

# *Massive Gravity in the Swampland*

Brando Bellazzini

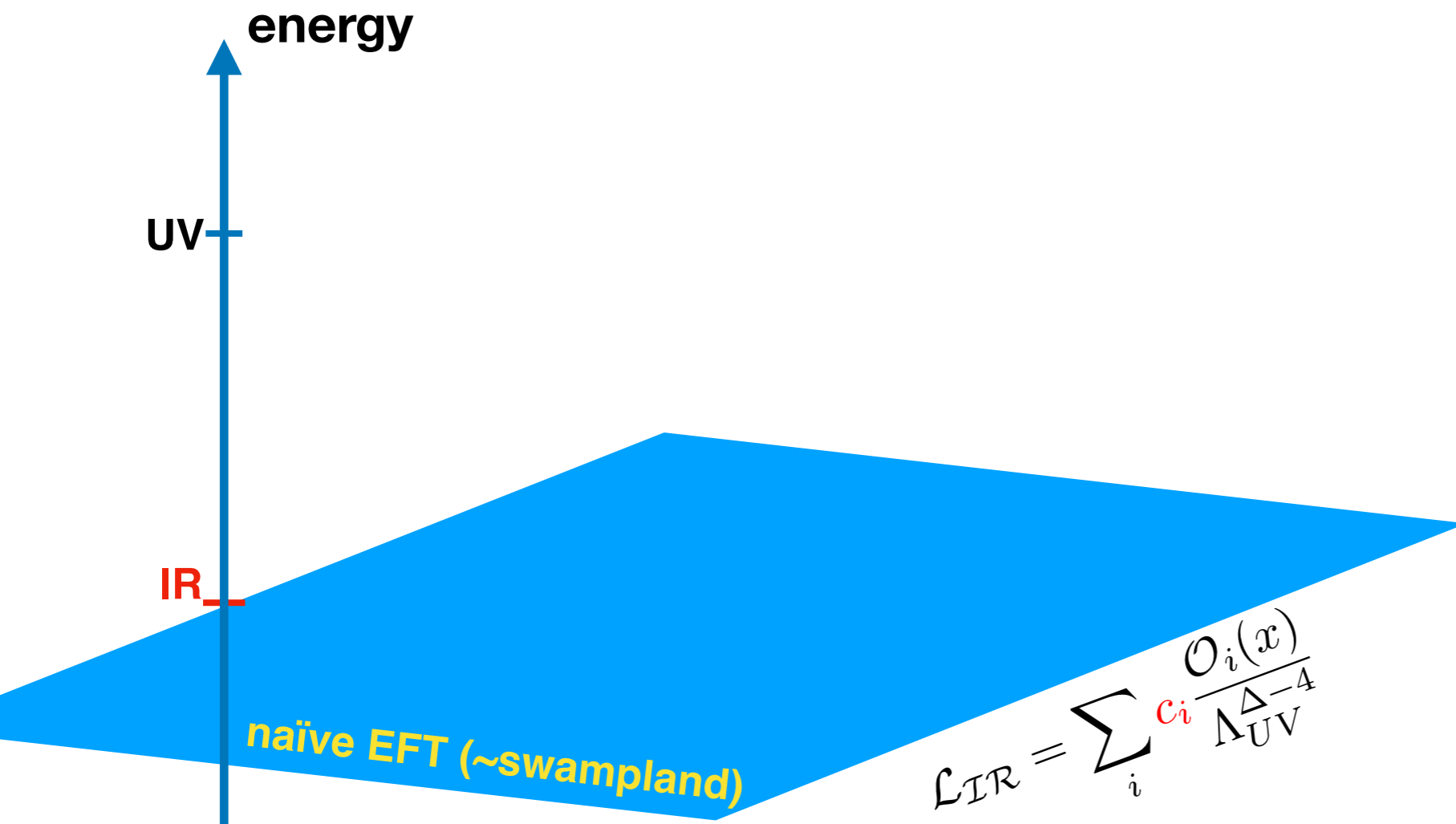


***based on arXiv: 2304.02550 by B.B., G. Isabella, S. Ricossa, F. Riva***

*XXIX XMass workshop @ IFT, Madrid. 13 Dec. 2023*

# QFT AND STRING SWAMPLAND

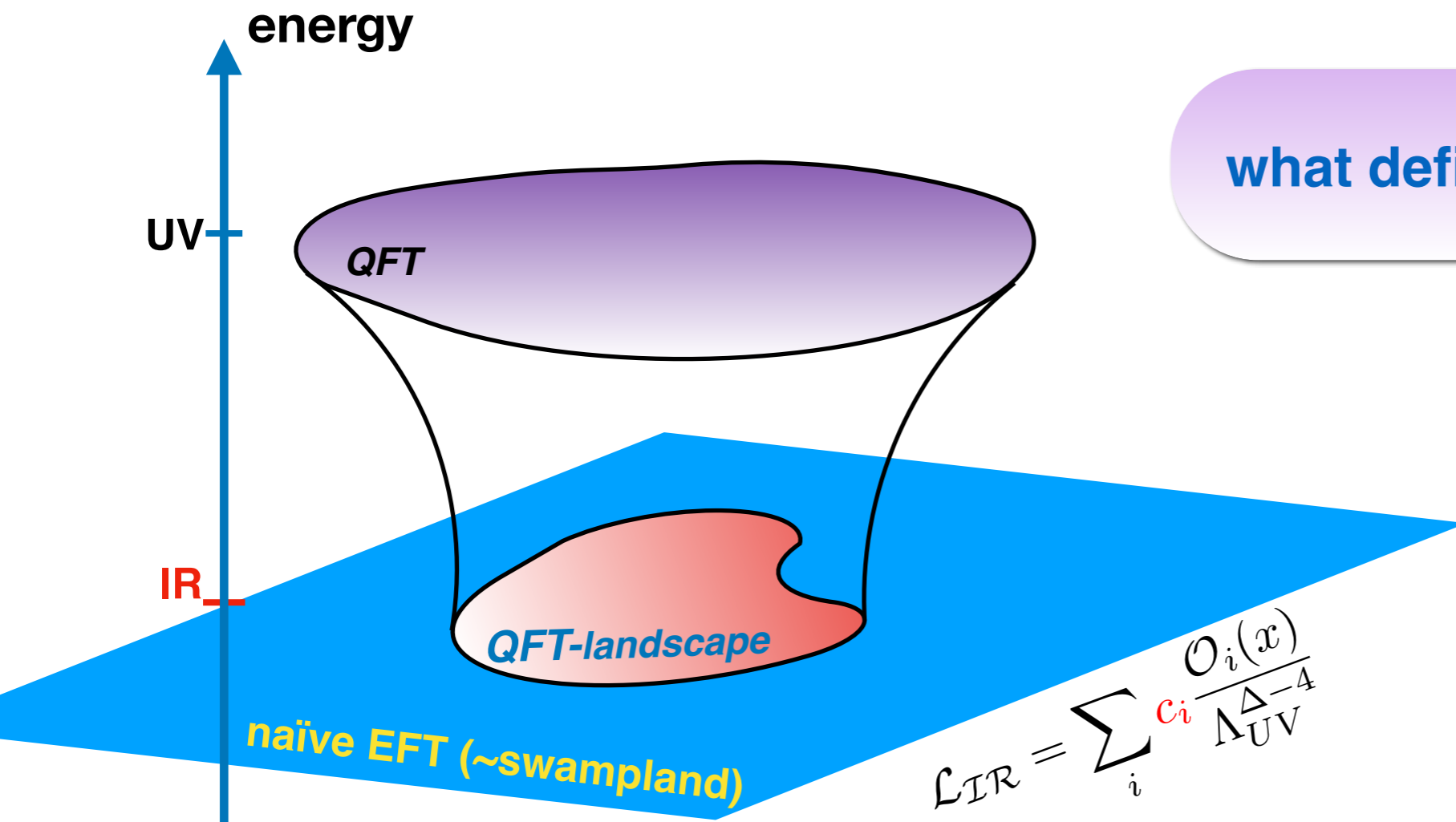
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what defines the boundary?

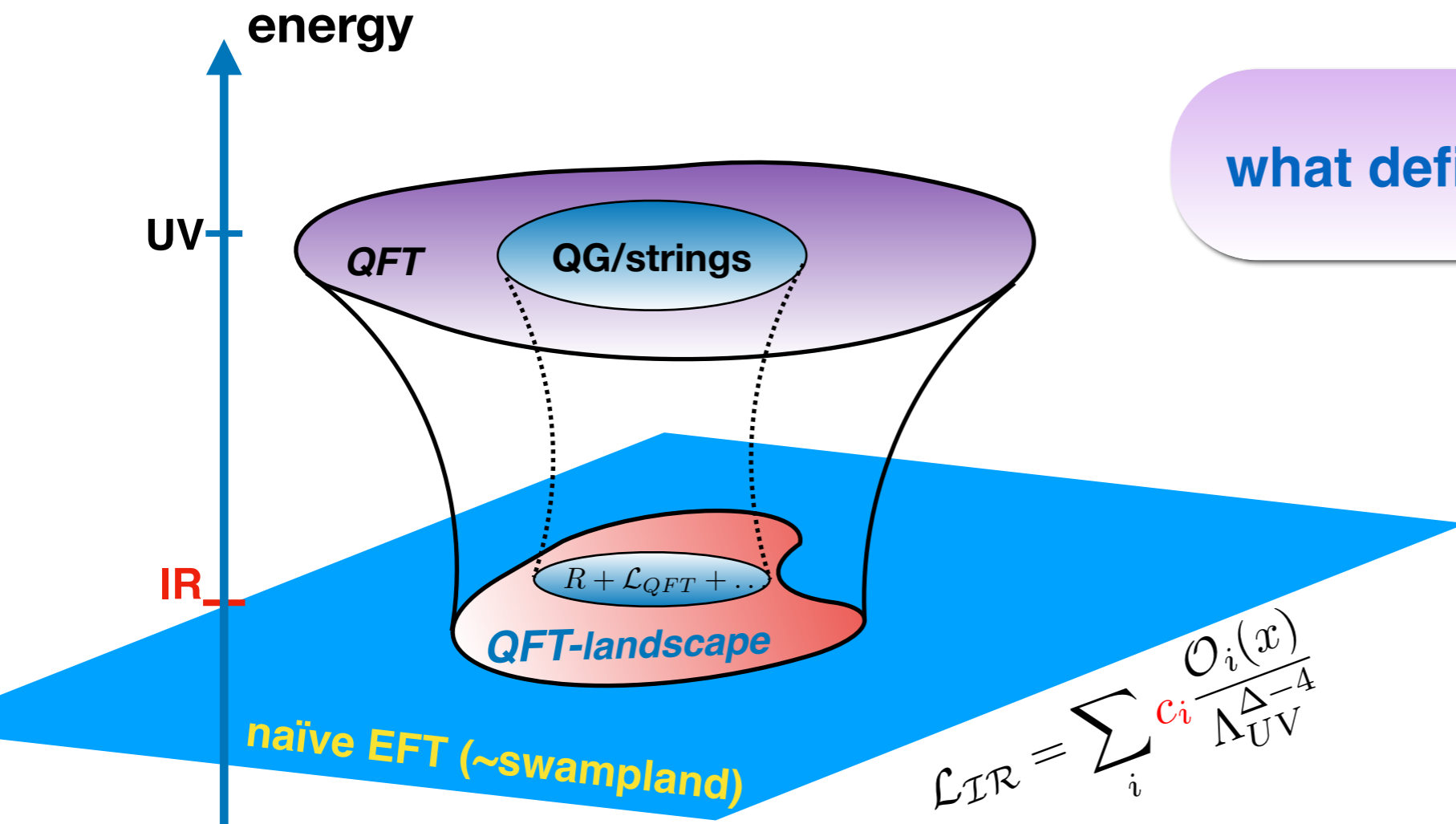


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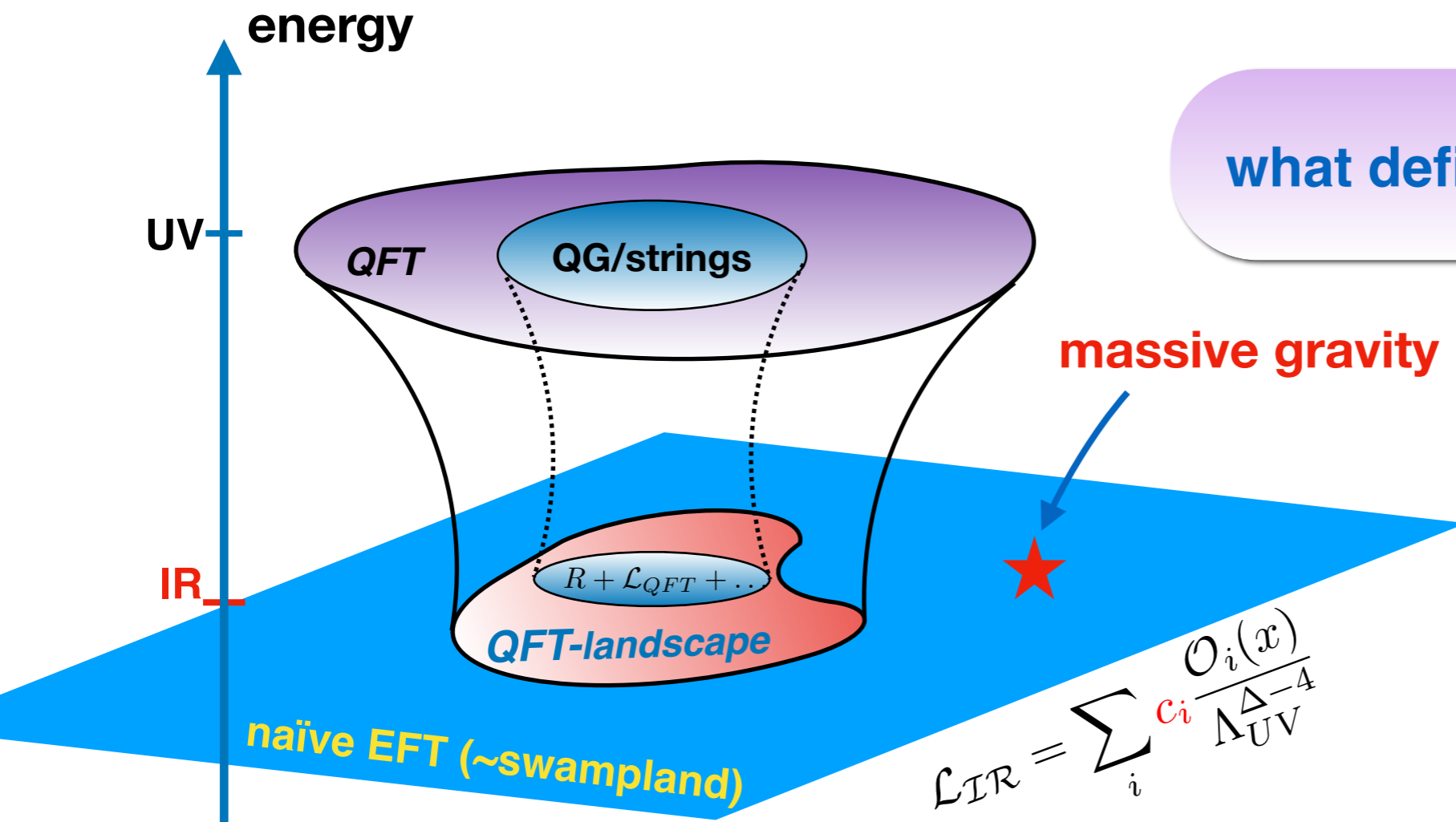
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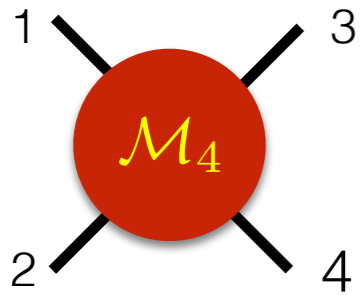
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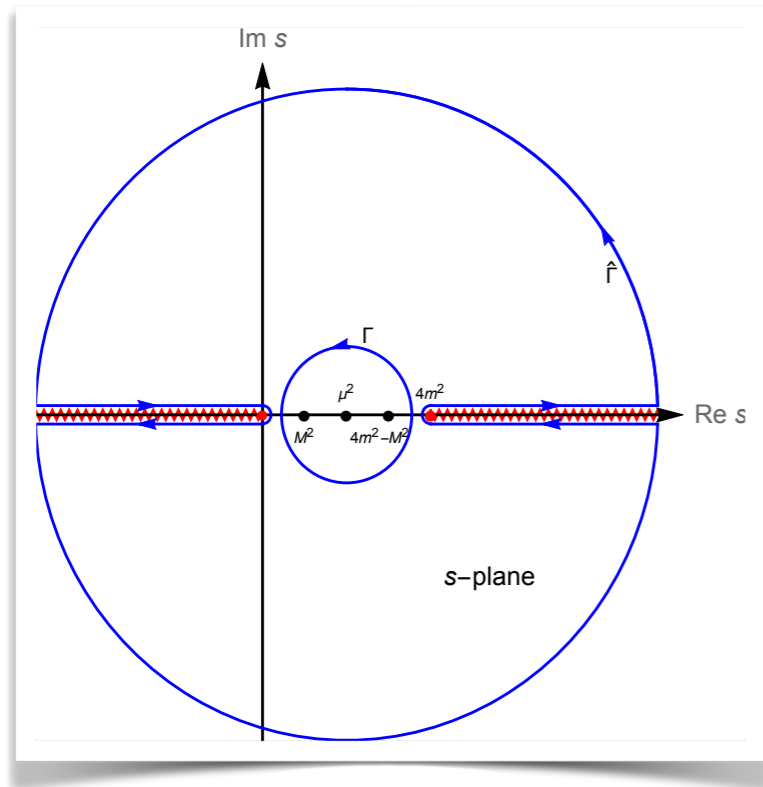
- Not every IR theory can be embedded in a consistent UV-theory
- This talk: massive gravity (and any isolated massive spin  $J \geq 2$ ) is in the swampland

