

# SMALL THINKS BIG Transfer learning in KM3NeT/ORCA with transformers

15/10/2024, Caen Iván Mozún Mateo On behalf of the KM3NeT collaboration









# **Overview**



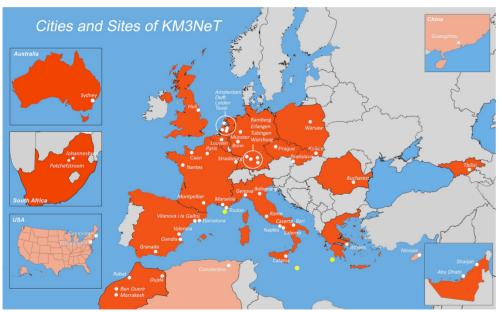
- 1. KM3NeT: neutrino telescopes
- 2. Need for data in large language models
- 3. Why transfer learning?
- 4. Multi-detector configuration and multi-task for KM3NeT/ORCA
- 5. Summary & The road ahead

# KM3NeT



#### KM3NeT is an **international collaboration**

- 22 countries
- 65 partner institutes
- ~250 members



#### Two undersea neutrino telescopes

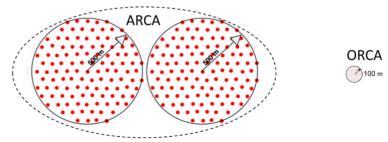
#### KM3NeT/ARCA

- Optimized for 1 TeV 10 PeV
- Identify high-energy neutrino sources in the Universe.
- 36m vertical spacing and 90m horizontal spacing

#### KM3NeT/ORCA

- Optimized for 1 100 GeV
- Determine the mass ordering of neutrinos.
- 9m vertical spacing and 20m horizontal spacing

Currently under construction: ORCA23 (20%), ARCA28 (12%)



# KM3NeT: neutrino telescopes

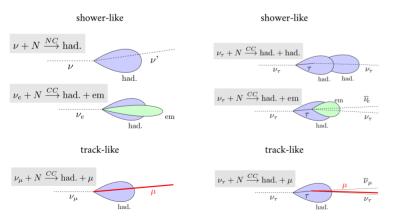


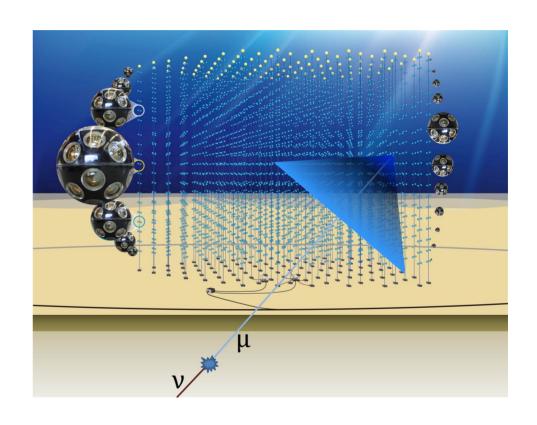
#### Same **technology**:

- 1 (2) building block(s) for ORCA (ARCA)
- 115 vertical detection units (DUs) per block
- 18 digital optical modules (DOMs) per DU
- 31" PMTs per DOM

#### Same **detection principle**:

Light collection from **Cherenkov radiation** emitted by particles traveling faster than the speed of light in water



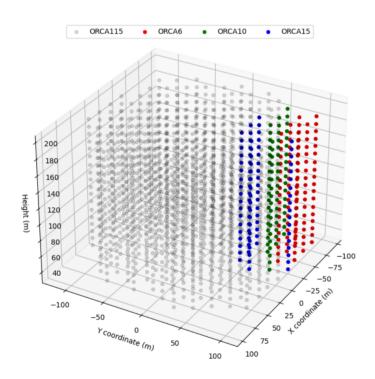


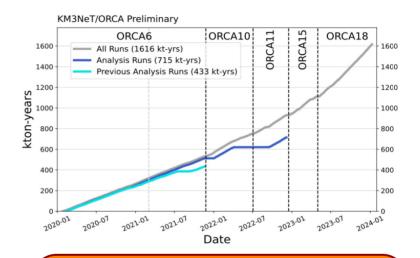
KM3NeT: Small thinks big

# **Building the detectors**



KM3NeT telescopes collect, process and analyze data as they are being built.





