

(GLOBAL) LEPTON NUMBER SYMMETRY AND NEUTRINO MASSES

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Alexander von Humboldt
Stiftung / Foundation

Accidental symmetries of the SM

The Standard Model has accidental perturbative symmetries, arising from:
gauge group + field content + renormalizability



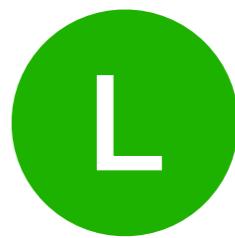
Baryon number

(Individual quark flavour numbers
are violated by CKM mixing)



Flavour numbers

$\alpha = e, \mu, \tau$



Lepton number

$$L = \sum_{\alpha} L_{\alpha}$$

Non perturbative effects violate both B and L, but preserve



$$\partial_{\mu} J_B^{\mu} = \partial_{\mu} J_L^{\mu} = \frac{N_f}{32\pi^2} \epsilon^{\mu\nu\sigma\tau} (-g_W^2 \text{Tr } W_{\mu\nu} W_{\sigma\tau} + g_Y^2 B_{\mu\nu} B_{\sigma\tau})$$

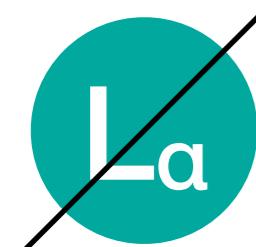
G. 't Hooft, Phys. Rev. Lett. 37 (1976) 8; Phys. Rev. D 14 (1976) 3432

Accidental symmetries: experimental status



**No evidence
of violation**

E.g. proton mean life $> 3.6 \times 10^{29}$ years CL=90%
PDG, Prog. Theor. Exp. Phys. 2022, 083C01 (2022)



**Violated in neutrino
oscillations**



New physics BSM

$$|U|_{3\sigma}^{\text{with SK-atm}} = \begin{pmatrix} 0.801 \rightarrow 0.845 & 0.513 \rightarrow 0.579 & 0.144 \rightarrow 0.156 \\ 0.244 \rightarrow 0.499 & 0.505 \rightarrow 0.693 & 0.631 \rightarrow 0.768 \\ 0.272 \rightarrow 0.518 & 0.471 \rightarrow 0.669 & 0.623 \rightarrow 0.761 \end{pmatrix}$$

I. Esteban, M. C. Gonzalez-Garcia, M. Maltoni, T. Schwetz and A. Zhou, arXiv:2007.14792 [hep-ph]



**No evidence
of violation**

Massive neutrinos violate it if
they are Majorana particles

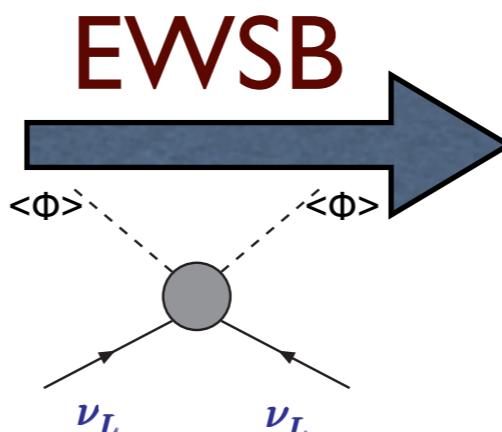
SM as an effective theory

Relaxing the renormalizability condition there is only one dim=5 gauge invariant operator
(Weinberg operator) S. Weinberg, Phys. Rev. Lett. 43 (1979) 1566

Lepton number violation

$$\frac{1}{2} \frac{c_{\alpha\beta}}{\Lambda} \left(\overline{l_L^c} \tilde{\Phi}^* \right) \left(\tilde{\Phi}^\dagger l_L^\beta \right) + h.c.$$

$$\Delta L = 2$$



Neutrino masses and mixing

$$\frac{v^2}{2} \frac{c_{\alpha\beta}}{\Lambda} \overline{\nu_{L\alpha}^c} \nu_{L\beta} + h.c.$$

New physics scale

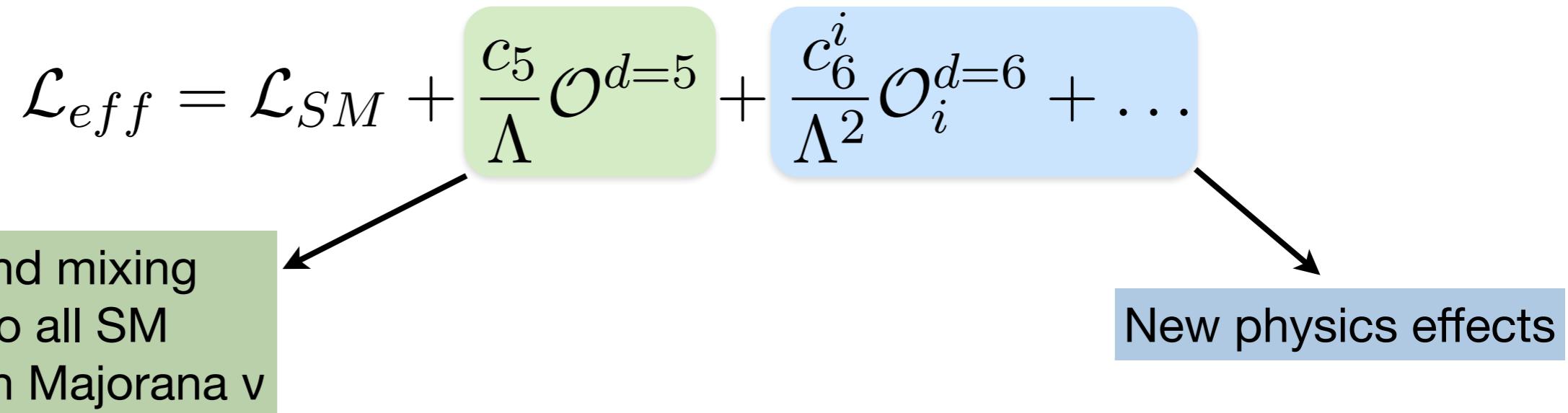
$$m_{\alpha\beta}^\nu = c_{\alpha\beta} \frac{v}{\Lambda} v \lesssim \text{eV} \ll v$$

Why are neutrinos so light?

Suppression mechanisms

- | | | |
|---|---------------------------|--------------------------|
| } | $\frac{v}{\Lambda} \ll 1$ | High NP scale |
| | $c_{\alpha\beta} \ll 1$ | Symmetry (Lepton number) |
| | $c_{\alpha\beta} \ll 1$ | Accidental cancellations |

Unveiling neutrino mass generation mechanism



If only Λ at work

$$\frac{c_6^i}{\Lambda^2} \approx \left(\frac{c_5}{\Lambda}\right)^2 \simeq \left(\frac{m_\nu}{v^2}\right)^2$$

New physics effects strongly suppressed by the ν mass scale

$$c_5 \ll 1 \quad \text{and} \quad c_6^{\text{LNV},i} \ll 1$$

If symmetry at work

$$c_6^{\text{LNC},i} \approx \mathcal{O}(1)$$

possible for L conserving operators

If accidental cancellation

$$c_5 \ll 1$$

$$c_6^i \approx \mathcal{O}(1)$$

possible for all operators

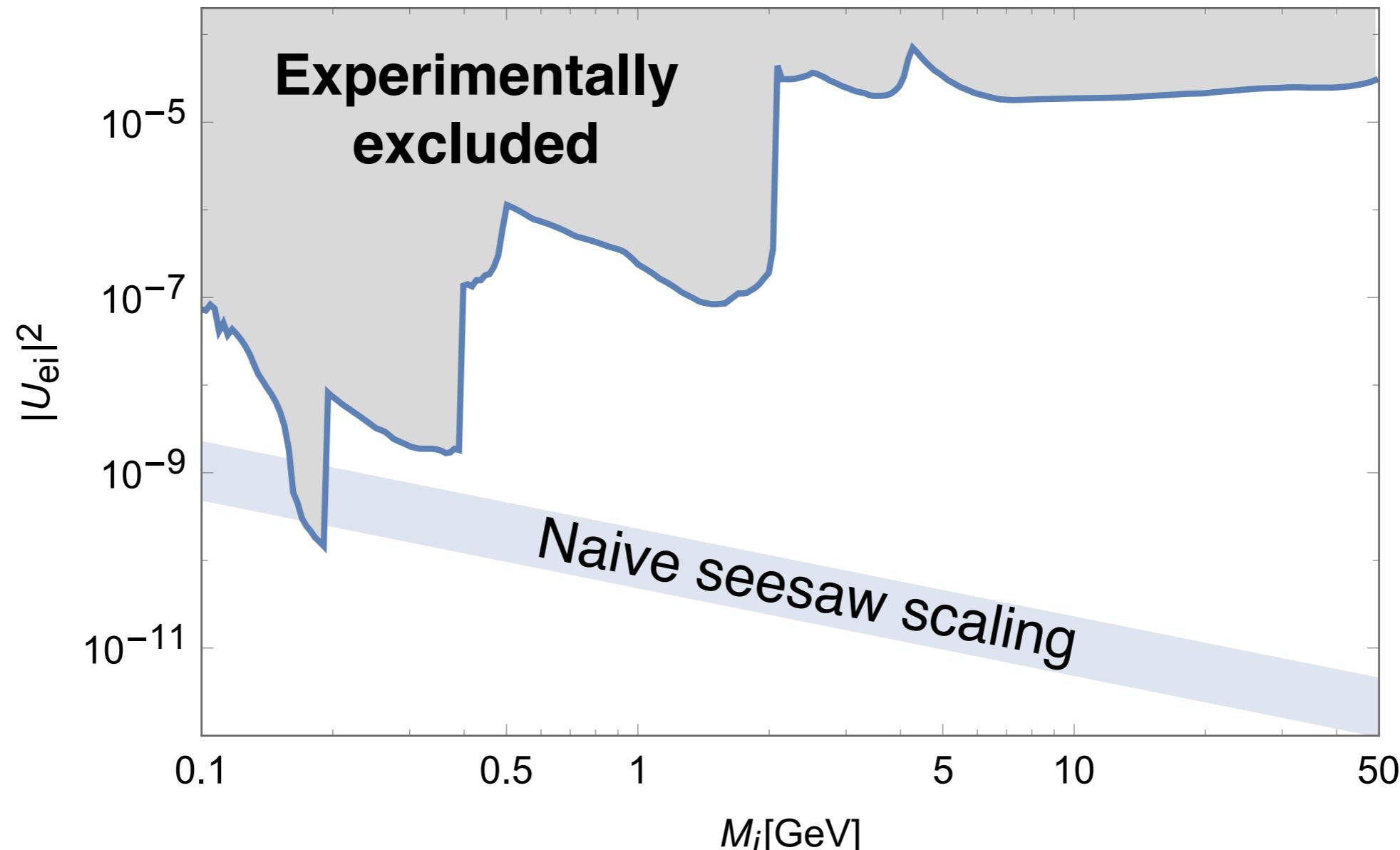
Λ suppression: naive Seesaw scaling

Seesaw scaling

$$m_\nu = -v^2 F \frac{1}{M} F^T$$

In the **absence** of any **structure** in the F and M matrices

$$|U_{\alpha i}| \lesssim \sqrt{\frac{m_\nu}{M}} \lesssim 10^{-5} \sqrt{\frac{\text{GeV}}{M}}$$



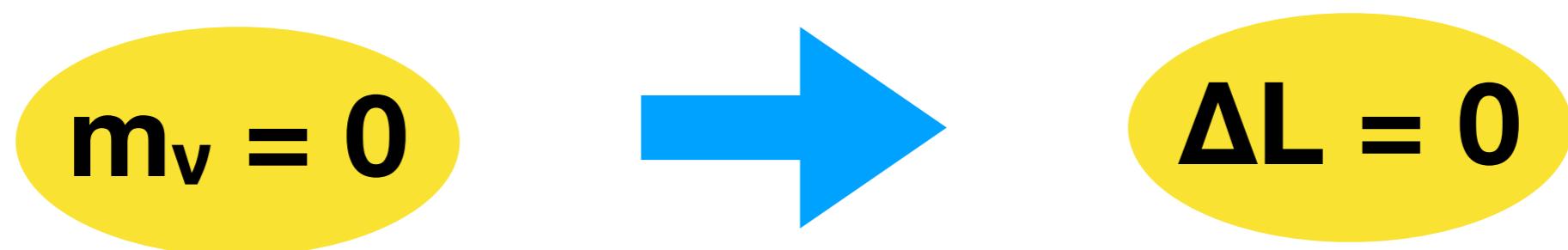
Symmetries: L number has a special role

Theorem: SM + fermionic gauge singlets

K. Moffat, S. Pascoli and C. Weiland, arXiv:1712.07611 [hep-ph]

“The most general gauge-singlet neutrino extensions of the SM with no cancellation between different orders of the seesaw expansion, no fine-tuned cancellations between different radiative orders and which lead to three massless neutrinos are lepton number conserving”

In the SM extended with fermionic gauge singlets (e.g. Right-Handed neutrinos)

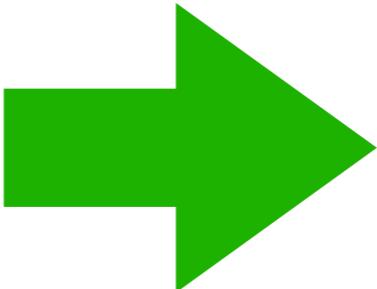


Unless there are accidental cancellations in m_v , the rate for Lepton number violating events is proportional to the small active neutrino masses

The theorem extends and generalises previous results: G. Ingelman and J. Rathsman, Z. Phys. C 60 (1993) 243; J. Gluza, hep-ph/0201002; J. Kersten and A. Y. Smirnov, arXiv:0705.3221 [hep-ph]

Accidental cancellations: quantify fine tuning

If a symmetry is present in the Lagrangian, it will be manifest at any order in perturbation theory



The **neutrino mass scale** is **stable** under **radiative corrections**

Compute neutrino masses m_ν at 1-loop, and quantify the level of fine-tuning of a solution as

$$f.t.(m_\nu) = \sqrt{\sum_{i=1}^3 \left(\frac{m_i^{\text{loop}} - m_i^{\text{tree}}}{m_i^{\text{loop}}} \right)^2}$$

m_i^{loop}

1-loop neutrino
mass spectrum

m_i^{tree}

tree-level neutrino
mass spectrum

Fermionic singlet extensions of the SM

SM + n gauge singlet fermions N_I

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + i\overline{N}_I \not{\partial} N_I - \left(F_{\alpha I} \overline{\ell}_L^\alpha \tilde{\phi} N_I + \frac{M_{IJ}}{2} \overline{N}_I^c N_J + h.c. \right)$$

↓
*3 x n matrix
Yukawa couplings*

 ↑
*n x n matrix
Majorana mass
couplings*

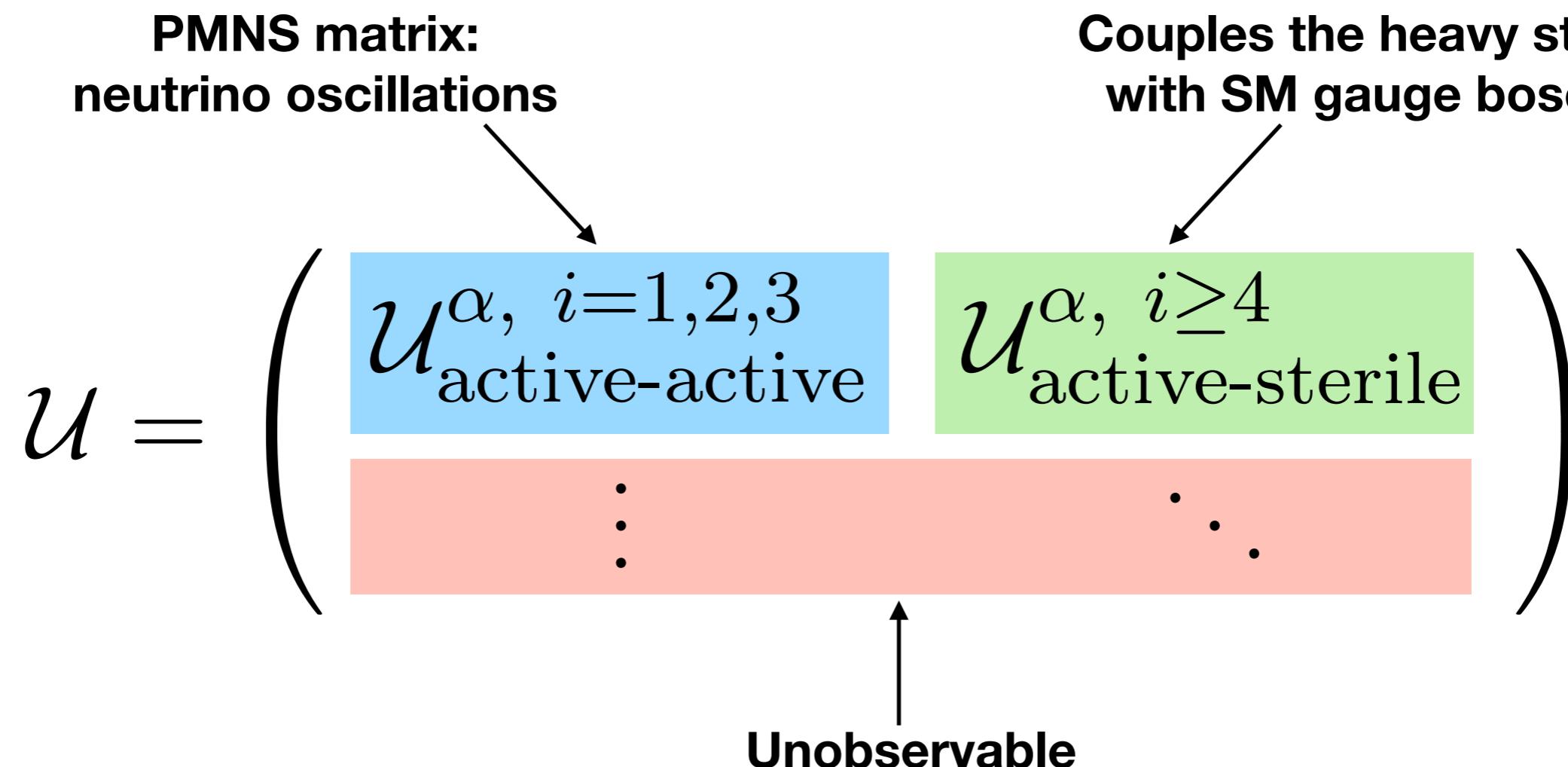
After electroweak phase transition $\langle \Phi \rangle = v \approx 174 \text{ GeV}$

$$-\mathcal{L}_m^\nu = \frac{1}{2} \begin{pmatrix} \overline{\nu}_L & \overline{N}^c \end{pmatrix} \underbrace{\begin{pmatrix} \delta m_\nu^{\text{loop}} & vF \\ vF^T & M \end{pmatrix}}_{\mathcal{M}} \begin{pmatrix} \nu_L^c \\ N \end{pmatrix} + h.c.$$

↓
*(3+n) dimensional
mass matrix*

Phenomenology of fermionic singlets

$$\mathcal{U}^T \mathcal{M} \mathcal{U} = \hat{\mathcal{M}}_{\text{diag}} \rightarrow \begin{cases} 3 \text{ light (mostly active) states} \\ n \text{ heavy (mostly sterile) states} \end{cases}$$



L symmetry and Majorana fields

Majorana fermions violate all global symmetries, including L

How to preserve lepton number with Majorana states?