# POSITIVITY BOUNDS

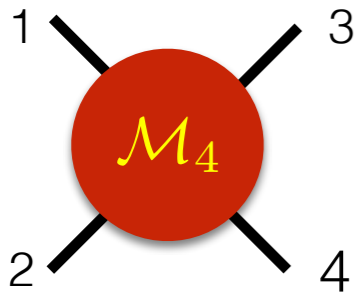
# UV-IR CONNECTION



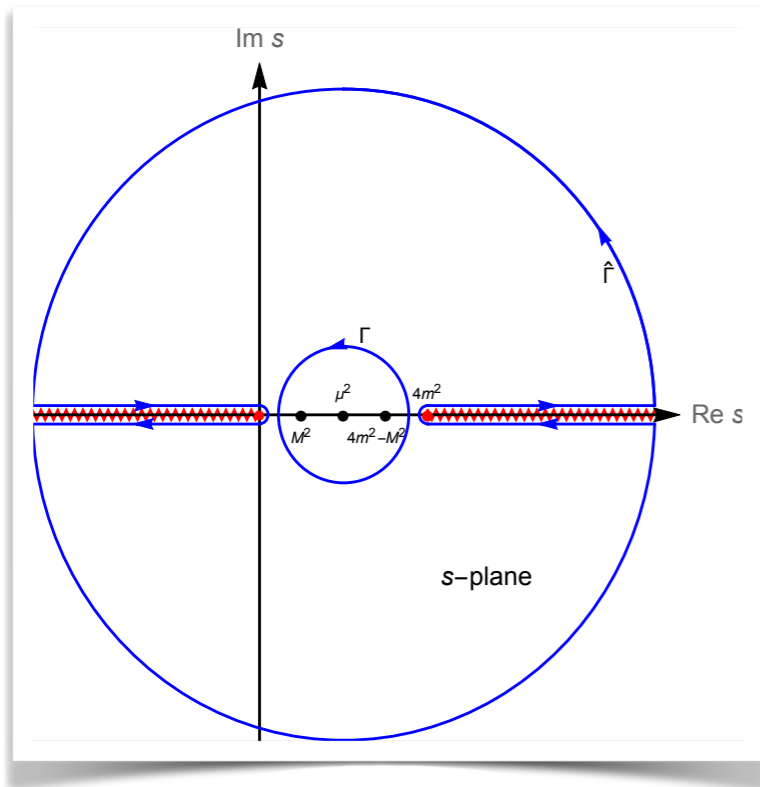
Analyticity, Crossing, Unitarity, Locality



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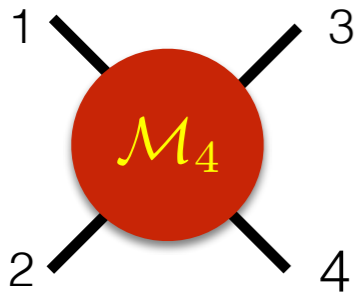
$$\mathcal{M}_{2 \rightarrow 2}^{(n)}|_{IR} = \int_0^\infty \frac{ds \operatorname{Im} \mathcal{M}_{2 \rightarrow 2}(s)}{s^{n+1}} > 0 \quad t = 0$$

IR-side UV-side

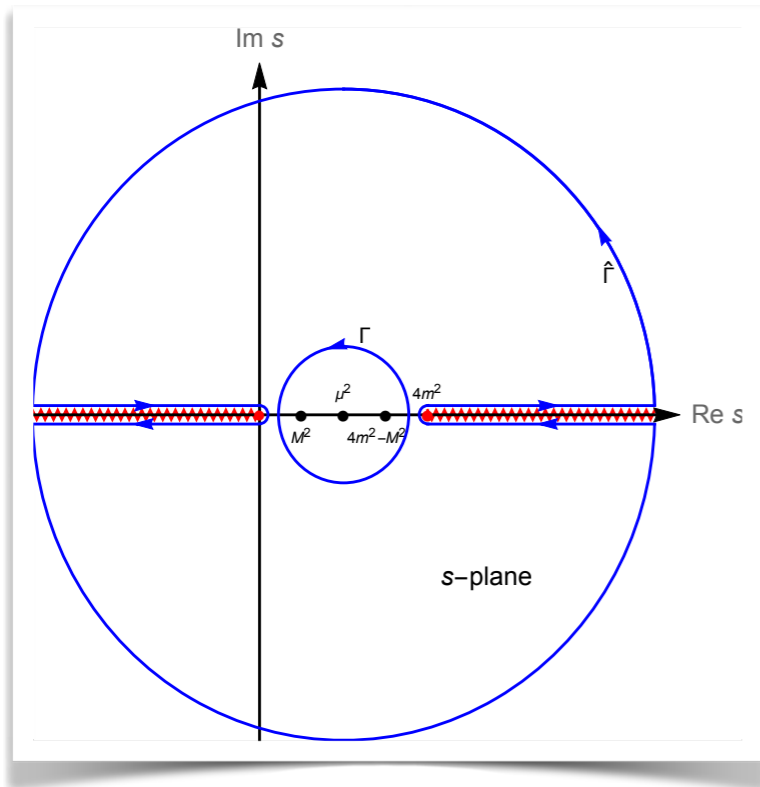
$s^2$ -terms are strictly positive

Adams, Arkani-Hamed, Dubovsky, Nicolis, Rattazzi hep-th/0602178  
BB 1605.06111

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paradigmatic example

$$\pi \rightarrow \pi + \text{const} \quad \mathcal{L} = \frac{1}{2}(\partial_\mu \pi)^2 + \frac{c}{\Lambda^4}(\partial_\mu \pi)^4 + \dots \rightarrow \mathcal{M}(\pi\pi \rightarrow \pi\pi)(s, t = 0) = cs^2 \rightarrow c > 0$$

$c < 0$  in the swampland

# POSITIVITY BOUNDS

$$\mathcal{M}(s)|_{\text{EFT}} = c_0 + c_2 s^2 + c_4 s^4 + \dots c_{21} s^2 t + \dots$$



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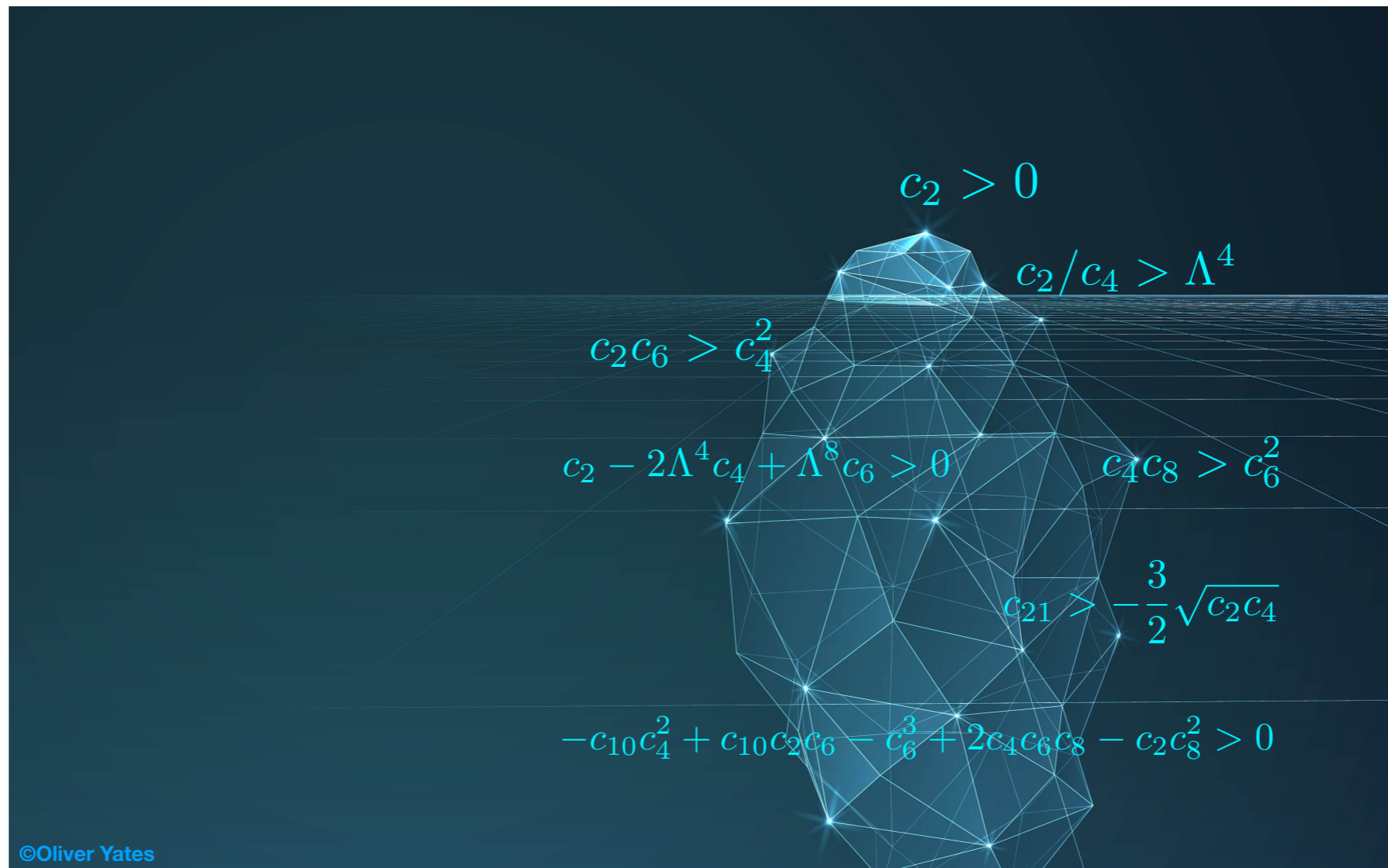
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what “theory”  
behind this landscape?

$$c_2 > 0$$

$$c_2/c_4 > \Lambda^4$$

$$c_2 c_6 > c_4^2$$

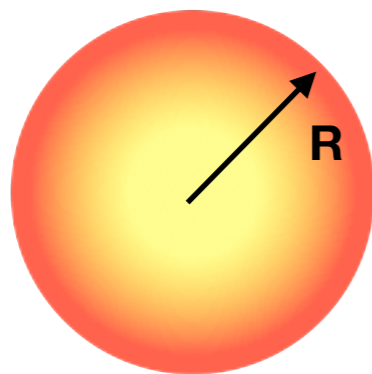
$$c_2 - 2\Lambda^4 c_4 + \Lambda^8 c_6 > 0$$

$$c_4 c_8 > c_6^2$$

$$c_{21} > -\frac{3}{2} \sqrt{c_2 c_4}$$

$$-c_{10} c_4^2 + c_{10} c_2 c_6 - c_6^3 + 2c_4 c_6 c_8 - c_2 c_8^2 > 0$$

# MOMENTS: AN ANALOGY WITH STARS



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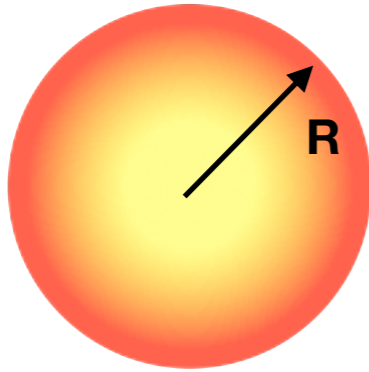
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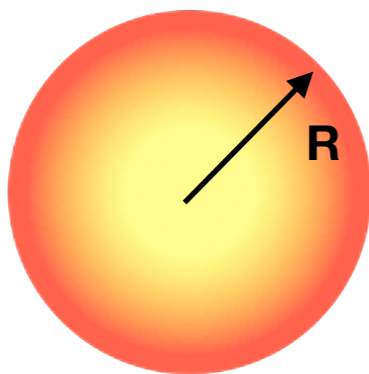
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double-sided bounds