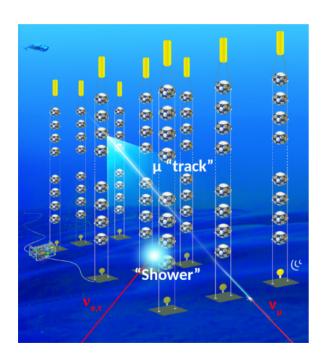
#### Foundation model in KM3NeT?

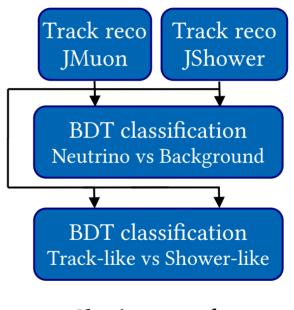
- learns as the detectors grow
- can handle multiple geometries
- can handle both KM3NeT/ORCA and KM3NeT/ARCA
- classification ↔ reconstruction

# Reconstructing neutrino physics



The **official KM3NeT pipeline for reconstruction and classification** relies on algorithms that are applied separately for track-like event or shower-like events. Then, simple BDTs are applied on the reconstructed variables for classification tasks.





Classic approach

# Reconstructing neutrino physics



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Track reco
JMuon

BDT classification
Neutrino vs Background

BDT classification
Track-like vs Shower-like

DL classification Neutrino vs Background

DL classification Track-like vs Shower-like

Track and shower reconstruction

**Deep Learning representation** 

Novel **deep learning** techniques use low-level information from the detector, i.e. light pulses to

- 1. Let the model decide the features to use
- 2. Generalise over a large input domain dimensions
- 3. Perform different tasks

Classification and reconstruction are performed **independently**, and for any type of event.

Large DL models needs **huge amounts of very diverse data** to generalize and interpolate, improving the performances of existing algorithms.

Classic approach

# **KM3NeT Deep Learning Outreach**



Various DL models tested. So far, no one is considered for official analysis.

#### **Convolutional Neural Networks**

- Event reconstruction for KM3NeT/ORCA using convolutional neural networks (M. Moser, KM3NeT)
- Event Classification and Energy Reconstruction for ANTARES using Convolutional Neural Networks (N. Geißelbrecht, ANTARES)
- Deep learning reconstruction in ANTARES (J. García-Méndez et al., ANTARES)
- Dark matter search towards the Sun using Machine Learning reconstructions of single-line events in ANTARES (J. García-Méndez et al., A NTARES)

#### **Deep Neural Networks**

• Deep Neural Networks for combined neutrino energy estimate with KM3NeT/ORCA6 (S. Peña Martínez, KM3NeT)

#### **Graph Neural Networks:**

- Development of detector calibration and graph neural network-based selection and reconstruction algorithms for the measurement of oscill ation parameters with KM3NeT/ORCA (D. Guderian, KM3NeT)
- Data reconstruction and classification with graph neural networks in KM3NeT/ARCA6-8 (F. Filippini et al., KM3NeT)
- Cosmic ray composition measurement using Graph Neural Networks for KM3NeT/ORCA (S. Reck, KM3NeT)
- Optimisation of energy regression with sample weights for GNNs in KM3NeT/ORCA (B. Setter, KM3NeT)
- Tau neutrino identification with Graph Neural Networks in KM3NeT/ORCA (L. Hennig, KM3NeT)

**More details here:** A Comprehensive Insight into Machine Learning Techniques in KM3NeT (J. Prado)

# Need for data in large language models

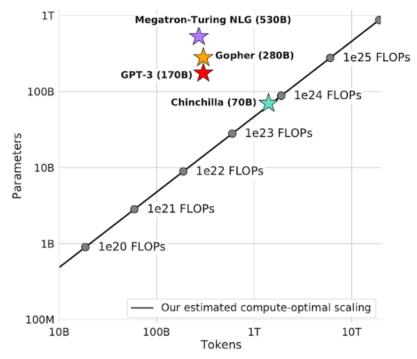


Neutrino telescope data is described as a set of spatial points with timing & charge information (point-cloud data), hence, most developed DL architectures are based on GNNs.