	Pair two states to form a Dirac state (equal masses, <i>maximal</i> <i>mixing</i> , opposite CP)	Decouple a state	Have a massless state
Exact symmetry	$M_1 = M_2$ $\mathcal{U}_{\alpha 1} = i \mathcal{U}_{\alpha 2}$	$\mathcal{U}_{\alpha i} = 0$	$M_i = 0$
Approximate symmetry	$\frac{M_2 - M_1}{M_1 + M_2} \ll 1$ $\mathcal{U}_{\alpha 1} \simeq i \mathcal{U}_{\alpha 2}$	$ \mathcal{U}_{\alpha,i} \ll \mathcal{U}_{\alpha,j \neq i} $	$M_i \ll M_{j \neq i}$

NEUTRINOLESS DOUBLE BETA DECAY

Double beta decay

2β decay: 2nd order weak process

$$\mathcal{N}(A, Z) \rightarrow \mathcal{N}(A, Z + 2) + 2e^- + 2\bar{\nu}_e$$

Only relevant when the single β decay is kinematically forbidden

$^{48}\text{Ca}, ^{76}\text{Ge}, ^{82}\text{Se}, ^{96}\text{Zr}, ^{100}\text{Mo}, ^{116}\text{Cd}, ^{130}\text{Te}, ^{136}\text{Xe}, ^{150}\text{Nd}$

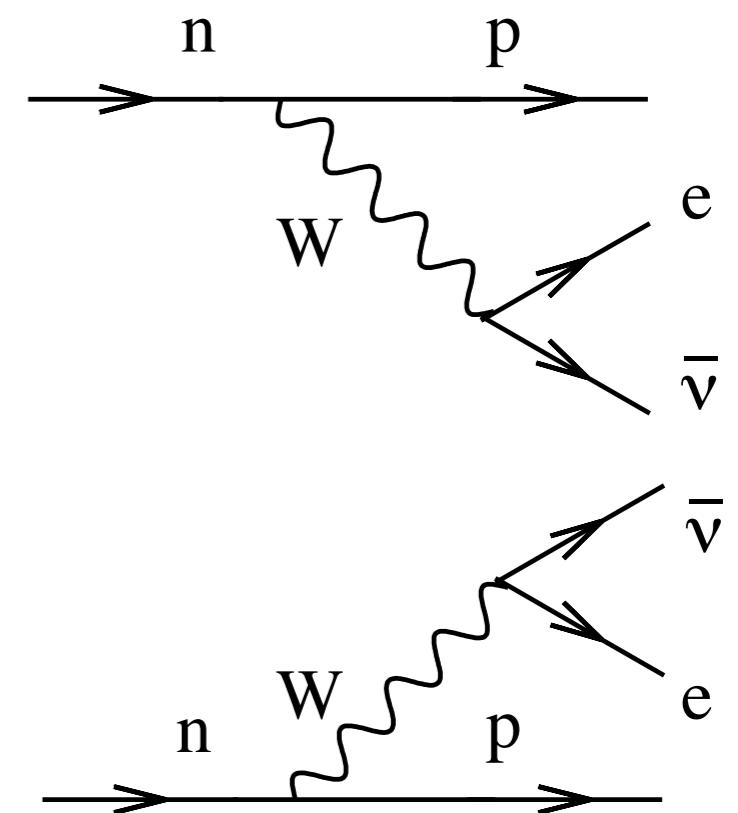
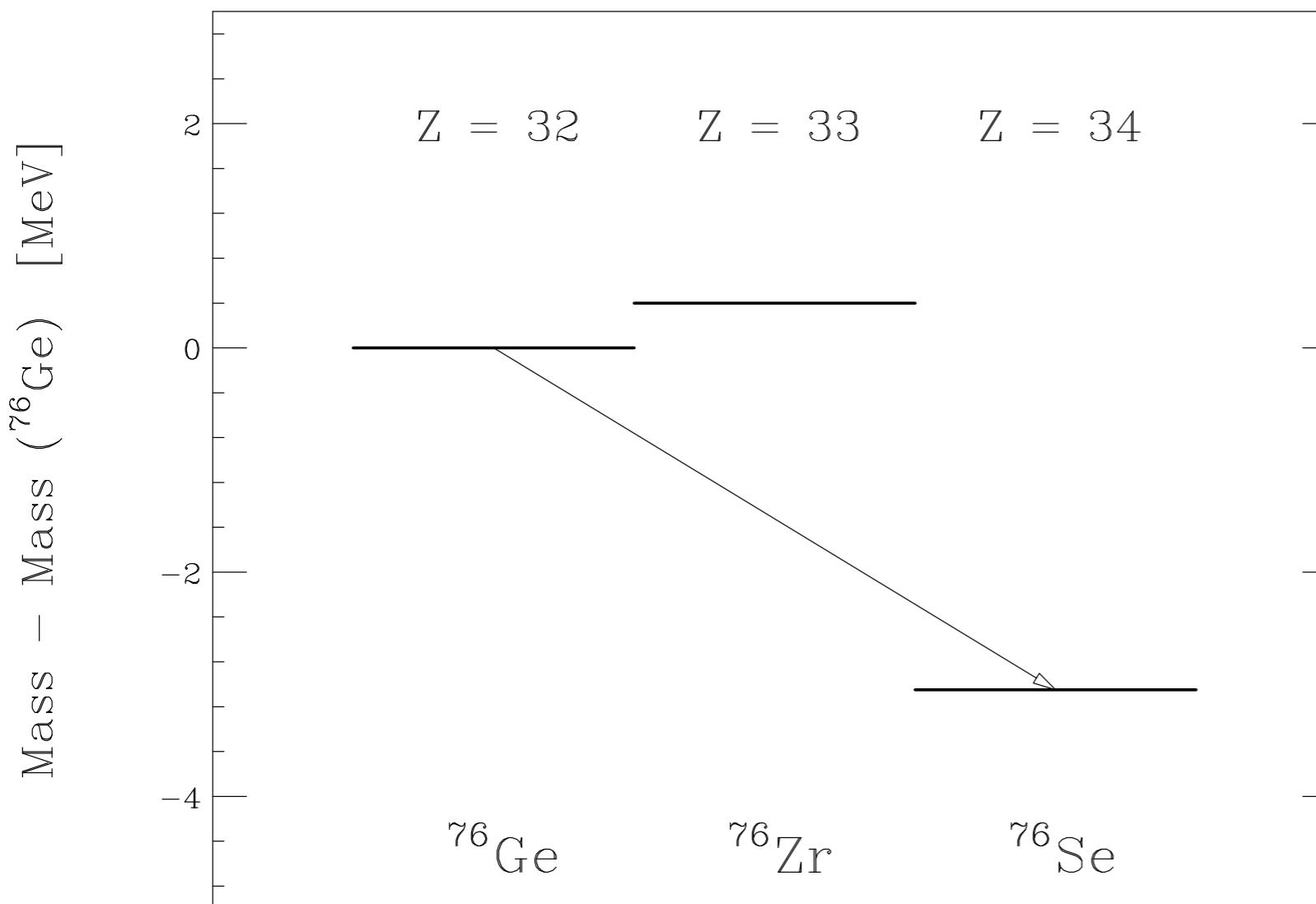


Figure from P. Lipari, Introduction to neutrino physics, in 2001 CERN-CLAF School of high-energy physics

Neutrinoless double beta decay: $\Delta L = 2$

W. H. Furry, Phys. Rev. 56 (1939) 1184

If neutrinos are Majorana particles $0\nu2\beta$ is possible

$$\mathcal{N}(A, Z) \rightarrow \mathcal{N}(A, Z + 2) + 2e^-$$

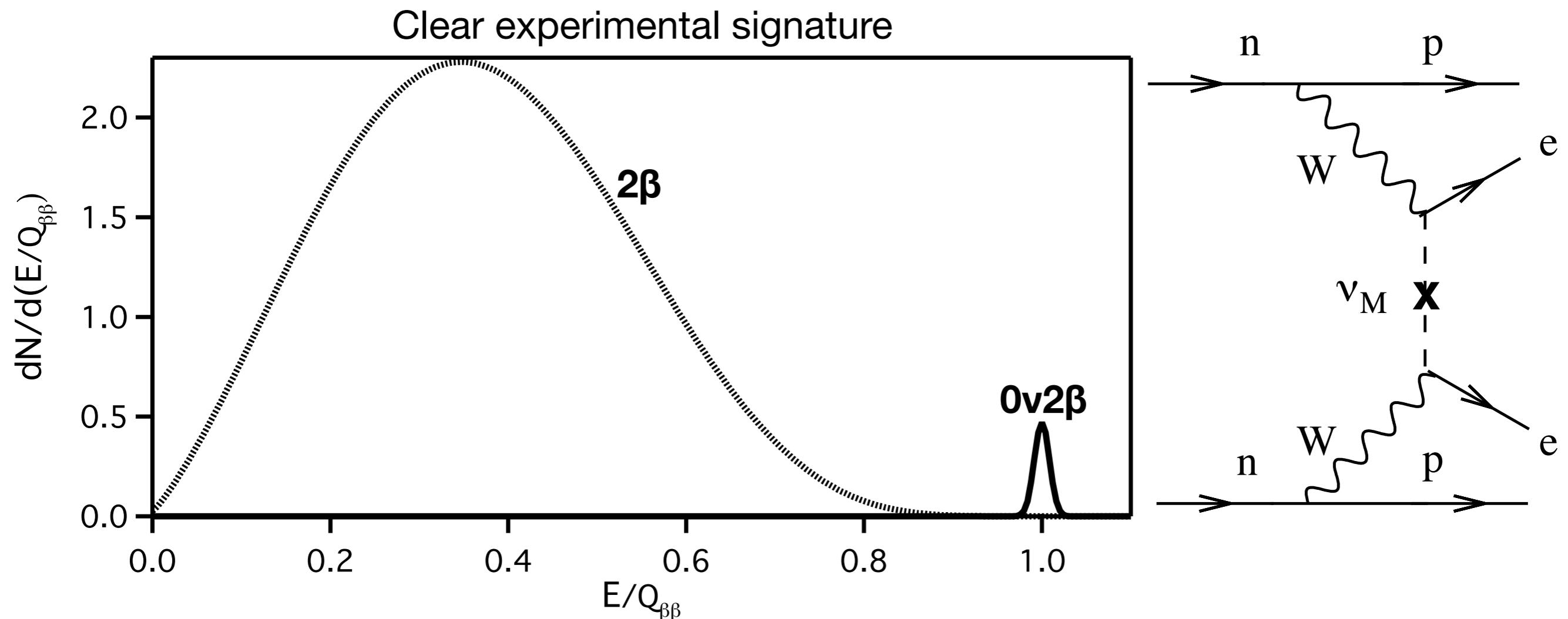


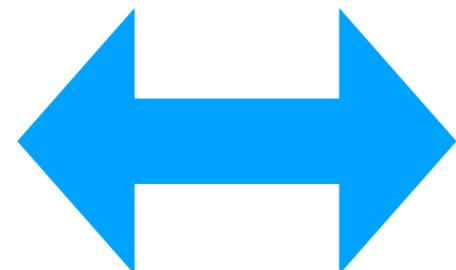
Figure modified from F. T. Avignone III, S. R. Elliott and J. Engel, arXiv:0708.1033 [nucl-ex]

The black box theorem

J. Schechter and J. W. F. Valle, Phys. Rev. D 25 (1982) 2951; E. Takasugi, Phys. Lett. 149B (1984) 372;
see also M. Duerr, M. Lindner and A. Merle, arXiv:1105.0901 [hep-ph]

Non-vanishing
0v2 β amplitude

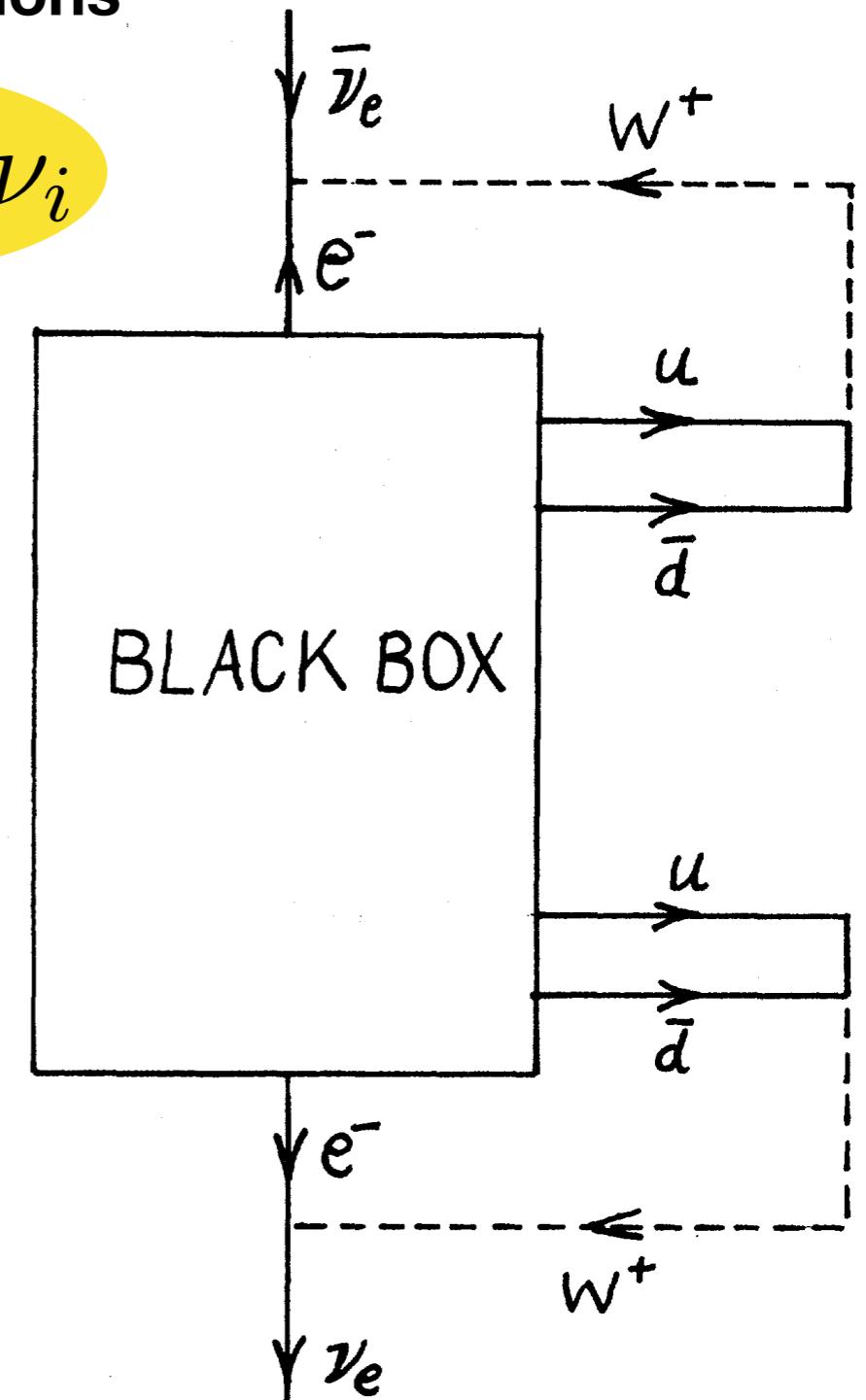
$$\Gamma_{0\nu 2\beta} \neq 0$$



Neutrinos are
Majorana fermions

$$\nu_i^c = e^{i\phi} \nu_i$$

Irrespectively of the underlying mechanism, a non-vanishing 0v2 β amplitude generates a Majorana mass term for the SM neutrinos

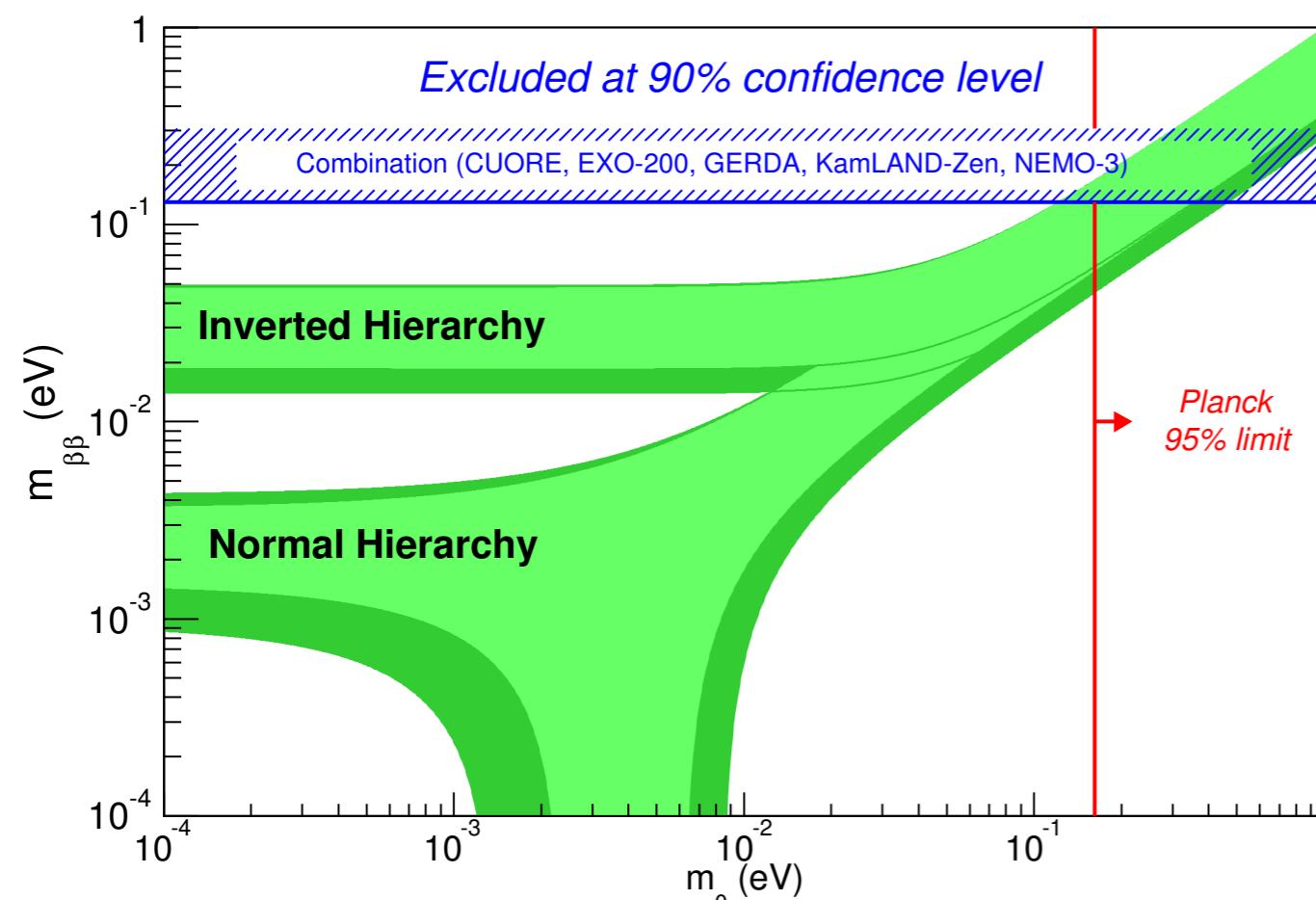


Experimental status: minimal SM

The amplitude for light neutrino exchange is proportional to

$$m_{2\beta} = \left| \sum_i U_{ei}^2 m_i \right|$$

From current knowledge on neutrino oscillation parameters it is possible to compute $m_{2\beta}$ as a function of unknown lightest neutrino mass, ordering and CP phases



Current bounds

Isotope	$T_{1/2}^{0\nu} (\times 10^{25} \text{ y})$	$\langle m_{\beta\beta} \rangle (\text{eV})$	Experiment
^{48}Ca	$> 5.8 \times 10^{-3}$	$< 3.5 - 22$	ELEGANT-IV
^{76}Ge	> 8.0	$< 0.12 - 0.26$	GERDA
^{82}Se	> 1.9	$< 0.24 - 0.52$	MAJORANA DEMONSTRATOR
	$> 3.6 \times 10^{-2}$	$< 0.89 - 2.43$	NEMO-3
^{96}Zr	$> 9.2 \times 10^{-4}$	$< 7.2 - 19.5$	NEMO-3
^{100}Mo	$> 1.1 \times 10^{-1}$	$< 0.33 - 0.62$	NEMO-3
^{116}Cd	$> 1.0 \times 10^{-2}$	$< 1.4 - 2.5$	NEMO-3
^{128}Te	$> 1.1 \times 10^{-2}$	—	—
^{130}Te	> 1.5	$< 0.11 - 0.52$	CUORE
^{136}Xe	> 10.7	$< 0.061 - 0.165$	KamLAND-Zen
^{150}Nd	> 1.8	$< 0.15 - 0.40$	EXO-200
	$> 2.0 \times 10^{-3}$	$< 1.6 - 5.3$	NEMO-3

Table from M. J. Dolinski, A. W. P. Poon and W. Rodejohann, arXiv:1902.04097 [nucl-ex]

Figure from P. Guzowski, L. Barnes, J. Evans, G. Karagiorgi, N. McCabe and S. Soldner-Rembold, arXiv:1504.03600 [hep-ex]