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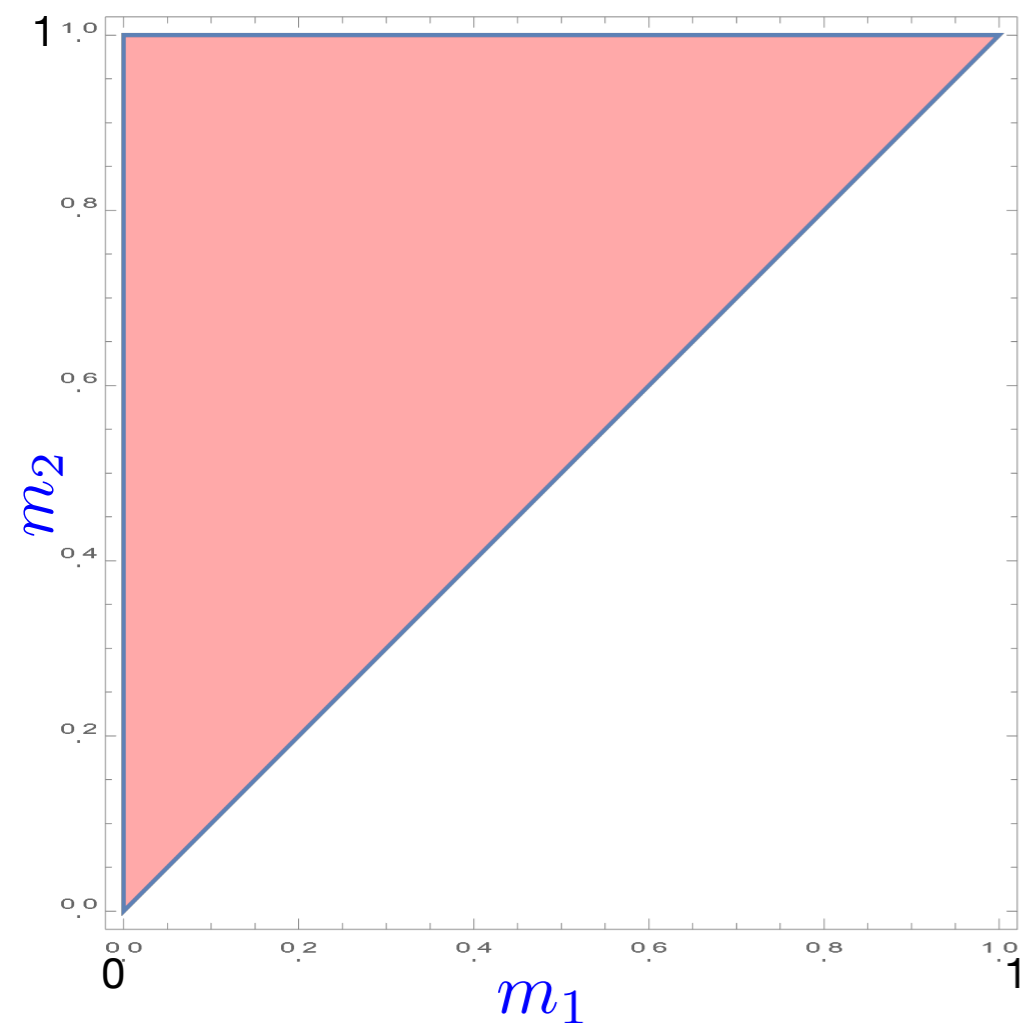
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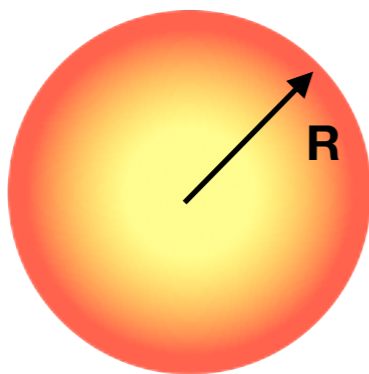
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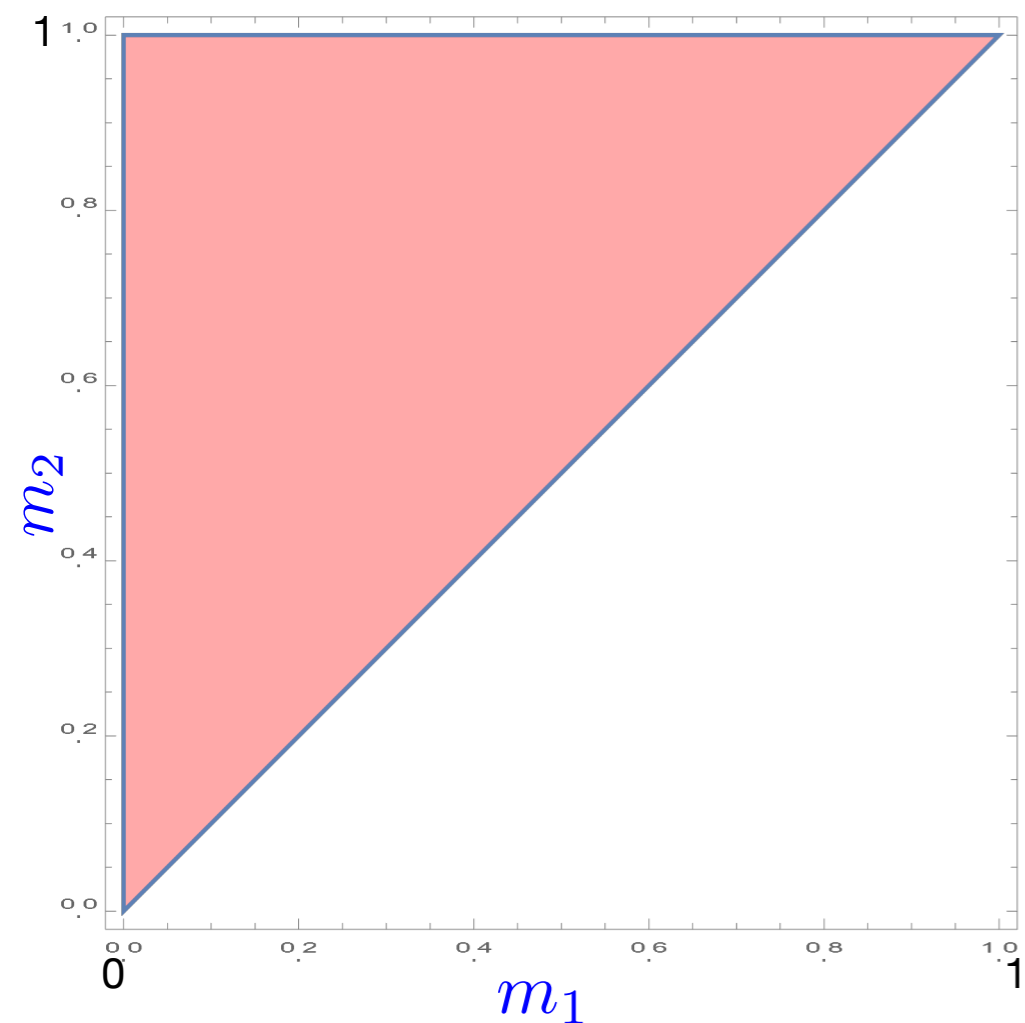
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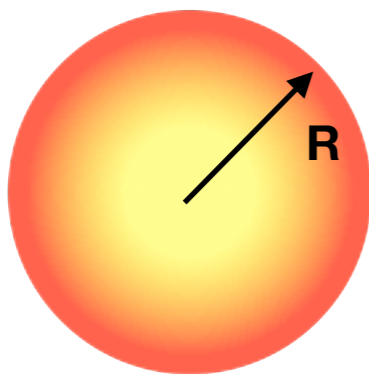
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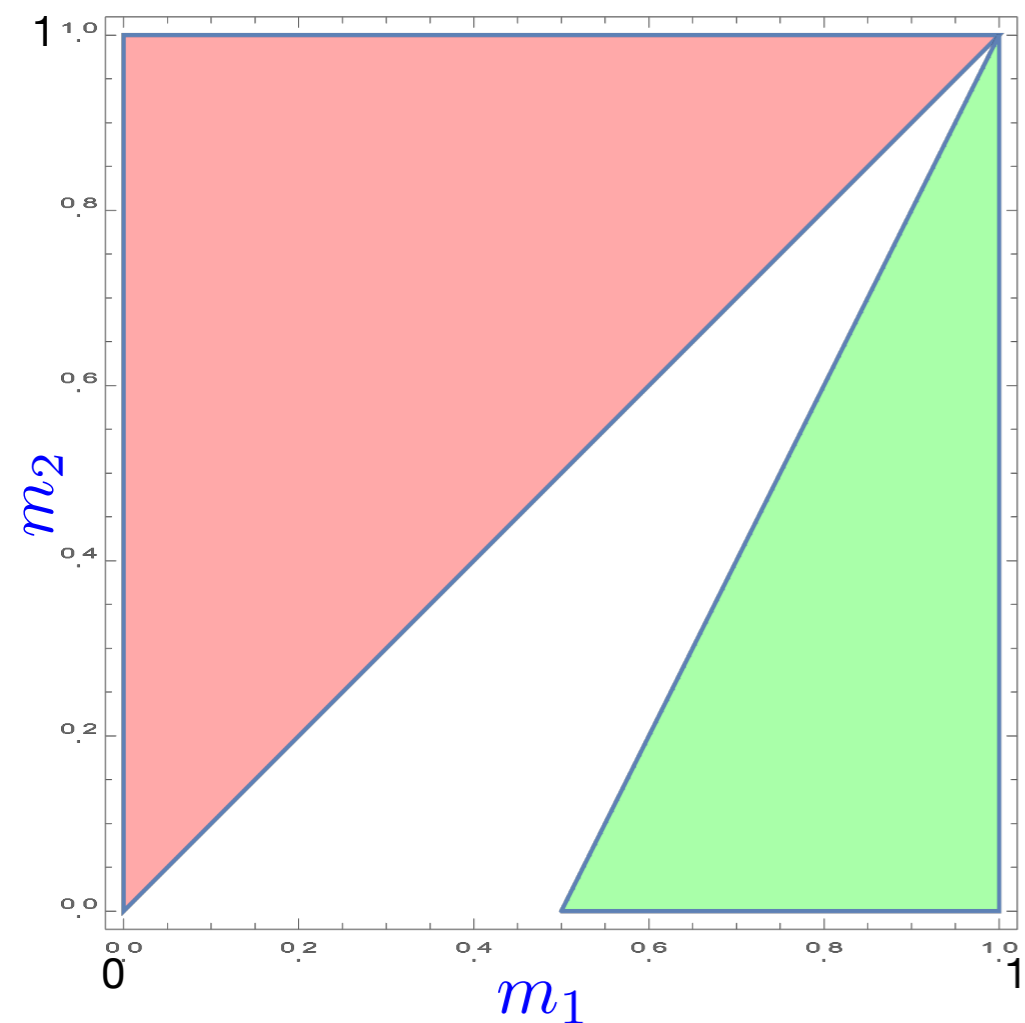
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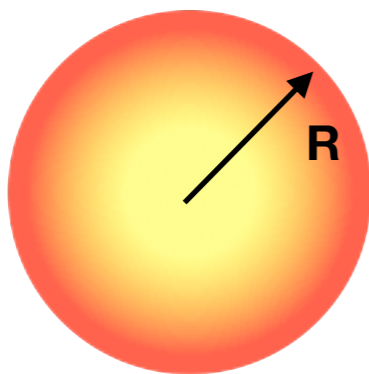
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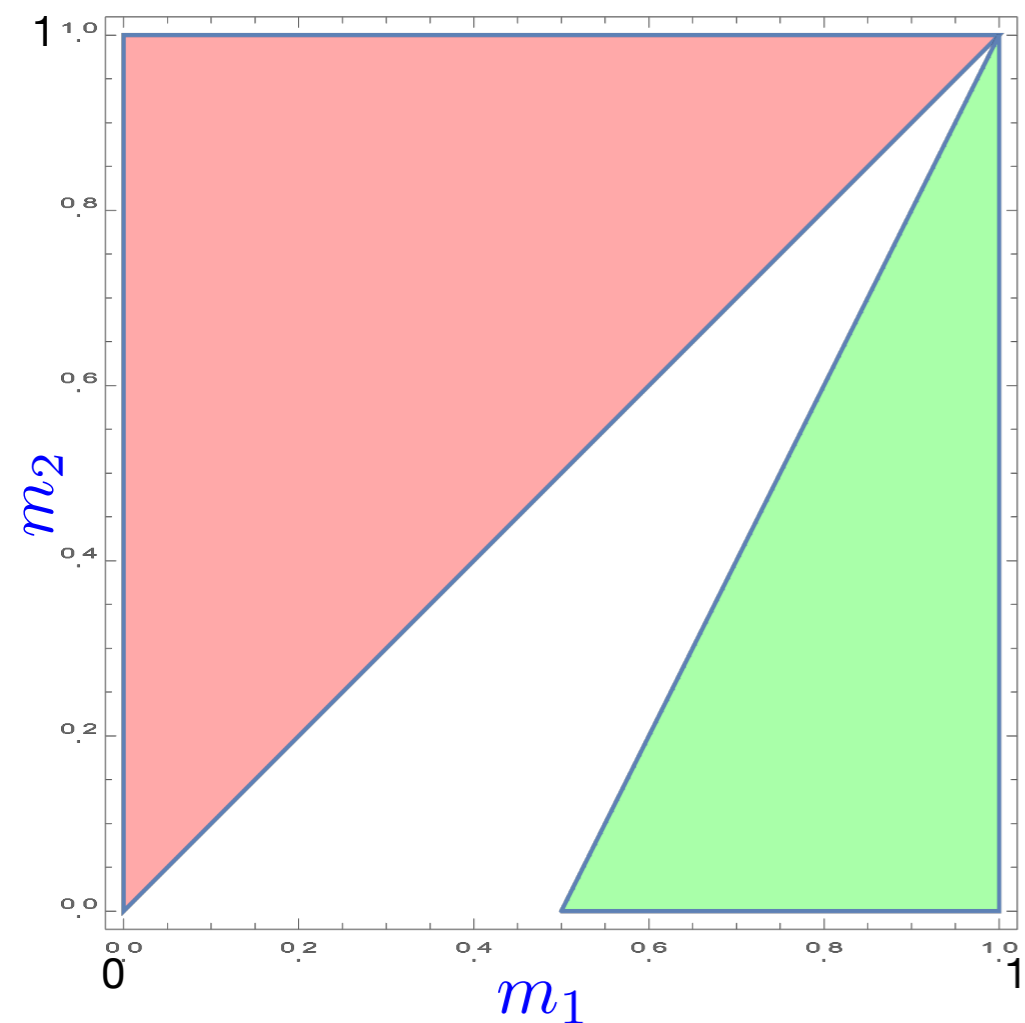
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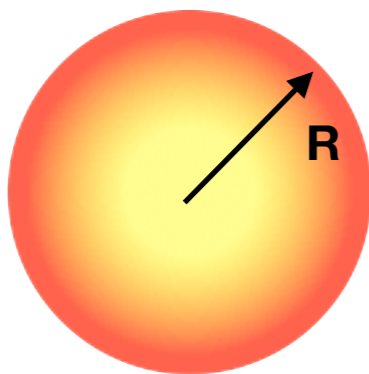