#### Language models are starting to overtake but...

- lot of trainable parameters
- lot of training data

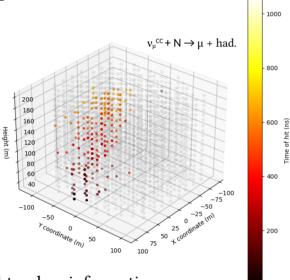
The data is too complex and requires a lot of computing resources to be produced and to encapsulate all the physics  $\rightarrow$  we must be efficient



Scaling law for trainable parameters and tokens for large language models arxiv.org/abs/2203.15556

# **Transformer architecture**

The input data is the low-level hit information that composes the light pattern detected in the telescope.

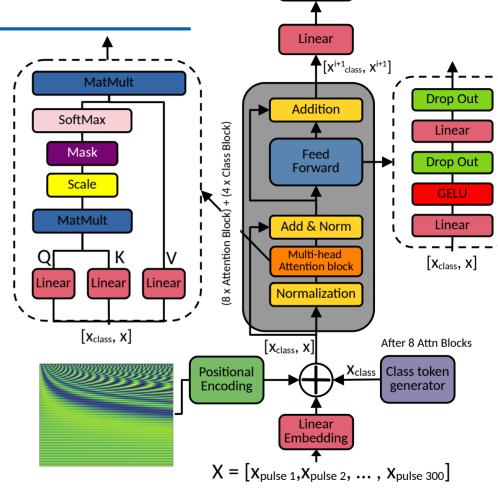


The light pulses information

$$X_{\text{pulse}} = [pos_x, pos_y, pos_z, dir_x, dir_y, dir_z, t, ToT]$$

is processed in parallel by the transformer and the highlevel information is extracted in the attention blocks.

Model has ~1.6M trainable parameters.



Prediction

# **Transfer learning studies**



#### Multiple tasks with a single model

- Classification and reconstruction done together
- Test the capacity of the model

#### Efficient use of data

- Run-by-Run: simulates MC runs based on data runs to reduce discrepancies
- Not enough data to train large models for every time the detector response is updated

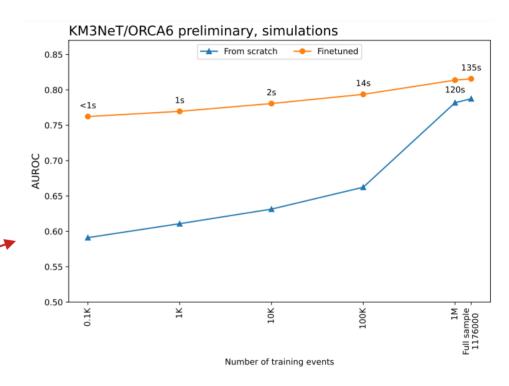
#### Missing detector information

- PMTs do not exist.
- PMTs correspond to DUs not deployed yet.

#### Efficient use of computing resources

- Saves time and increases performance
- The information is propagated across detectors

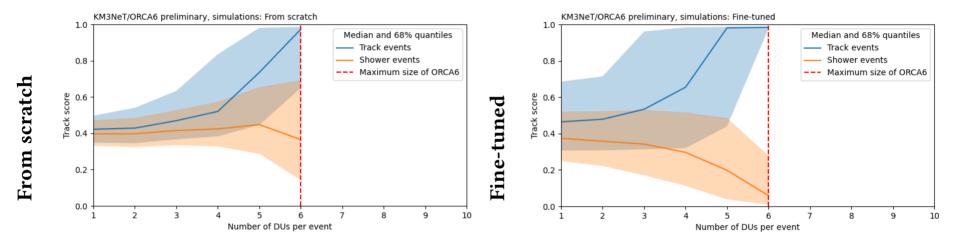




AUROC value for track-shower classification with KM3NeT/ORCA6 data. The AUROC curves are shown as function of the training size sample.



