

AI GOES MAD²

Machine Learning in KM₃NeT

AI goes MAD-2

15 october 2024

Jorge Prado González (jprado@km3net.de)

Instituto de Física Corpuscular, Valencia, Spain

Work in collaboration with Iván Mozún Mateo



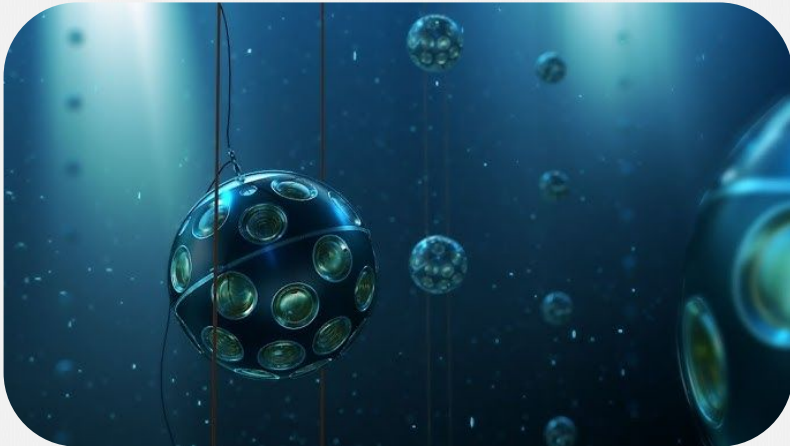
Index

- Neutrino telescope: KM3NeT.
- Machine Learning in KM3NeT.
- Model performance comparison.



PART I

Neutrino Telescope: KM3NeT

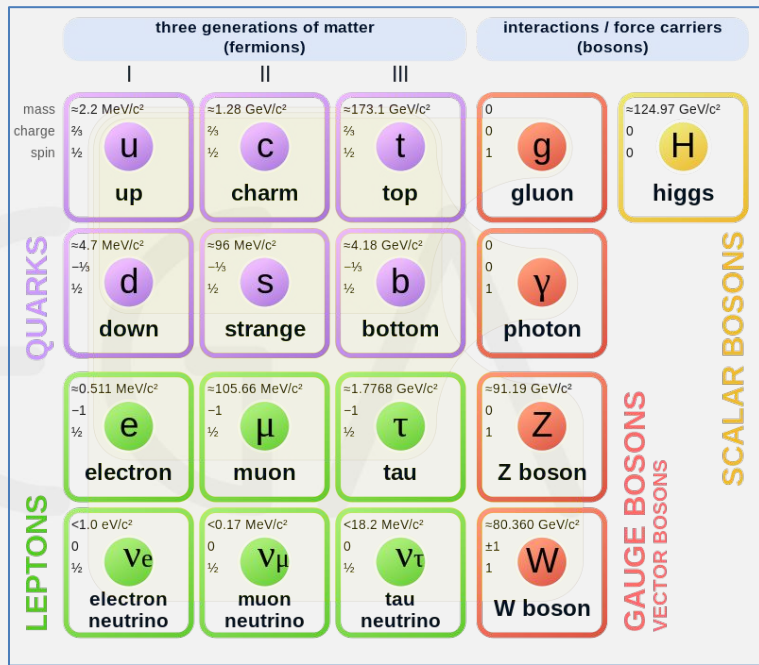


Neutrinos

- Neutrinos are **particles of the Standard Model**.
- Neutrinos are **the second most abundant particles** in the universe (after photons).
- They only interact via weak force with extremely small cross sections.
- Neutrinos can enlighten the path and help **answering some unsolved questions about particle physics**, as well as to **understand some astrophysical objects**.