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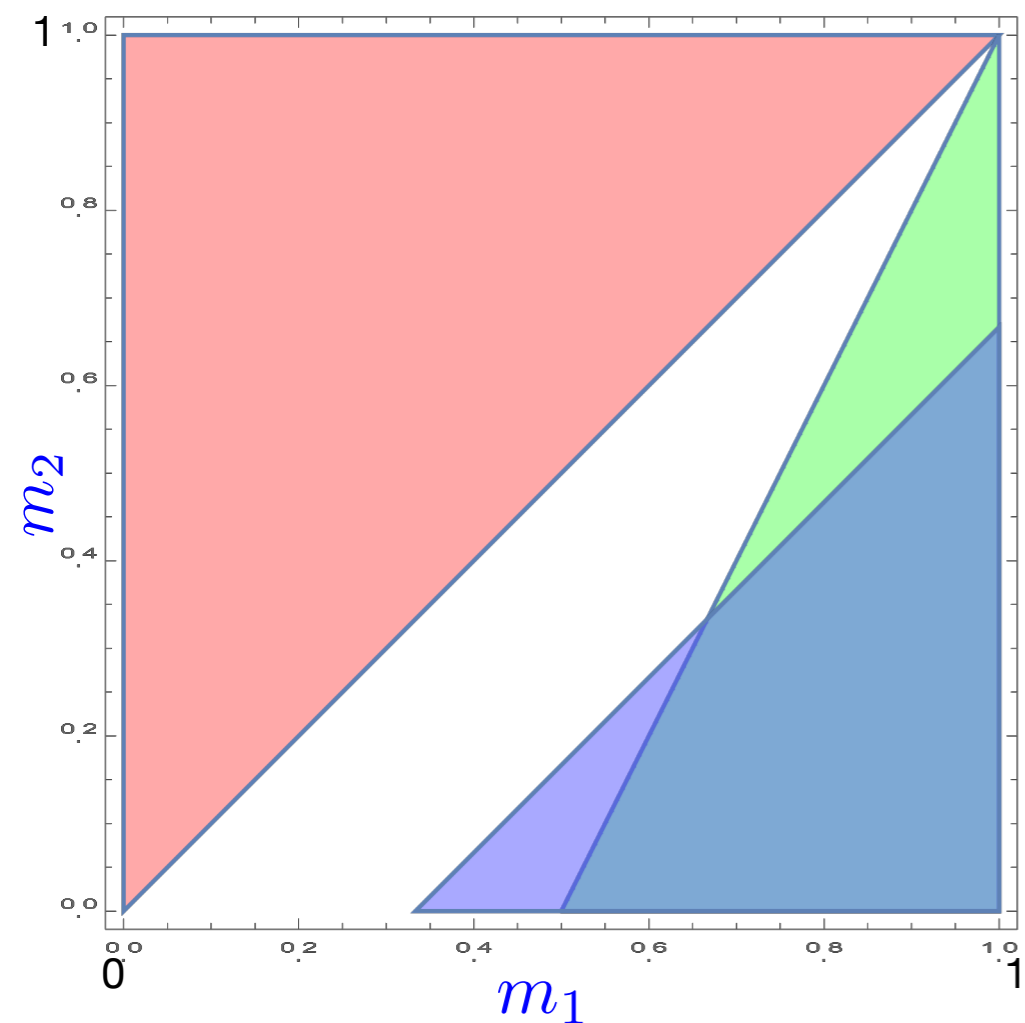
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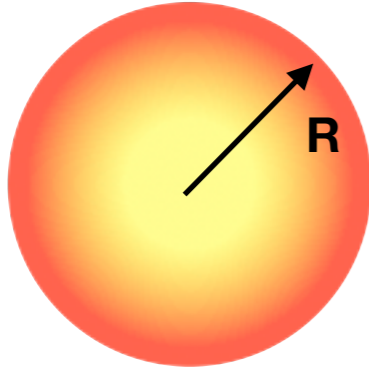
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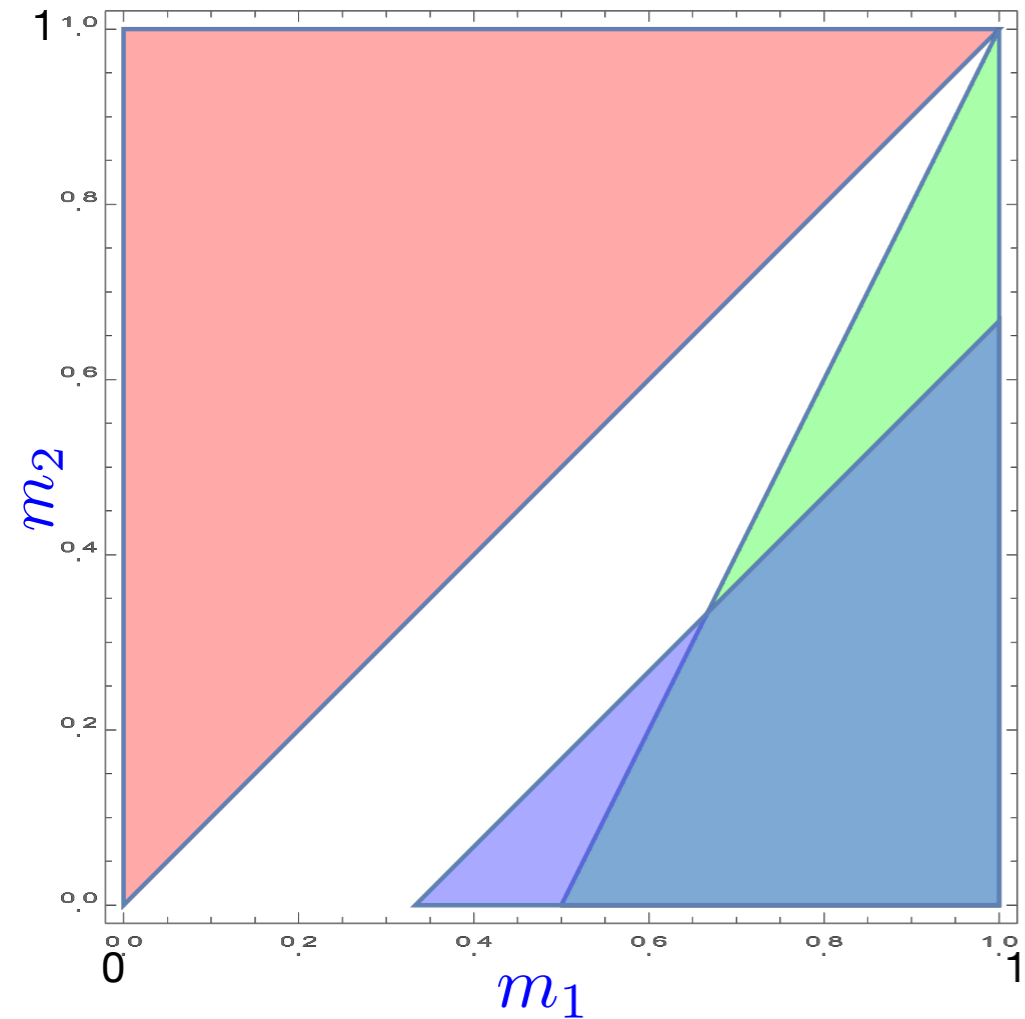
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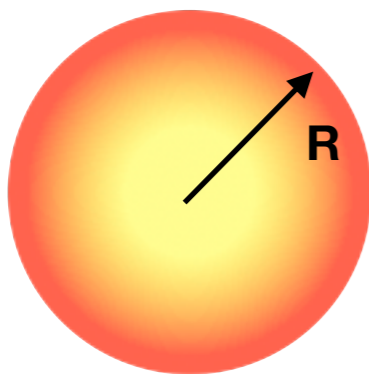
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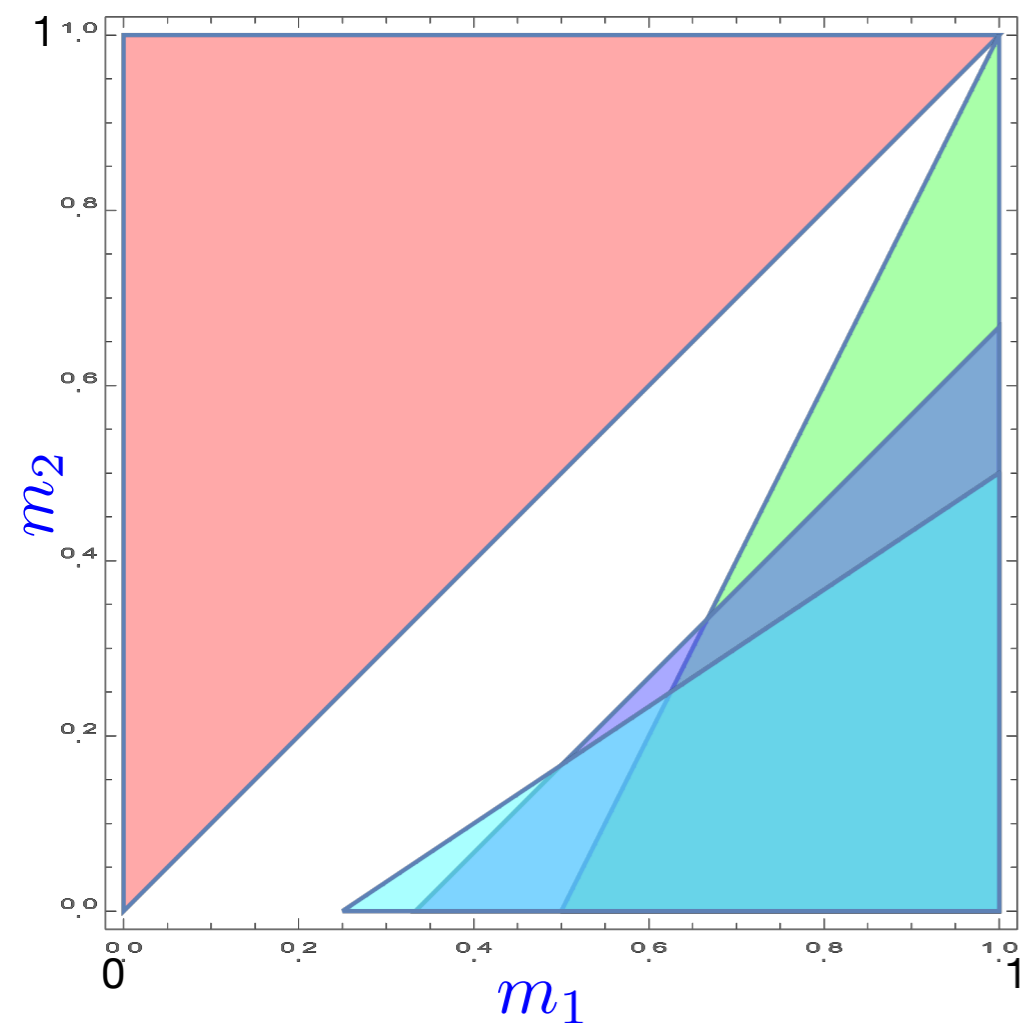
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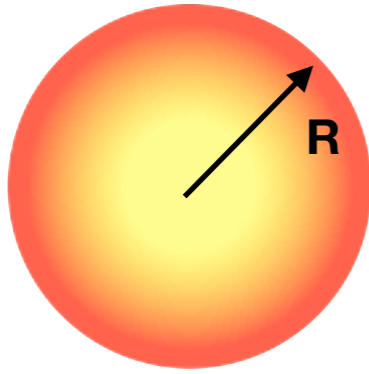
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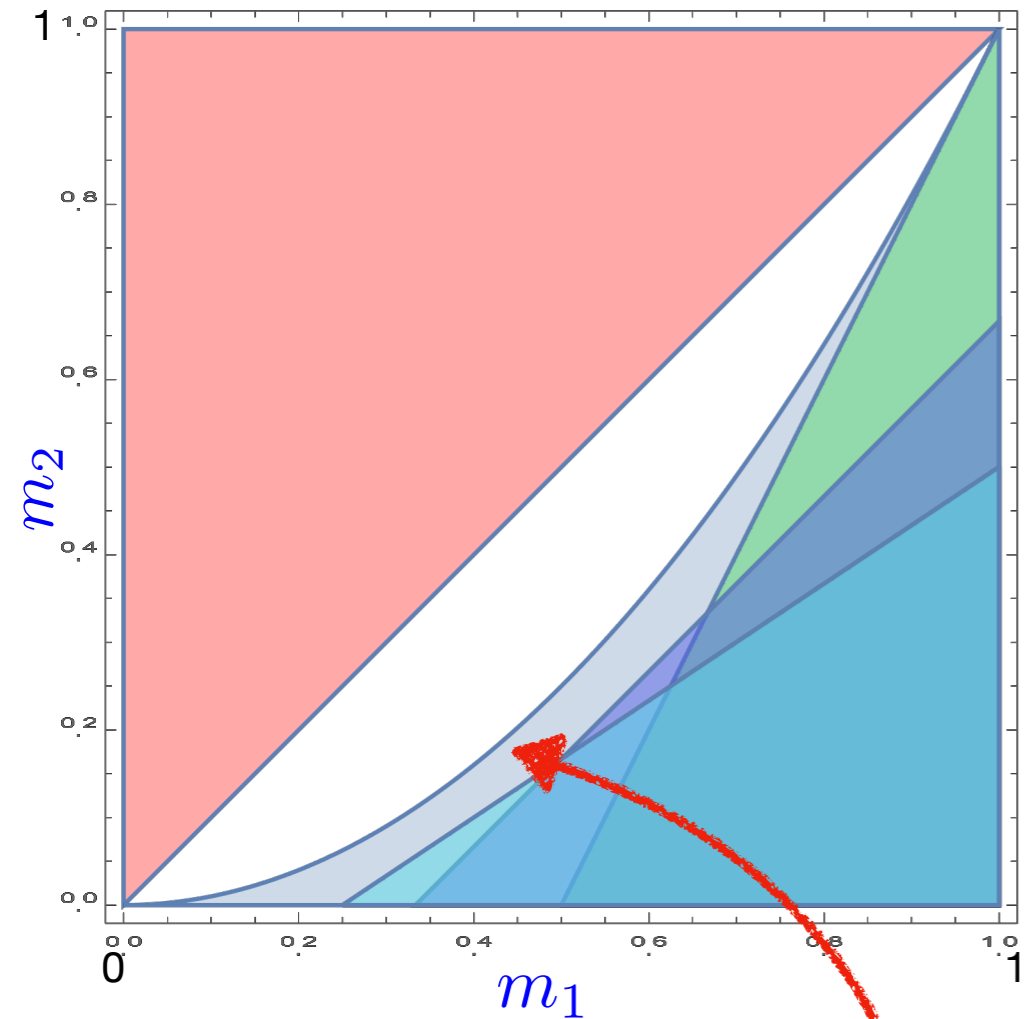
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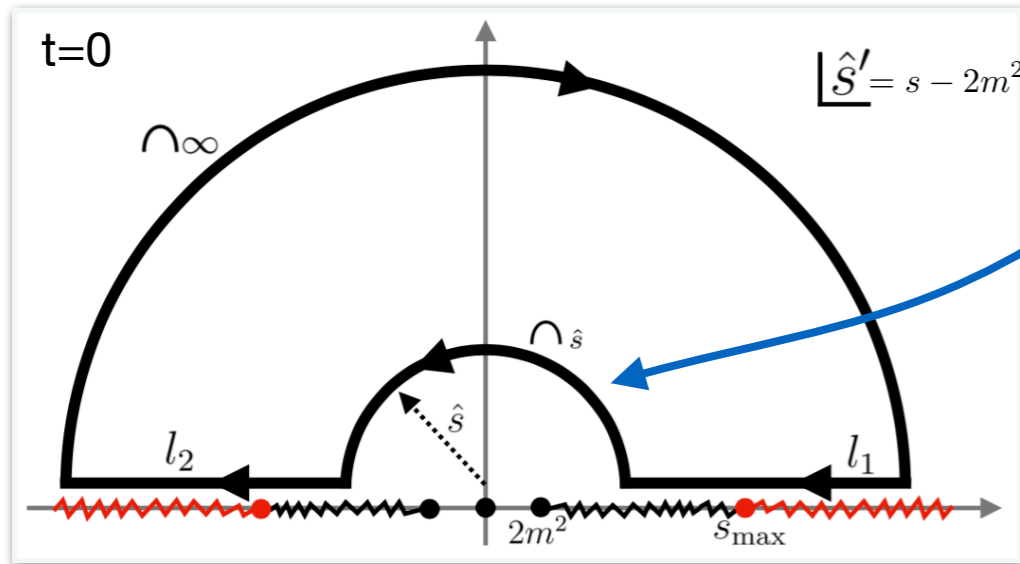
non-linear bounds (EFT-hedron approach) = inf-many linear bounds