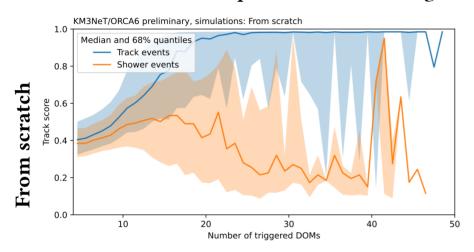
The model is able to **interpolate** to **non-existing DUs information** because it pre-learned the full geometry.

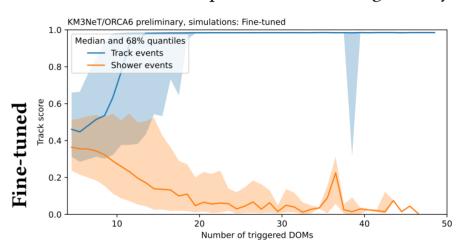


- Limited data: 200k events for a detector with 6 lines is not enough to do a proper separation
- Performance: fine-tuned model works way better than the scratch one
- High dependence on event geometry: not enough discrimination with few lines



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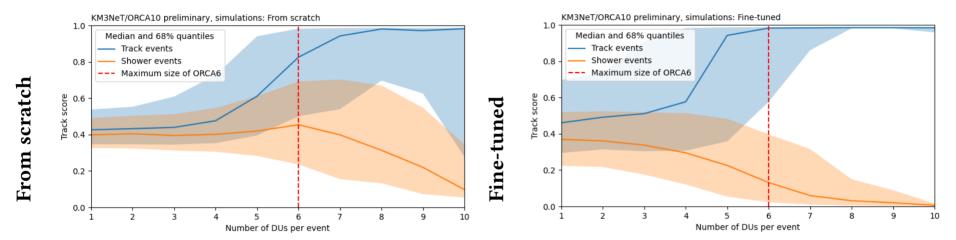




- Limited data: 200k events for a detector with 6 lines is not enough to do a proper separation
- Performance: fine-tuned achieves separation in events with above 10 triggered DOMs
- Peak at ~40 triggered DOMs: fine-tuned model compensates the low statistics



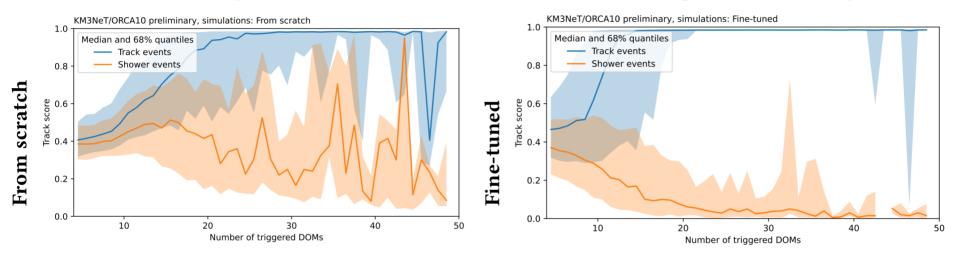
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- Limited data: 200k events for a detector with 6 lines is not enough to do a proper separation
- Performance: fine-tuned model works way better than the scratch one
- High dependence on event geometry: not enough discrimination with few lines
- Major improvement with increasing detector size ↔ better event containment