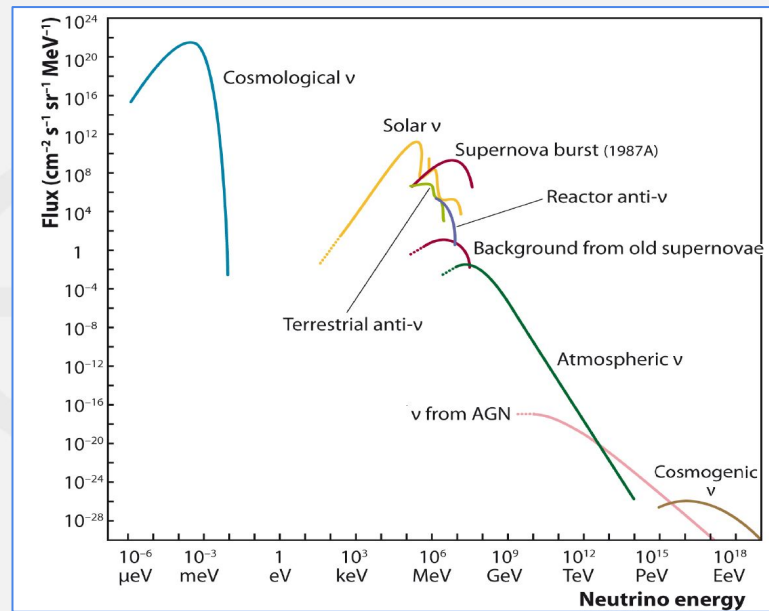


100 trillion ν/s

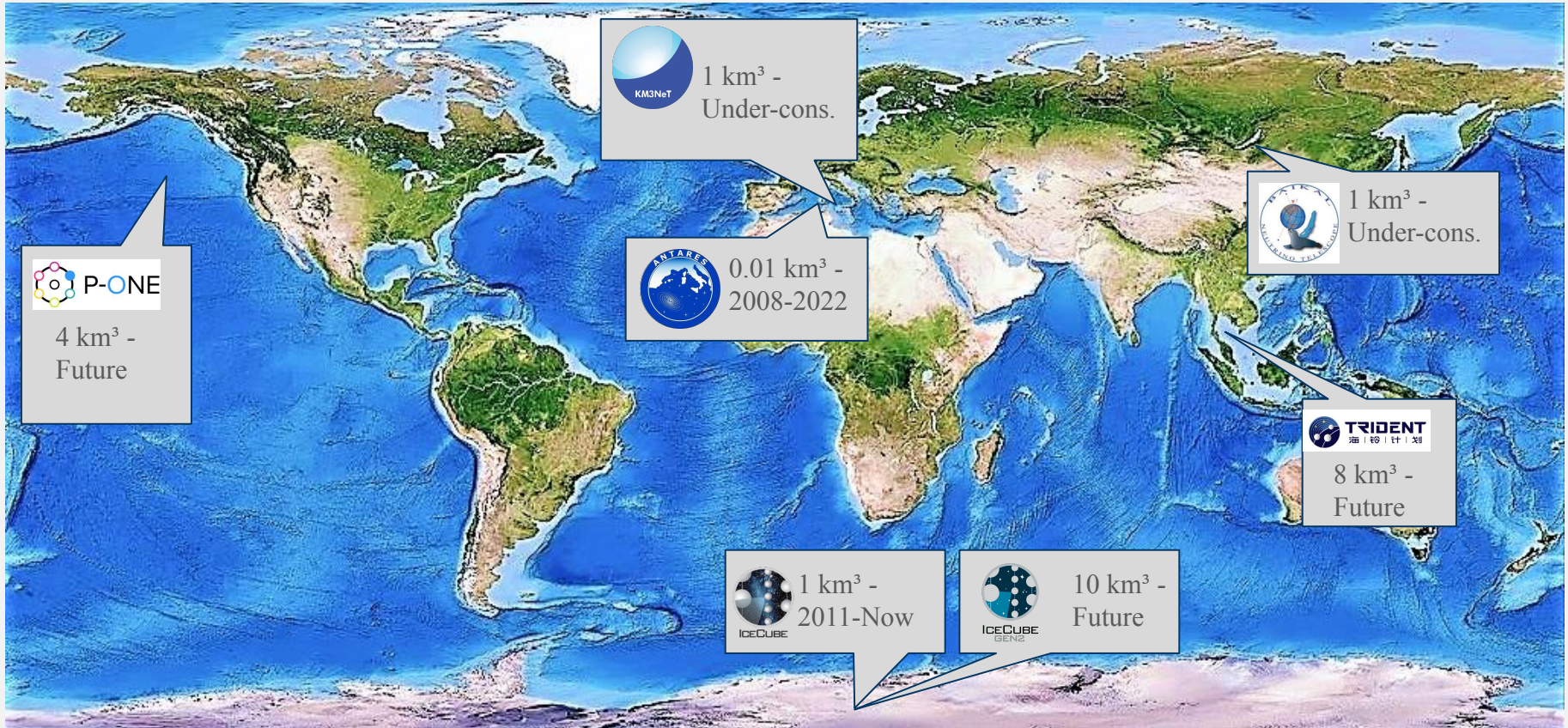


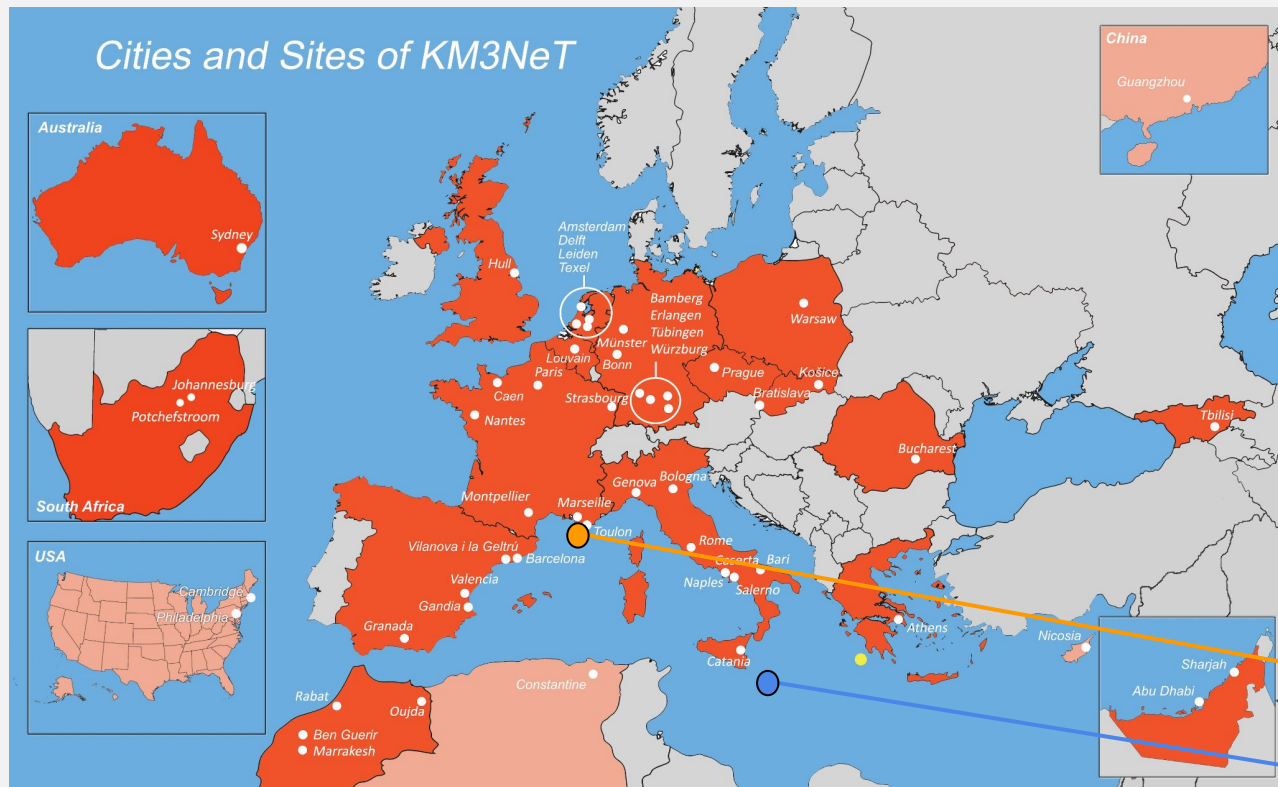
Neutrinos

- Neutrinos are **particles of the Standard Model**.
- Neutrinos are **the second most abundant particles** in the universe (after photons).
- They only interact via weak force with extremely small cross sections.
- Neutrinos can enlighten the path and help **answering some unsolved questions about particle physics**, as well as to **understand some astrophysical objects**.
- Neutrinos coming from very different sources:
 - The Sun.
 - Supernovae.
 - AGNs.
 - Nuclear reactors.
 - Atmospheric neutrinos.