**main lessons so far:**

- 1) all bounds  $\leftrightarrow$  all positive functions in  $[0,R]$**
- 2) it's possible to project on a finite subspace**
- 3) non-linear bounds= infinitely many linear**

**back to the amplitude**

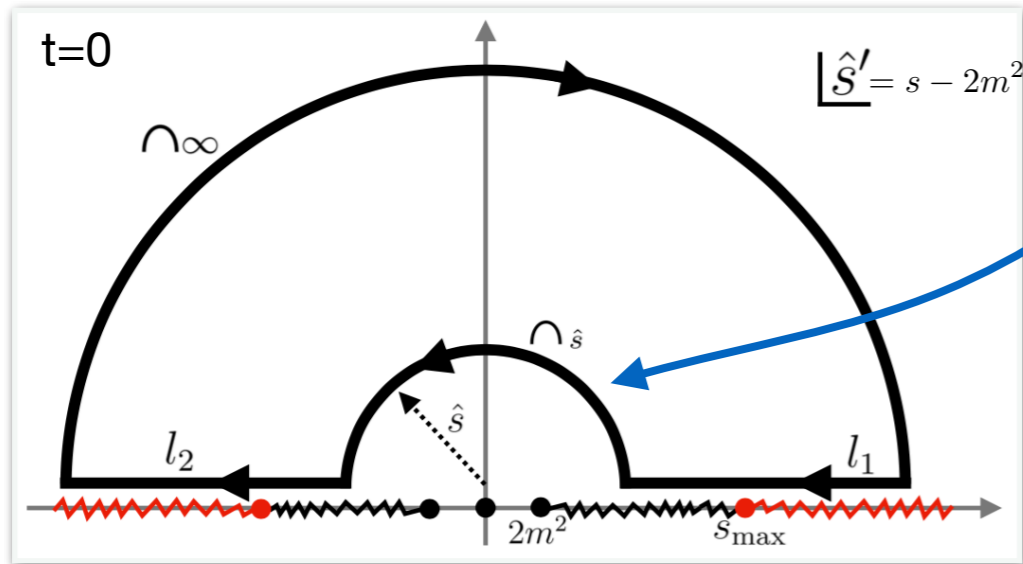
# MOMENTS=ARCS



$$a_n(\hat{s}) \equiv \int_{\mathcal{C}_{\hat{s}}} \frac{d\hat{s}'}{\pi i} \frac{\hat{\mathcal{M}}(\hat{s}')}{\hat{s}'^{2n+3}}$$

**“Arcs”** IR-rep.  $n \in \mathbb{N}$

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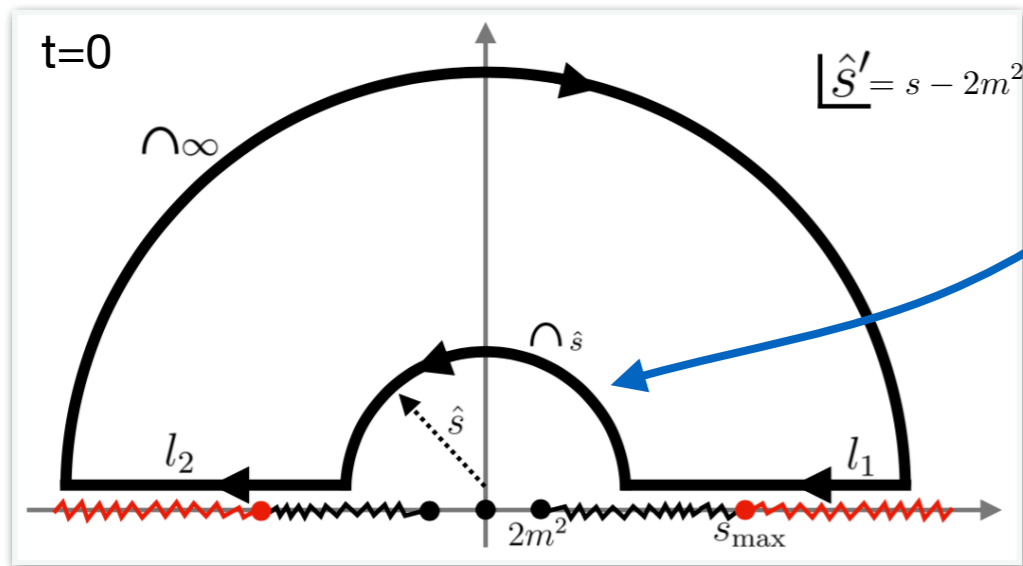
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theory space = most general positive function in  $[0,1]$

$$\int_0^1 \underbrace{f(x) d\mu(x)}_{>0}$$

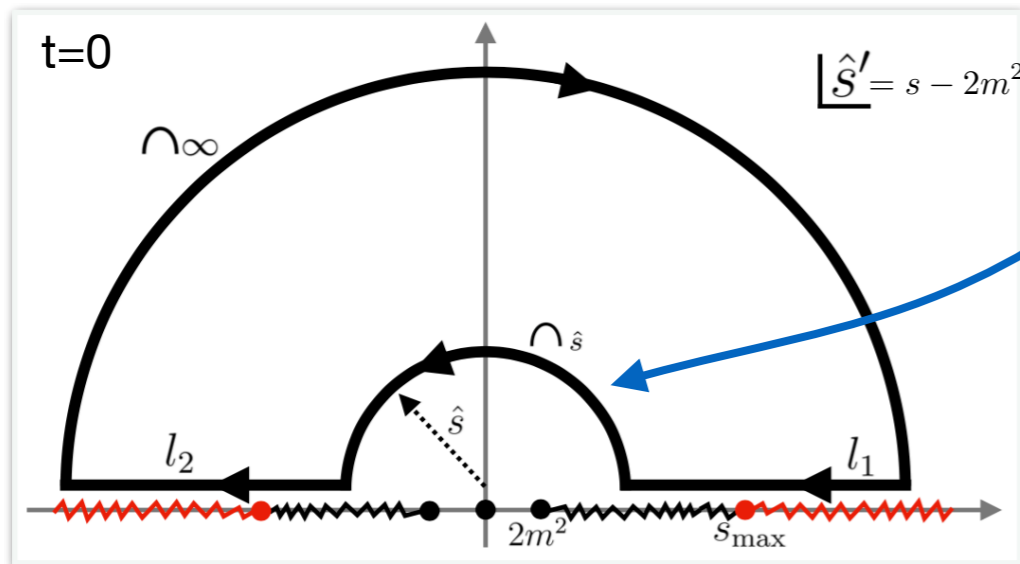


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Bernstein pol.



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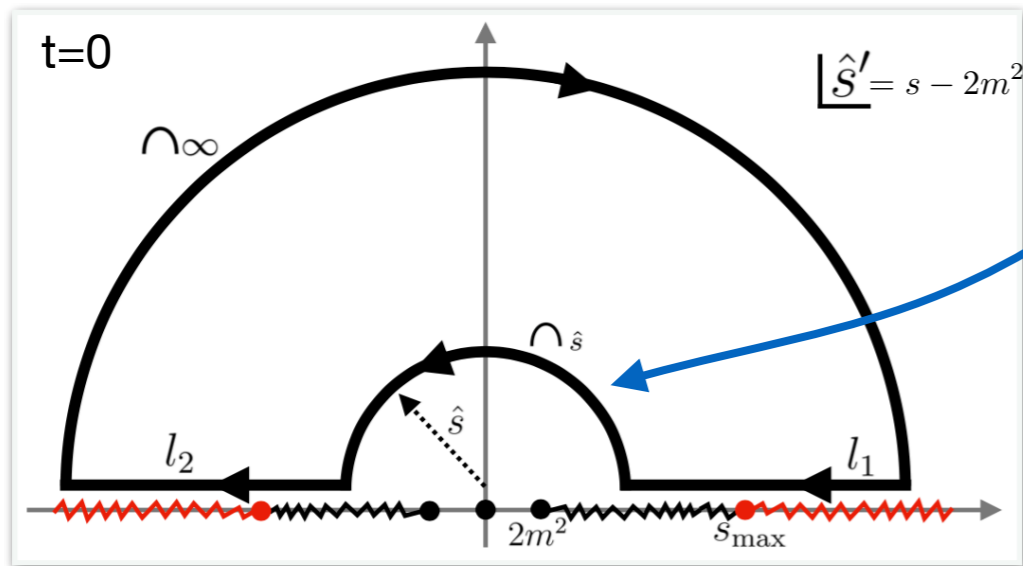
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**All bounds:**

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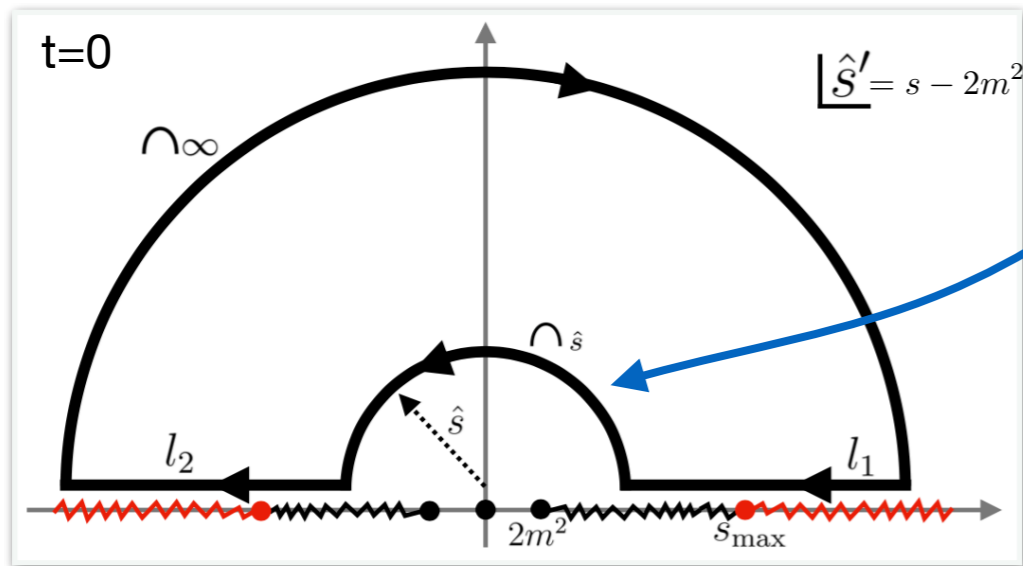
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# MOMENTS=ARCS



$$a_n(\hat{s}) \equiv \int_{\cap_{\hat{s}}} \frac{d\hat{s}'}{\pi i} \frac{\hat{\mathcal{M}}(\hat{s}')}{\hat{s}'^{2n+3}}$$

**“Arcs”** IR-rep.  $n \in \mathbb{N}$

$$= \frac{2}{\pi} \int_{\hat{s}}^{\infty} d\hat{s}' \frac{\text{Im} \hat{\mathcal{M}}(\hat{s}')}{\hat{s}'^{2n+3}} = \frac{1}{\hat{s}^{2n+2}} \int_0^1 x^n d\mu(x, \hat{s})$$

UV-rep.

theory space = most general positive function in  $[0,1]$

$$\int_0^1 \underbrace{f(x) d\mu(x)}_{>0}$$



$$f(x) = \sum_{n,m} \underbrace{x^m (1-x)^{n-m}}_{>0} \underbrace{c_{nm}}_{>0}$$

Bernstein pol.