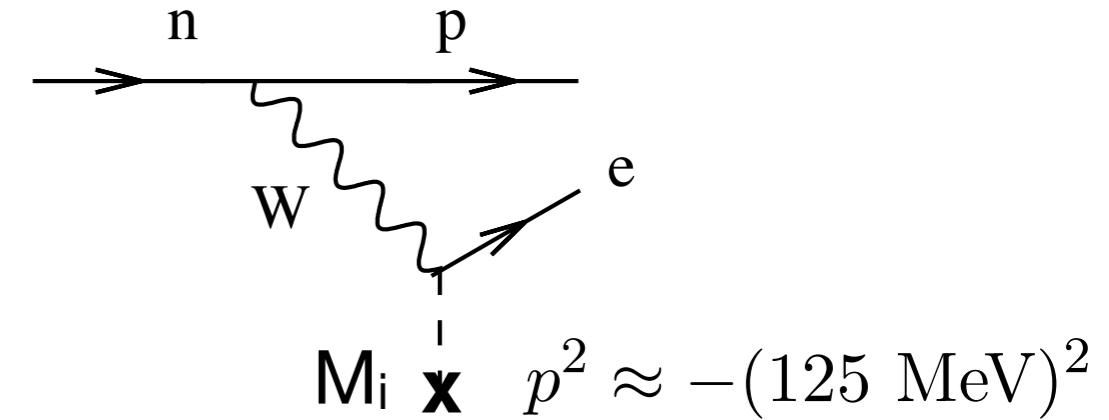
Contribution of heavy neutrinos

Heavy Majorana neutrinos contribute as well to $0\nu2\beta$ amplitude

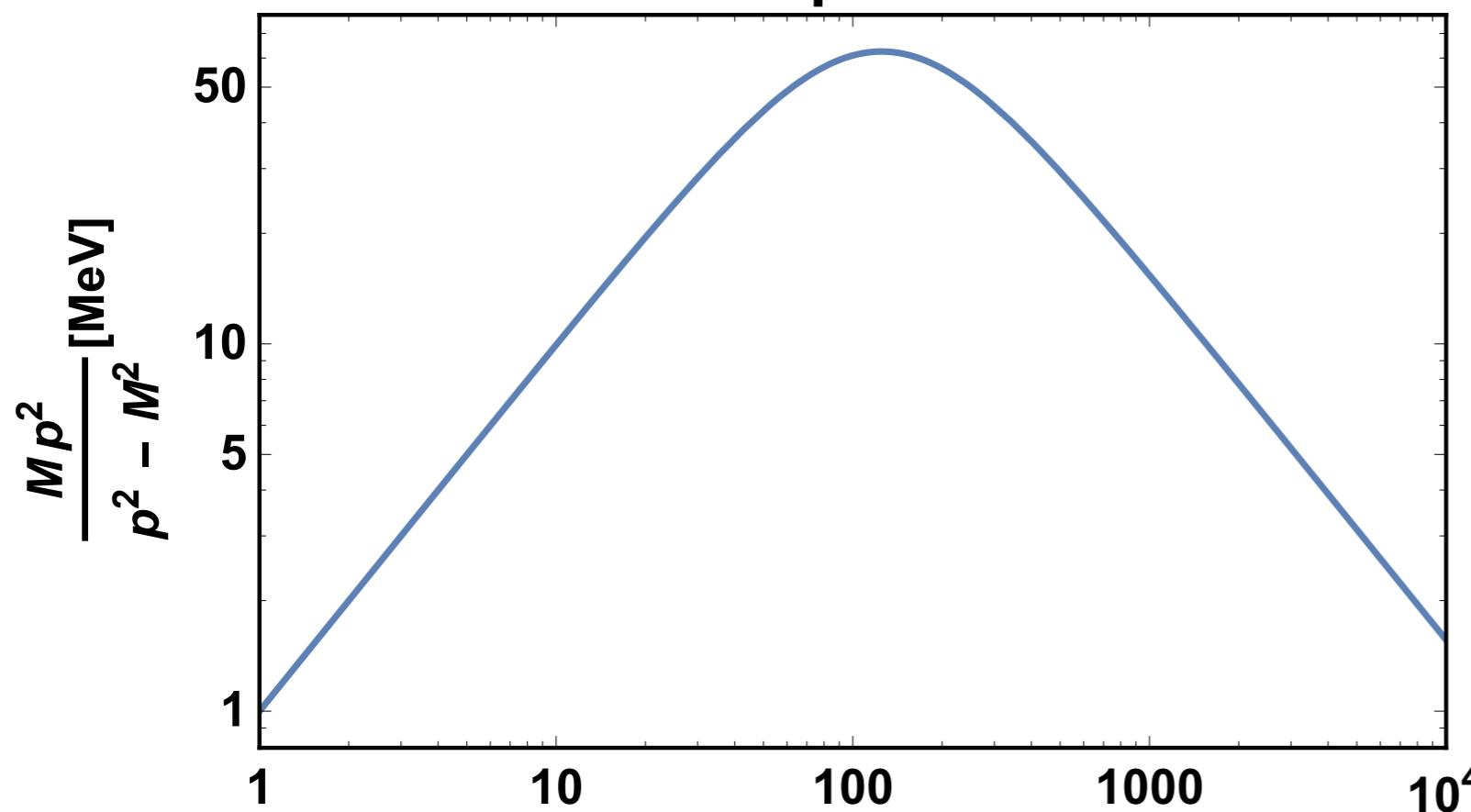
F. L. Bezrukov, hep-ph/0505247; M. Blennow, E. Fernandez-Martinez, J. Lopez-Pavon and J. Menendez, arXiv:1005.3240 [hep-ph]; A. Abada and M.L., arXiv:1401.1507 [hep-ph]; A. Faessler, M. González, S. Kovalenko and F. Šimkovic, arXiv:1408.6077 [hep-ph]; A. Abada, V. De Romeri, M.L., A. M. Teixeira and T. Toma, arXiv:1712.03984 [hep-ph]; A. Babič, S. Kovalenko, M. I. Krivoruchenko and F. Šimkovic, arXiv:1804.04218 [hep-ph]

$$\mathcal{A}^{0\nu2\beta} \propto \sum_i M_i \mathcal{U}_{ei}^2 M^{0\nu2\beta}(M_i)$$

$$M^{0\nu2\beta}(M_i) \simeq M^{0\nu2\beta}(0) \frac{p^2}{p^2 - M_i^2}$$



Mass dependence



If pseudo-Dirac

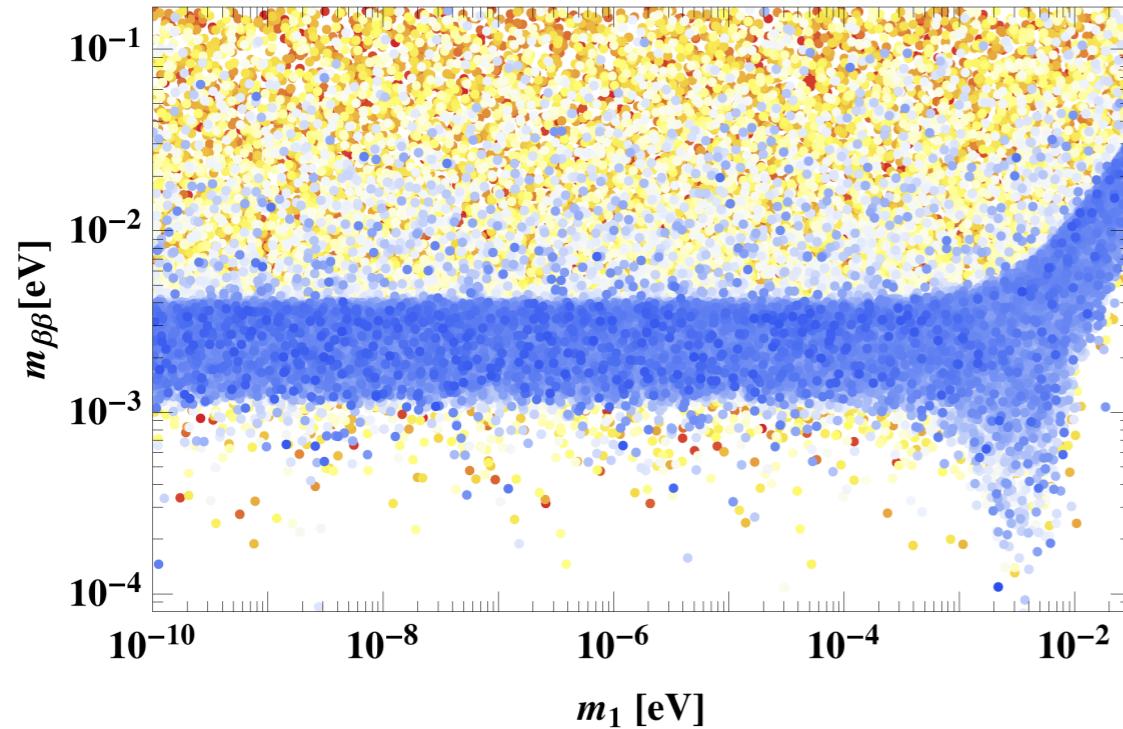
$$M_1 \simeq M_2$$
$$\mathcal{U}_{e1} \simeq i \mathcal{U}_{e2}$$

*cancellation between contributions
of single Majorana states*

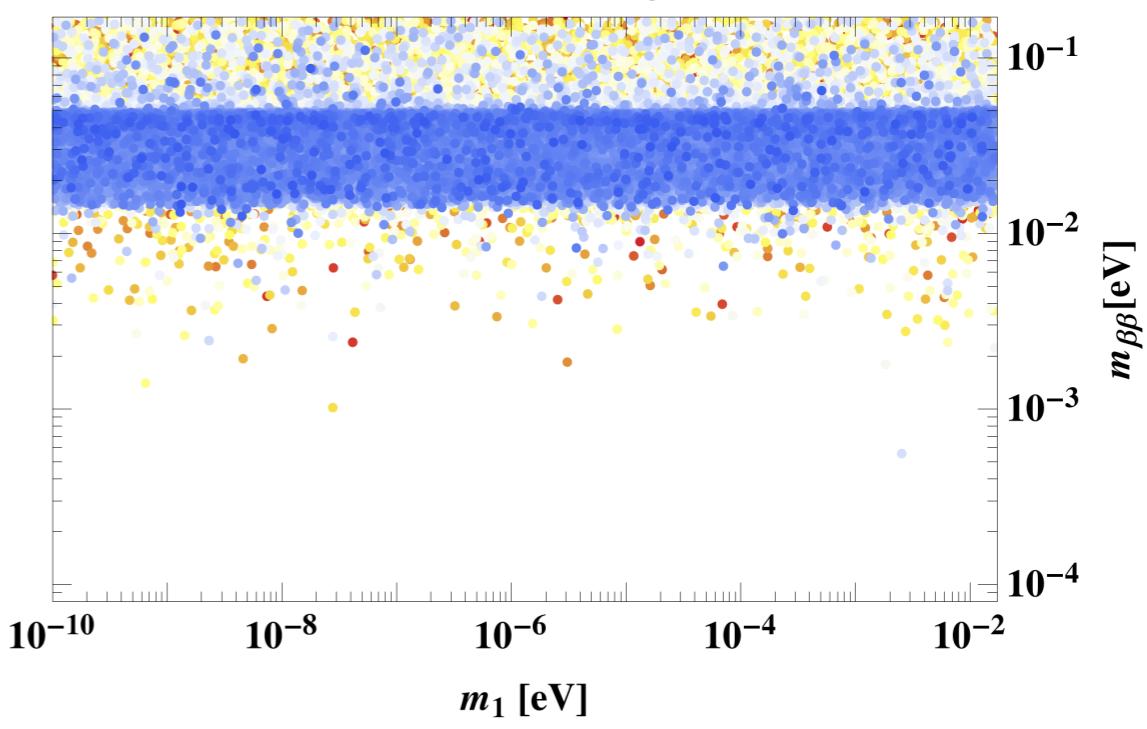
Heavy neutrinos at GeV scale

Blue points: not fine tuned  **Red points: fine tuned**

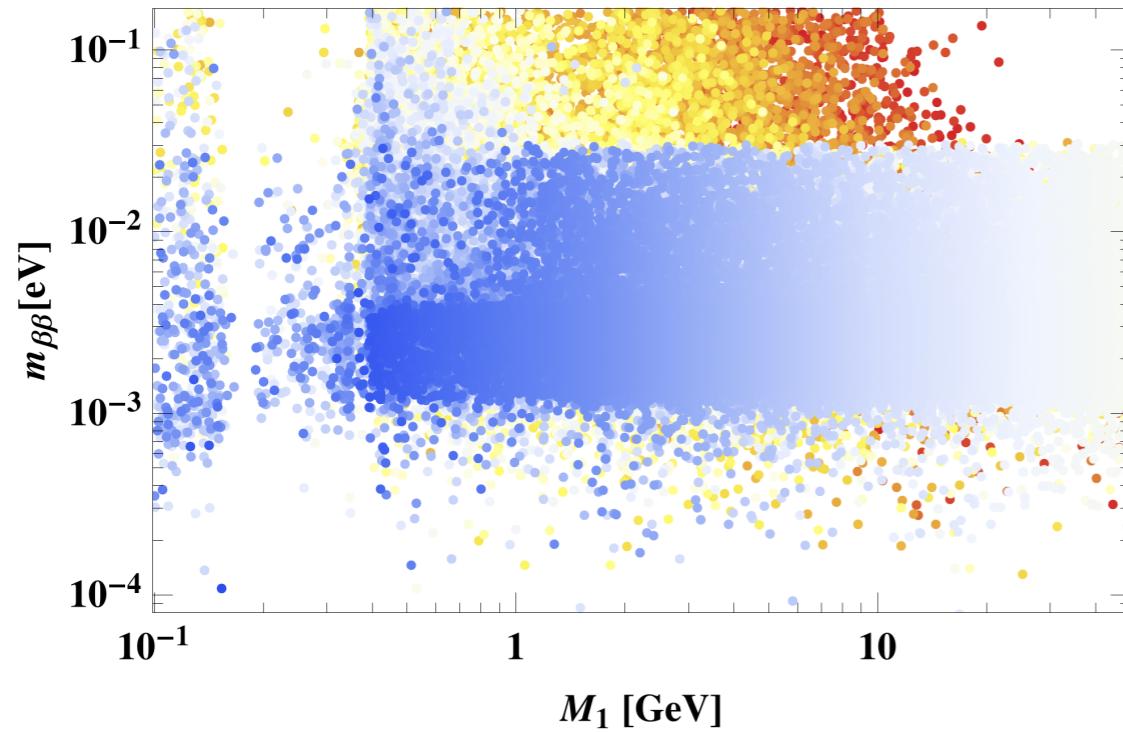
Normal Ordering



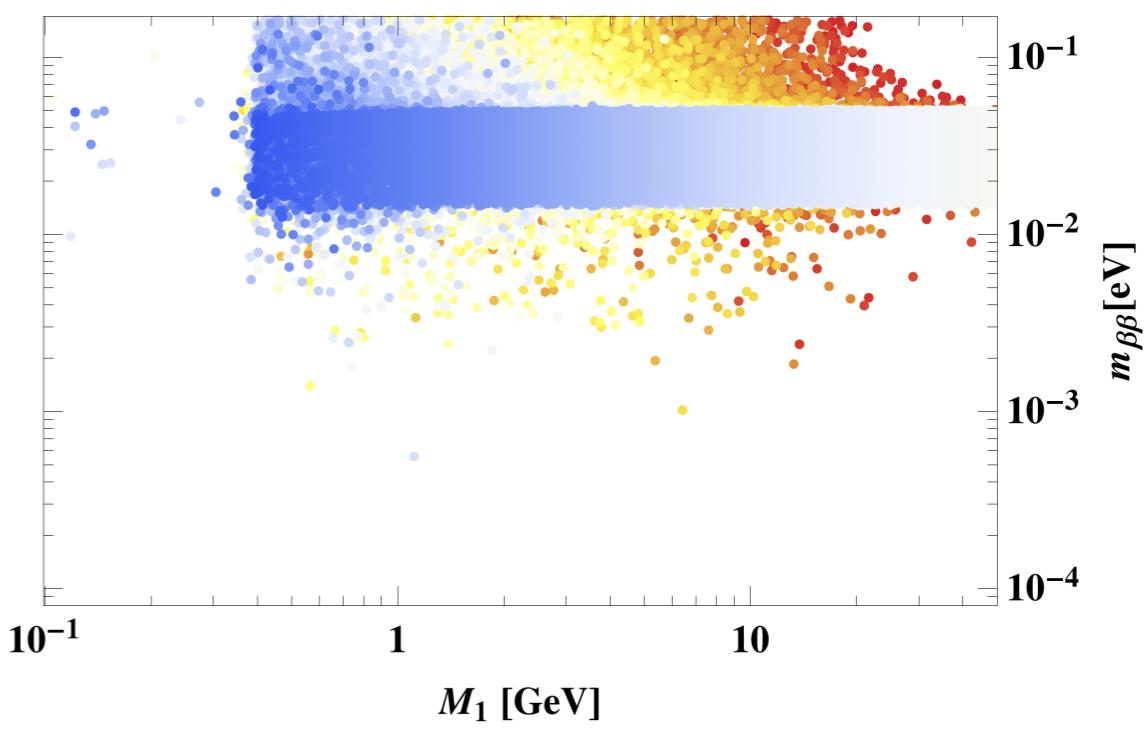
Inverted Ordering



m_1 [eV]



m_1 [eV]



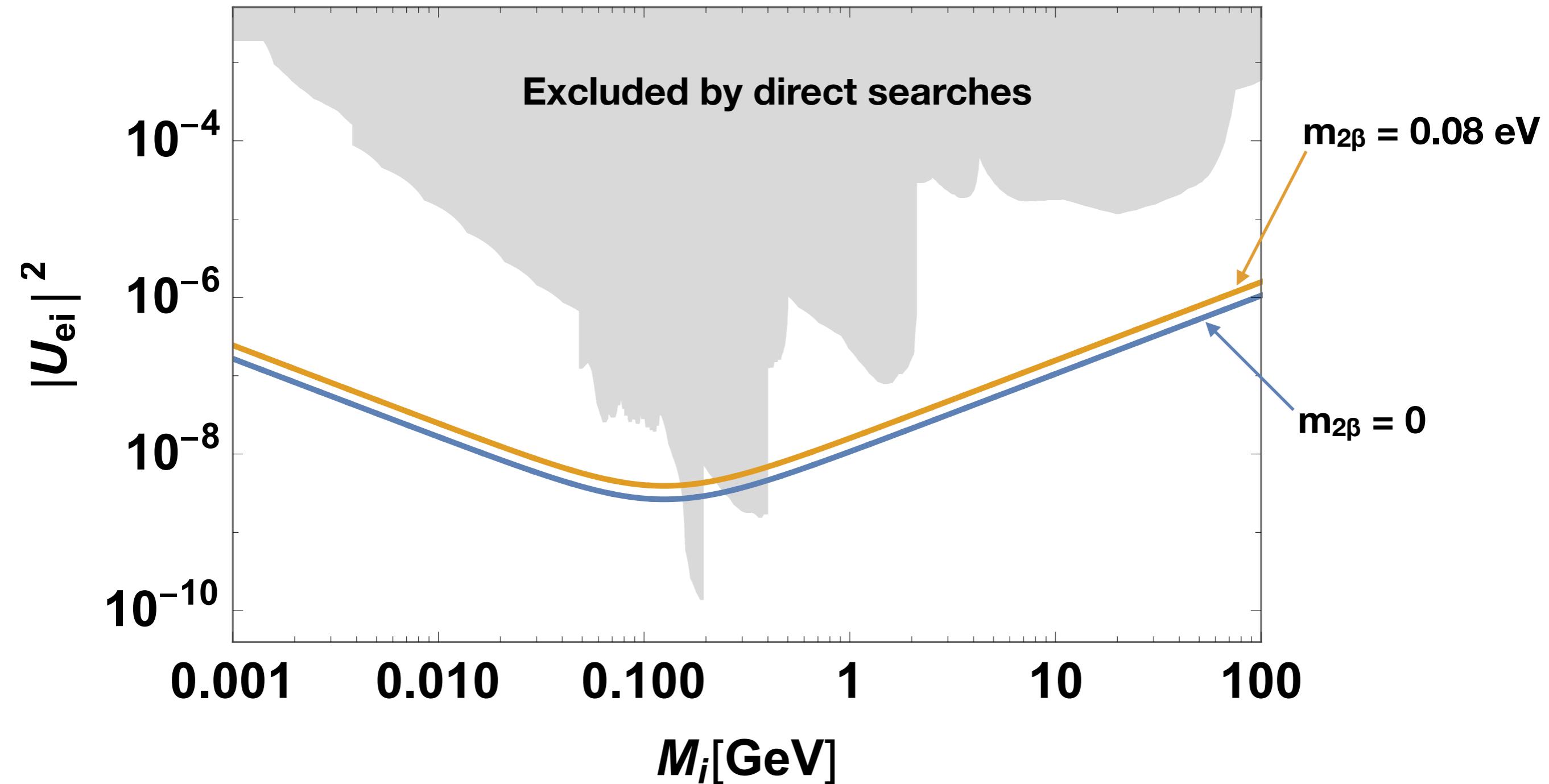
M_1 [GeV]