Neutrino telescopes





- International collaboration with
 - ~250 members.
 - 65 partner institutes.
 - Over 22 countries.
- Two detectors in different sites: **KM3NeT/ORCA** and **KM3NeT/ARCA**:
 - Same technology.
 - Same data processing.
 - Same software and common dataformats.
 - Different size and granularity.

KM3NeT/ORCA

KM3NeT/ARCA

KM3NeT - ARCA and ORCA

• KM3NeT/ORCA:

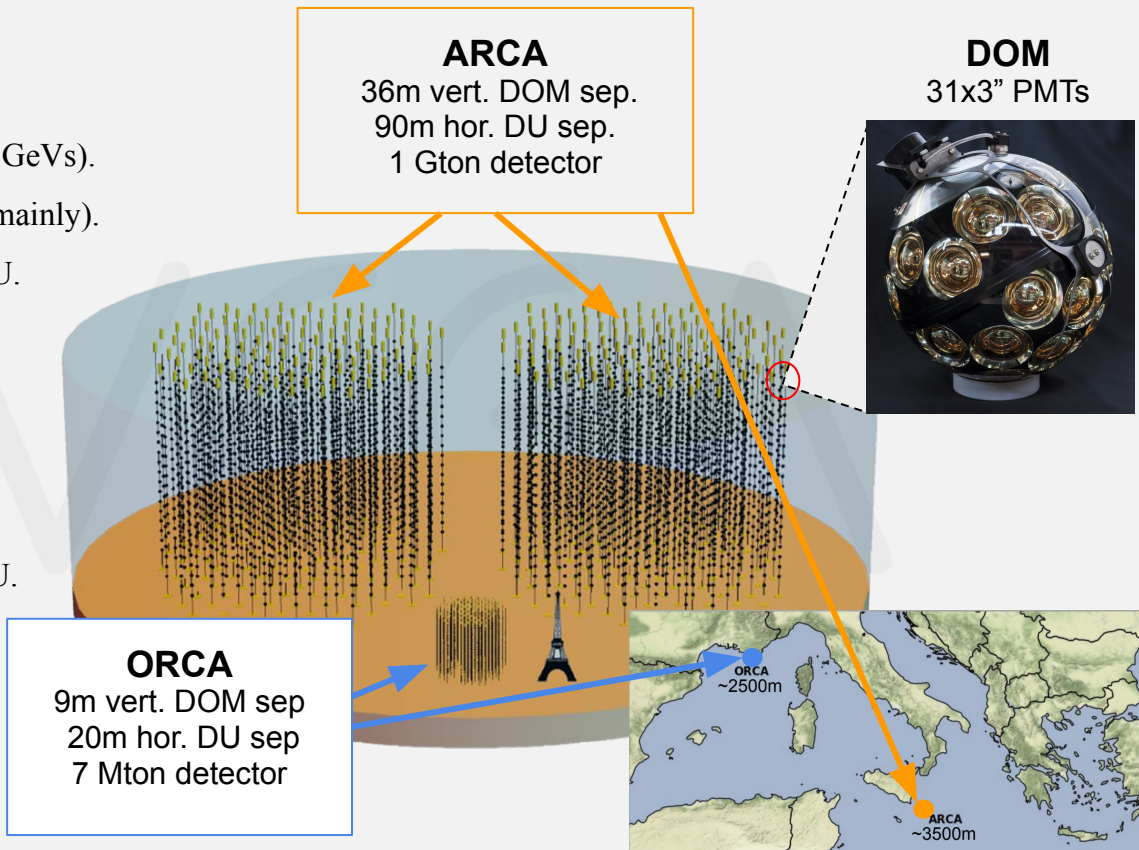
- Low energies (~few GeV to hundreds of GeVs).
- Fundamental neutrino property studies (mainly).
- **Full ORCA:** 115 DUs, 18 DOMs per DU.
- **Current ORCA:** 23 DUs deployed.