**All bounds:**

$$\frac{1}{\hat{s}^{2n+2}} \int_0^1 x^n (1-x)^k d\mu(x, \hat{s}) > 0$$

**“Hausdorff’s moment theorem”**

**“complete monotonic series”**

$$\Delta a_n \equiv \hat{s}^2 a_{n+1} - a_n \stackrel{=}{=} (-\Delta)^k a_n > 0$$

# EXAMPLE @ TREE-LEVEL

## all bounds

$$\Delta^0 a_n = a_n > 0$$

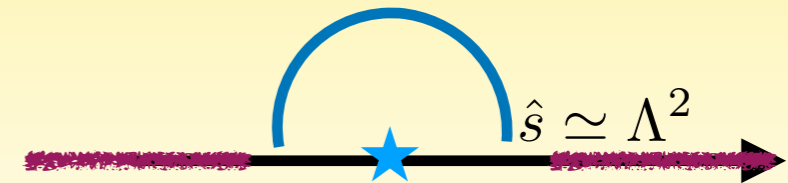
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⋮

## tree-level

$$\frac{\mathcal{M}(s)|_{\text{EFT}}}{s^{2n+3}} = \frac{c_0 + c_2 s^2 + c_4 s^4 + \dots}{s^{2n+3}}$$



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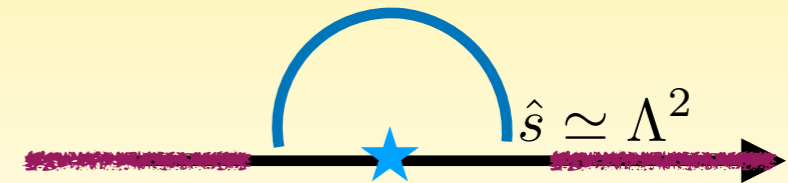
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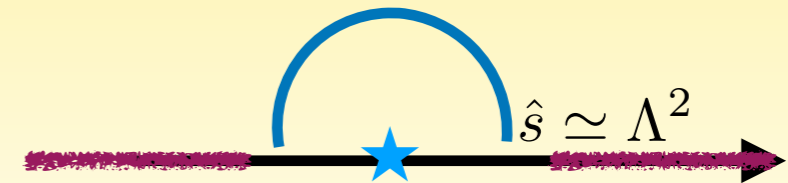
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sharp upper bound on the cutoff

$$\Lambda^4 < \frac{c_n}{c_{n+2}}$$

even for very weakly coupled theory  $c \ll 1$   
much more stringent strong-coupling cutoff

$$\mathcal{M}|_{\text{EFT}} = g^2 (1 + \bar{c}_2 s^2 + \bar{c}_4 s^4 + \dots) \quad \text{bound independent of } g \ll 1$$

EXAMPLE: MASSIVE HIGHER SPINS

# HIGHER SPINS $J > 2$

$$\begin{array}{l}
 \epsilon_{\mu_1 \mu_2 \dots \mu_J}^{(0)} \sim (E/m)^J \\
 \epsilon_{\mu_1 \mu_2 \dots \mu_J}^{(\pm 1)} \sim (E/m)^{J-1} \\
 \vdots \\
 \epsilon_{\mu_1 \mu_2 \dots \mu_J}^{(\pm J)} \sim (E/m)^0
 \end{array}
 \rightarrow
 \left\{
 \begin{array}{l}
 \mathcal{M}_{\text{long.}}^{(J-\text{even})} = g^2(m_J/\Lambda_J) \left[ \left(\frac{E}{m_J}\right)^{3J} + \dots + \frac{E^4}{m_J^4} + \dots \right] \\
 \mathcal{M}_{\text{long.}}^{(J-\text{odd})} = g^2(m_J/\Lambda_J) \left[ \left(\frac{E}{m_J}\right)^{3J+1} + \dots + \frac{E^4}{m_J^4} + \dots \right]
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 \right.$$

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**Main Lesson:** EFTs isolated Massive Higher spin ( $J > 2$ ) in the swampland

1903.08664 B.B., F. Riva, J. Serra, F. Sgarlata

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**caveat:** trivial forward amplitude  $\mathcal{M}(t \rightarrow 0) \approx 0$

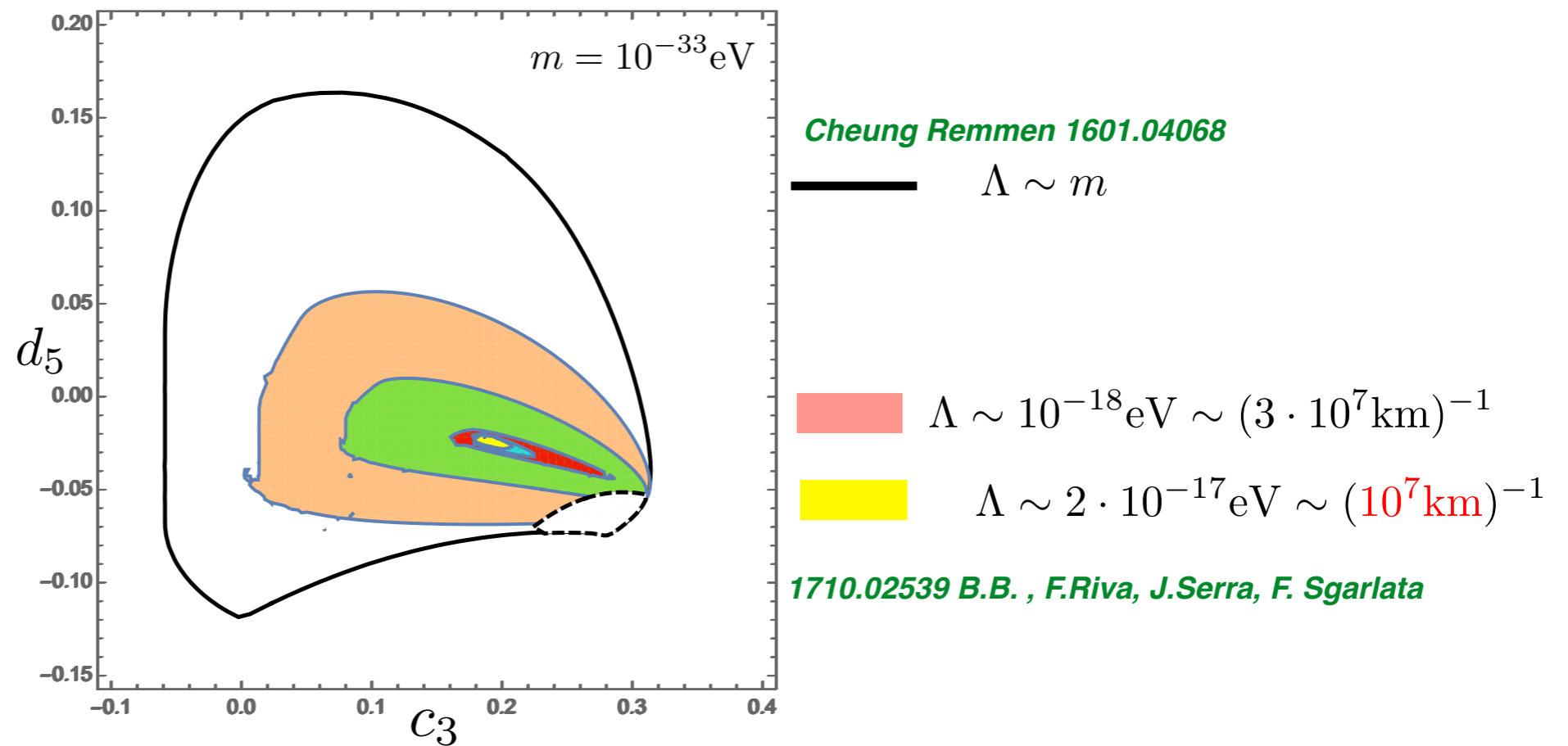
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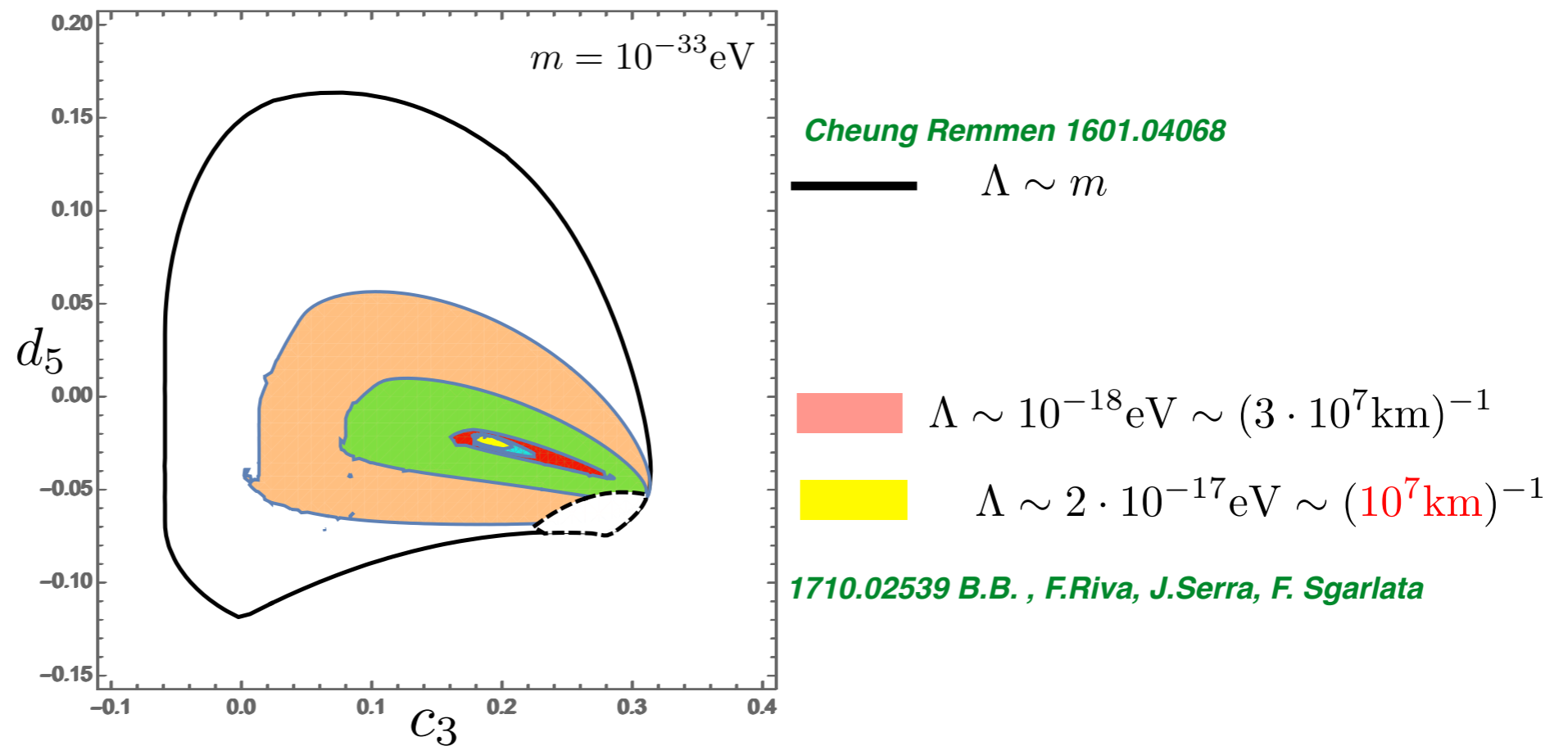
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need to go beyond forward limit

BEYOND FORWARD

# FINITE-T

$$\mathcal{M}(s)|_{\text{EFT}} = c_0 + c_2 s^2 + c_4 s^4 + \dots + c_{21} s^2 t + c_{22} s^2 t^2 + \dots$$

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“Arcs finite T  $\leftrightarrow$  2D-moments”

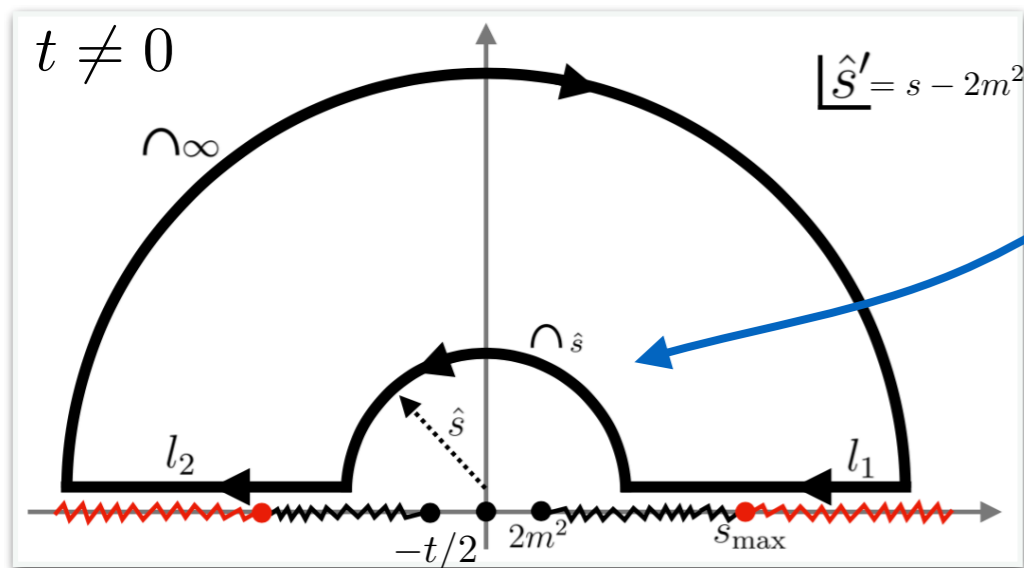
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$$a_n(\hat{s}, t) = \frac{2}{\pi} \int_{\hat{s}}^{\infty} d\hat{s}' \frac{\text{Im} \hat{\mathcal{M}}(\hat{s}', t)}{(\hat{s}' + \frac{t}{2})^{2n+3}}, \quad n \in \frac{\mathbb{N}}{2} \quad \text{UV-rep.}$$

$$= \sum_p t^p c_p^{n,m} \mu_{n,m}$$

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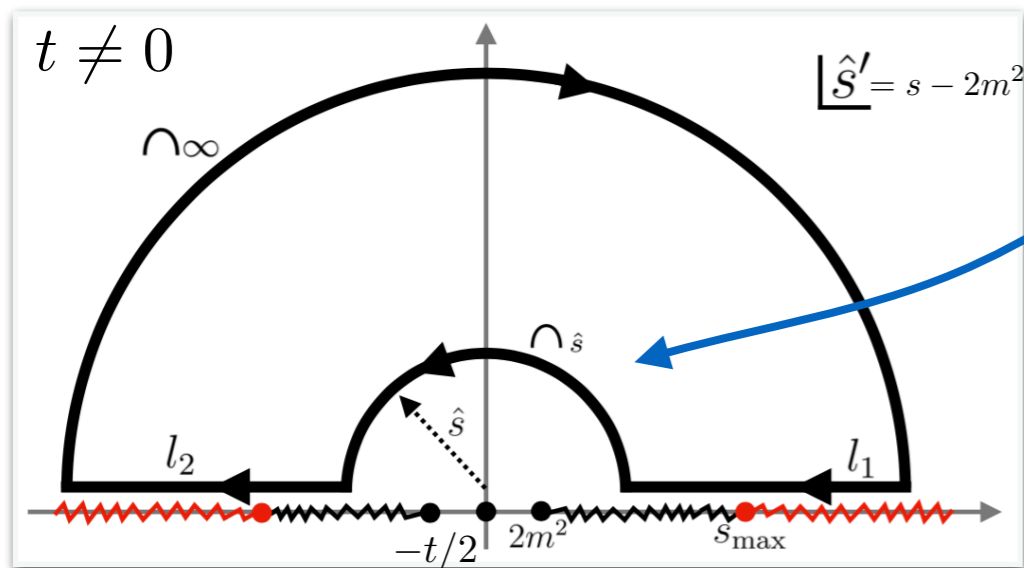
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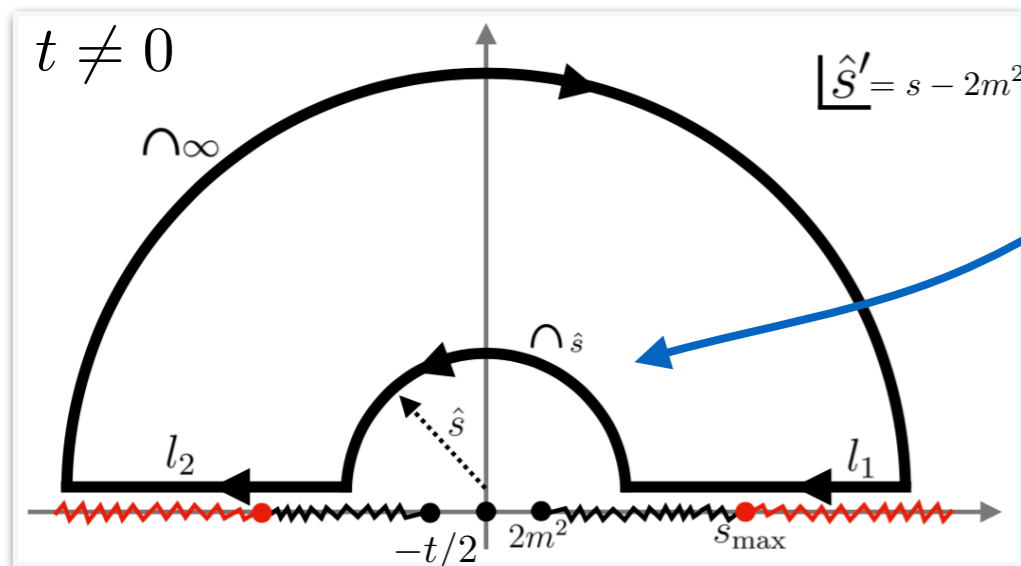
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$$\boxed{-\frac{3}{2} < \frac{c_{21} \hat{s}}{c_2} < 5.3}$$