**Figures from A. Abada, G. Arcadi, V. Domcke, M. Drewes, J. Klaric and M. Lucente, arXiv:1810.12463 [hep-ph];
see also J. Lopez-Pavon, S. Pascoli and C. f. Wong, arXiv:1209.5342 [hep-ph]; J. Lopez-Pavon, E. Molinaro
and S. T. Petcov, arXiv:1506.05296 [hep-ph]**

Extracting constraints on heavy neutrinos

$$\mathcal{A}^{0\nu 2\beta} \propto \sum_{i=1}^{3+n} M_i |U_{ei}|^2 M^{0\nu 2\beta}(M_i)$$

0v2 β constraints depend on the full mass spectrum (light + heavy)



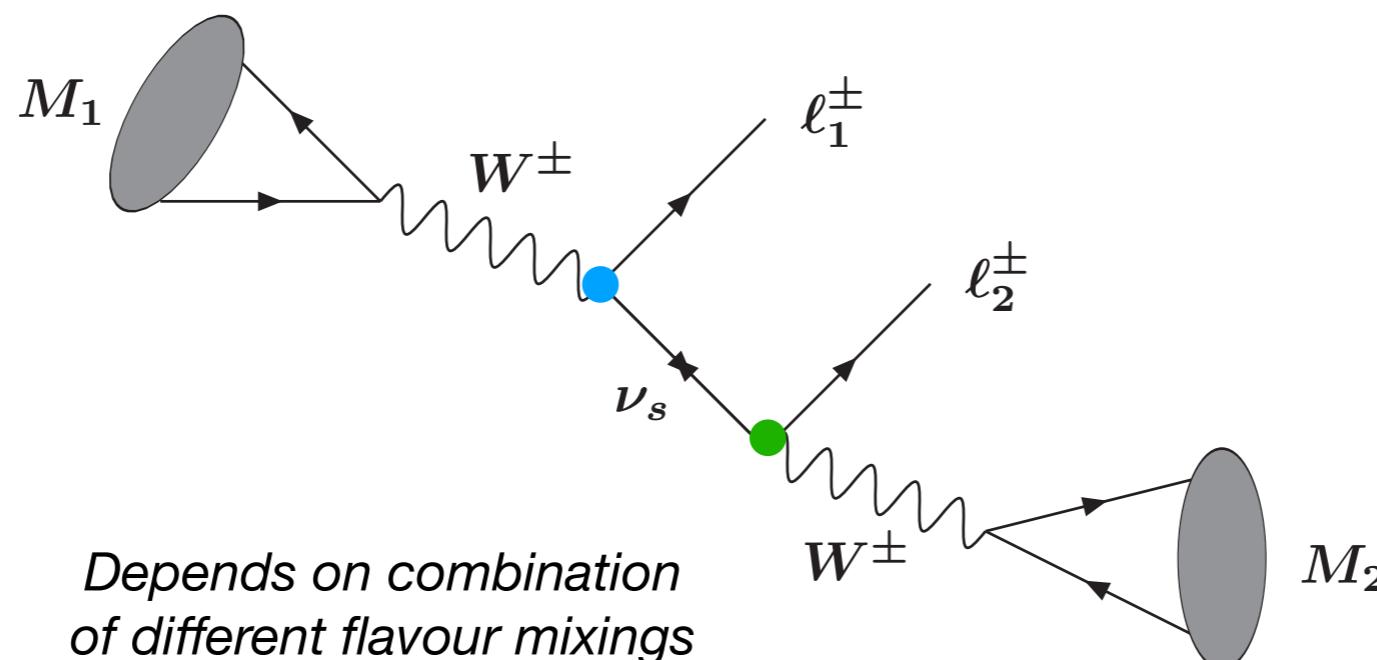
These constraints do not apply to (pseudo-)Dirac particles

TAU AND MESON DECAY

L-violating τ and meson decay

Heavy Majorana neutrinos can mediate L-violating decays of pseudo-scalar mesons and τ lepton

$$M_1(p, m_{M_1}) \rightarrow \ell_\alpha(k_1, m_{\ell_\alpha}) \ell_\beta(k_2, m_{\ell_\beta}) M_2(k_3, m_{M_2})$$



$$i\mathcal{M}_P \equiv i\mathcal{M}_{P1} + i\mathcal{M}_{P2} = 2i G_F^2 V_{M_1} V_{M_2} U_{\ell_\alpha 4} U_{\ell_\beta 4} m_4 f_{M_1} f_{M_2} \left[\frac{\bar{u}(k_1) k_3 \not{p} P_R v(k_2)}{m_{31}^2 - m_4^2 + im_4 \Gamma_4} + \frac{\bar{u}(k_1) \not{p} k_3 P_R v(k_2)}{m_{23}^2 - m_4^2 + im_4 \Gamma_4} \right]$$

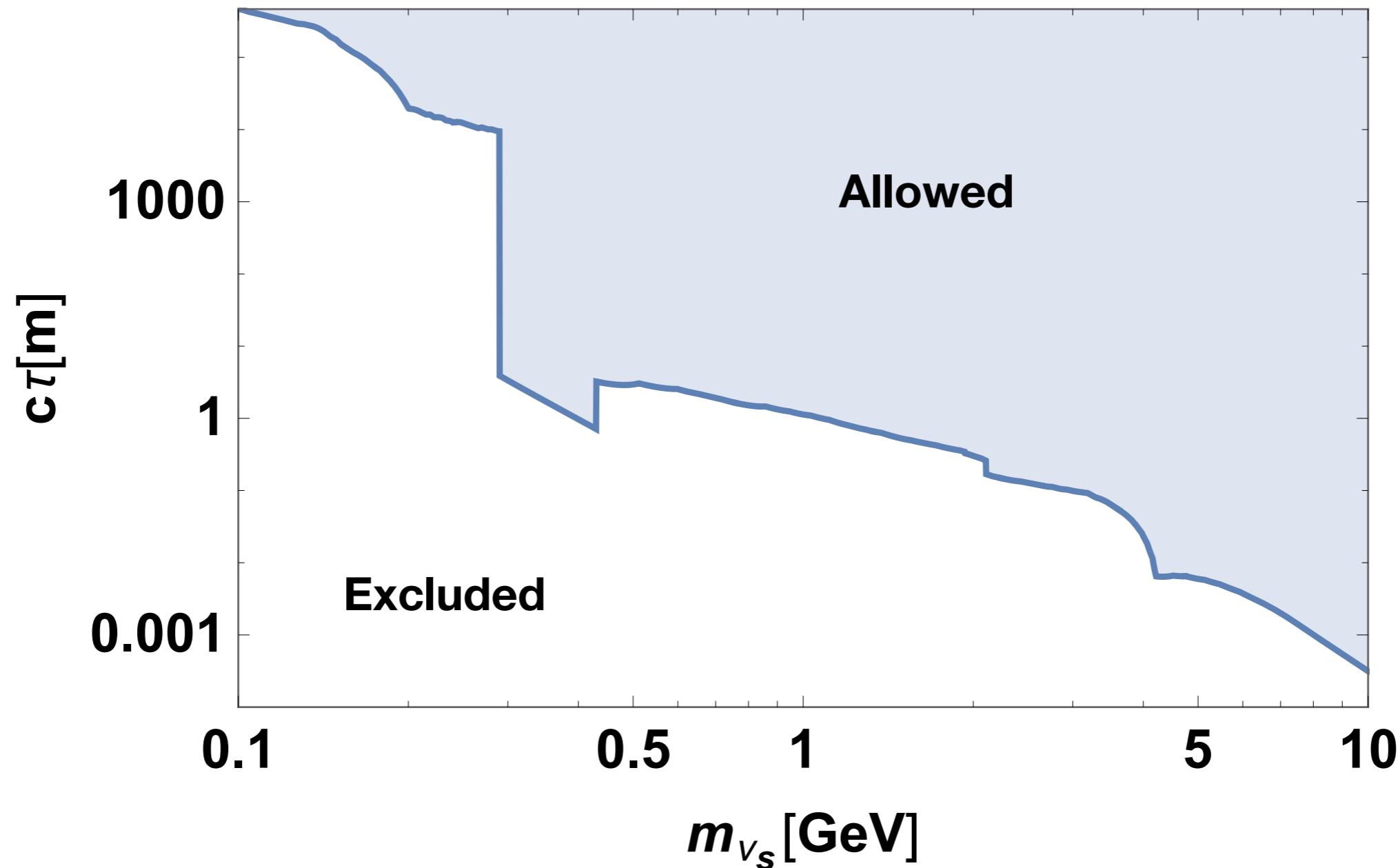
Negligible amplitude unless the intermediate state can go on-shell

$$\frac{1}{(m_{ij}^2 - m_4^2)^2 + m_4^2 \Gamma_4^2} \rightarrow \frac{\pi}{m_4 \Gamma_4} \delta(m_{ij}^2 - m_4^2)$$

Lifetime limitations

In the resonant regime $i\mathcal{M} \propto \frac{M_{\nu_s}}{\Gamma_{\nu_s}} \equiv M_{\nu_s} \tau_{\nu_s}$

But too long-lived heavy neutrinos decay outside the detector



Asking for observable (inside detector) decays imposes a further constraint

Current bounds

Tables (and list of references) from A. Abada, V. De Romeri, M.L., A. M. Teixeira and T. Toma, arXiv:1712.03984 [hep-ph]

Meson decay

LNV decay	Current bound		
	$\ell_\alpha = e, \ell_\beta = e$	$\ell_\alpha = e, \ell_\beta = \mu$	$\ell_\alpha = \mu, \ell_\beta = \mu$
$K^- \rightarrow \ell_\alpha^- \ell_\beta^- \pi^+$	6.4×10^{-10} [41]	5.0×10^{-10} [41]	1.1×10^{-9} [41]
$D^- \rightarrow \ell_\alpha^- \ell_\beta^- \pi^+$	1.1×10^{-6} [41]	2.0×10^{-6} [78]	2.2×10^{-8} [79]
$D^- \rightarrow \ell_\alpha^- \ell_\beta^- K^+$	9.0×10^{-7} [78]	1.9×10^{-6} [78]	1.0×10^{-5} [78]
$D^- \rightarrow \ell_\alpha^- \ell_\beta^- \rho^+$	—	—	5.6×10^{-4} [41]
$D^- \rightarrow \ell_\alpha^- \ell_\beta^- K^{*+}$	—	—	8.5×10^{-4} [41]
$D_s^- \rightarrow \ell_\alpha^- \ell_\beta^- \pi^+$	4.1×10^{-6} [41]	8.4×10^{-6} [78]	1.2×10^{-7} [79]
$D_s^- \rightarrow \ell_\alpha^- \ell_\beta^- K^+$	5.2×10^{-6} [78]	6.1×10^{-6} [78]	1.3×10^{-5} [78]
$D_s^- \rightarrow \ell_\alpha^- \ell_\beta^- K^{*+}$	—	—	1.4×10^{-3} [41]
$B^- \rightarrow \ell_\alpha^- \ell_\beta^- \pi^+$	2.3×10^{-8} [80]	1.5×10^{-7} [81]	4.0×10^{-9} [82]
$B^- \rightarrow \ell_\alpha^- \ell_\beta^- K^+$	3.0×10^{-8} [80]	1.6×10^{-7} [81]	4.1×10^{-8} [83]
$B^- \rightarrow \ell_\alpha^- \ell_\beta^- \rho^+$	1.7×10^{-7} [81]	4.7×10^{-7} [81]	4.2×10^{-7} [81]
$B^- \rightarrow \ell_\alpha^- \ell_\beta^- D^+$	2.6×10^{-6} [84]	1.8×10^{-6} [84]	6.9×10^{-7} [85]
$B^- \rightarrow \ell_\alpha^- \ell_\beta^- D^{*+}$	—	—	2.4×10^{-6} [41]
$B^- \rightarrow \ell_\alpha^- \ell_\beta^- D_s^+$	—	—	5.8×10^{-7} [41]
$B^- \rightarrow \ell_\alpha^- \ell_\beta^- K^{*+}$	4.0×10^{-7} [81]	3.0×10^{-7} [81]	5.9×10^{-7} [81]
LNV matrix m_ν	m_ν^{ee}	$m_\nu^{e\mu}$	$m_\nu^{\mu\mu}$

τ decay

LNV decay	Current bound	
	$\ell = e$	$\ell = \mu$
$\tau^- \rightarrow \ell^+ \pi^- \pi^-$	2.0×10^{-8}	3.9×10^{-8}
$\tau^- \rightarrow \ell^+ \pi^- K^-$	3.2×10^{-8}	4.8×10^{-8}
$\tau^- \rightarrow \ell^+ K^- K^-$	3.3×10^{-8}	4.7×10^{-8}
LNV matrix m_ν	$m_\nu^{e\tau}$	$m_\nu^{\mu\tau}$

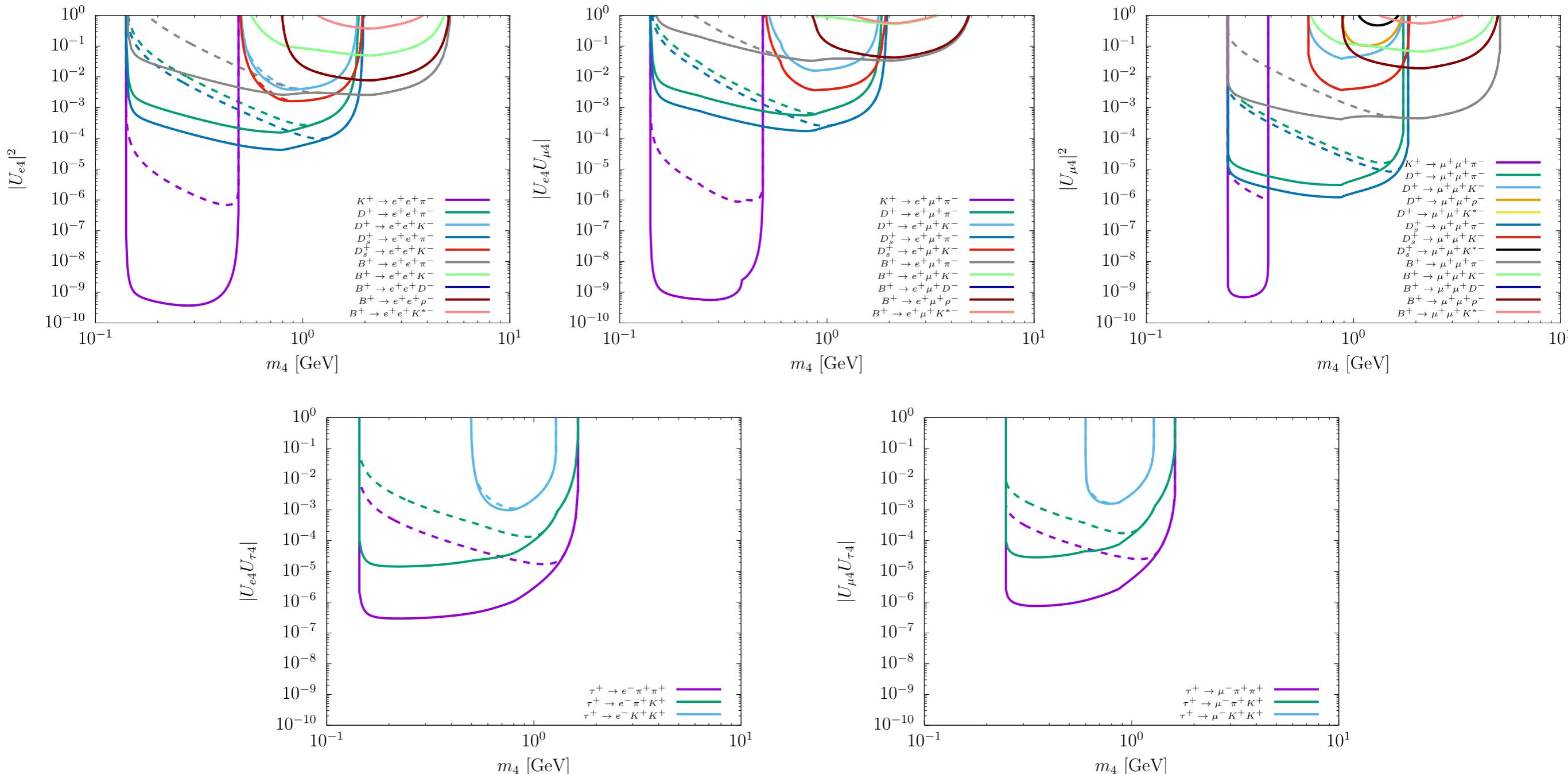
Results from

Belle [84],
BABAR [78,80,81] and
LHCb [79,82,83,85];
summarised in PDG [41]