• KM3NeT/ARCA:

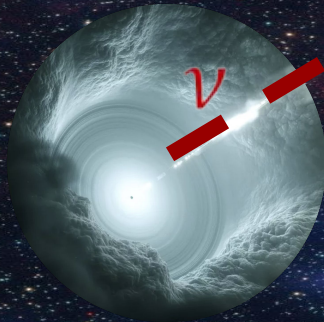
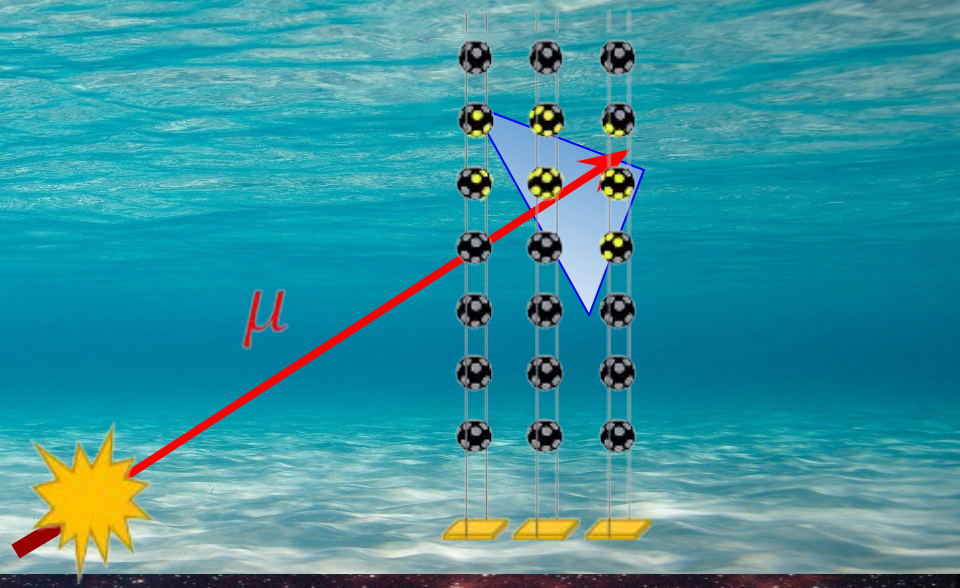
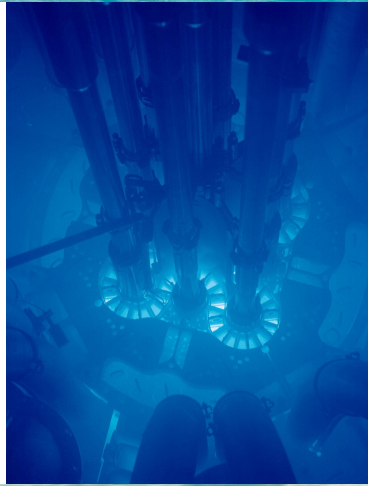
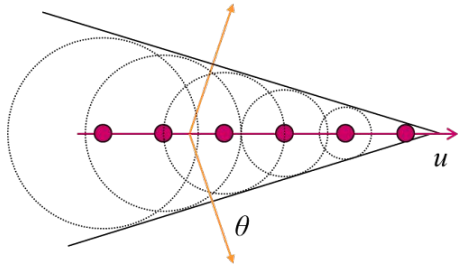
- High energies (sub-TeV to few PeV).
- Astrophysical studies (mainly).
- **Full ARCA:** 230 DUs, 18 DOMs per DU.
- **Current ARCA:** 28 DUs deployed.

DU: Detection Unit. String of 18 DOMs.

DOM: Digital Optical Module.



