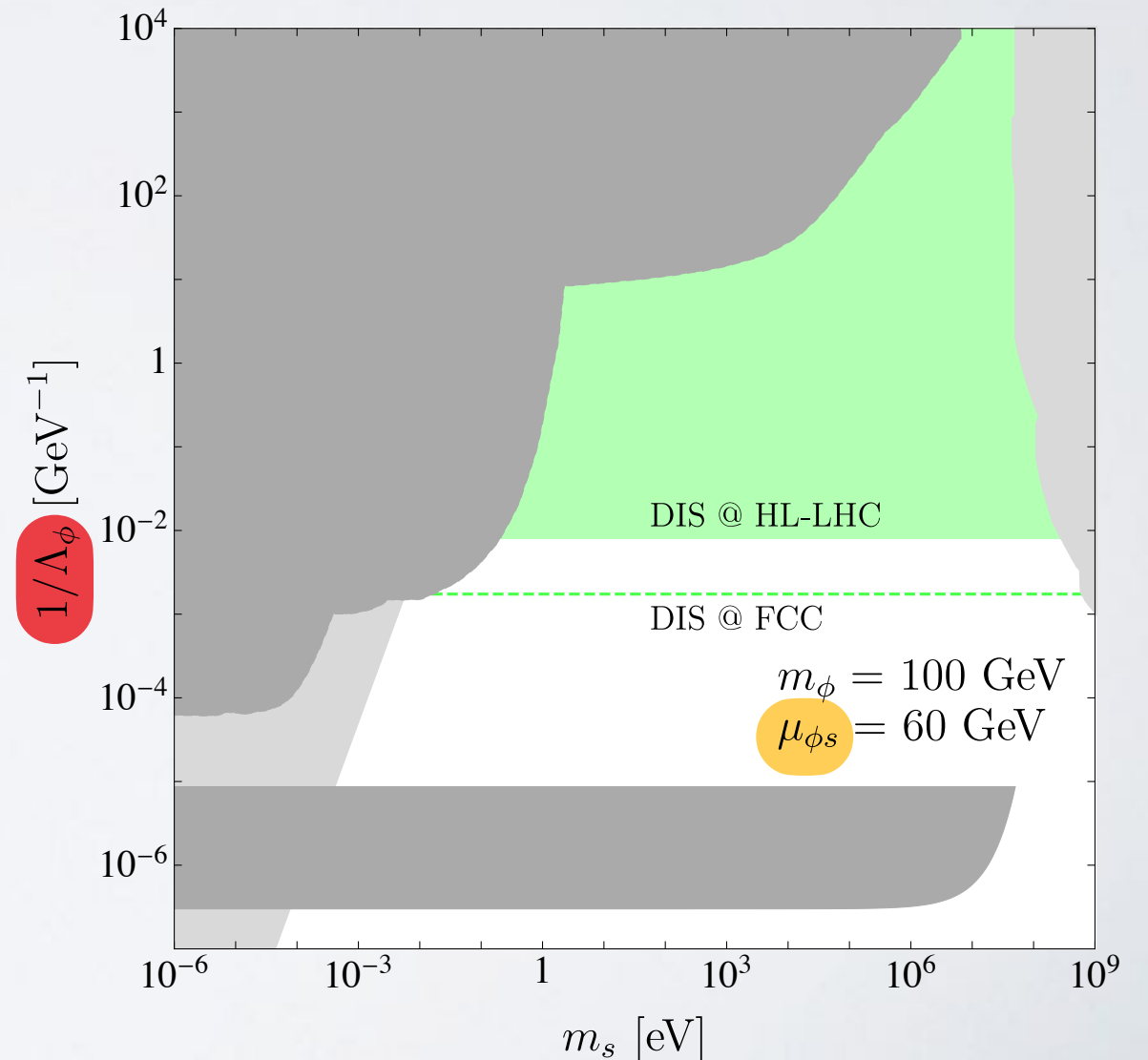
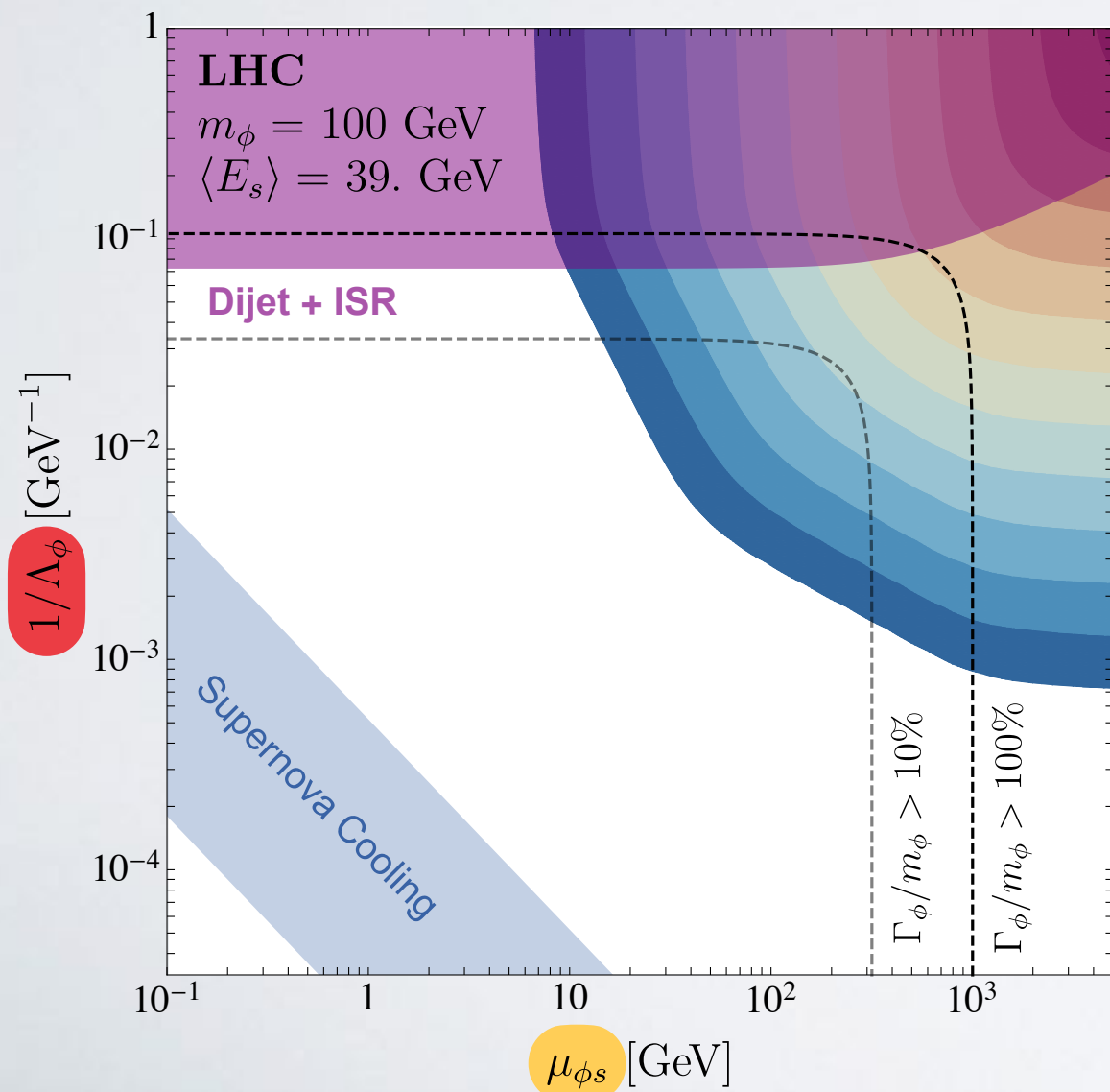
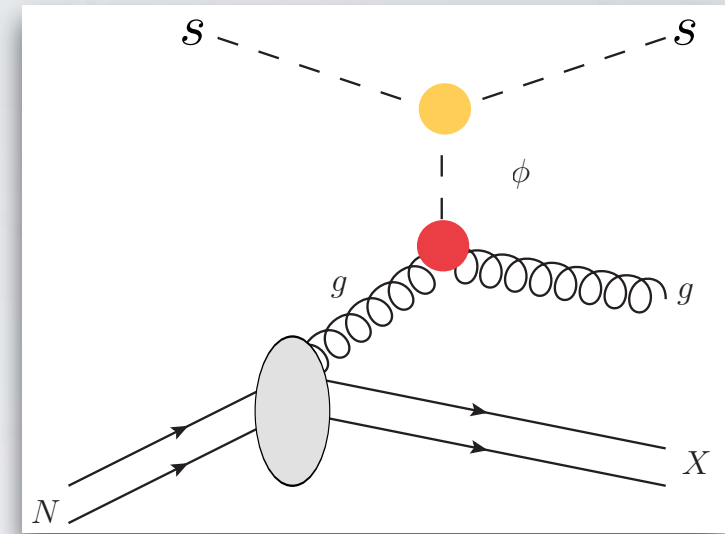
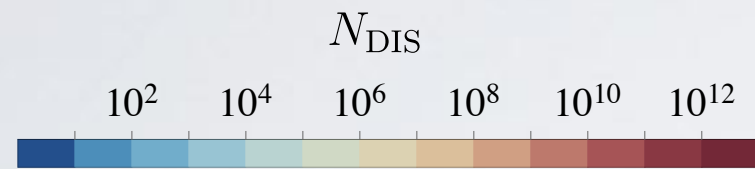