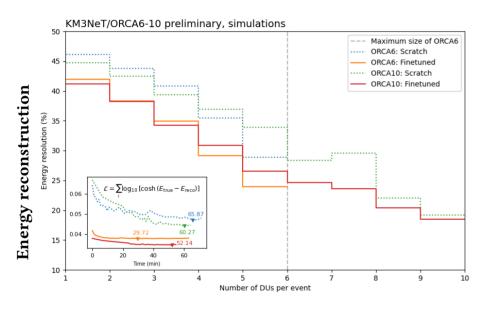
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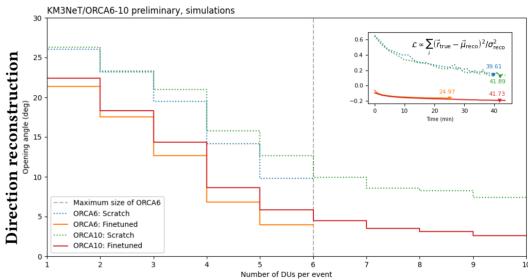


- Limited data: 200k events for a detector with 6 lines is not enough to do a proper separation
- Performance: fine-tuned achieves separation in events with above 10 triggered DOMs
- Multiple peaks: fine-tuned model compensates the low statistics



The model is able to **interpolate** to **non-existing DUs information** because it pre-learned the full geometry.





#### **Energy and direction reconstruction:**

- Loss curves reveal fine-tuning's performance boost
- Similar resolution in energy reconstruction, but in less time!
- Direction reconstruction improved significantly

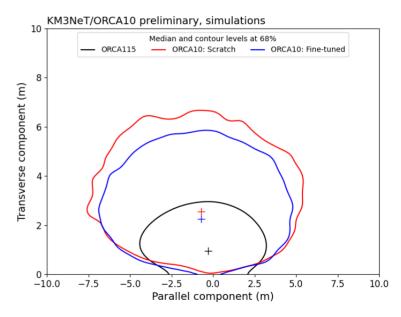


#### **Interaction reconstruction vertex:**

- The hardest task
- Dynamic detector coordinates from rbr approach
- Small fiducial volume burdens the reconstruction

Parallel component (m)

Better data representation?
Build a latent space to effectively accommodate dynamic coordinates and detector conditions & geometries

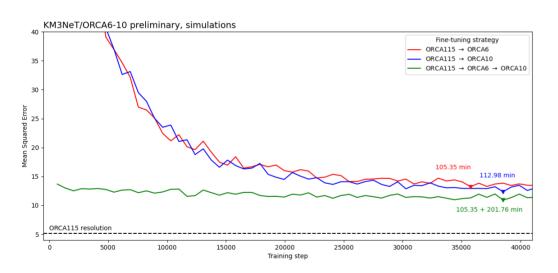


Interaction vertex reconstruction at KM3NeT/ORCA6 (left) and KM3NeT/ORCA10 (right) projected over the neutrino direction for 1-100 GeV atmospheric neutrinos.



#### **Interaction reconstruction vertex:**

- The hardest task
- Dynamic detector coordinates from rbr approach
- Small fiducial volume burdens the reconstruction

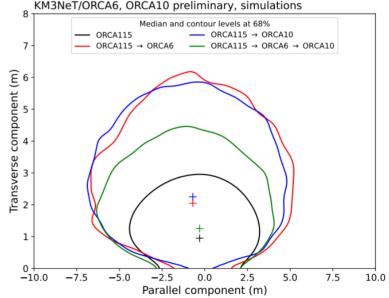


#### From big to small...

 $ORCA115 \rightarrow ORCA6$ , ORCA10, etc.

#### ...and vice versa

 $ORCA115 \rightarrow ORCA6 \rightarrow ORCA10 \rightarrow etc.$ 

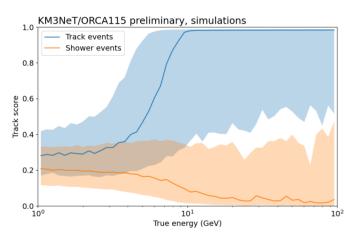


We need to propagate the knowledge between detectors!