2011.02957 S. Caron-huot, van Duong

2011.02400 A. Tolley, Z. Wang, S.Y. Zhou

2102.08951 Caron-huot, Mazac, Rastelli, Simmons-Duffin

2112.12561 B.B., Riembau, Riva

Huang, Rodina, Zhiboedov, Arkani-Hamed, Pomarol.... etc

$|\text{INELASTIC}| < \text{ELASTIC}$

&

FINITE-T AND CROSSING

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**Massive particle crossing: very complicated**

$(2J+1)^4$  helicities mix!

$$\mathcal{M}_{\lambda_1 \lambda_2}^{\lambda_3 \lambda_4}(s, t) = \sum_{\lambda'_i = -S}^S X_{\lambda_1 \lambda_2 \lambda_3 \lambda_4}^{\lambda'_1 \lambda'_2 \lambda'_3 \lambda'_4}(s, t) \mathcal{M}_{\lambda'_1 \lambda'_4}^{\lambda'_3 \bar{\lambda}'_2}(u, t)$$

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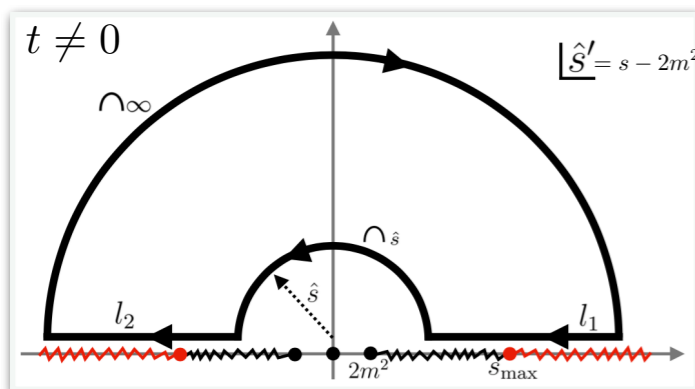
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EFTs with gap  $M$  to new states  $m^2 \ll |t| \ll s \lesssim M^2$

Crossing nice, error bounded by EFT-calculable quantity

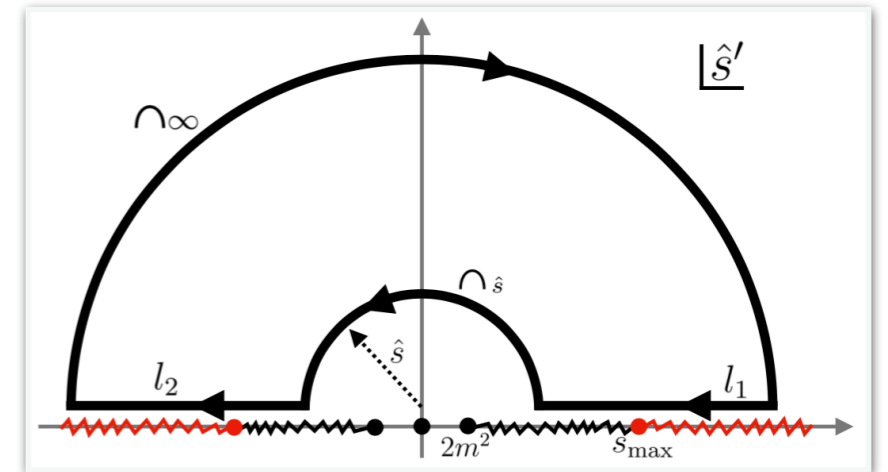
$$|\text{error}| < \frac{\sqrt{|t|m}}{M^2} c_{\lambda'_1 \lambda'_2}^{\lambda_1 \lambda_2} \mathcal{A}_{\lambda'_1 \lambda'_2}(0) / \mathcal{A}_{\lambda_1 \lambda_2}(0)$$



# | INELASTIC | < ELASTIC

$$A_{\lambda_1 \lambda_2}^{(2)} = \int_{M^2}^{\infty} \frac{ds}{2\pi} \frac{\langle 3^{\lambda_1} 4^{\lambda_2} | \mathcal{M}^\dagger \mathcal{M} | 1^{\lambda_1} 2^{\lambda_2} \rangle + \langle 3^{\lambda_1} 4^{\bar{\lambda}_2} | \mathcal{M}^\dagger \mathcal{M} | 1^{\lambda_1} 2^{\bar{\lambda}_2} \rangle}{(s - 2m^2 + t/2)^3} \quad t \neq 0$$

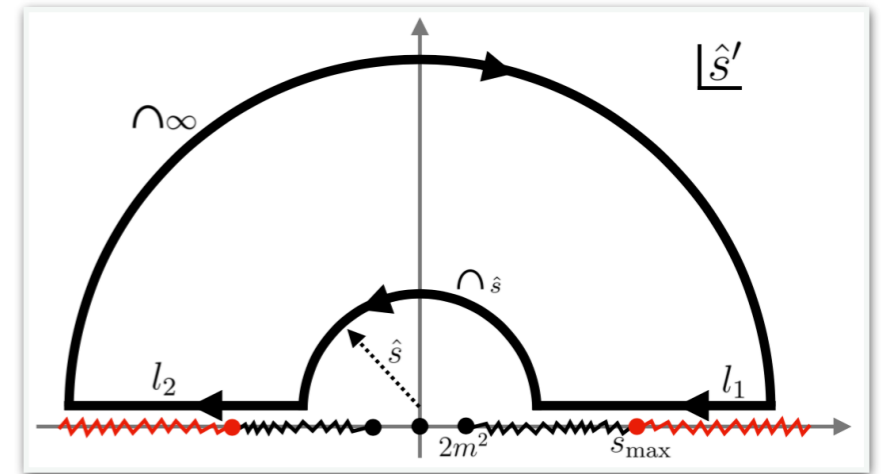
lack of positivity at t-finite?



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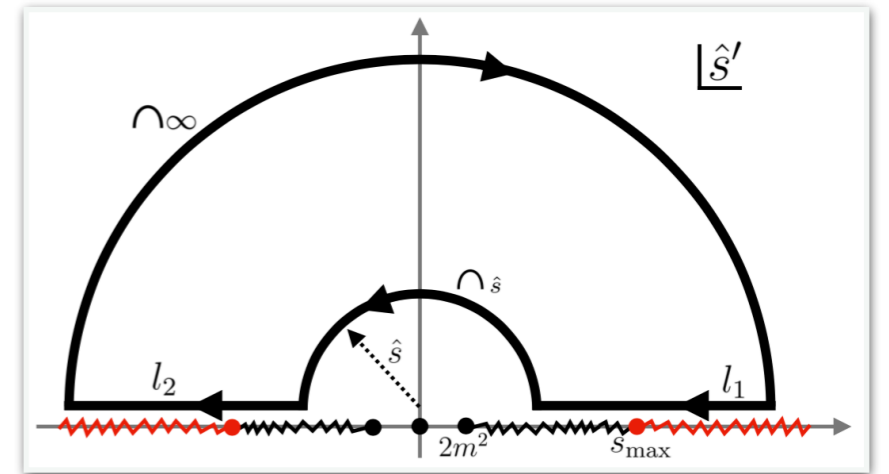
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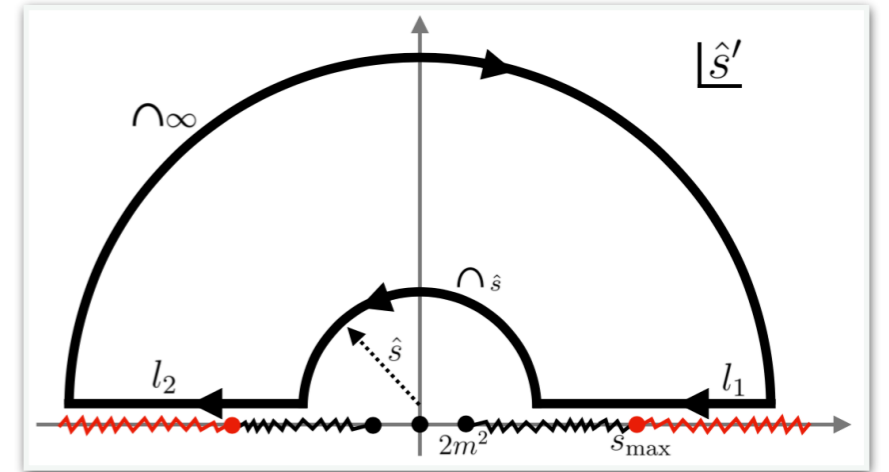
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$$\begin{array}{c}
 |\rightarrow\leftarrow\rangle \quad |\nearrow\swarrow\rangle \\
 \text{positive definite matrix} \\
 \mathcal{M}^\dagger \mathcal{M} \succ 0
 \end{array}
 \implies
 \left| \langle 3^{\lambda_1} 4^{\lambda_2} | \mathcal{M}^\dagger \mathcal{M} | 1^{\lambda_1} 2^{\lambda_2} \rangle \right| \leq \langle 1^{\lambda_1} 2^{\lambda_2} | \mathcal{M}^\dagger \mathcal{M} | 1^{\lambda_1} 2^{\lambda_2} \rangle$$

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positive definite matrix

finite-t arc bounded by t=0:

$$\frac{|A_{\lambda_1 \lambda_2}^{(2)}(t)|}{A_{\lambda_1 \lambda_2}^{(2)}(0)} \leq \frac{1}{\left(1 + \frac{t/2}{M^2 - 2m^2}\right)^3}$$

# BACK TO MASSIVE GRAVITY

$$\mathcal{L} = \sqrt{-g} \frac{m_{\text{Pl}}^2}{2} \left[ R - \frac{m^2}{4} (h^2 + \dots + c_3 h^3 + \dots + d_5 h^4) \right] \quad 5 \text{ d.o.f, } 2 \text{ free parameters}$$

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$$\begin{aligned} \mathcal{A}_{00} &\xrightarrow{m^2 \ll |t|} \frac{t}{6\Lambda_3^6} (1 - 4c_3 + 36c_3^2 + 64d_5) \\ \mathcal{A}_{0+} &\xrightarrow{m^2 \ll |t|} \frac{t}{96\Lambda_3^6} (1 + 24c_3 + 144c_3^2 + 384d_5) \\ \mathcal{A}_{++} &\xrightarrow{m^2 \ll |t|} \frac{9t}{64\Lambda_3^6} (1 - 4c_3)^2, \end{aligned}$$

**vs**

$$\begin{aligned} \mathcal{A}_{00} &\xrightarrow{t=0} \frac{2m^2}{9\Lambda_3^6} (7 - 6c_3 - 18c_3^2 + 48d_5) \\ \mathcal{A}_{0+} &\xrightarrow{t=0} \frac{m^2}{48\Lambda_3^6} (91 - 312c_3 + 432c_3^2 + 384d_5) \\ \mathcal{A}_{++} &\xrightarrow{t=0} \frac{m^2}{8\Lambda_3^6} (7 - 24c_3^2 + 48d_5). \end{aligned}$$

# BACK TO MASSIVE GRAVITY

$$\mathcal{L} = \sqrt{-g} \frac{m_{\text{Pl}}^2}{2} \left[ R - \frac{m^2}{4} (h^2 + \dots + c_3 h^3 + \dots + d_5 h^4) \right] \quad 5 \text{ d.o.f, 2 free parameters}$$

$$h_{\mu\nu} \rightarrow h_{\mu\nu}^{(\pm 2)} + \partial_{(\mu} A_{\nu)}^{(\pm 1)} + \partial_\mu \partial_\nu \pi^{(0)}$$

$$\Lambda_3 = (m^2 M_{\text{Pl}})^{1/3} \approx 10^{-13} \text{ eV} \approx (1000 \text{ km})^{-1}$$

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**TENSION!**

$$\frac{|\mathcal{A}_{\lambda_1 \lambda_2}(t)|}{\mathcal{A}_{\lambda_1 \lambda_2}(0)} \leq \frac{1}{\left(1 + \frac{t/2}{M^2 - 2m^2}\right)^3}$$

# BACK TO MASSIVE GRAVITY

$$\mathcal{L} = \sqrt{-g} \frac{m_{\text{Pl}}^2}{2} \left[ R - \frac{m^2}{4} (h^2 + \dots + c_3 h^3 + \dots + d_5 h^4) \right] \quad \text{5 d.o.f, 2 free parameters}$$