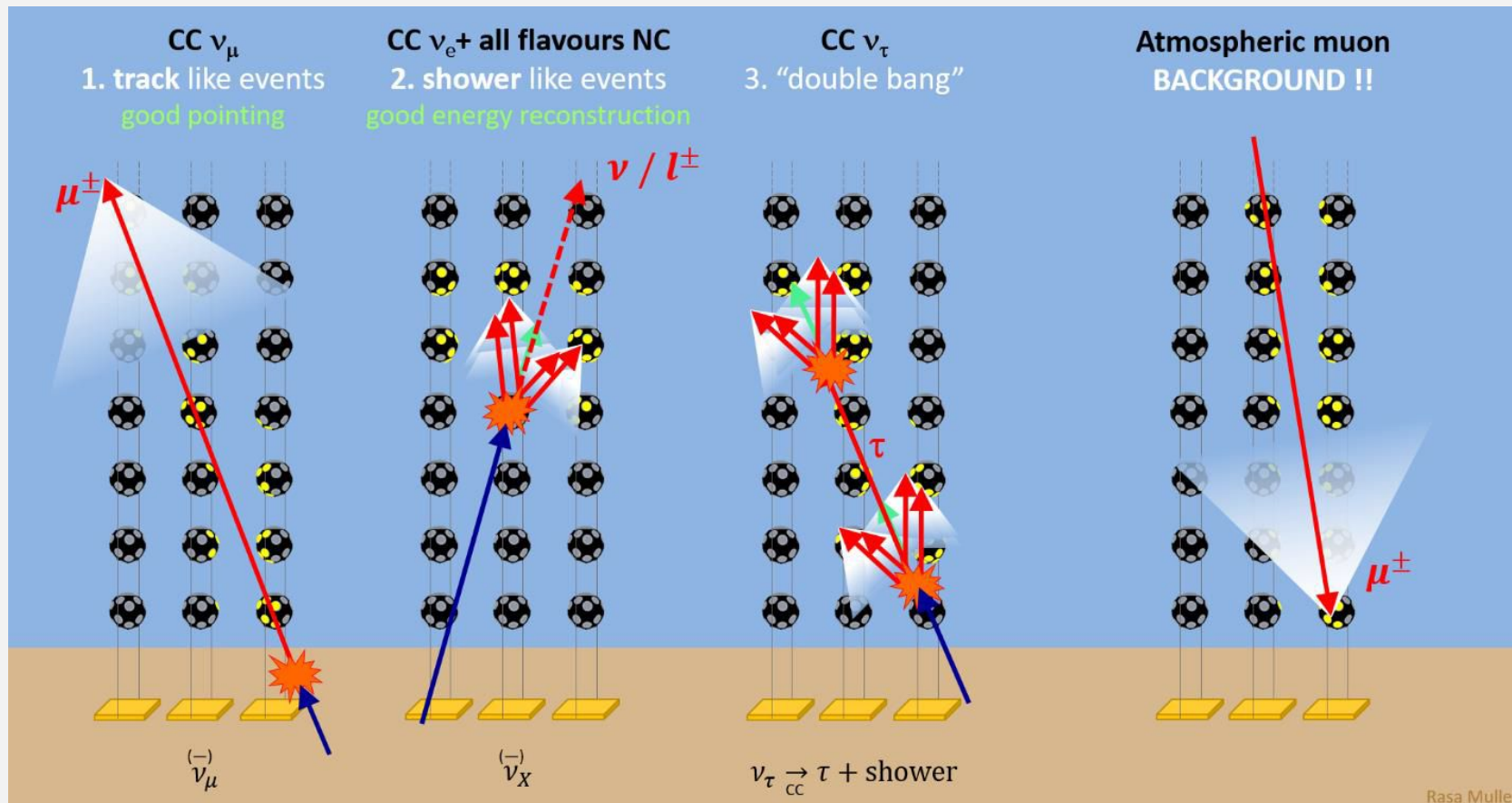
Cherenkov light



We can detect neutrinos!