upper bounds from the Belle

Constraints: single intermediate state

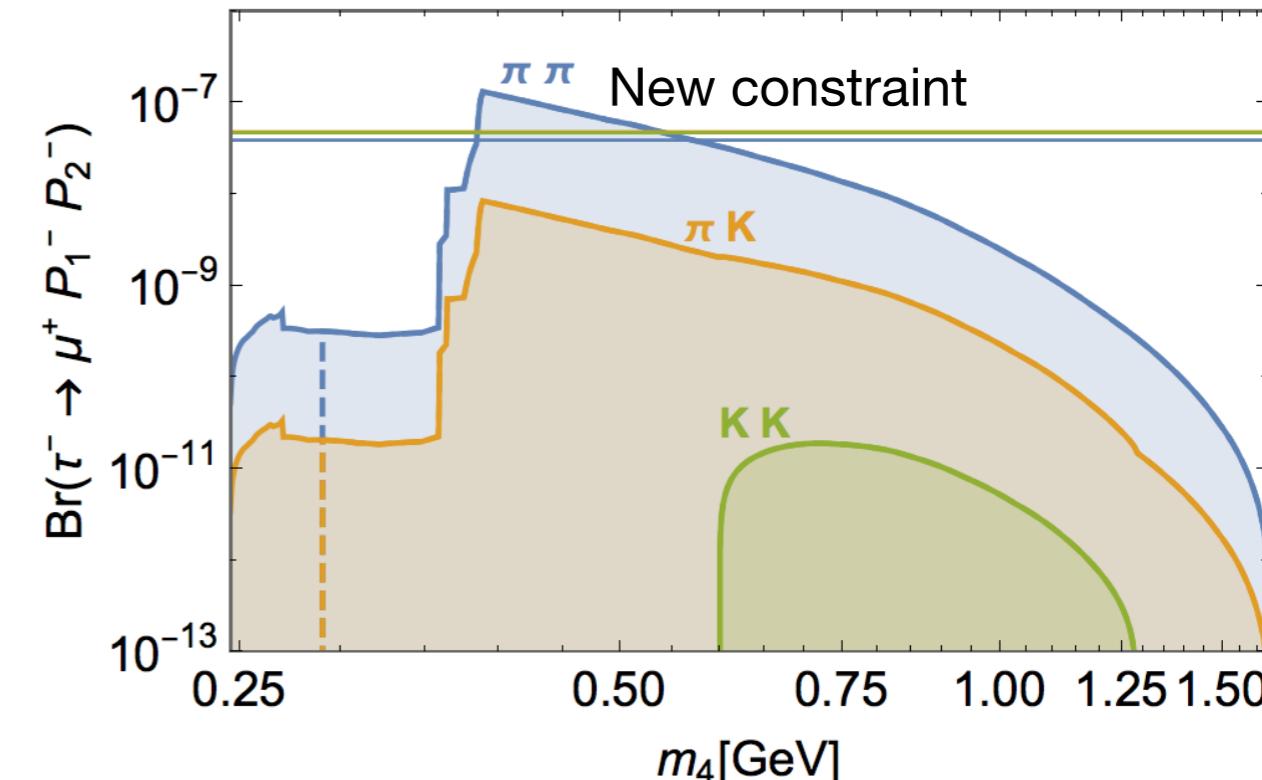
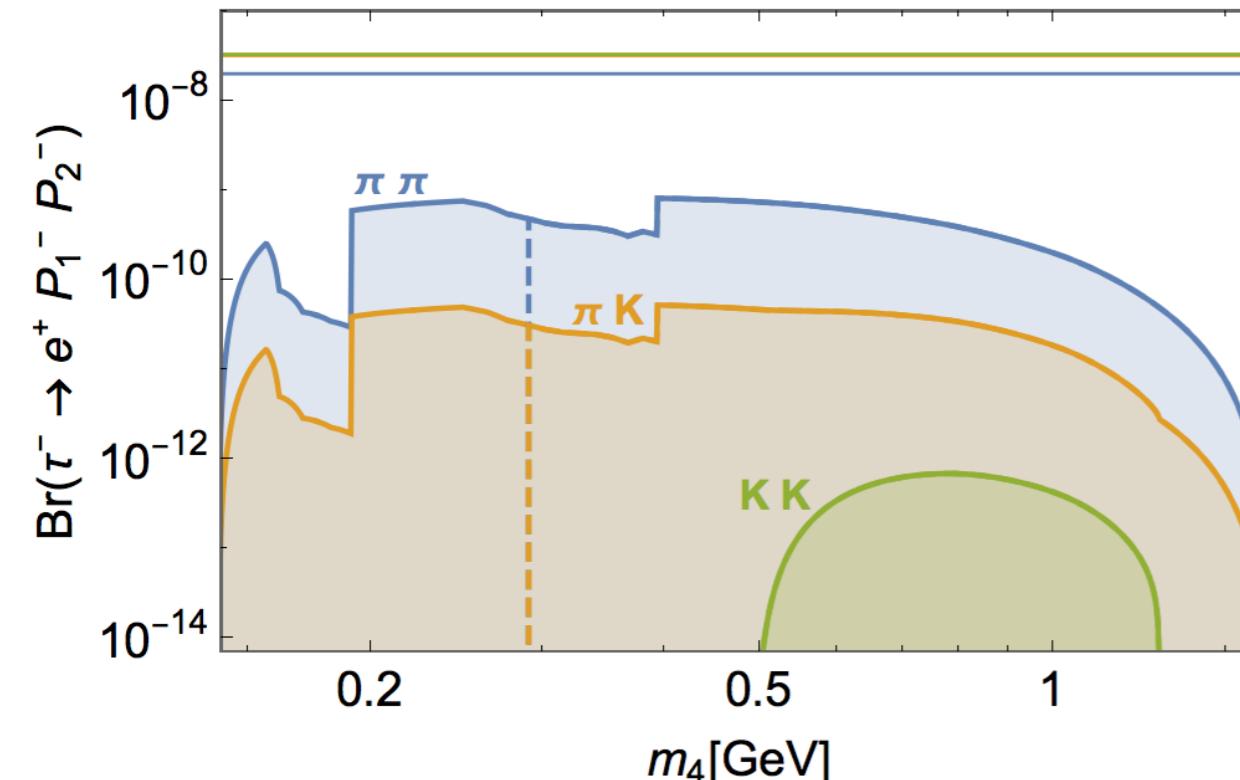
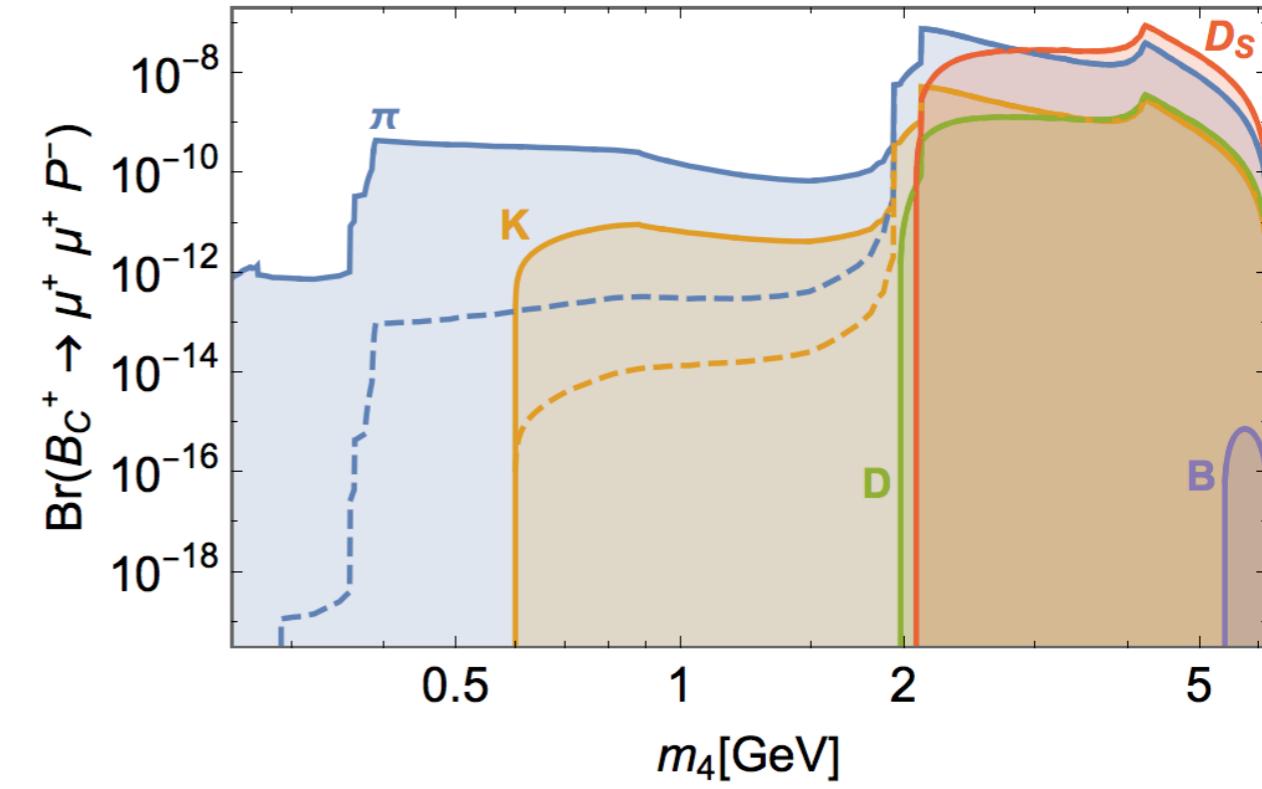
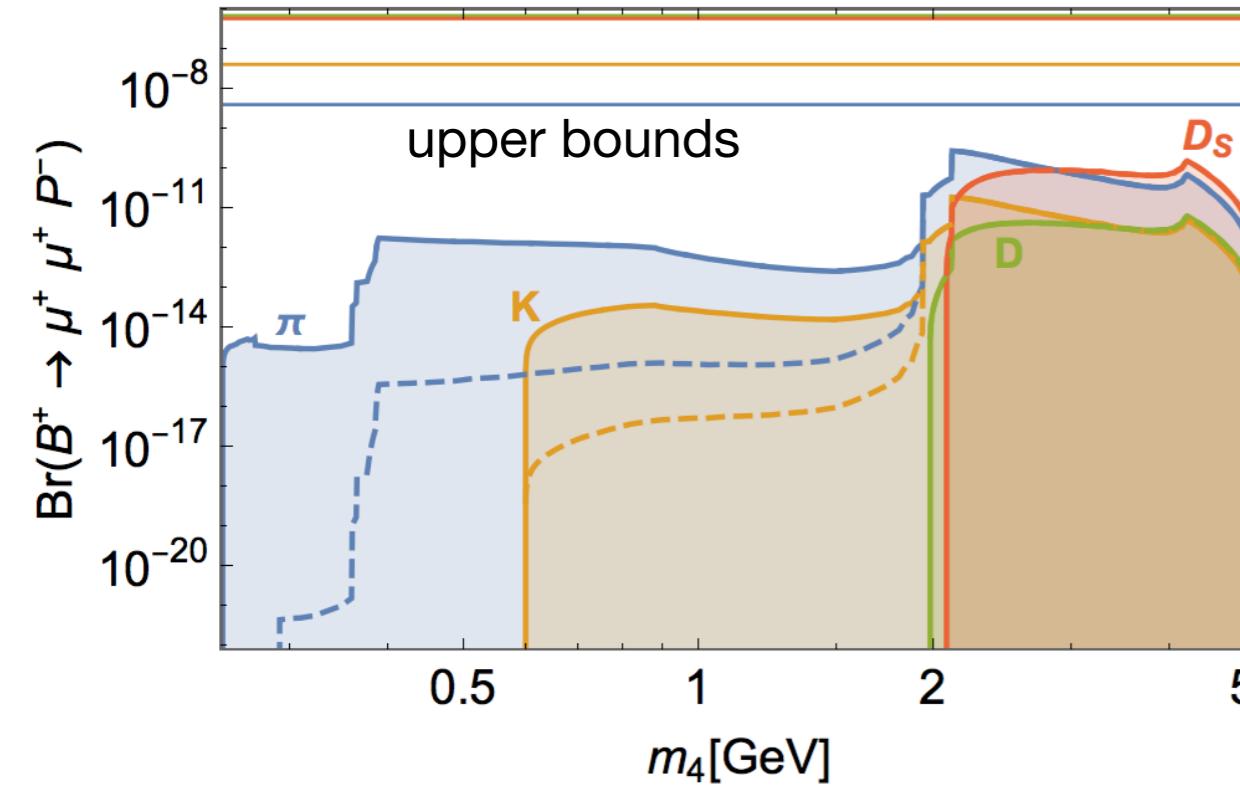
Figures from A. Abada, V. De Romeri, M.L., A. M. Teixeira and T. Toma, arXiv:1712.03984 [hep-ph];
 see also A. Atre, T. Han, S. Pascoli and B. Zhang, arXiv:0901.3589 [hep-ph]



Dashed lines: the on-shell heavy neutrino travels for less than 10 m

Some predictions: single intermediate state

Comprehensive analysis for τ and pseudo-scalar mesons in 1712.03984
(all possible initial and 3-body final states)



Multiple intermediate states: interference

A. Abada, C. Hati, X. Marcano and A. M. Teixeira, arXiv:1904.05367 [hep-ph]

If more than one heavy neutrino mediate the process, and

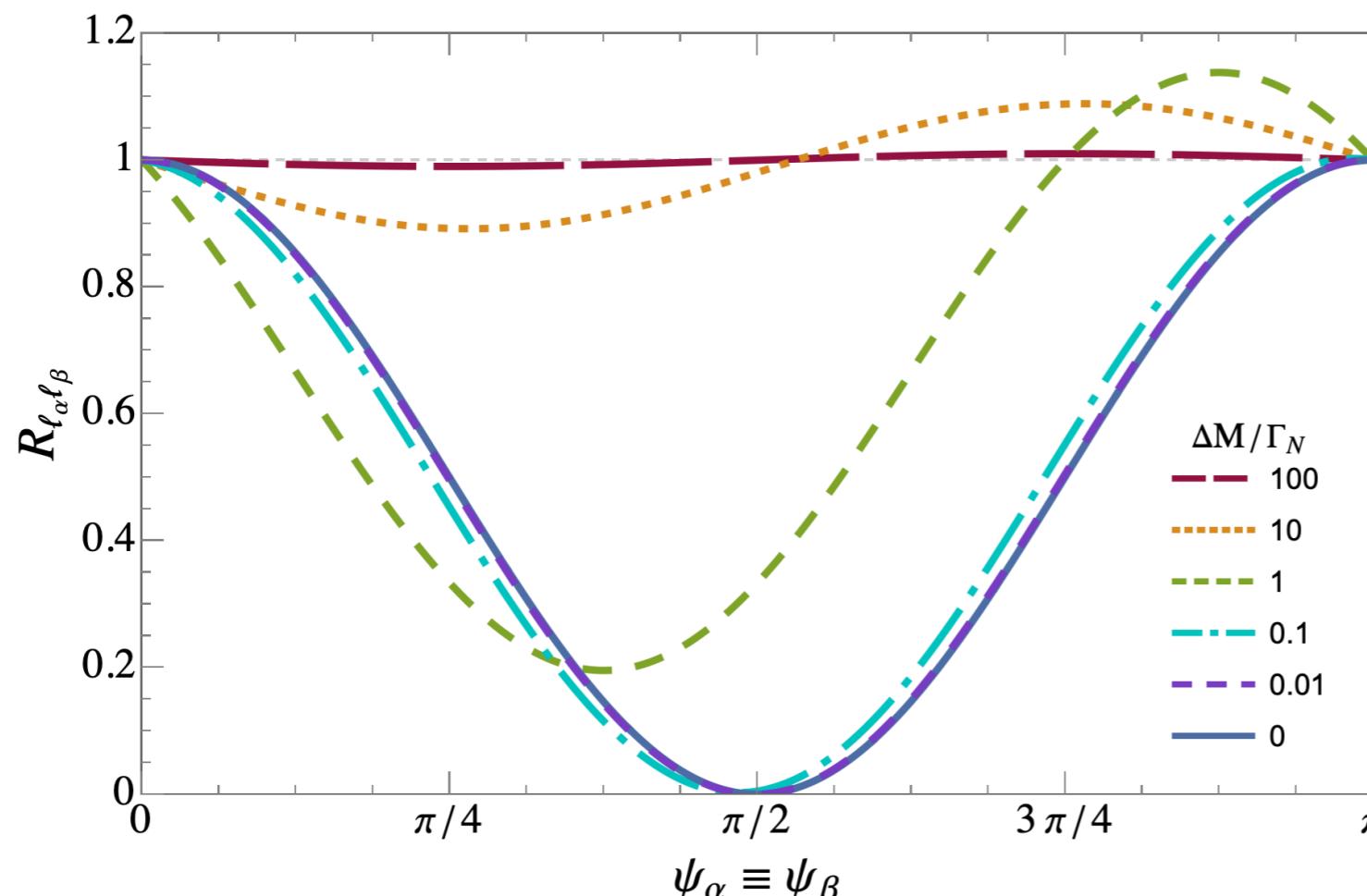
$$\Delta M \ll M \quad \text{and} \quad \Delta M < \Gamma_N$$

interference effects arise due to the CP-violating phases

$$\left| \mathcal{A}_{M \rightarrow M' \ell_\alpha^+ \ell_\beta^-}^{\text{LNC}} \right|^2 \propto \left| U_{\alpha 4} U_{\beta 4}^* g(m_4) + U_{\alpha 5} U_{\beta 5}^* g(m_5) \right|^2,$$

$$\left| \mathcal{A}_{M \rightarrow M' \ell_\alpha^+ \ell_\beta^+}^{\text{LNV}} \right|^2 \propto \left| U_{\alpha 4} U_{\beta 4} f(m_4) + U_{\alpha 5} U_{\beta 5} f(m_5) \right|^2,$$

$$R_{\ell_\alpha \ell_\beta} \equiv \frac{\Gamma_{M \rightarrow M' \ell_\alpha^\pm \ell_\beta^\pm}^{\text{LNV}}}{\Gamma_{M \rightarrow M' \ell_\alpha^\pm \ell_\beta^\mp}^{\text{LNC}}}$$



$$U_{\alpha i} = e^{-i\phi_{\alpha i}} |U_{\alpha i}|$$

$$\psi_\alpha \equiv \phi_{\alpha 5} - \phi_{\alpha 4}$$

Dirac limit

$\frac{\Delta M}{\Gamma_N}$	=	0
ψ_α	=	$\frac{\pi}{2}$

LHC SEARCHES

LNV at LHC

Heavy neutrinos in pp collisions produced through a variety of mechanisms

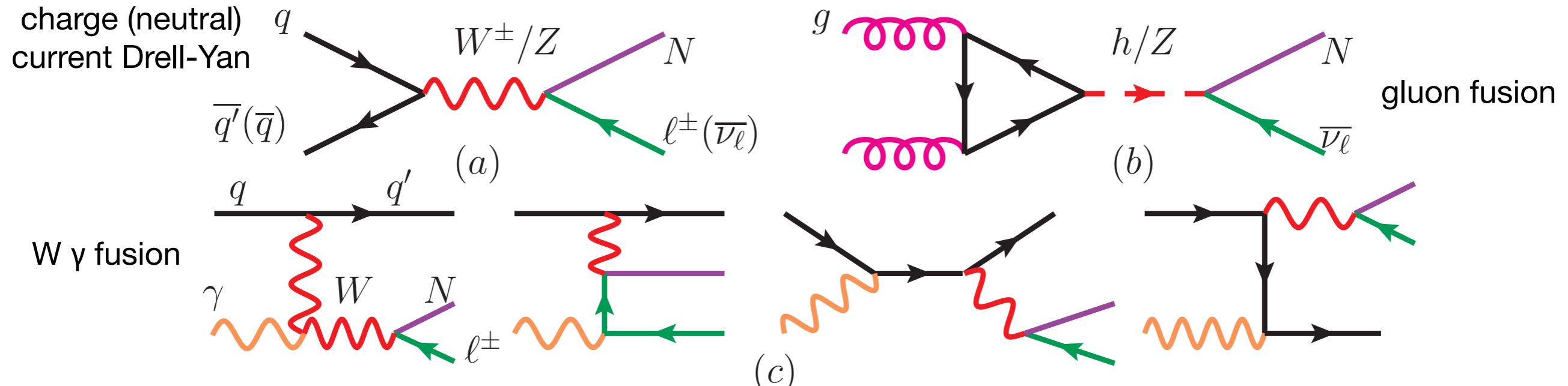


Figure from C. Degrande, O. Mattelaer, R. Ruiz and J. Turner, arXiv:1602.06957 [hep-ph];
see also Y. Cai, T. Han, T. Li and R. Ruiz, arXiv:1711.02180 [hep-ph]

LNV can manifest with clean experimental signatures:
e.g. two same-sign leptons (any flavour combination of e and μ) and at least one jet