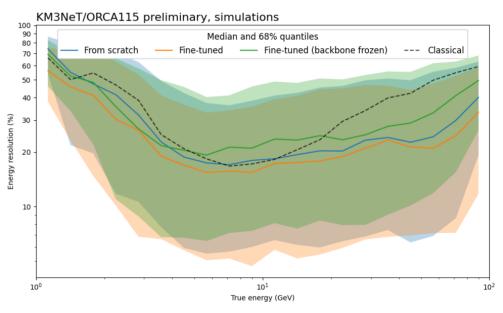
#### Classification to energy reconstruction

- A model that can handle multiple tasks
- A dataset 1 unrelated to dataset 2 (different detectors or water properties or atmospheric muons), helps a model into performing another task and makes it more robust



Track-score in function of neutrino energy in KM3NeT/ORCA115 for **dataset 1** 

# Multi-task study ORCA115 dataset 1: track-shower ORCA115 dataset 2: energy 850k tracks & 850k showers each

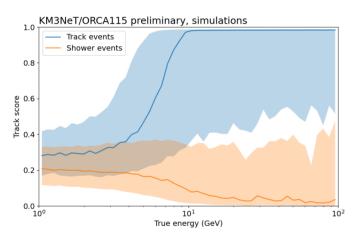


Energy resolution in function of neutrino energy in KM3NeT/ORCA115 for tracks from **dataset 2** 



#### Classification to energy reconstruction

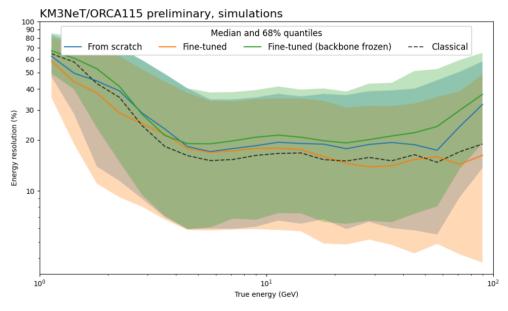
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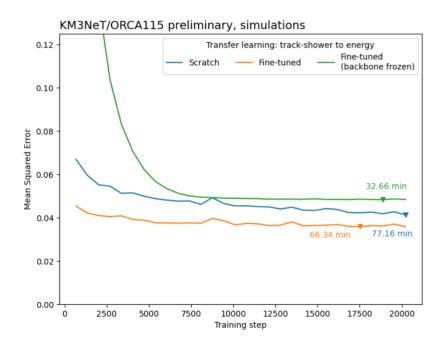
Energy resolution in function of neutrino energy in KM3NeT/ORCA115 for showers from **dataset 2** 



#### Classification to energy reconstruction

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Fine-tuned model shows **faster convergence and efficiency**.

Freezing the backbone has a **trade-off** between training speed-up and accuracy → suboptimal feature representation

Overall, the three cases achieve show improvements with respect to classical reconstruction methods.

# **Summary**



#### Transfer Learning in multiple-detectors

- Transformers are particularly effective to deal with small detectors and very limited data
- Further optimization is still needed in vertex reconstruction

#### **Transfer Learning for multi-task**

- Speeds up training and boosts model robustness
- Leverages knowledge from different tasks

#### The road ahead

- From simulations to data: ensure consistency and accuracy when transitioning to real detector data
- Robustness tests & uncertainties: validate model reliability across different conditions and detectors
- Estimate improvements as the detector grows to optimize scalability
- Develop **common benchmark** with state-of-the-art models  $\rightarrow$  On the way, see Jorge's talk
- Implement any deep learning reconstruction in the **official data processing pipeline** → Almost there
- Start testing pre-training models (BERT-like, GPT-like) with neutrino telescope data

# Thank you for your attention!



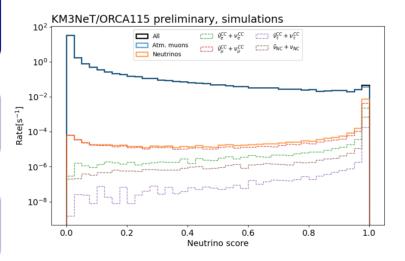
**Motivation:** the transformer is a language model

- KM3NeT/ORCA115 is the final detector, having all the possible neutrino physics encapsulated
- We can think of other configurations as similar languages to learn
- The information about KM3NeT/ORCA115 is used to understand our current detector

DL classification Neutrino vs Background

DL classification Track-like vs Shower-lik

Track and shower reconstruction



Event rate for neutrino score (0 for atmospheric muons, 1 for neutrinos).