[Bauer, PF, Reimitz, Plehn, 2005.13551]

# DIRECT DETECTION AT THE LHC

## 2. Scalar DM with new scalar mediator

$$\frac{d^2 \hat{\sigma}_{\text{DIS}}}{dx dy} = \frac{\alpha_s^2}{4\pi \hat{s}} \left( \frac{\mu_{\phi s}}{\Lambda_\phi} \right)^2 \frac{Q^4}{(Q^2 + m_\phi^2)^2}$$



# DIRECT DETECTION AT THE LHC

## 3. ALP DM with Higgs mediator

$$\frac{d^2 \hat{\sigma}_{\text{DIS}}}{dx dy} = \frac{g_{hgg}^2}{16\pi \hat{s}} \frac{Q^4}{\Lambda_{ha}^4} \left( \frac{Q^2 + 2m_a^2}{Q^2 + m_h^2} \right)^2$$

- With Higgs coupling

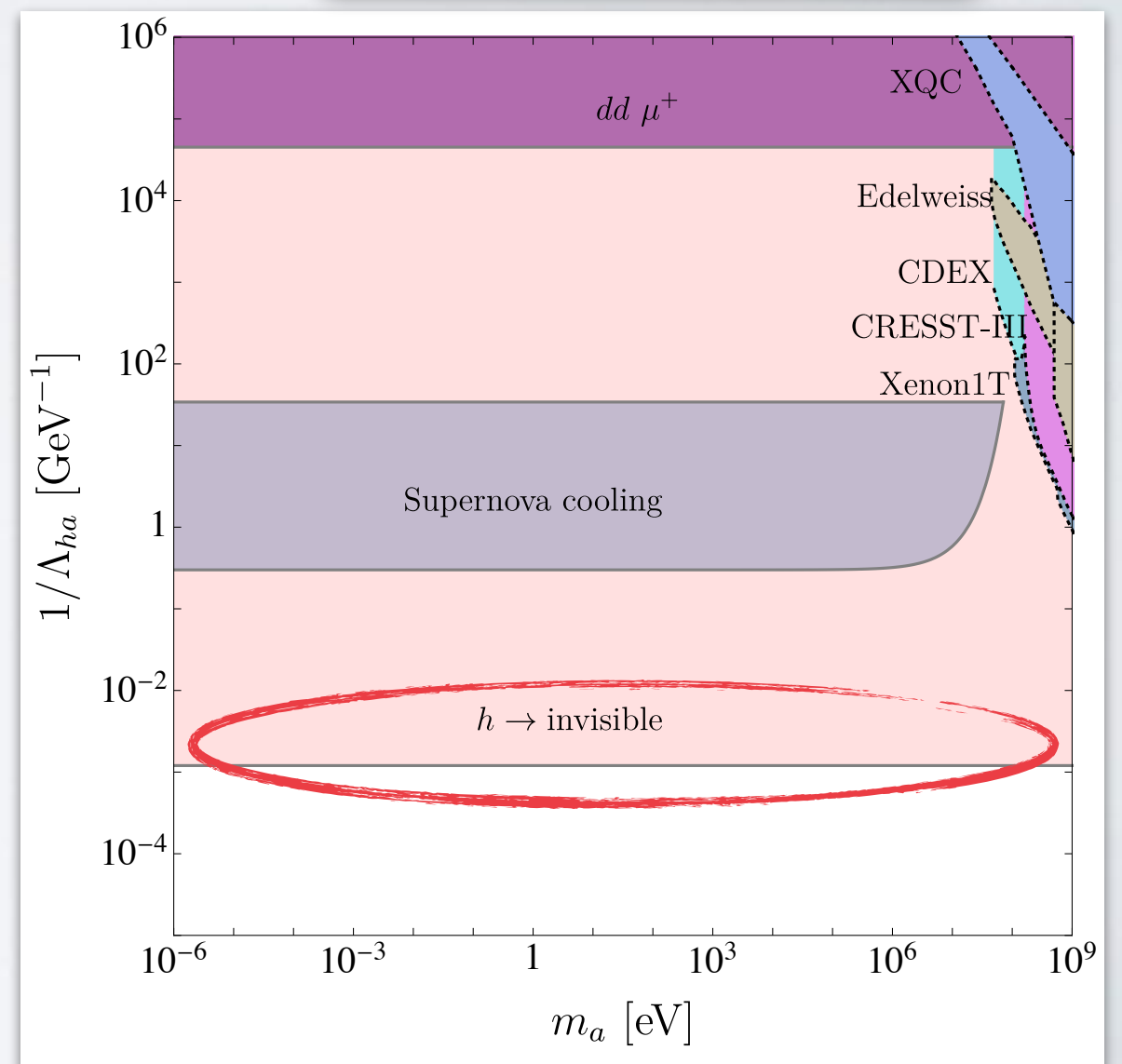
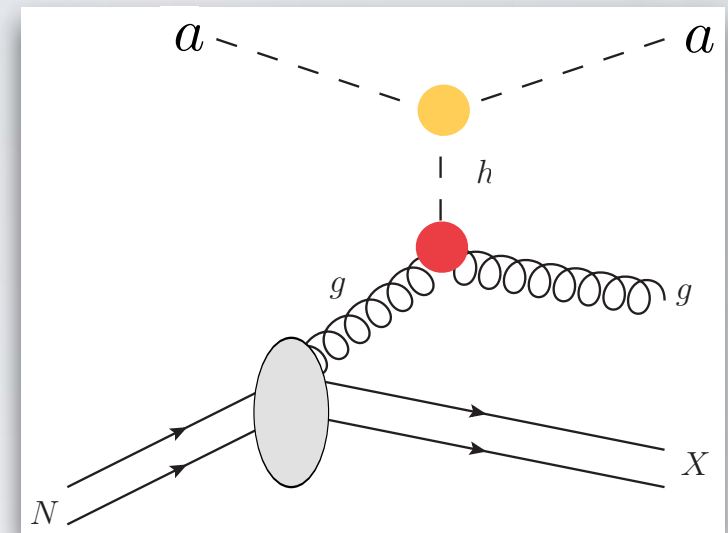
$$g_{hgg} = \alpha_s / (12\pi)$$

and the Higgs invisible constraint

$$\Lambda_{ha} \gtrsim 832 \text{ GeV}$$

we can estimate

$$P_{\text{DIS}} = 1 - e^{-L_E n_{\text{Pb}} \sigma_{\text{Pb}}} e^{-L_H n_{\text{Fe}} \sigma_{\text{Fe}}} \approx 10^{-23} \text{ ⚡}$$