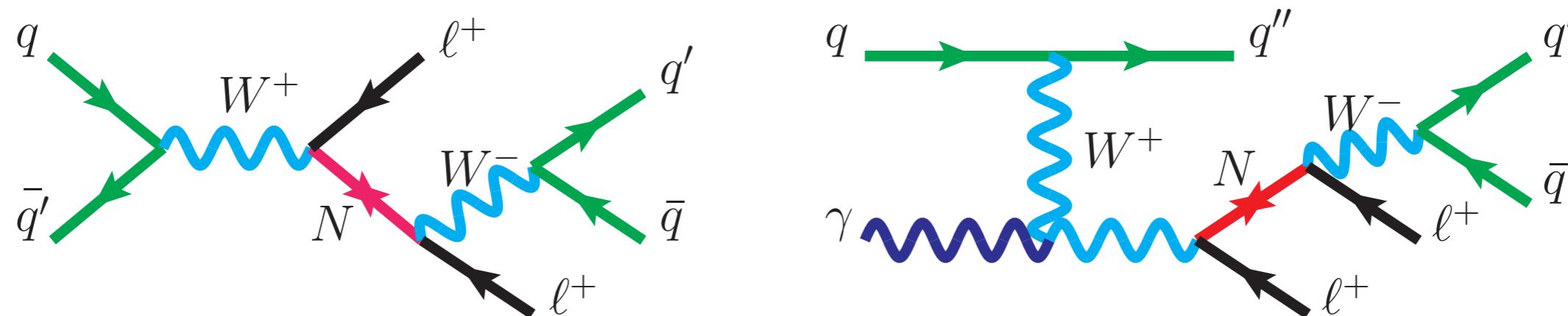
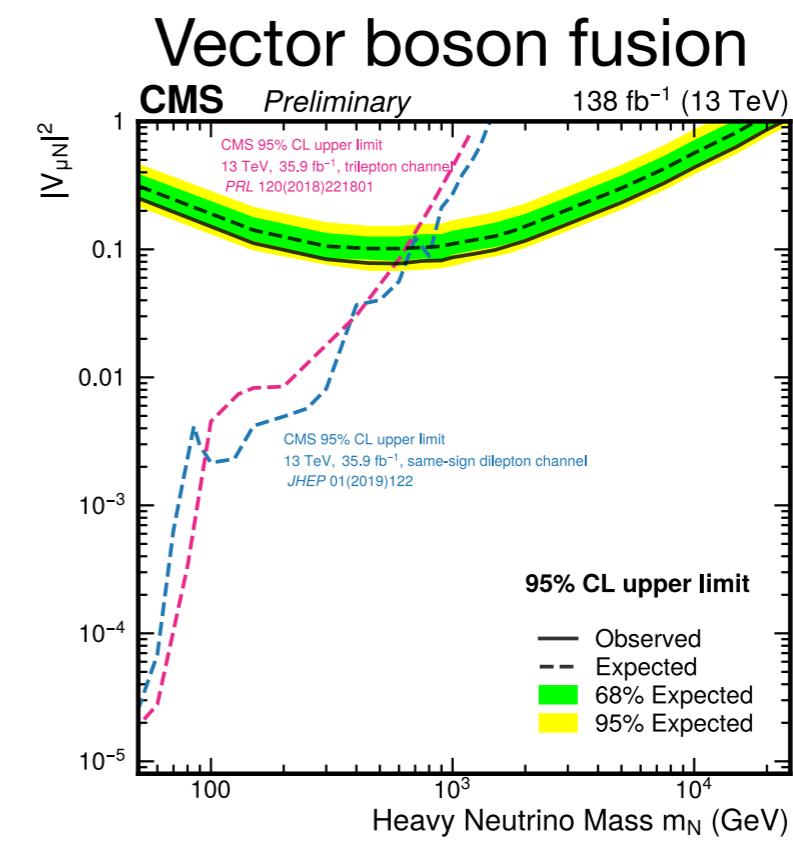
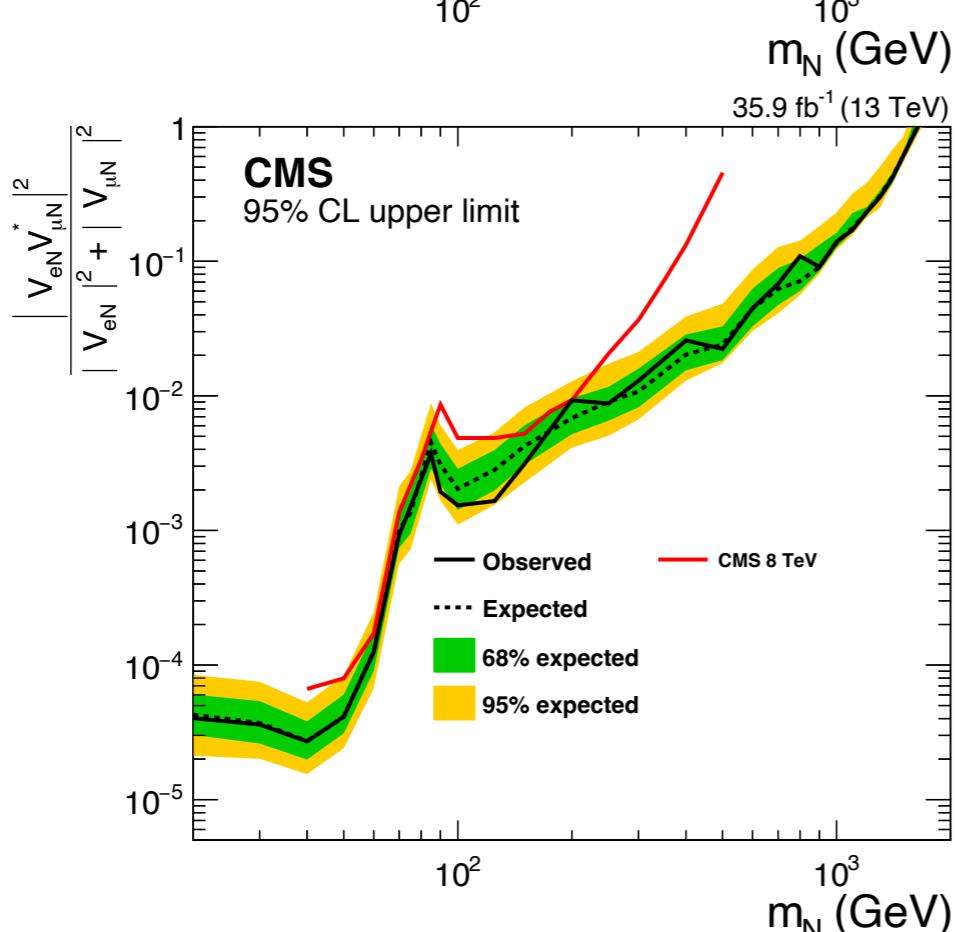
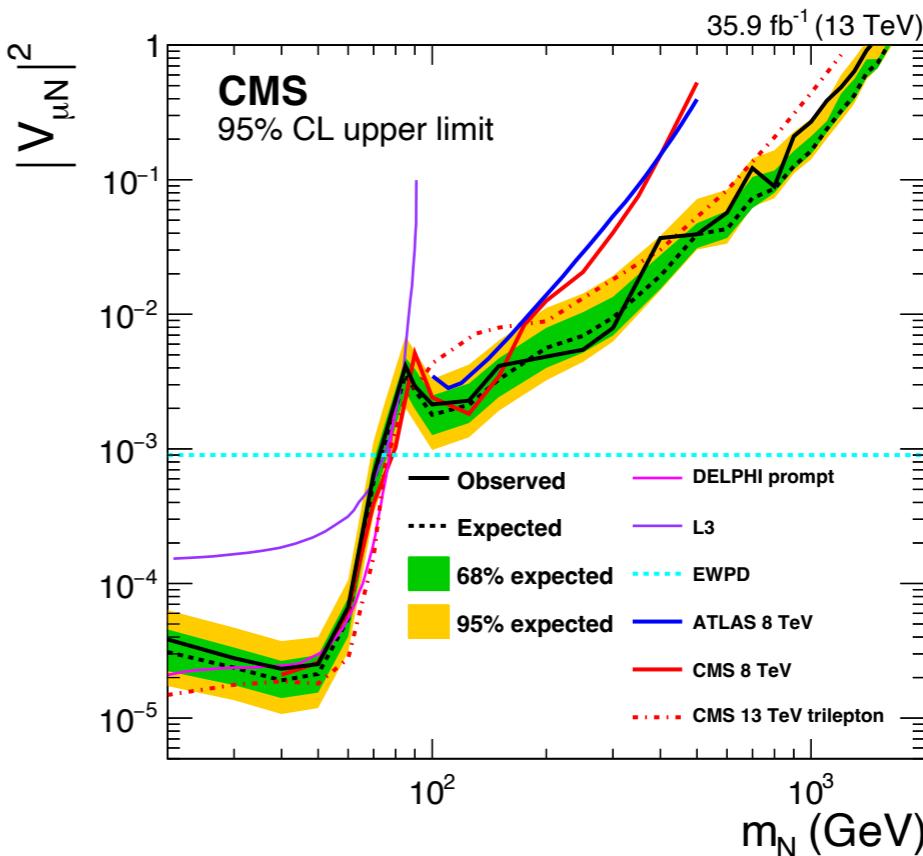
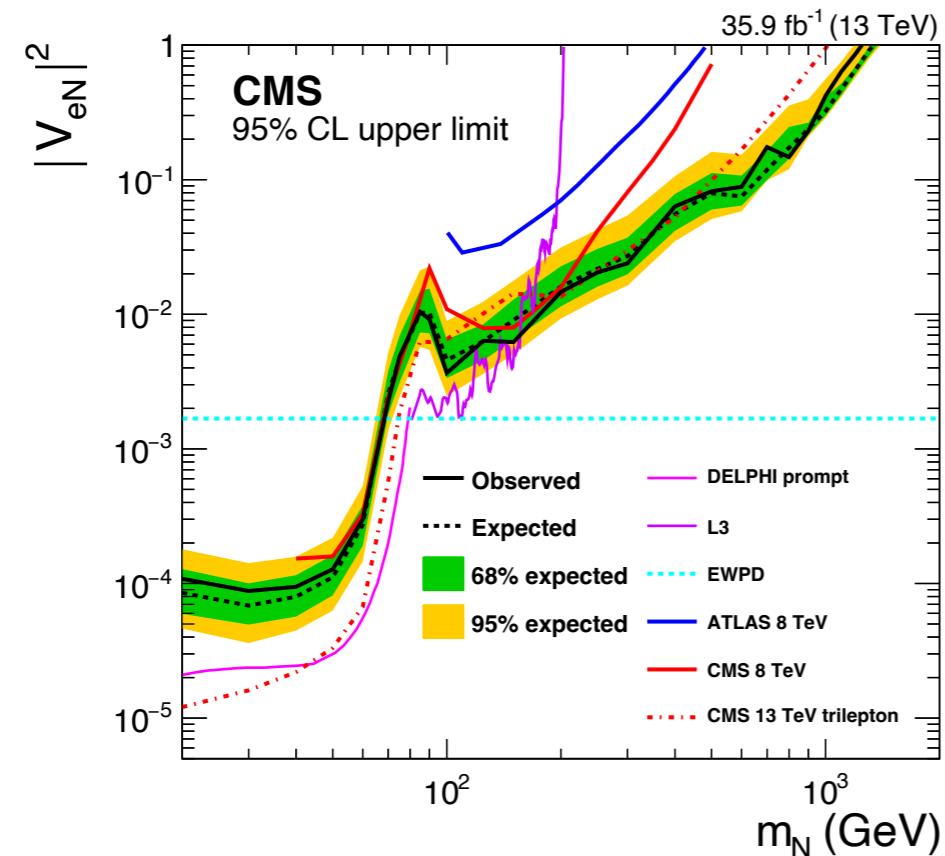


Figure from CMS Collaboration, arXiv:1806.10905 [hep-ex]

Current bounds: single mediator

CMS Collaboration, arXiv:1806.10905 [hep-ex]; see also ATLAS Collaboration, arXiv:1506.06020 [hep-ex]



LNV/LNC oscillations

Y. Nir, Conf. Proc. C9207131, 81 (1992); G. Anamiati, M. Hirsch and E. Nardi, arXiv:1607.05641 [hep-ph]

Flavour eigenstate = coherent superposition of mass eigenstates

$$\left\{ \begin{array}{l} N_\ell = \frac{1}{\sqrt{2}}(N_+ - iN_-) \\ N_{\bar{\ell}} = \frac{1}{\sqrt{2}}(N_+ + iN_-) \end{array} \right. \xrightarrow{\text{evolution}} \left\{ \begin{array}{l} N_\ell(t) = g_+(t)N_\ell + g_-(t)N_{\bar{\ell}} \\ N_{\bar{\ell}}(t) = g_-(t)N_\ell + g_+(t)N_{\bar{\ell}} \end{array} \right.$$

$$g_+(t) = e^{-iMt}e^{-\frac{\Gamma}{2}t} \cos\left(\frac{\Delta M}{2}t\right)$$

$$g_-(t) = i e^{-iMt}e^{-\frac{\Gamma}{2}t} \sin\left(\frac{\Delta M}{2}t\right)$$

$$\Delta M = M^+ - M^-$$

$\Delta M \gg \Gamma$ **decay after decoherence (Majorana limit)**

$\Delta M \approx \Gamma$ **oscillations**

$\Delta M \ll \Gamma$ **oscillations do not develop (Dirac limit)**

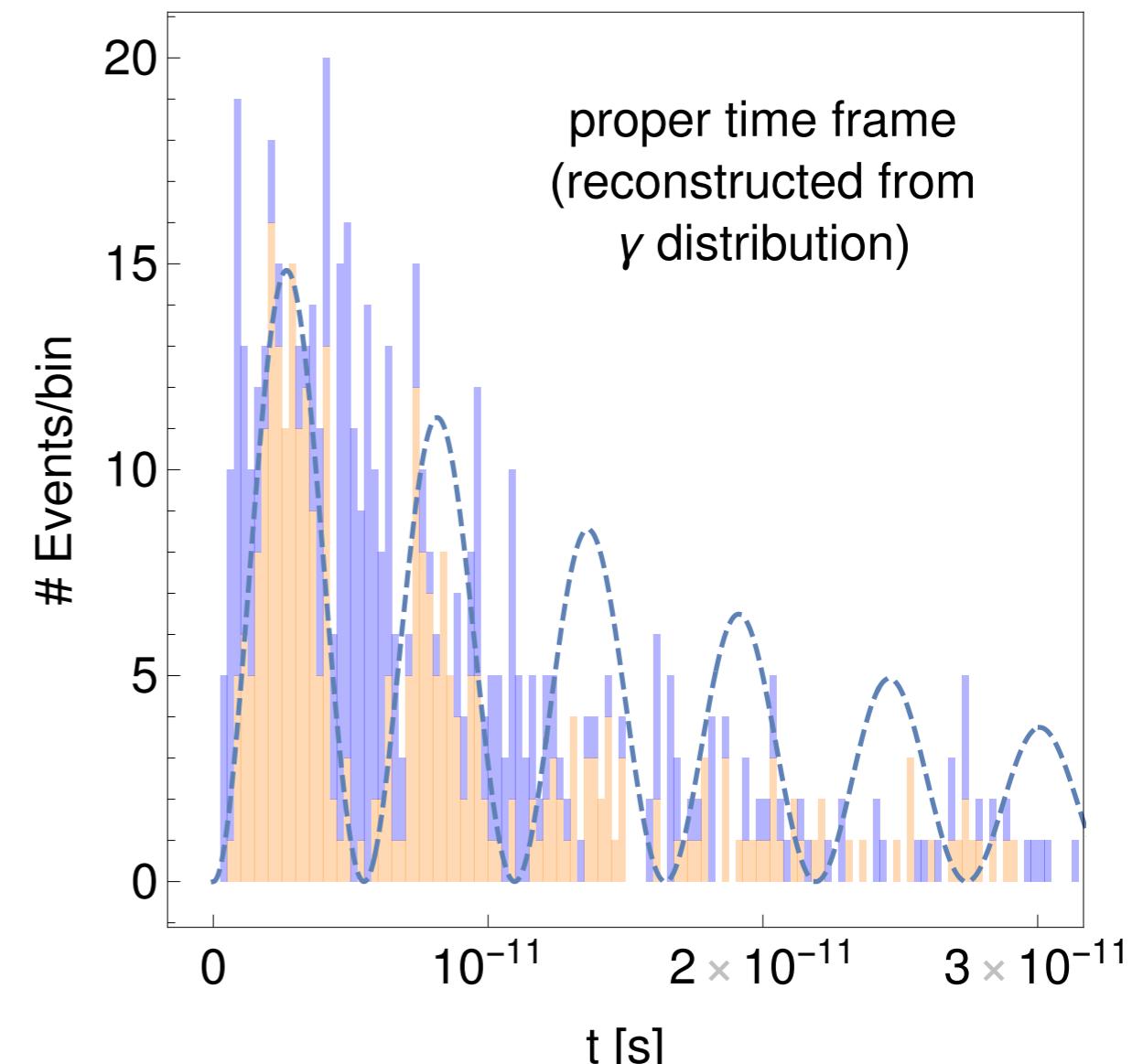
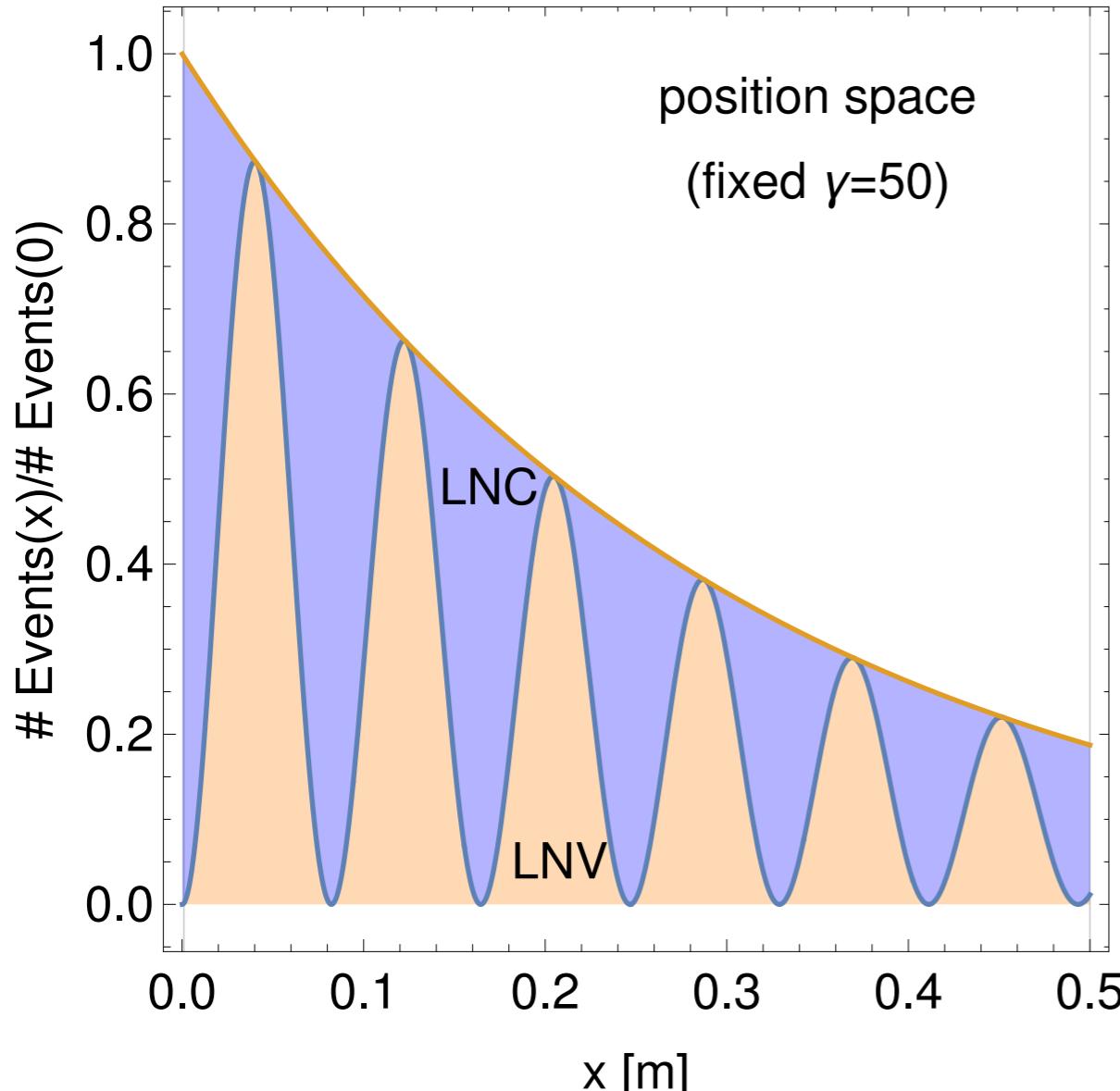
$$R_{\ell\ell}(t_1, t_2) = \frac{\int_{t_1}^{t_2} |g_-(t)|^2 dt}{\int_{t_1}^{t_2} |g_+(t)|^2 dt} = \frac{\#(\ell^+\ell^+) + \#(\ell^-\ell^-)}{\#(\ell^+\ell^-)}$$

$$R_{ll}(0, \infty) = \frac{\Delta M^2}{2\Gamma^2 + \Delta M^2}$$

Are these oscillations observable?

S. Antusch, E. Cazzato and O. Fischer, arXiv:1709.03797 [hep-ph]

E.g. LHCb experiment for
Linear Seesaw with $M = 7 \text{ GeV}$, $U^2 = 10^{-5}$, Inverted Ordering



However, for heavy neutrinos with $\gamma=50$

- very forward rapidity
- very small track separation of decay products

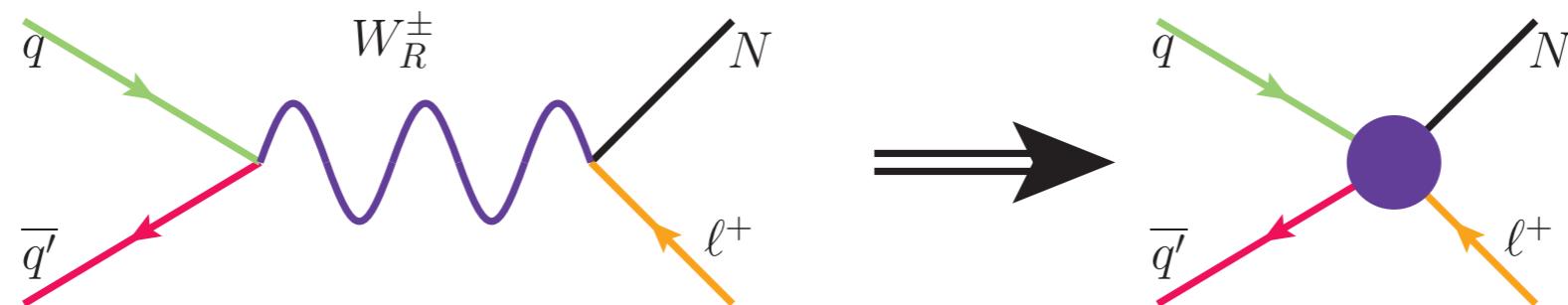
Richard Ruiz, private communication

Why to look for LNV if $m_v \approx 0$?