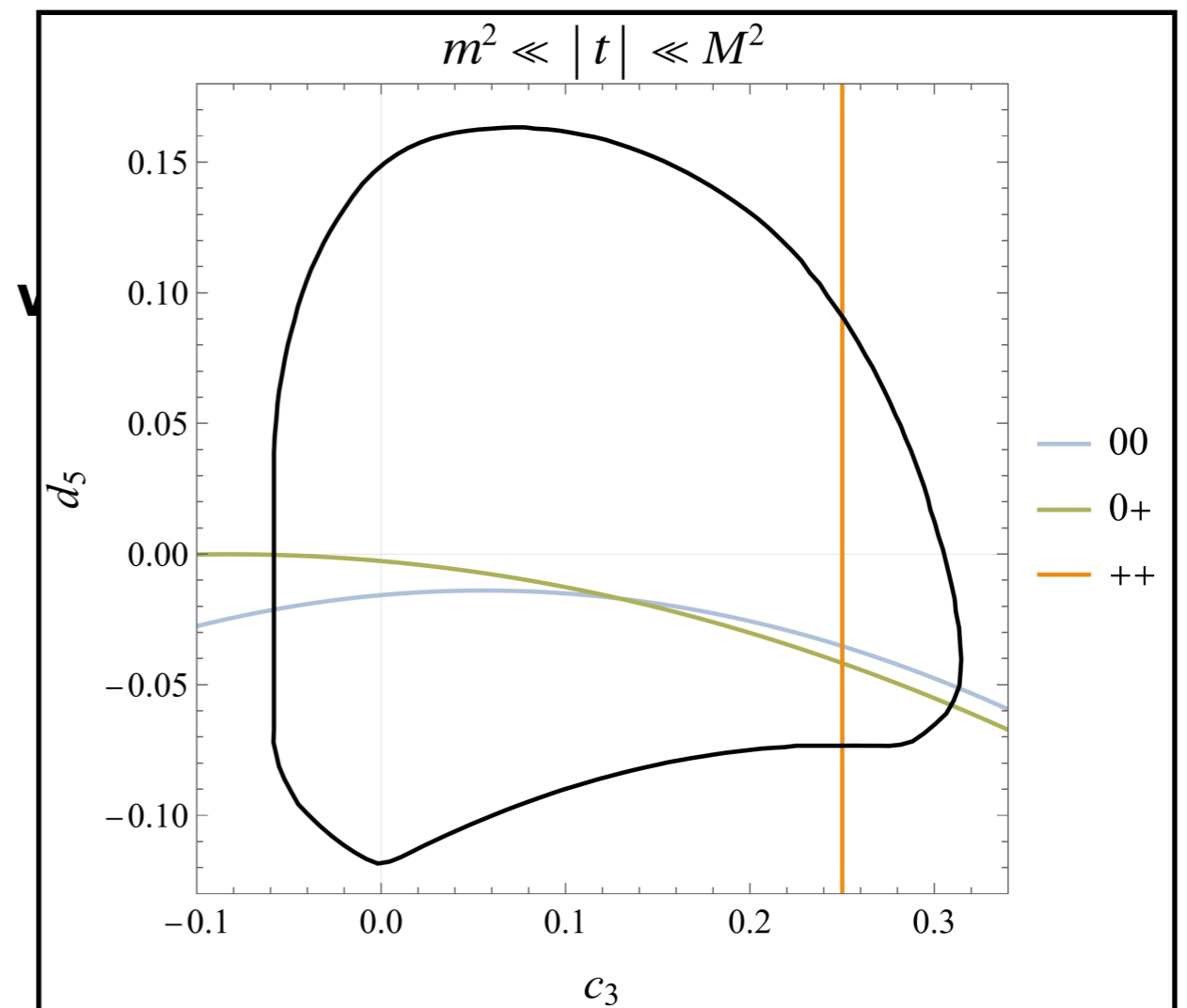
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**TENSION!**

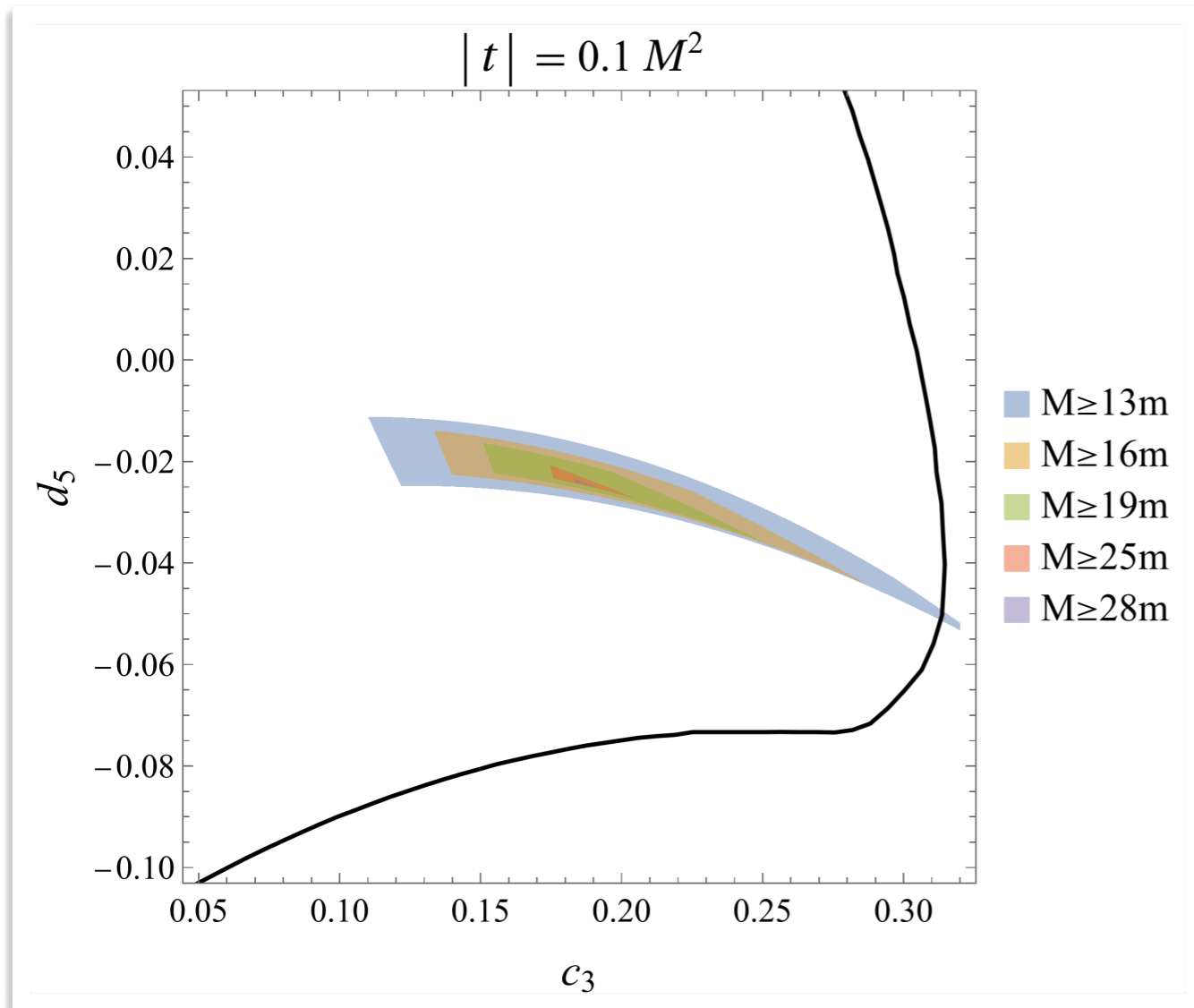
$$\left| \frac{\mathcal{A}_{\lambda_1 \lambda_2}(t)}{\mathcal{A}_{\lambda_1 \lambda_2}(0)} \right| \leq \frac{1}{\left( 1 + \frac{t/2}{M^2 - 2m^2} \right)^3}$$



**no intersection** → **M bounded above by few m**  
(assuming analyticity in s cut-plane for |t| within EFT)



# MASSIVE GRAVITY CUTOFF



2304.02550 B.B., Isabella, Ricossa, Riva

$$\frac{|\mathcal{A}_{\lambda_1 \lambda_2}(t)|}{\mathcal{A}_{\lambda_1 \lambda_2}(0)} \leq \frac{1}{\left(1 + \frac{t/2}{M^2 - 2m^2}\right)^3}$$

no parametric separation with  $m$

$$\Rightarrow M < 30m \sim o(10)H$$

teeny tiny range of validity  
(60 orders worse than GR!)

no physical system in the universe where this EFT can be used

# ISOLATED $J > 2$ SPINS?

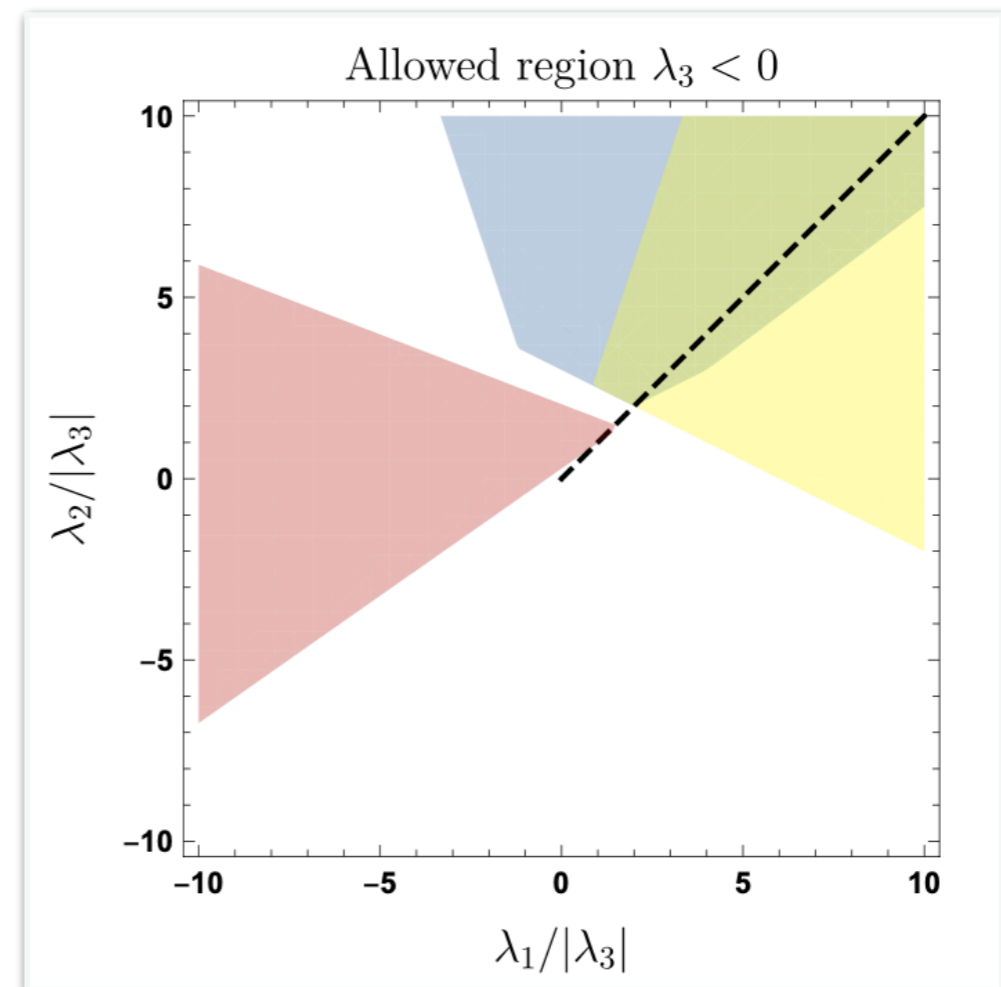
suffer of the same problem:  
longitudinal-helicities grow too fast with energy, either in s or t or both

$$\mathcal{M}_{\text{long.}}^{(J\text{-even})} = g^2(m_J/\Lambda_J) \left[ \left(\frac{E}{m_J}\right)^{3J} + \dots + \frac{E^4}{m_J^4} + \dots \right]$$

$$\mathcal{M}_{\text{long.}}^{(J\text{-odd})} = g^2(m_J/\Lambda_J) \left[ \left(\frac{E}{m_J}\right)^{3J+1} + \dots + \frac{E^4}{m_J^4} + \dots \right]$$

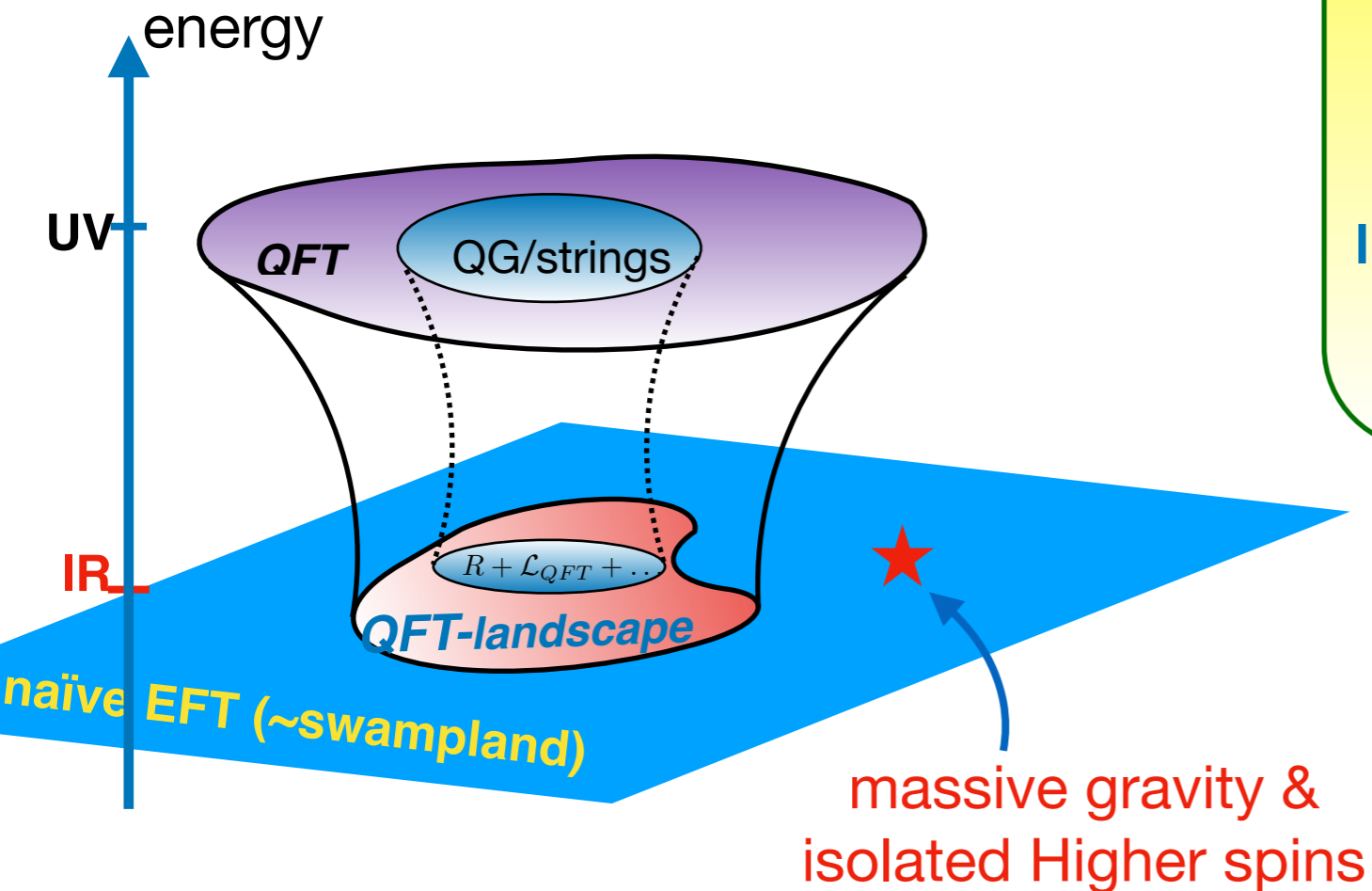
example  $J=3$ : fails positive already @  $t=0$

1903.08664 B.B., F. Riva, J. Serra, F. Sgarlata



# CONCLUSIONS

# CONCLUSIONS



- I. positivity bounds shape the swampland-landscape boundary
- II. isolated ( $M \gg m$ ) higher-spins can't be UV completed (corollary: massive gravity is not positive!)

## Open Question for Christmas:

do massive higher-spins only come in towers of infinitely many states ?  
e.g. KK tower, Regge states, QCD resonances,... (no parametric separation)

thank you!

# IMAGE CREDITS

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