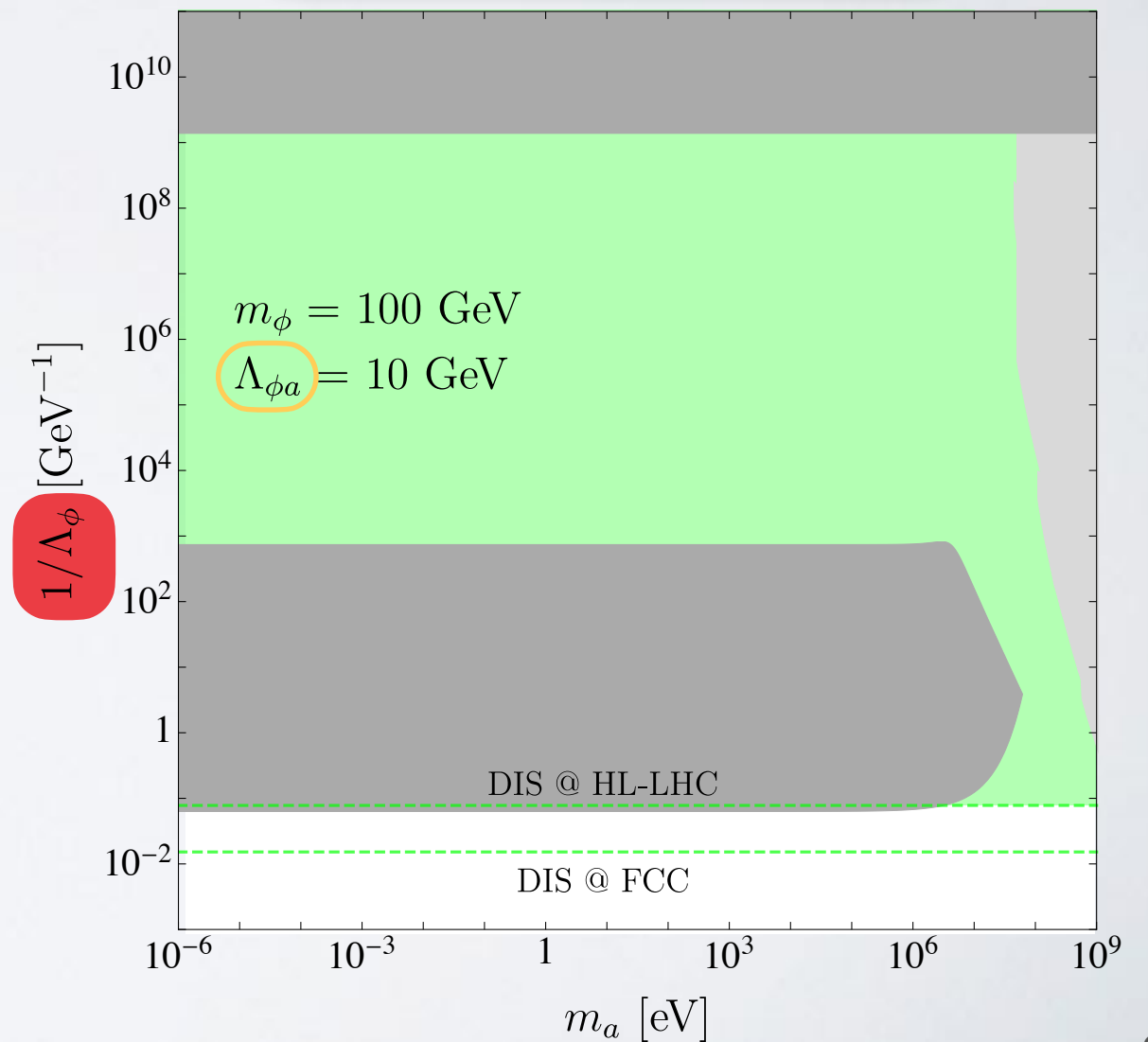
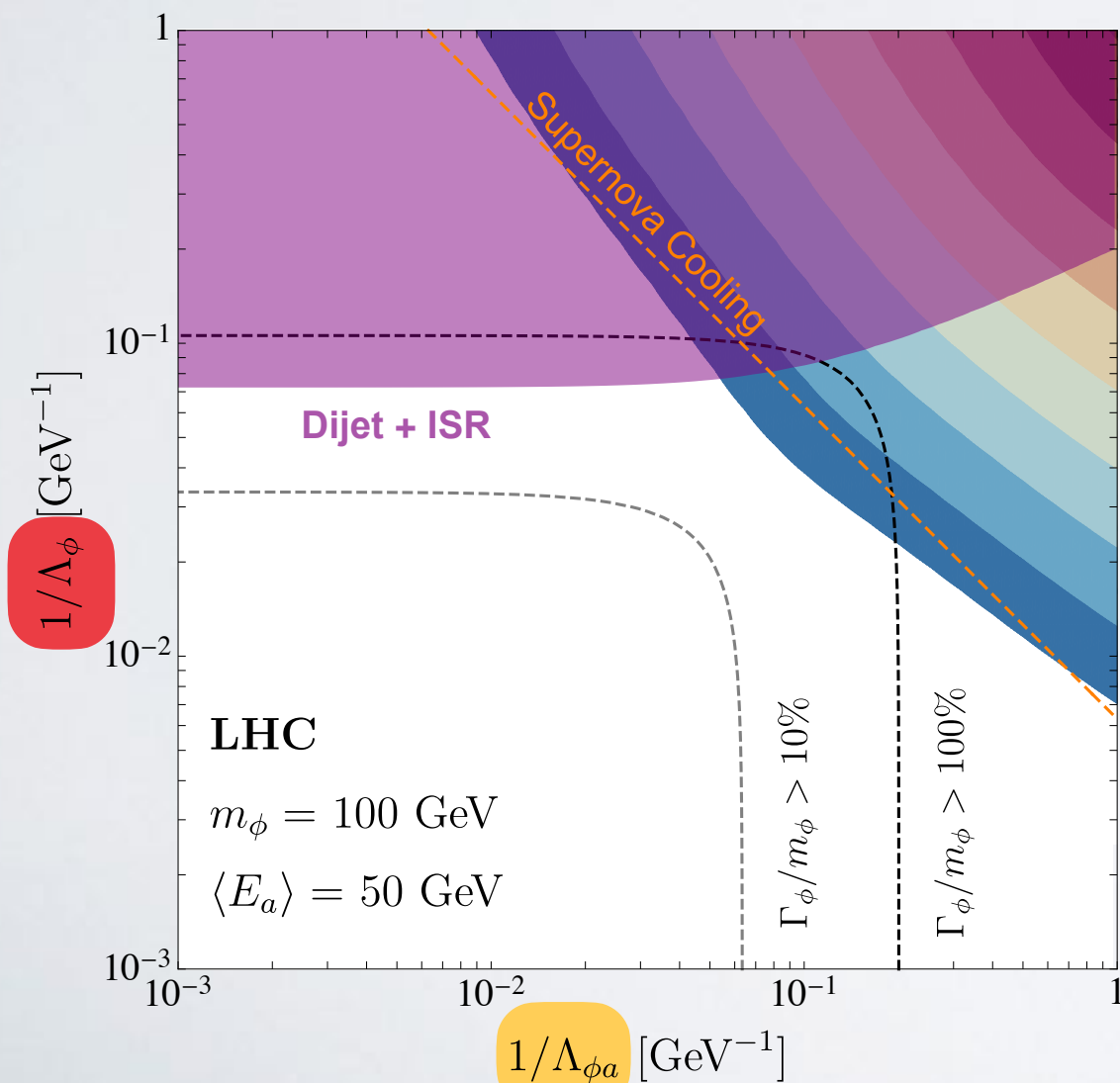
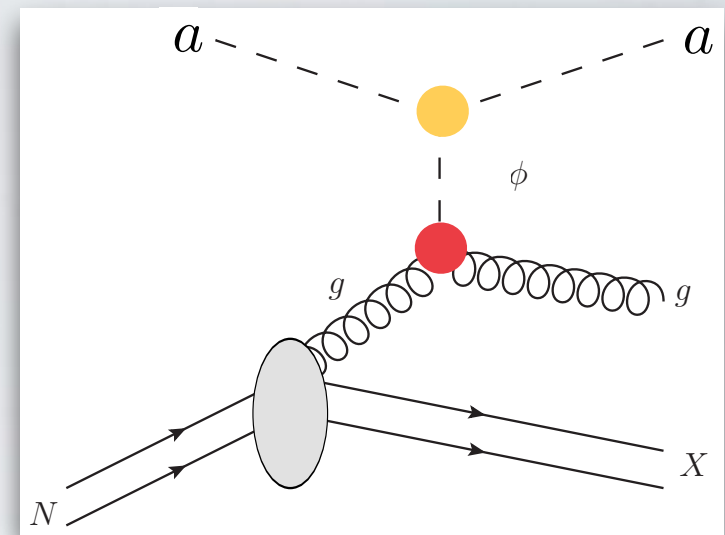
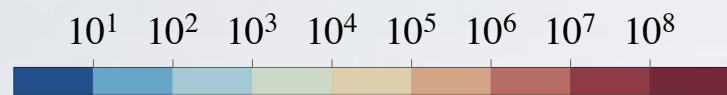


# DIRECT DETECTION AT THE LHC

## 4. ALP DM with scalar mediator

$$\frac{d^2 \hat{\sigma}_{\text{DIS}}}{dx dy} = \frac{\alpha_s^2}{16\pi \hat{s}} \underbrace{\Lambda_{\phi a}^2}_{\text{yellow}} \underbrace{\Lambda_{\phi}^2}_{\text{red}} \left( \frac{Q^2 + 2m_a^2}{Q^2 + m_{\phi}^2} \right)^2$$