Equivalence between L conservation and massless neutrinos only holds in SM + singlet fermions

E.g. Left-right symmetric model

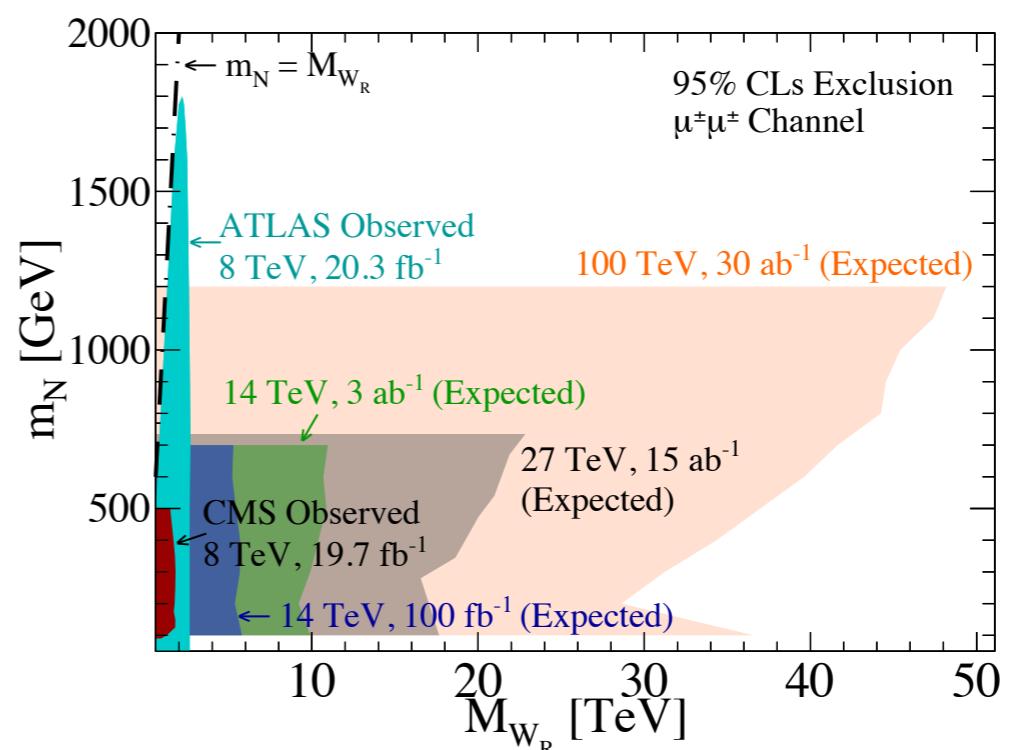
If new gauge mediators are too heavy, light N are still accessible



Courtesy of
Richard Ruiz

When $M_{W_R} \gg \sqrt{\hat{s}}$ but $m_N \lesssim \mathcal{O}(1)$ TeV, $pp \rightarrow N\ell + X$ production in the LRSM and minimal Type I Seesaw are not discernible¹¹

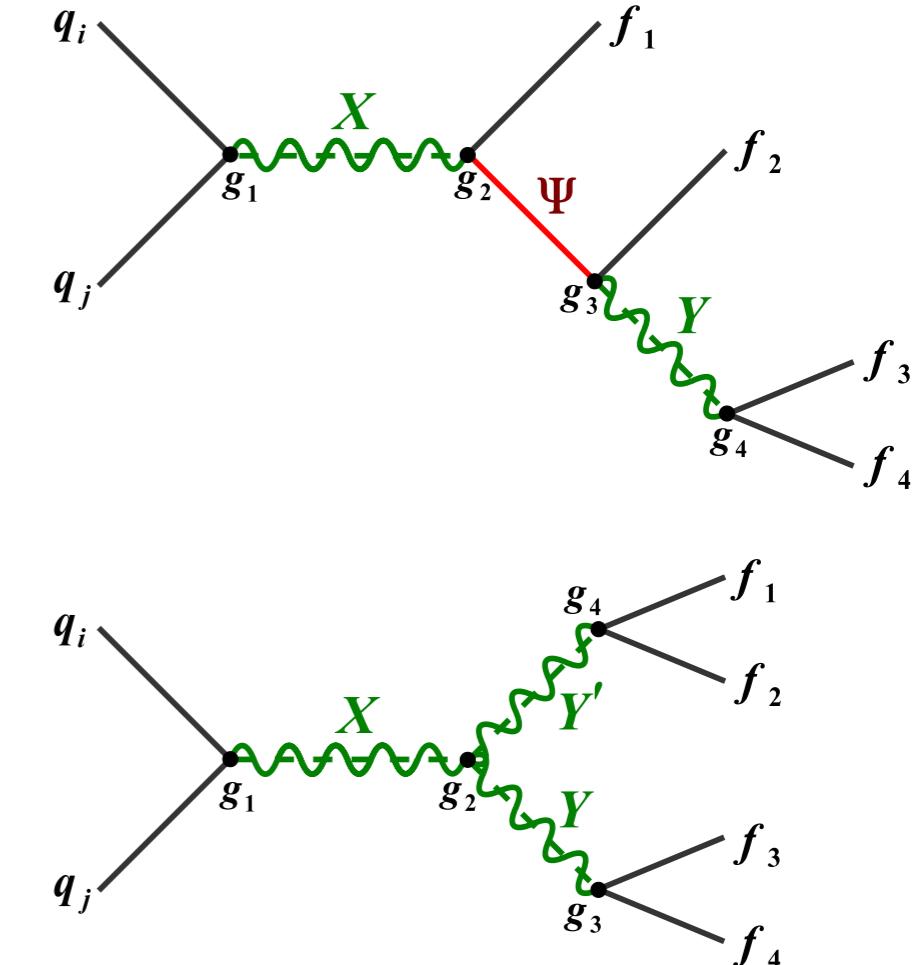
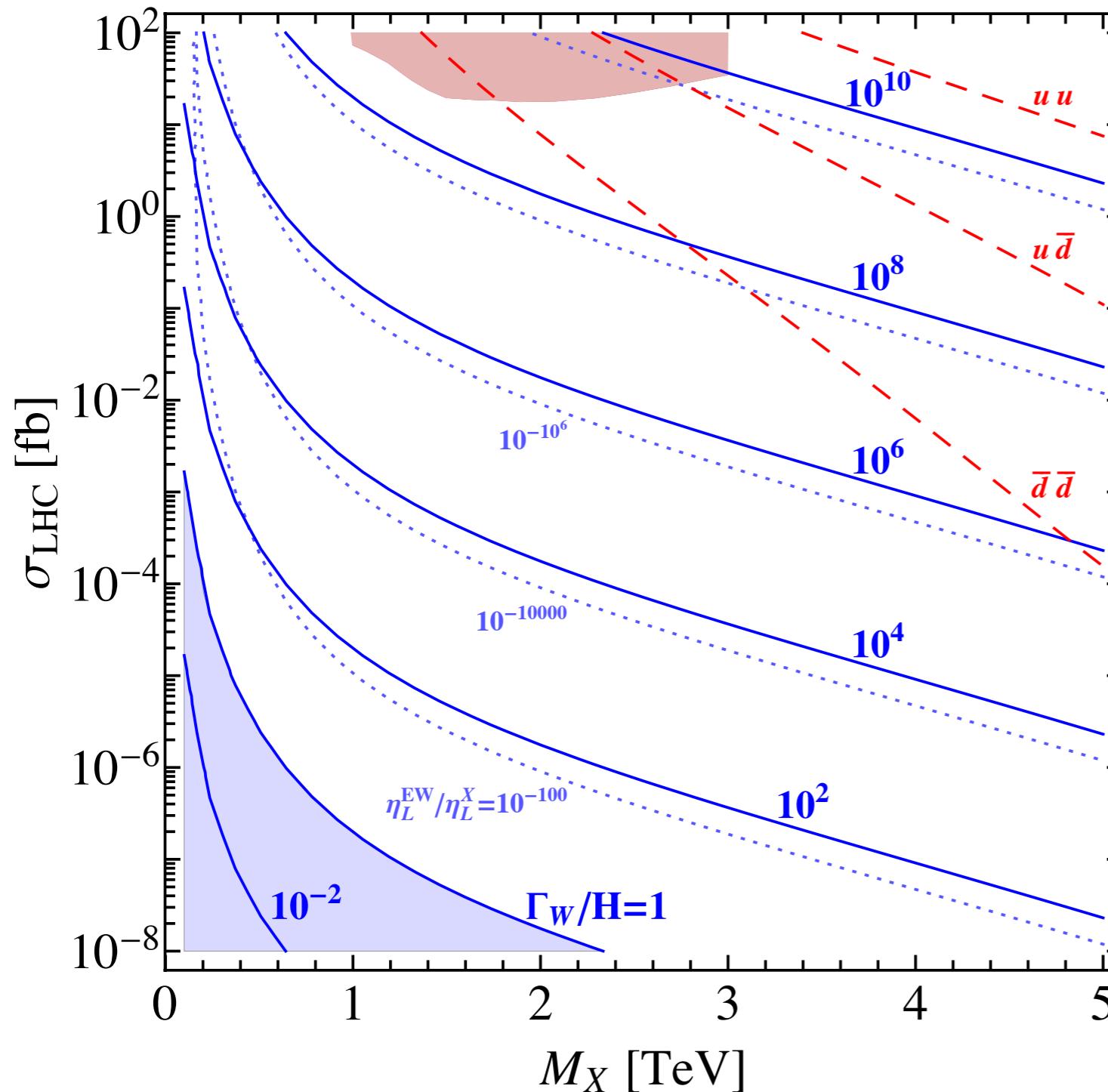
- **Signature:** $pp \rightarrow \ell^\pm \ell^\pm + nj + X + p_T^\ell \gtrsim \mathcal{O}(m_N)$ + no MET
 - At 14 (100) TeV with $\mathcal{L} = 1$ (10) ab^{-1} , $M_{W_R} \lesssim 9$ (40) TeV probed
 - **DO NOT STOP SEARCHING FOR TYPE I LNV**



¹¹Han, Lewis, RR, Si, [1211.6447]; RR, [1703.04669]

Falsify high-scale leptogenesis with LNV

J. M. Frere, T. Hambye and G. Vertongen, arXiv:0806.0841 [hep-ph];
 F. F. Deppisch, J. Harz and M. Hirsch, arXiv:1312.4447 [hep-ph]



$$\frac{\Gamma_W}{H} = \frac{0.028}{\sqrt{g_*}} \frac{M_P M_X^3}{T^4} \frac{K_1(M_X/T)}{f_{q_1 q_2} (M_X/\sqrt{s})} \times (s \sigma_{\text{LHC}})$$

A LNV observation at LHC likely falsifies high-scale leptogenesis

EARLY UNIVERSE

Low Scale Leptogenesis with 2 RHN

The approximate L conservation forces the **HNL** to be **degenerate in mass**

$$M_0 = \begin{pmatrix} 0 & \frac{1}{\sqrt{2}}Yv & 0 \\ \frac{1}{\sqrt{2}}Yv & 0 & \Lambda \\ 0 & \Lambda & 0 \end{pmatrix} \quad \rightarrow \quad m_\nu = 0, \quad M_{1,2} = \sqrt{|\Lambda|^2 + \frac{1}{2}|Yv|^2}$$

$$\Delta M_{ISS} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & \xi \Lambda \end{pmatrix} \quad \Delta M_{LSS} = \begin{pmatrix} 0 & 0 & \frac{\epsilon}{\sqrt{2}} Y'v \\ 0 & 0 & 0 \\ \frac{\epsilon}{\sqrt{2}} Y'v & 0 & 0 \end{pmatrix} \quad \rightarrow \quad m_\nu \simeq 2\epsilon \frac{m_D^2}{M_{1,2}}, \quad \Delta m^2 \simeq 2\xi M_{1,2}^2$$

This allows to account for BAU via freeze-in leptogenesis with low-scale NHL

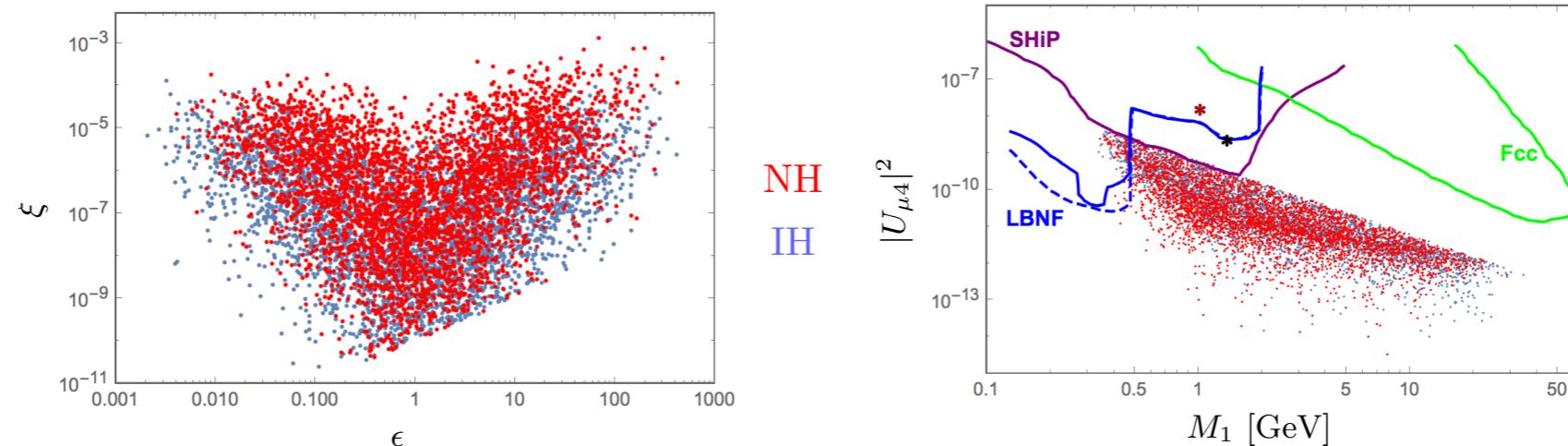
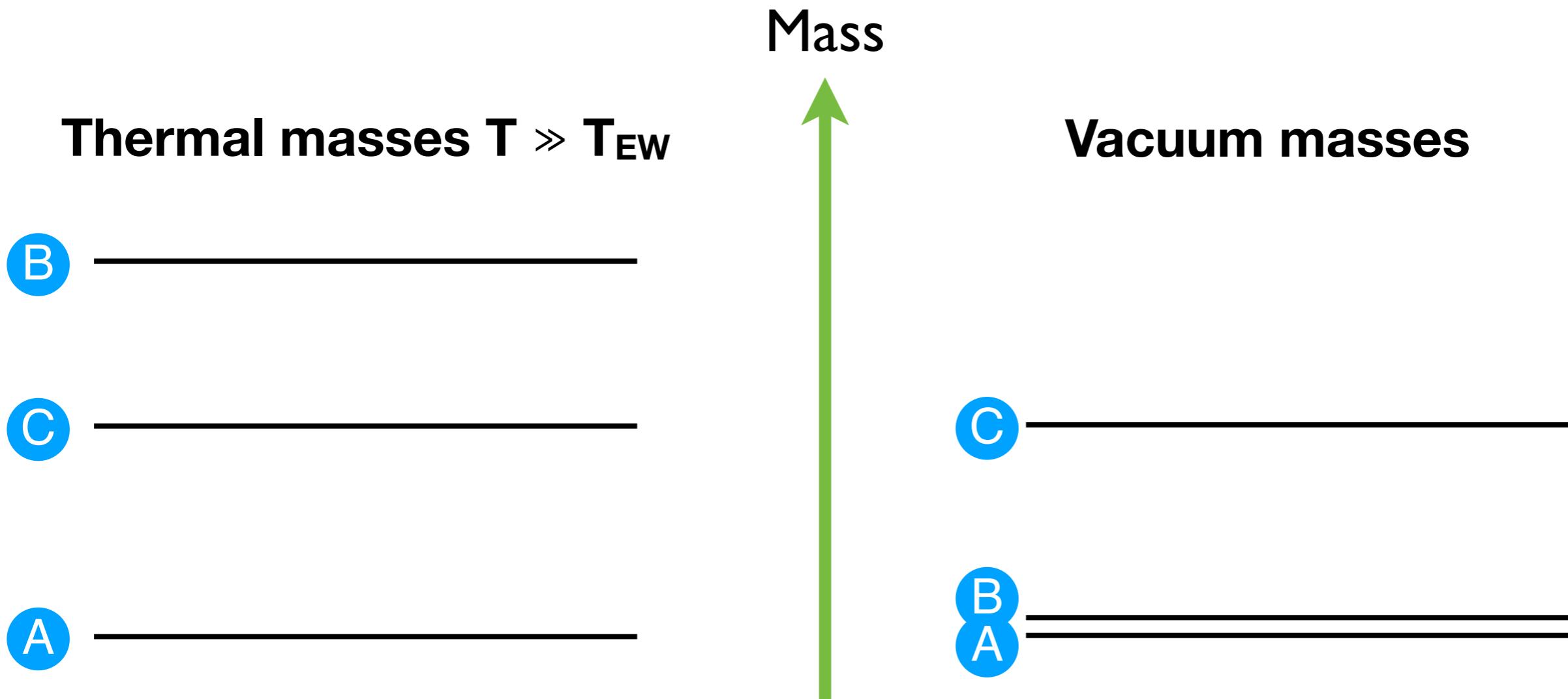


Figure from A. Abada, G. Arcadi, V. Domcke and M. Lucente, arXiv:1507.06215 [hep-ph]

Many studies, e.g. E. K. Akhmedov, V. A. Rubakov and A. Y. Smirnov, hep-ph/9803255; T. Asaka and M. Shaposhnikov, hep-ph/0505013; M. Shaposhnikov, arXiv:0804.4542 [hep-ph]; T. Asaka and H. Ishida, arXiv:1004.5491 [hep-ph]; T. Asaka, S. Eijima and H. Ishida, arXiv:1112.5565 [hep-ph]; L. Canetti, M. Drewes and M. Shaposhnikov, arXiv:1204.4186 [hep-ph]; L. Canetti, M. Drewes, T. Frossard and M. Shaposhnikov, arXiv:1208.4607 [hep-ph]; P. Hernández, M. Kekic, J. López-Pavón, J. Racker and N. Rius, arXiv:1508.03676 [hep-ph]...