**Purpose:** reject background data (atm. muon) from neutrino signal.

Atmospheric muons are more energetic, having their starting & ending points in most of the cases, out of the fiducial volume.

The model easily isolates **neutrino events** as they are mostly **fully contained** in the detector.



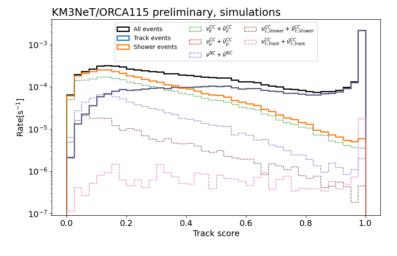
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Event rate for track score (0 for showers, 1 for tracks) for 1-100 GeV atmospheric neutrinos.

**Purpose:** separate the two neutrino event topologies, track-like and shower-like.

Enough separation power below 10 GeV (AUROC = 0.82)

High separation power above 10 GeV (AUROC = 0.91).

Low energy events do not contain enough pulses to properly separate these two categories



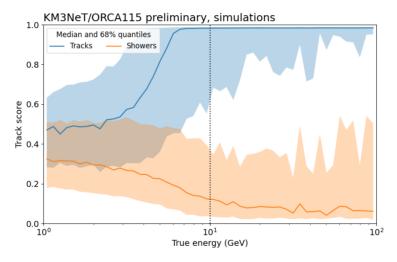
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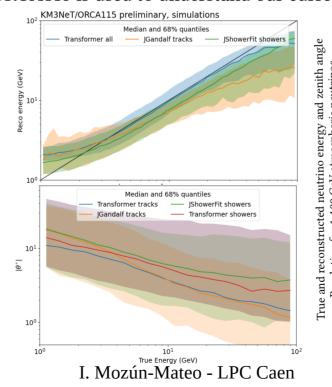
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Track and shower reconstruction



**Purpose:** reconstruct neutrino energy and neutrino direction.

Reconstruction done simultaneously for both track-like and shower-like events.

Saturation at high energies due to event containment.

Underestimation at low energies due to limited number of pulses.



Thanks to the modular structure of GraphNeT: different detector configuration are easily handled.