$N_{\text{DIS}}$



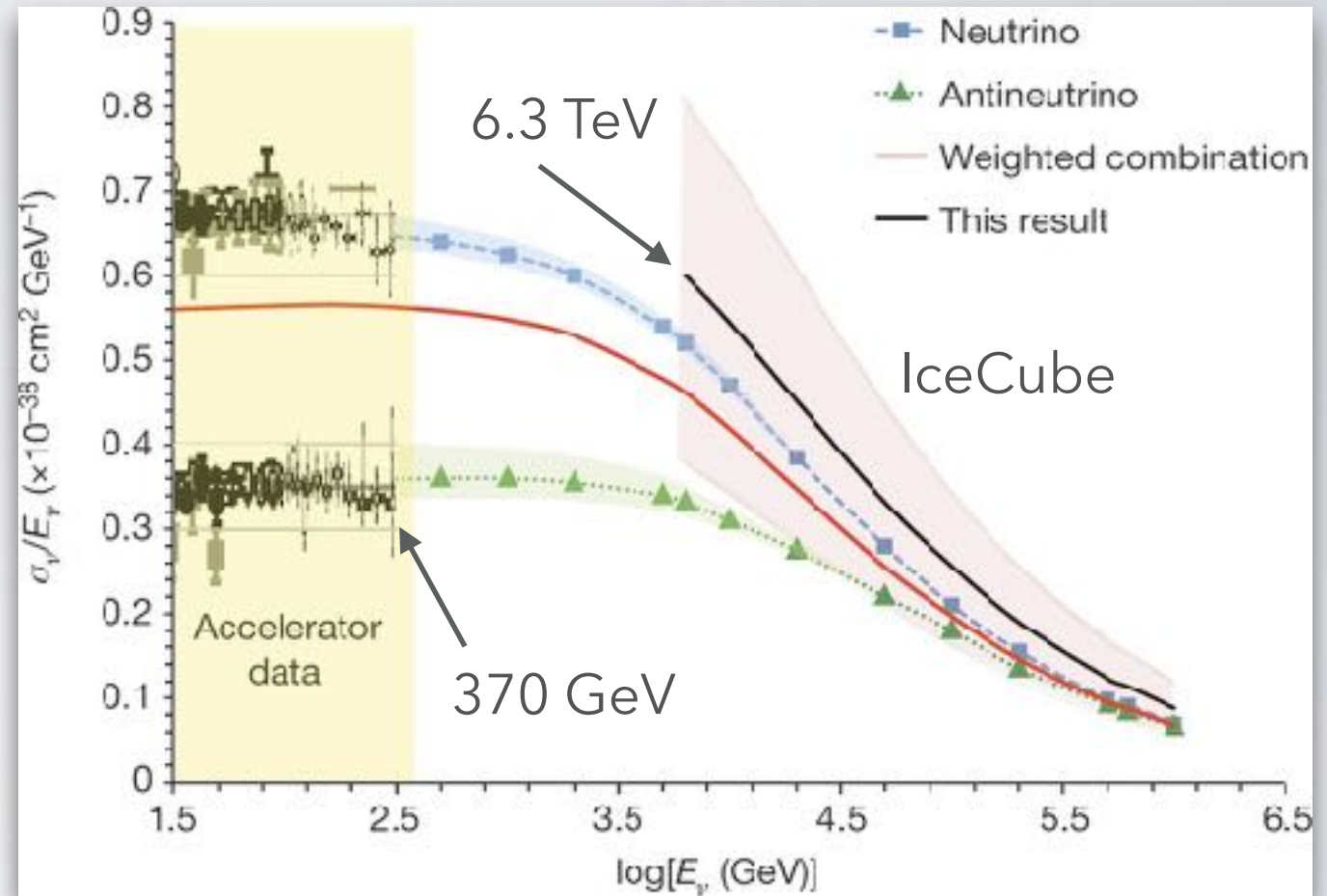
# NEUTRINOS

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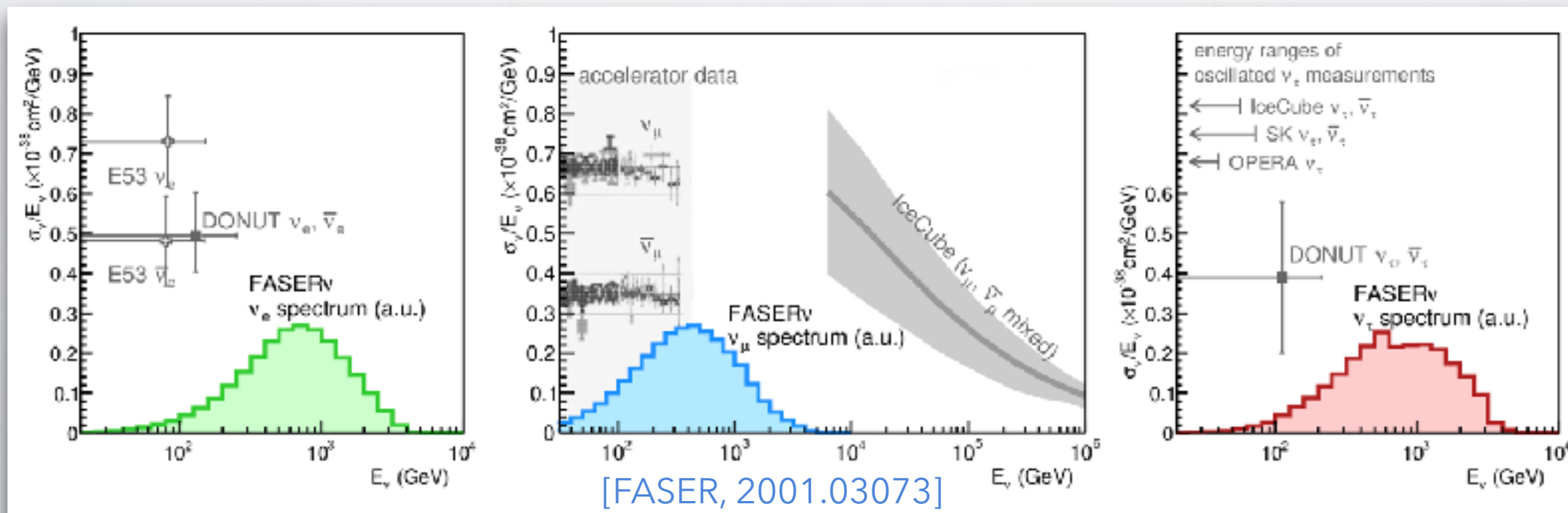


# NEUTRINO CROSS SECTIONS

- Neutrinos still least understood particles of SM: CP violation, masses, Majorana vs. Dirac
- Gap in knowledge of neutrino cross section between 370 GeV and 6.3 TeV!
- New LHC forward experiments (FASER, SND)!



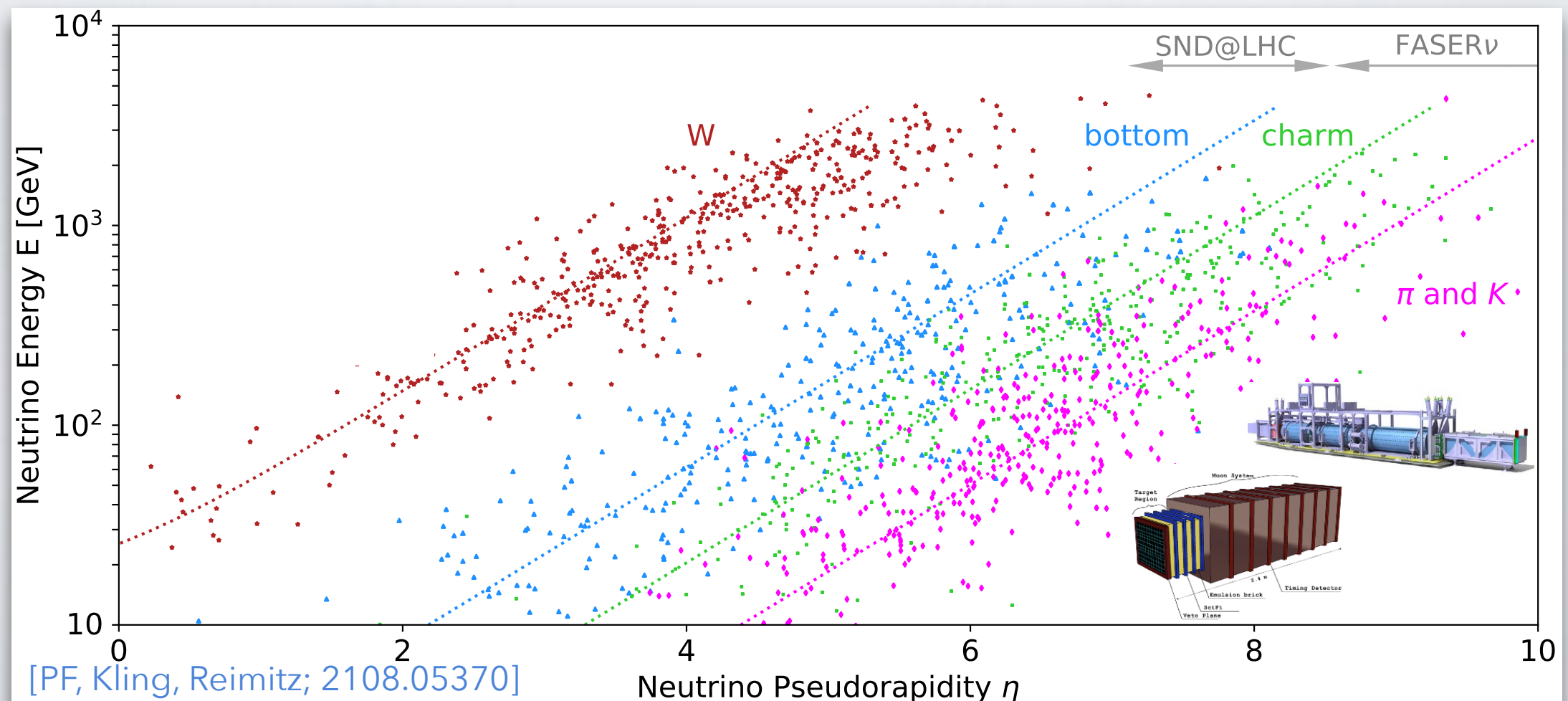
[IceCube, *Nature* 551 (2017) 596-600]