Low Scale Leptogenesis with 3 RHN

Mass spectrum with 3 right-handed neutrinos and B - L symmetry



If the vacuum mass of the decoupled state is heavier than the pseudo-Dirac one, there is **necessarily** a level crossing at some finite temperature

A. Abada, G. Arcadi, V. Domcke, M. Drewes, J. Klaric and M. Lucente, arXiv:1810.12463 [hep-ph]

Level crossing: resonant asymmetry production

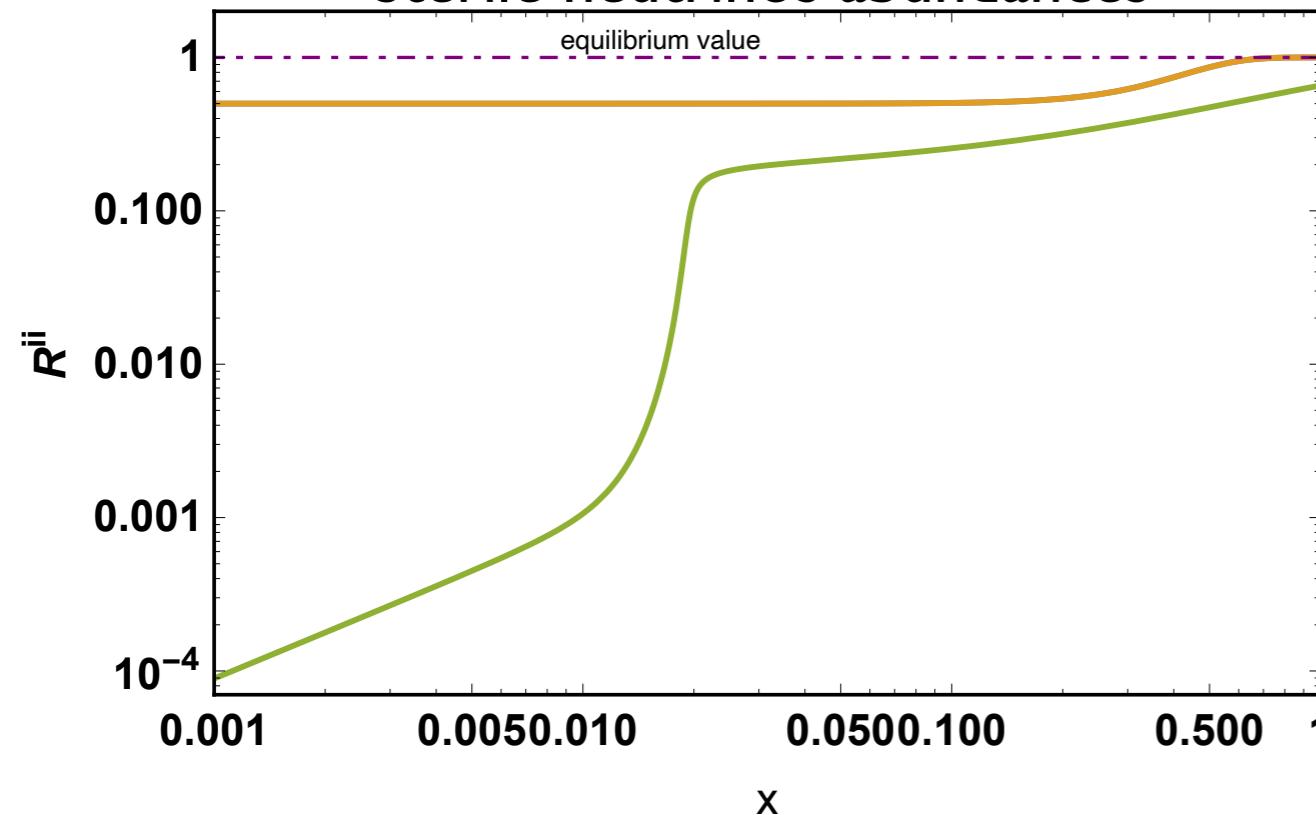
$$x = \frac{T}{T_{EW}}$$

$T_{EW} = 140$ GeV

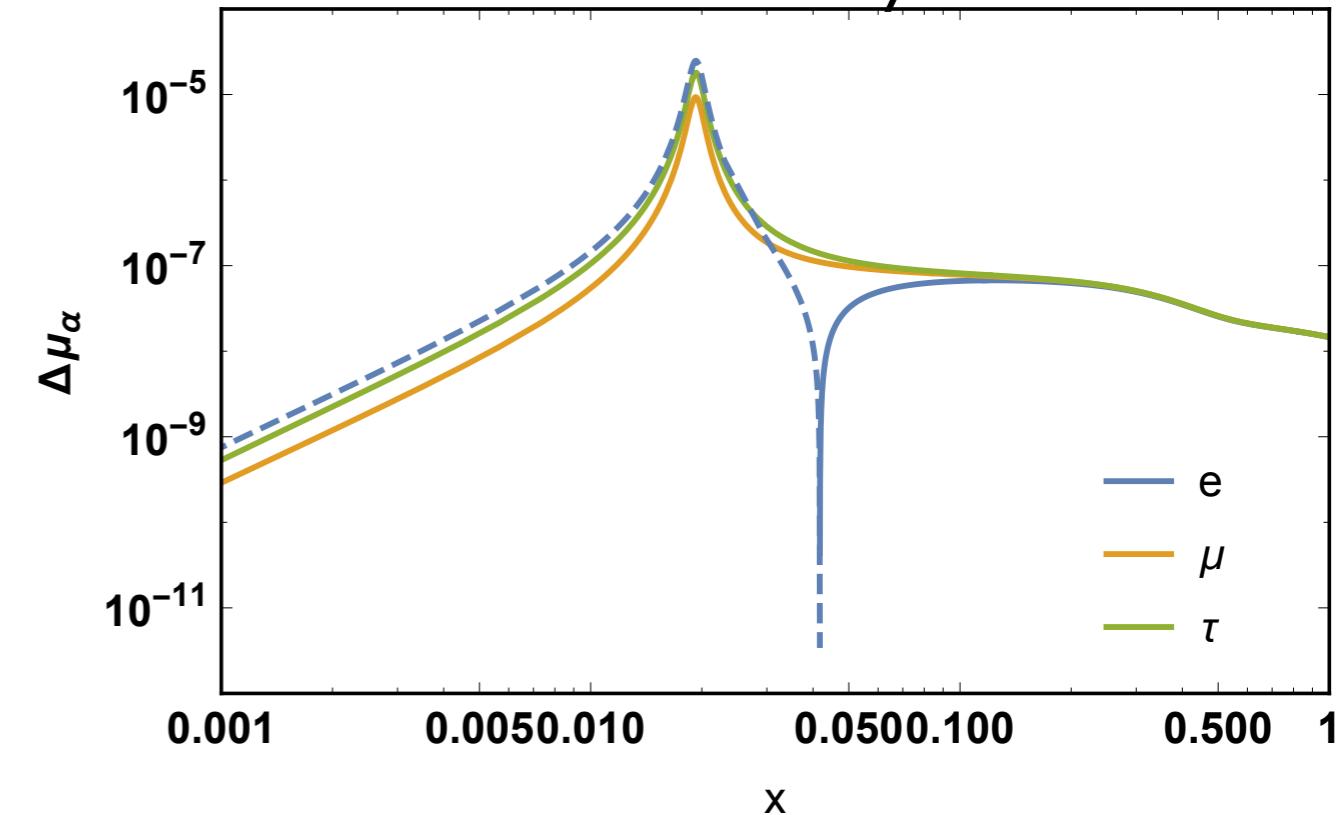
R : sterile neutrinos density matrix

μ_α : active flavours chemical potentials

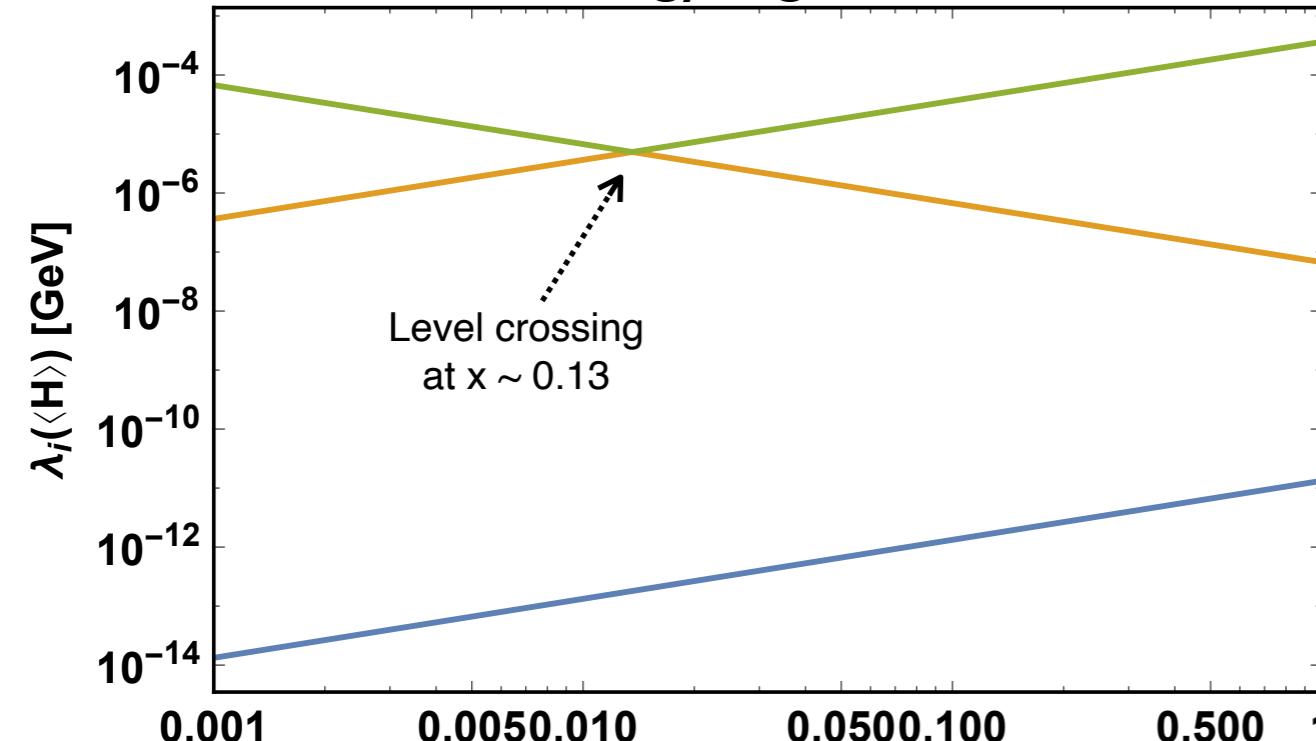
Sterile neutrinos abundances



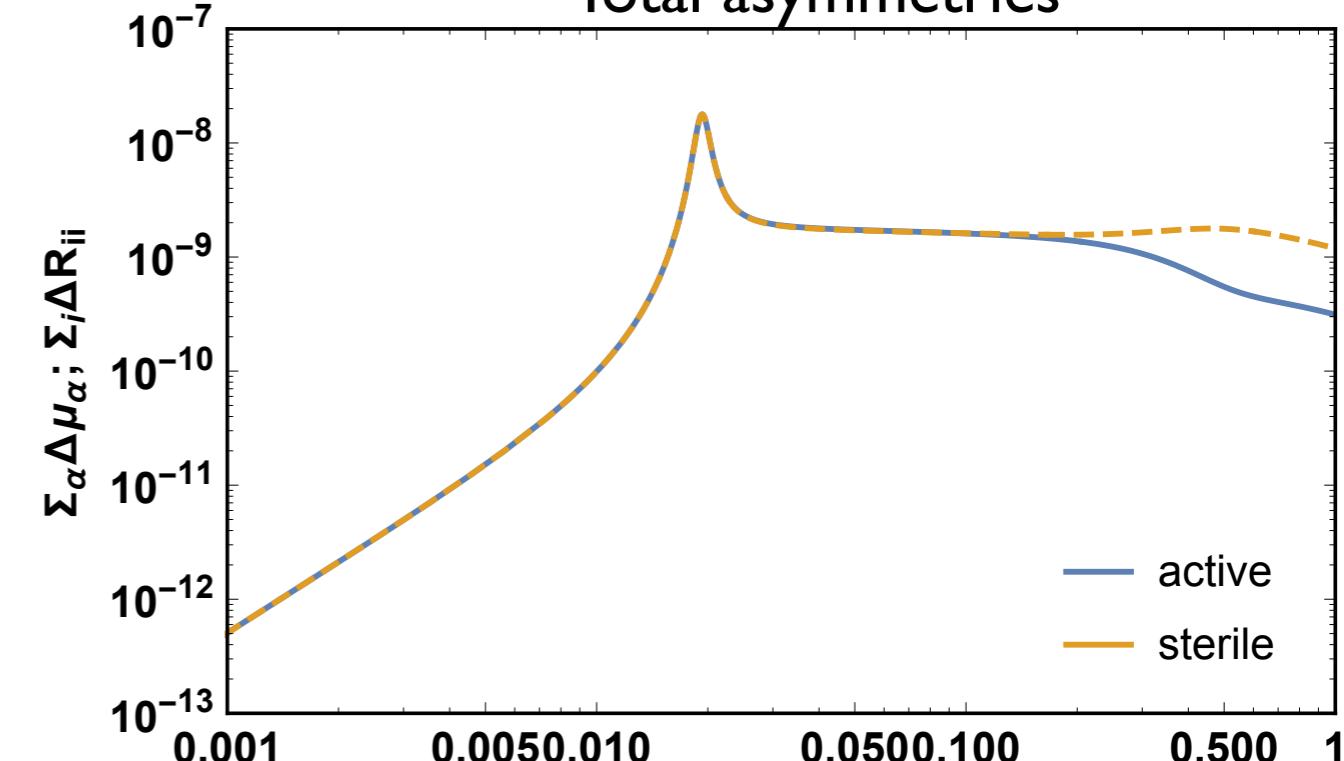
Active flavours asymmetries



Energy eigenvalues



Total asymmetries



Conclusion

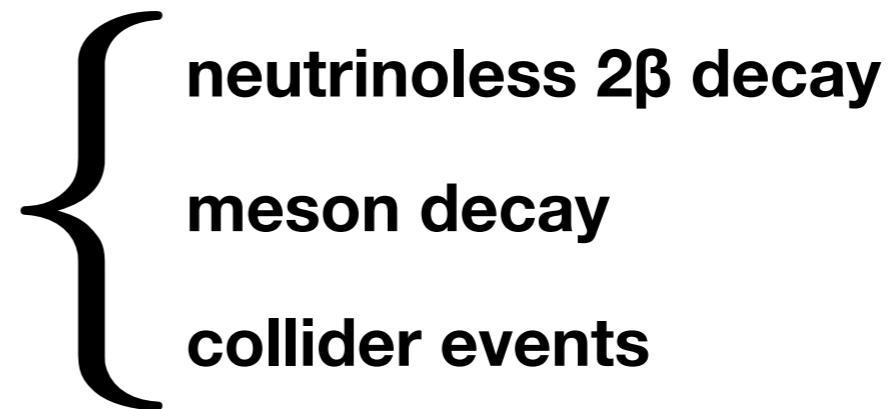
Lepton number is an accidental symmetry of the SM: test its conservation!

It is violated by non-renormalizable SM operators: EFT generation of v masses

LNV phenomenology is generally connected with v mass generation mechanism

LNV rates depend in general on the interference of multiple virtual states

Possible to look for LNV in e.g.



LNV observation could signal the existence of new gauge bosons and/or falsify high-scale leptogenesis

A global B-L symmetry has non-trivial consequences in the early Universe

Backup

Temperature of level crossing

The level crossing temperature in the SM + 3 RHN can be estimated at

$$T_{\text{crossing}} \approx \frac{2\sqrt{2}\bar{M}\sqrt{\mu'^2 - 1}}{\sqrt{\sum_a |F_a|^2}} = 2.8 \times 10^5 \text{ GeV} \left(\frac{\bar{M}}{\text{GeV}} \right) \frac{\sqrt{\mu'^2 - 1}}{\sqrt{\sum_a |(F_a/10^{-5})|^2}}$$

where the Majorana and Yukawa matrices are parameterised as

$$M_M = \bar{M} \begin{pmatrix} 1 - \mu & 0 & 0 \\ 0 & 1 + \mu & 0 \\ 0 & 0 & \mu' \end{pmatrix} \quad F = \frac{1}{\sqrt{2}} \begin{pmatrix} F_e(1 + \epsilon_e) & iF_e(1 - \epsilon_e) & F_e\epsilon'_e \\ F_\mu(1 + \epsilon_\mu) & iF_\mu(1 - \epsilon_\mu) & F_\mu\epsilon'_\mu \\ F_\tau(1 + \epsilon_\tau) & iF_\tau(1 - \epsilon_\tau) & F_\tau\epsilon'_\tau \end{pmatrix}$$

Momentum averaged effective Hamiltonian for the heavy neutrinos

$$\langle H \rangle = \langle H_0 + V_N \rangle = \frac{\pi^2}{36\zeta(3)} \left(\frac{\text{diag}(0, M_2^2 - M_1^2, M_3^2 - M_1^2)}{T} + \frac{T}{8} F^\dagger F \right)$$

Quantum kinetic equations for freeze-in leptogenesis

$$\begin{aligned}
\frac{dR_N}{dt} = & -i [\langle H \rangle, R_N] - \frac{1}{2} \langle \gamma^{(0)} \rangle \left\{ F^\dagger F, R_N - I \right\} - \frac{1}{2} \langle \gamma^{(1b)} \rangle \left\{ F^\dagger \mu F, R_N \right\} + \langle \gamma^{(1a)} \rangle F^\dagger \mu F + \\
& - \frac{1}{2} \langle \tilde{\gamma}^{(0)} \rangle \left\{ M_M F^T F^* M_M, R_N - I \right\} + \frac{1}{2} \langle \tilde{\gamma}^{(1b)} \rangle \left\{ M_M F^T \mu F^* M_M, R_N \right\} + \\
& - \langle \tilde{\gamma}^{(1a)} \rangle M_M F^T \mu F^* M_M , \\
\frac{d\mu_{\Delta a}}{dt} = & - \frac{9 \zeta(3)}{2 N_D \pi^2} \left\{ \langle \gamma^{(0)} \rangle \left(F R_N F^\dagger - F^* R_{\bar{N}} F^T \right) - 2 \langle \gamma^{(1a)} \rangle \mu F F^\dagger + \right. \\
& + \langle \gamma^{(1b)} \rangle \mu \left(F R_N F^\dagger + F^* R_{\bar{N}} F^T \right) \\
& + \langle \tilde{\gamma}^{(0)} \rangle \left(F^* M_M R_{\bar{N}} M_M F^T - F M_M R_N M_M F^\dagger \right) - 2 \langle \tilde{\gamma}^{(1a)} \rangle \mu F^* M_M^2 F^T \\
& \left. + \langle \tilde{\gamma}^{(1b)} \rangle \mu \left(F^* M_M R_{\bar{N}} M_M F^T + F M_M R_N M_M F^\dagger \right) \right\}_{aa} ,
\end{aligned}$$

G. Sigl and G. Raffelt, Nucl. Phys. B406 (1993) 423; E. K. Akhmedov, V. A. Rubakov and A. Yu. Smirnov, 9803255; T. Asaka and M. Shaposhnikov, 0505013; L. Canetti, M. Drewes, T. Frossard and M. Shaposhnikov, 1208.4607; T. Asaka, S. Eijima and H. Ishida, 1112.5565; P. Hernandez, M. Kekic, J. Lopez-Pavon, J. Racker and J. Salvado, 1606.06719; S. Antusch, E. Cazzato, M. Drewes, O. Fischer, B. Garbrecht, D. Gueter et al., 1710.03744; A. Abada, G. Arcadi, V. Domcke, M. Drewes, J. Klaric and M. Lucente, 1810.12463

Some SM extensions with L symmetry

Linear seesaw

E. K. Akhmedov, M. Lindner, E. Schnapka and J. W. F. Valle, hep-ph/9507275 and hep-ph/9509255; S. M. Barr, hep-ph/0309152; M. Malinsky, J. C. Romao and J. W. F. Valle, hep-ph/0506296; M. B. Gavela, T. Hambye, D. Hernandez and P. Hernandez, arXiv:0906.1461 [hep-ph]

Inverse seesaw

D. Wyler and L. Wolfenstein, Nucl. Phys. B 218 (1983) 205; R. N. Mohapatra, Phys. Rev. Lett. 56 (1986) 561; R. N. Mohapatra and J. W. F. Valle, Phys. Rev. D 34 (1986) 1642; J. Bernabeu, A. Santamaria, J. Vidal, A. Mendez and J. W. F. Valle, Phys. Lett. B 187 (1987) 303; M. C. Gonzalez-Garcia and J. W. F. Valle, Phys. Lett. B 216 (1989) 360; F. Deppisch and J. W. F. Valle, hep-ph/0406040; A. Abada and M. Lucente, arXiv:1401.1507 [hep-ph]

Supersymmetry with R-parity violation

J. R. Ellis, G. Gelmini, C. Jarlskog, G. G. Ross and J. W. F. Valle, Phys. Lett. 150B (1985) 142; G. G. Ross and J. W. F. Valle, Phys. Lett. 151B (1985) 375; J. C. Romao, M. A. Diaz, M. Hirsch, W. Porod and J. W. F. Valle, hep-ph/9907499; A. Abada and M. Losada, hep-ph/9908352; M. Hirsch, M. A. Diaz, W. Porod, J. C. Romao and J. W. F. Valle, Phys. Rev. D62 (2000) 113008 [hep-ph/0004115]; A. Abada, S. Davidson and M. Losada, hep-ph/0111332

Scale invariance

V. V. Khoze and G. Ro, arXiv:1307.3764 [hep-ph]

Technicolor-inspired

T. Appelquist and R. Shrock, hep-ph/0204141; T. Appelquist and R. Shrock, hep-ph/0301108

vMSM

M. Shaposhnikov, hep-ph/0605047

Low-scale seesaw realisations

A. Ibarra, E. Molinaro and S. T. Petcov, arXiv:1007.2378 [hep-ph]; A. Ibarra, E. Molinaro and S. T. Petcov, arXiv:1103.6217 [hep-ph]; D. N. Dinh, A. Ibarra, E. Molinaro and S. T. Petcov, arXiv:1205.4671 [hep-ph]