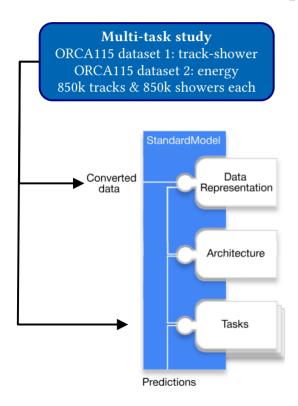
```
Multi-detector study
                ORCA6 (Feb20 - Nov21): 4075 runs
               ORCA10 (Dec21 – May22): 1889 runs
                          100k tracks & 100k showers
                    configuration = detector config['path'],
                                         detector_config['selection']['ORCA115'],
detector_config['selection']['ORCAX'],
                    shift coordinates = True
data_definition = HitsSequence(
                                detector = detector,
                                node definition = NodesAsHitsTimeSeries(
                                                                                input feature names = features.
                                                                               max_hits = config["data"]["max_hits"],
trig name = config["data"]["trig name"
dataset = SQLiteDataset(
                            path = detector_config['path'],
truth_table = config['data']['truth_table_name'],
pulsemaps = config['data']['pulsemap'],
                             features = features,
                            graph_definition = data_definition
                             selection = config['selection'],
rain dataset, val dataset = train test split(
                                                  train size = 1 - config['data']['validation size'],
                                                 test_size = config['data']['validation_size'],
random state = config['training']['seed'],
```

```
ass ORCA(Detector):
  """ Detector class for ORCA."""
     configuration: Optional[str],
     du_selection: Optional[Tuple[List[int], List[int]]],
     shift coordinates: Optional[bool] = True,
     self.configuration = os.path.join(KM3NeT GEOMETRY TABLE DIR. configuration)
     self.du selection ORCAll5 = du selection[0]
     self.du selection ORCAX = du selection[1]
     self.shift = shift coordinates
         KM3NeT GEOMETRY TABLE DIR, "ORCA115.parquet"
     if self.shift:
  geometry table path = os.path.join(
         KM3NeT GEOMETRY TABLE DIR. "ORCA115.parquet"
  string id column = "DU id"
  floor id column = "floor id"
  sensor id column = "dom id"
  def shift to ORCA115(self):
     ORCAll5 df = pd.read parquet(self.geometry_table_path)
     ORCAX df = pd.read parquet(self.configuration)
     ORCA115 DUs = ORCA115 df[ORCA115 df['DU id'].isin(self.du selection ORCA115)]
     ORCAX DUs = ORCAX df[ORCAX df['DU id'].isin(self.du selection ORCAX)]
     x shift = ORCAX DUs['pos x'].mean() - ORCA115 DUs['pos x'].mean()
     y shift = ORCAX DUs['pos y'].mean() - ORCA115 DUs['pos y'].mean()
     z shift = ORCAX DUs['pos z'].mean() - ORCA115 DUs['pos z'].mean()
     return (x shift, y shift, z shift)
```

```
eature map(self) -> Dict[str, Callable]:
    ""Map standardization functions to each dimension of input data."""
   feature map = {
   return feature map
def dom x(self, x: torch.tensor) -> torch.tensor:
   if self.shift:
def dom y(self, x: torch.tensor) -> torch.tensor:
   if self shift:
def dom z(self, x: torch.tensor) -> torch.tensor:
   return (x - 117.5) / 7.75
lef _dom_time(self, x: torch.tensor) -> torch.tensor:
   return (x - 1800) / 180
lef tot(self, x: torch.tensor) -> torch.tensor:
   return (x - 75) / 7.5
def _dir_xy(self, x: torch.tensor) -> torch.tensor:
lef dir z(self, x: torch.tensor) -> torch.tensor:
```



Thanks to the modular structure of GraphNeT: fine-tuning between tasks is as well possible.



```
backbone = Transformer(
                           seg length = config["backbone"]["seg length"].
                           n features = config["backbone"]["n features"].
                           position encoding = config["backbone"]["position_encoding"],
                           emb dims = config["backbone"]["emb dims"],
                           num heads = config["backbone"]["num heads"],
                           dropout attn = config["backbone"]["dropout attn"],
                           hidden dim = config["backbone"]["hidden_dim"],
                           dropout FFNN = config["backbone"]["dropout FFNN"].
                           no hits blocks = config["backbone"]["no hits blocks"],
                           no evt blocks = config["backbone"]["no evt blocks"],
task = BinaryClassificationTask(
                                   hidden size = backbone.nb outputs,
                                    target labels = config["task"]["target"],
                                    loss function = BinaryCrossEntropyLoss(),
model = StandardModel(
                        graph definition = data definition,
                        backbone = backbone.
                        tasks = [task].
                       optimizer class = AdamW.
                       optimizer kwargs = config["optimizer"]["parameters"],
                       scheduler class = None,
                       scheduler kwargs = None,
                        scheduler config = None.
  config['training']['fine tune']:
   backbone weights = torch.load(config["pretrained"])['state dict']
    model.load state dict(backbone weights, strict = False)
    if config['training']['freeze backbone']:
       for name, param in model.named parameters():
           if name.startswith('backbone'):
               param.requires grad = False
```