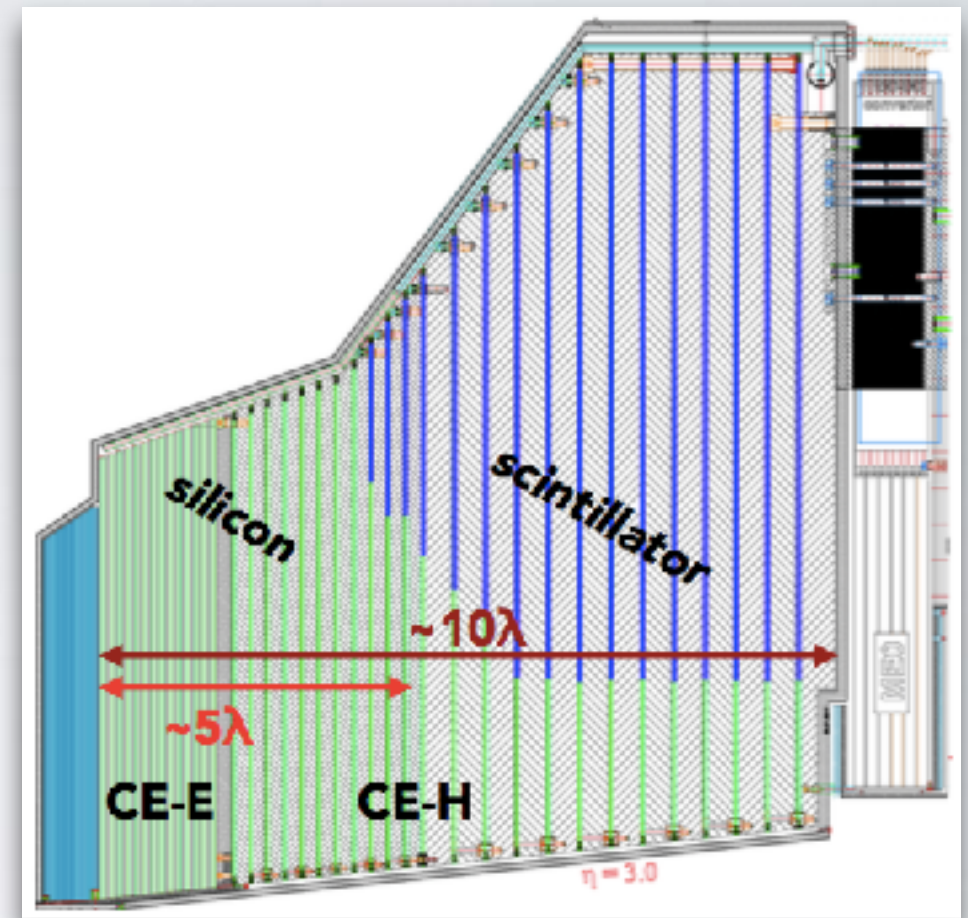
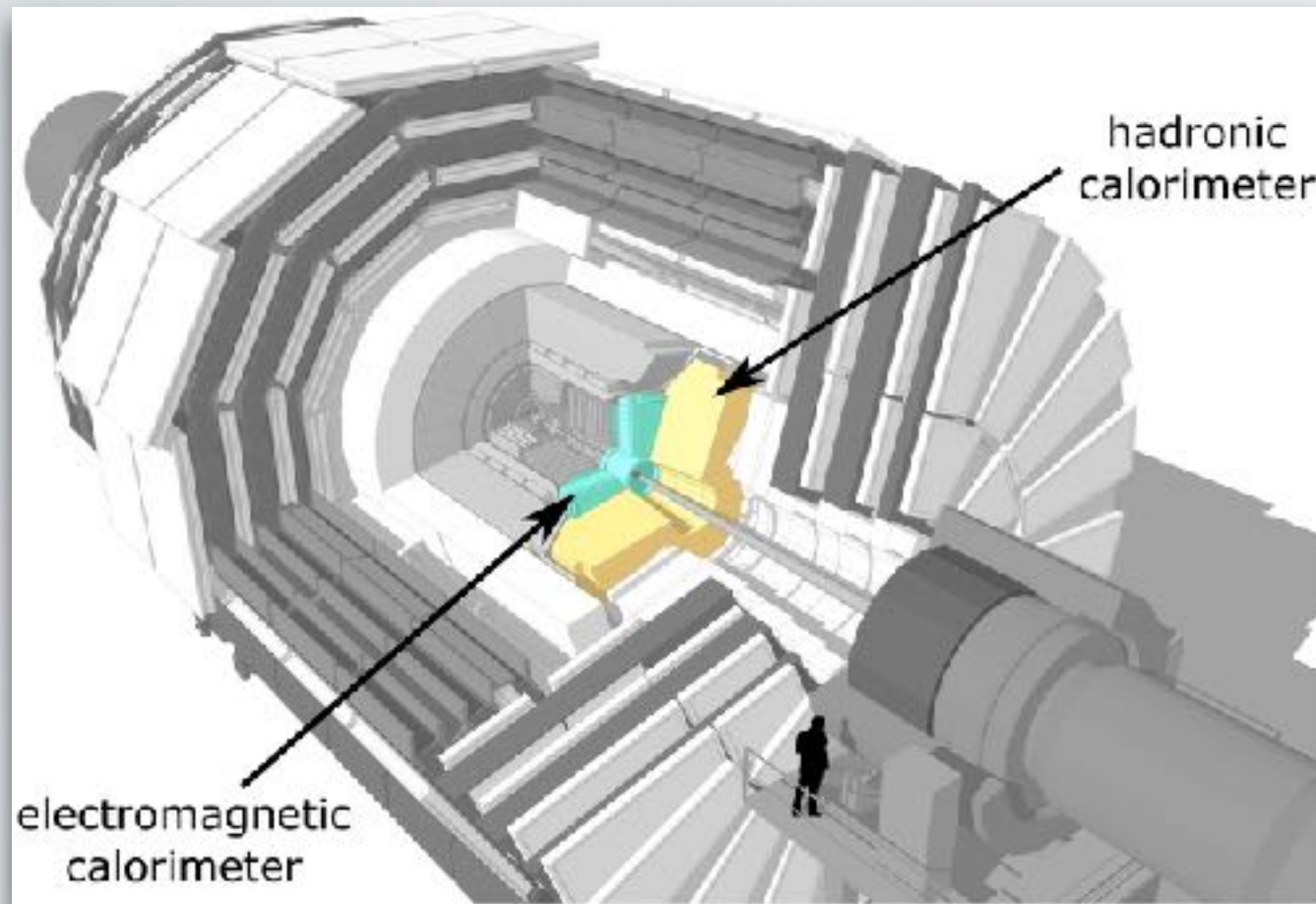
[FASER, 2001.03073]

# APPLICATION TO NEUTRINOS

- Displaced jet signature interesting for **light particles** coupled to **weak scale mediator**  $\Rightarrow$  highly energetic **neutrinos @ LHC**
- Neutrino production typically peaks in very forward direction  $\Rightarrow$  dedicated forward experiments have excellent sensitivity to neutrinos from meson decays
- Large **unused potential** of high energy **W-produced neutrinos!**



# CMS ENDCAP CALORIMETER



[CMS; Nucl.Instrum.Meth.A 978 (2020) 164428]

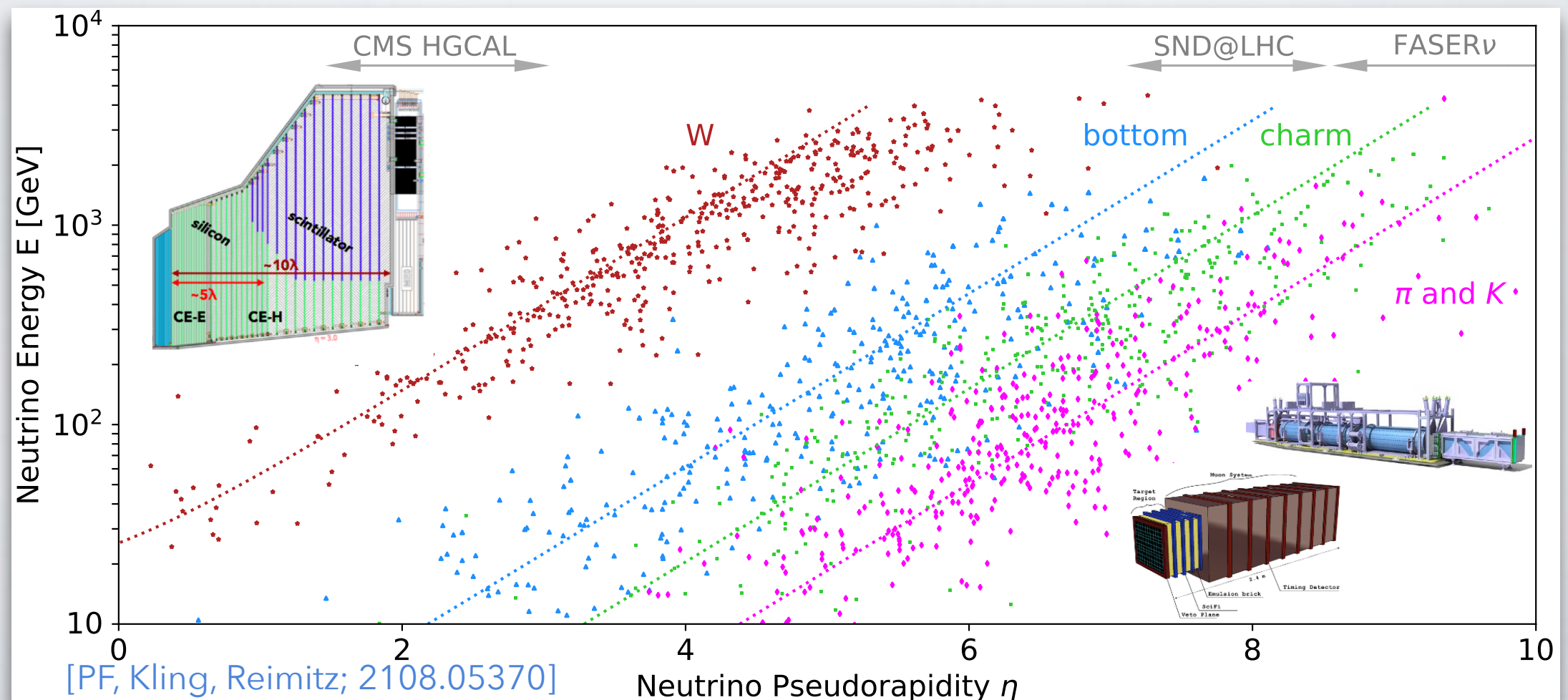
- Upgraded CMS high-granularity endcap calorimeter ideally suited to search for forward neutrino scattering
- Angular coverage in the forward region between  $1.5 \leq |\eta| \leq 3.0$
- High cell granularity (0.5-1) cm<sup>2</sup> allows for high resolution measurement of lateral shower development and good two-shower separation!



# APPLICATION TO NEUTRINOS

- Displaced jet signature interesting for **light particles** coupled to **weak scale mediator**  $\Rightarrow$  highly energetic **neutrinos @ LHC**
- CMS high-granularity endcap calorimeter upgrade (**HGCAL**) can access high-energy neutrinos ( $E_\nu \gtrsim \mathcal{O}(100)$  GeV) from W production!

$\Rightarrow$  How do we search for those neutrinos at CMS?





# NEUTRINOS FROM W DECAY

- **PROBLEM:**  
Huge background of neutral hadron due to pile-up!

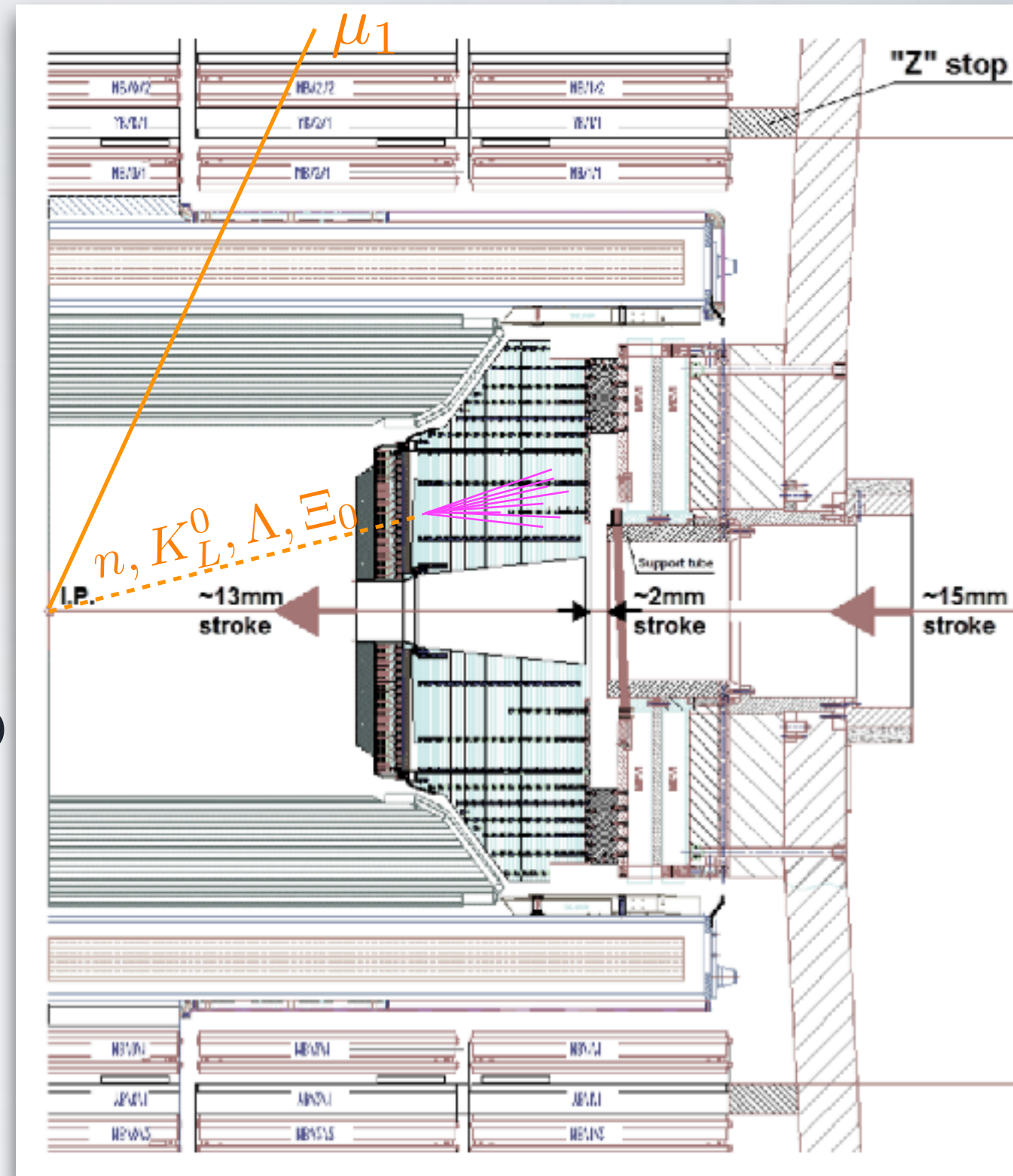
$$qq' \rightarrow W + \text{QCD} \rightarrow \mu_1 + \text{QCD}$$

and heavy hadron decays

$$qq' \rightarrow b/c + \text{QCD} \rightarrow \mu_1 + \text{QCD}$$

- Scattering of neutral hadron can fake neutrino jet:

neutral hadron +  $N \rightarrow \text{jet}$



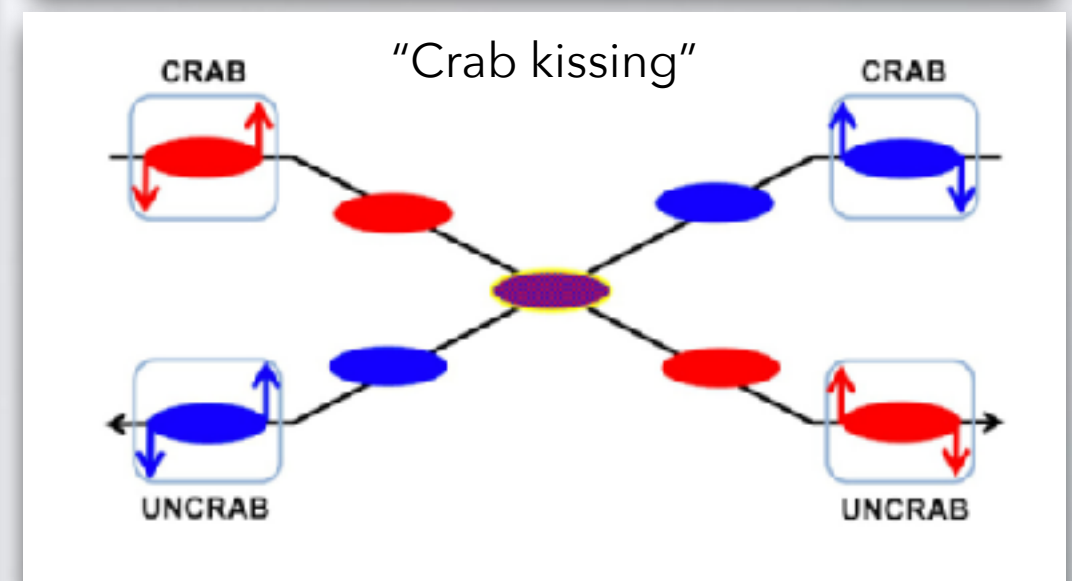
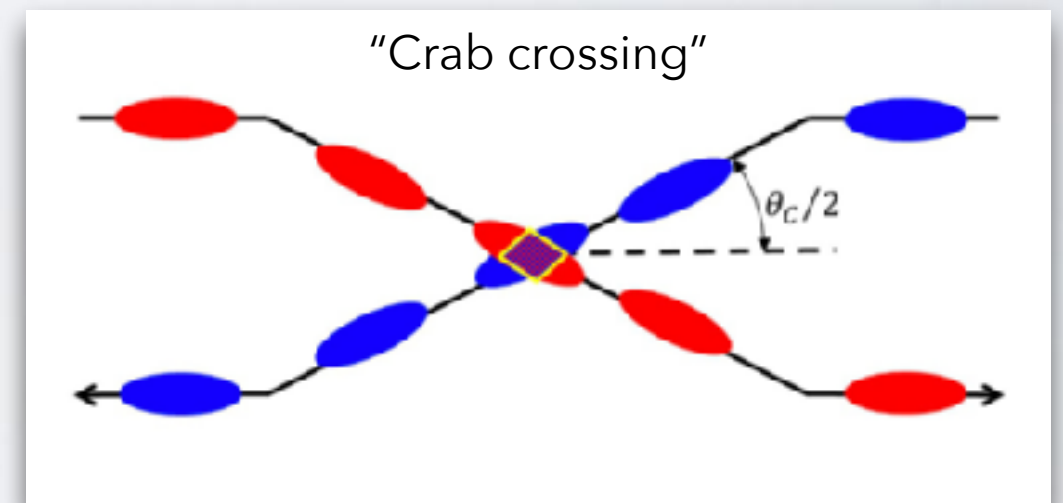
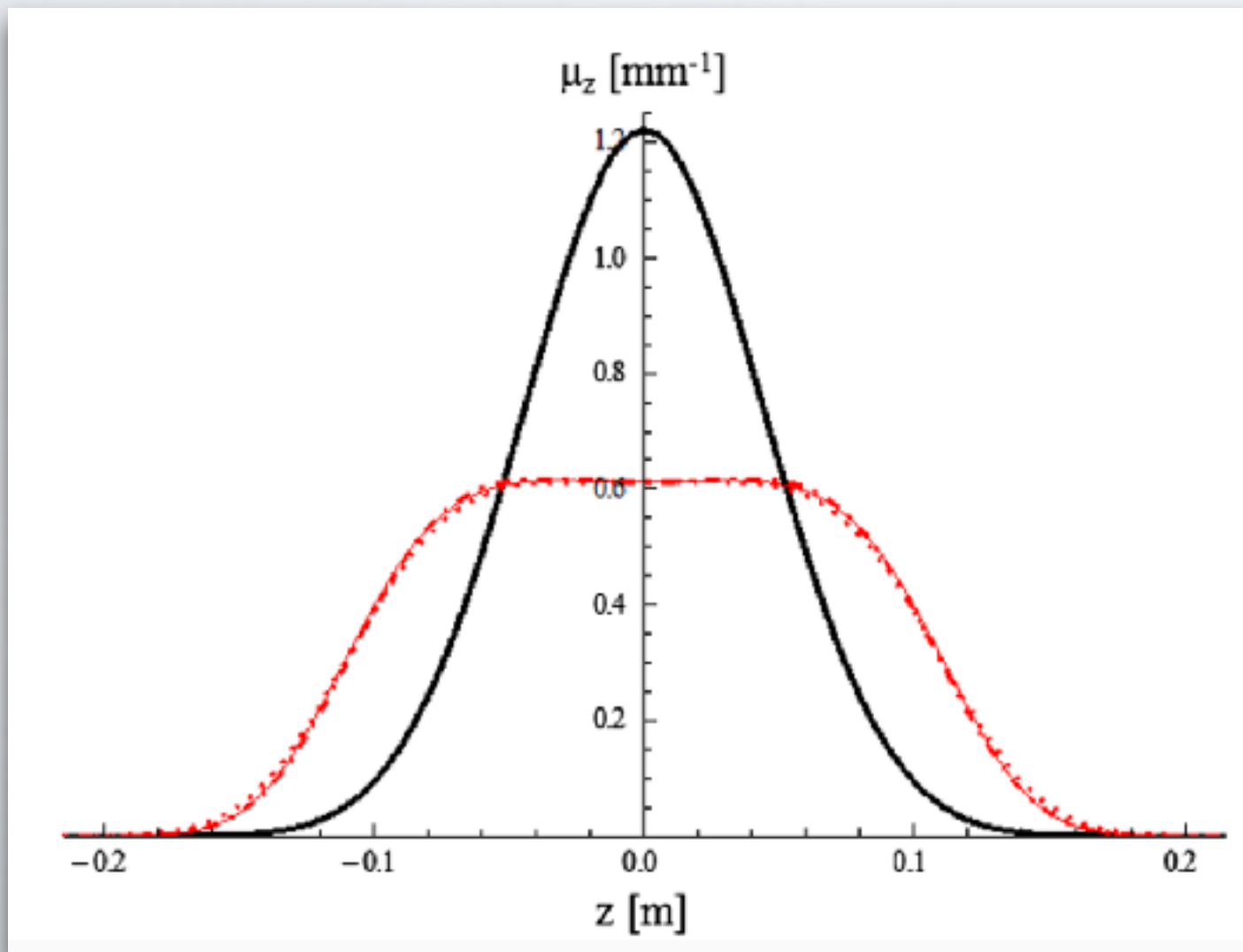


# PILE-UP MITIGATION

- HL-LHC: average of 130-200 pile-up events per bunch crossing ( $\sim 40$  now)
  - *Crab kissing*: Novel collision technique stretches pile-up over  $\sim 31.4$  cm
  - **HGCAL has excellent timing** window of  $\sim 90$  ps/ $\Delta_l \sim 2.7$  cm
- ➔ Can **reduce pile-up to  $\sim 11$**  per bunch crossing

[Verdu-Andres et al.; *Nucl.Part.Phys.Proc.* 273-275 (2016) 193-197]

[Fartoukh; *Phys. Rev. ST Accel. Beams* **17**, 111001]



# NEUTRINOS FROM W DECAY

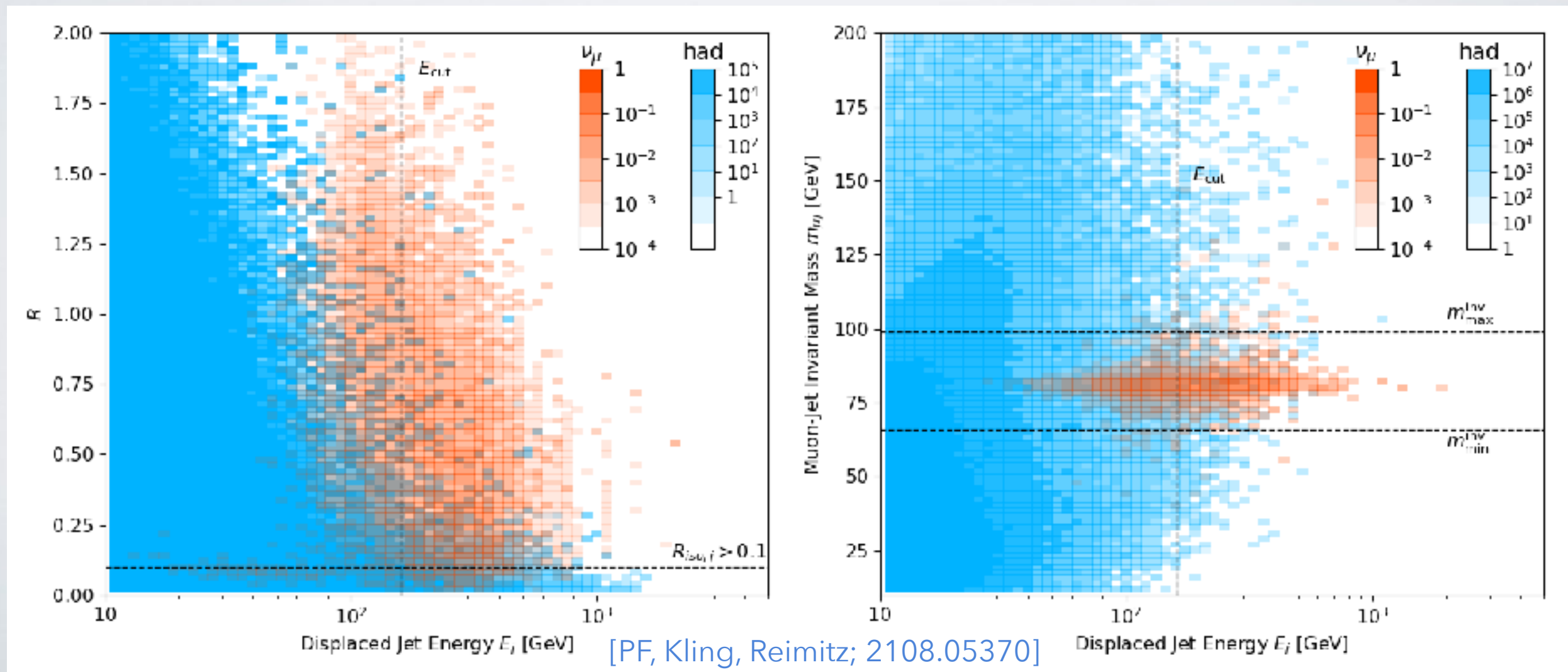
A. Isolated primary muon:  $R_{\text{iso},\mu_1} > 0.1$ ,  $p_{T,\mu_1} > 20 \text{ GeV}$ ,  $|\eta_{\mu_1}| < 2.4$

B. Isolated jet:  $R_{\text{iso},j} > 0.1$

C. W mass cut on invariant mass:  $66 \text{ GeV} < m_{\mu\nu} < 99 \text{ GeV}$

D. Displaced jet energy cut:  $E_{\text{cut}} > 160 \text{ GeV}$

background  
still  
dominates!



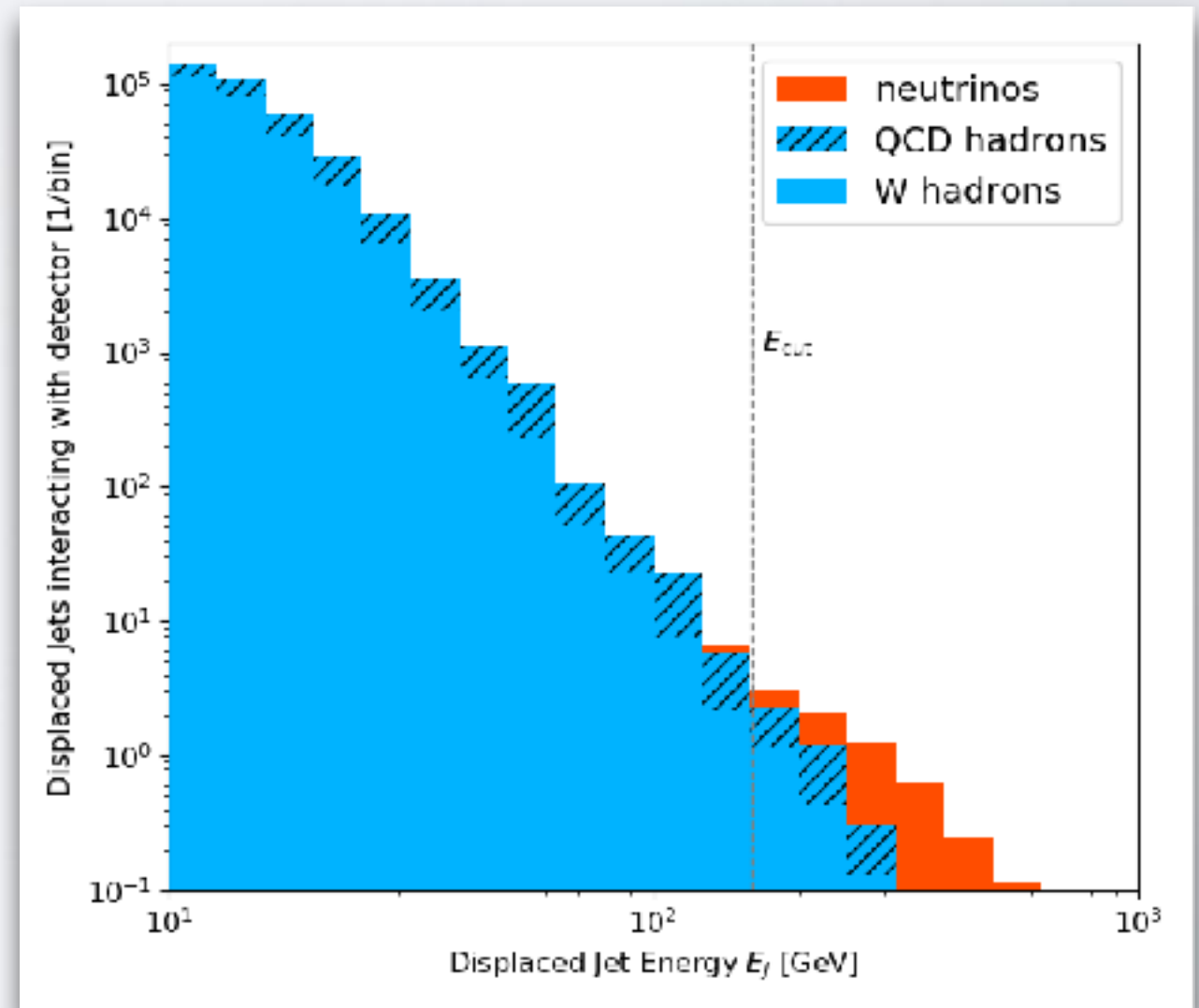
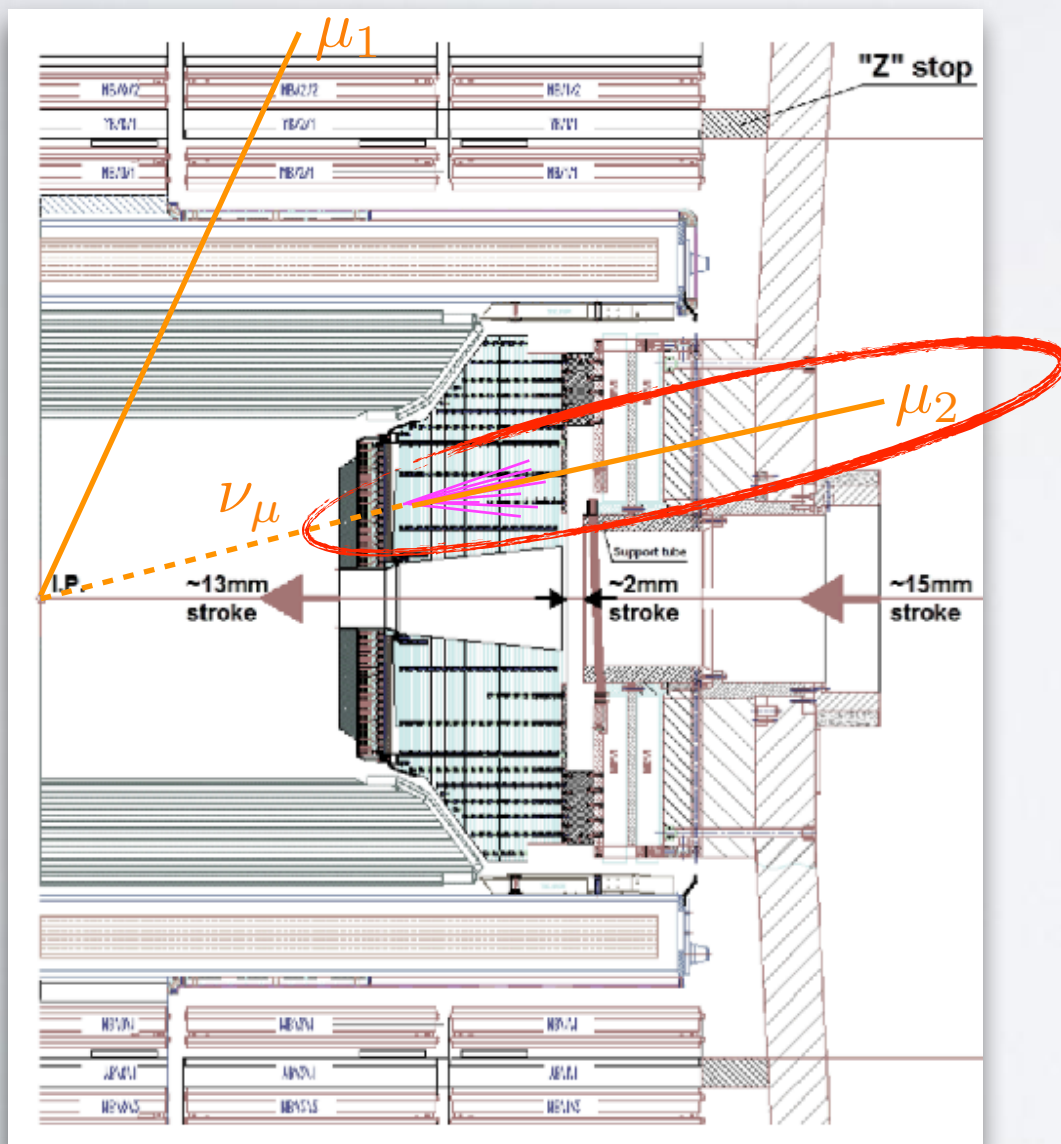
# NEUTRINOS FROM W DECAY

[PF, Kling, Reimitz; 2108.05370]

- Require highly energetic secondary muon:

$$E_{\mu_2} / E_j > 0.33$$

Cuts	Hadrons	Neutrinos
isolated muon	$1.02 \cdot 10^{11}$	7.59
isolated jet	$8.63 \cdot 10^{10}$	7.05
$W$ mass	$1.92 \cdot 10^9$	6.55
secondary muon	$3.49 \cdot 10^5$	5.48
$E_j > 160$ GeV	3.52	3.60



# CONCLUSIONS

- **Displaced recoil jets** are promising signature to search for **light physics at the LHC**
- Complementary to existing direct detection or ULDM probes!  
**Promising for momentum-suppressed interactions!**
- Promising signature to detect **neutrino scattering** at the LHC in **CMS HGCAL**
- Hadronic background suppression via **highly energetic secondary muon**
- More improvements and generalisations:
  - central muon station events
  - sterile neutrino
  - meson decays
  - shower development/jet variables
  - b/c neutrinos