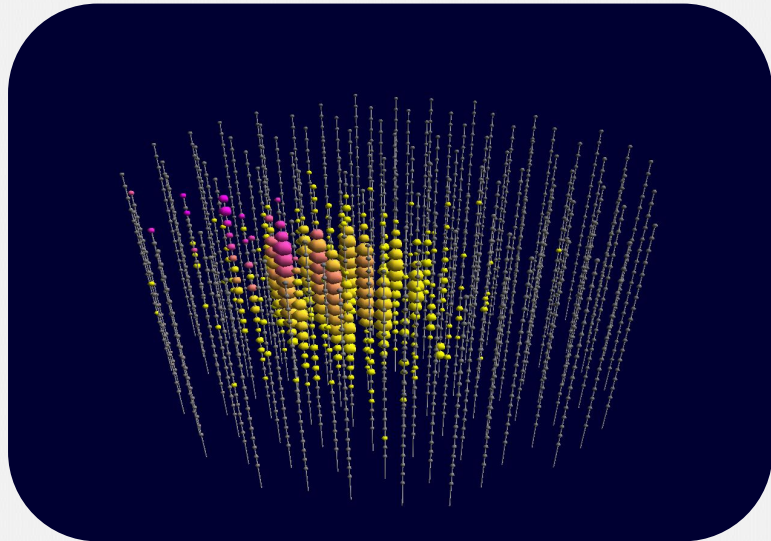
[Video](#)

KM3NeT - Detection principle. Event topology



PART II

Machine Learning in KM3NeT



Deep Learning Projects in KM3NeT

CNNs:

- [Event reconstruction for KM3NeT/ORCA using convolutional neural networks](#) (M. Moser, KM3NeT)
- [Event Classification and Energy Reconstruction for ANTARES using Convolutional Neural Networks](#) (N. Geißelbrecht, ANTARES)
- [Deep learning reconstruction in ANTARES](#) (J. García-Méndez et al., ANTARES)
- [Dark matter search towards the Sun using Machine Learning reconstructions of single-line events in ANTARES](#) (J. García-Méndez et al., ANTARES)

Fully-connected NNs:

- [Deep Neural Networks for combined neutrino energy estimate with KM3NeT/ORCA6](#) (S. Peña Martínez, KM3NeT)

Several different ML-based projects being already part of physics analyses (BDTs, RFs...):

- [ParamPID: t/s, nu/noise and nu/mu classifier with XGBoost](#) (A. Lazo & L. Maderer, KM3NeT)
- CR composition measurement: Atm. muon bundle reconstruction using RFs (P. Kalaczynski, KM3NeT)
- BoostTauID: identify GeV tau neutrinos in ORCA with XGBoost/ParamPID (N. Geißelbrecht, KM3NeT)

Deep Learning Projects in KM3NeT

GNNs:

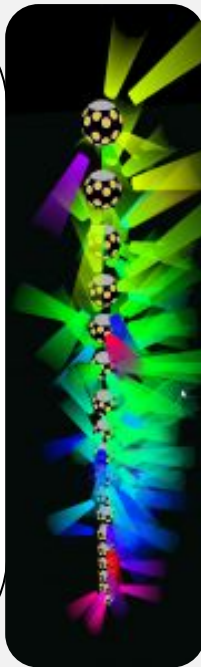
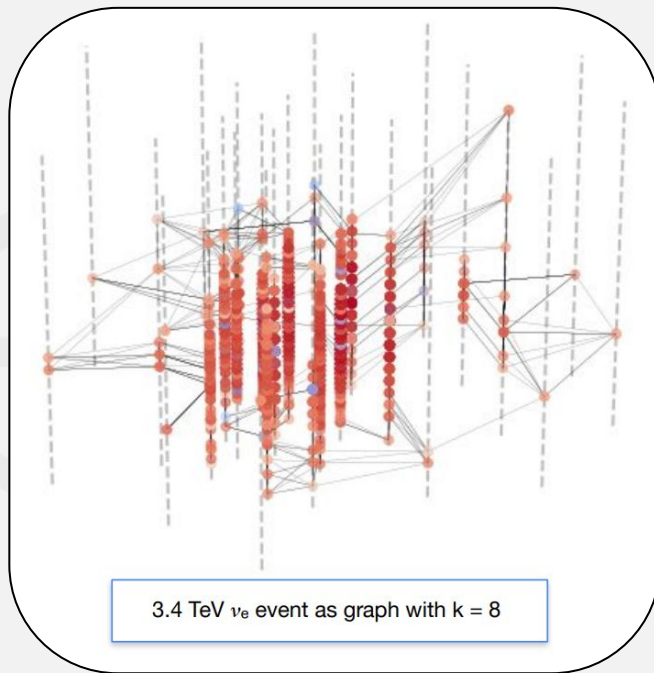
- [Development of detector calibration and graph neural network-based selection and reconstruction algorithms for the measurement of oscillation parameters with KM3NeT/ORCA](#) (D. Guderian, KM3NeT)
- [Data reconstruction and classification with graph neural networks in KM3NeT/ARCA6-8](#) (F. Filippini et al., KM3NeT)
- [Cosmic ray composition measurement using Graph Neural Networks for KM3NeT/ORCA](#) (S. Reck, KM3NeT)
- [Optimisation of energy regression with sample weights for GNNs in KM3NeT/ORCA](#) (B. Setter, KM3NeT)
- [Tau neutrino identification with Graph Neural Networks in KM3NeT/ORCA](#) (L. Hennig, KM3NeT)
- Energy reconstruction in ARCA21 using GNNs (E. Tragia, P. Gkotsis, E. Drakopoulou, KM3NeT)
- Particle ID classification, energy, direction and interaction vertex position reconstruction in KM3NeT/ORCA using Dynedg (J. Prado, KM3NeT)
- Neutrino Selection using GNNs for ARCA28 (A. Veutro, KM3NeT)
- Heavy neutral lepton signal identification using DYENDGE in KM3NeT/ORCA (J. Prado, KM3NeT)

Transformers/Foundation models:

- Transformer based classification and reconstruction in KM3NeT/ORCA (I. Mozún, KM3NeT)
- Transfer learning in KM3NeT/ORCA with transformers (I. Mozún, KM3NeT)

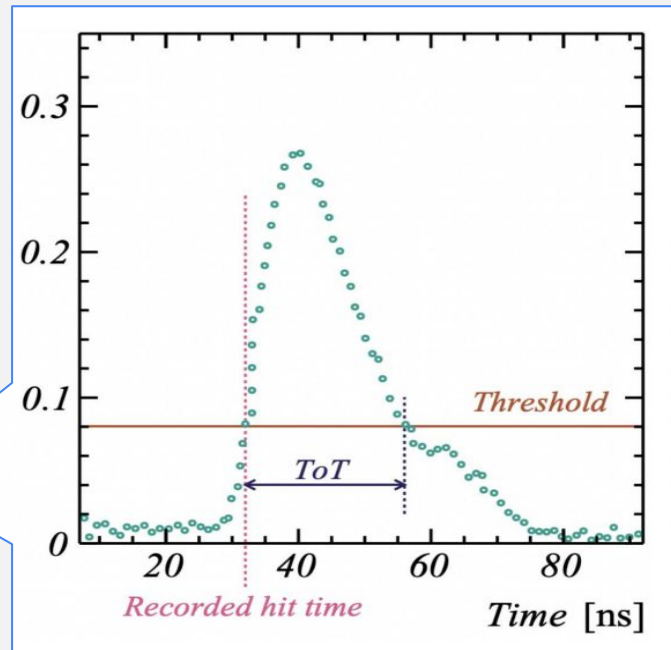
Representing data with graphs

- “It is preferable not to shape the problem to the tool, but the tool to the problem” [My ML professor].
- GNNs using graphs as input **capture the irregular geometries** of our events with **no underlying assumption of on the geometry**.
- Using **nodes** with inputs:
 - Position where the hit happens (x, y, z)
 - The direction of the PMT collecting the hit (dx, dy, dz)
 - The time when the hit happens (t)
 - The time that the PMT is collecting over 3 PE (**ToT**)
- Connected to its N-nearest neighbors in an euclidean space.



Representing data with graphs

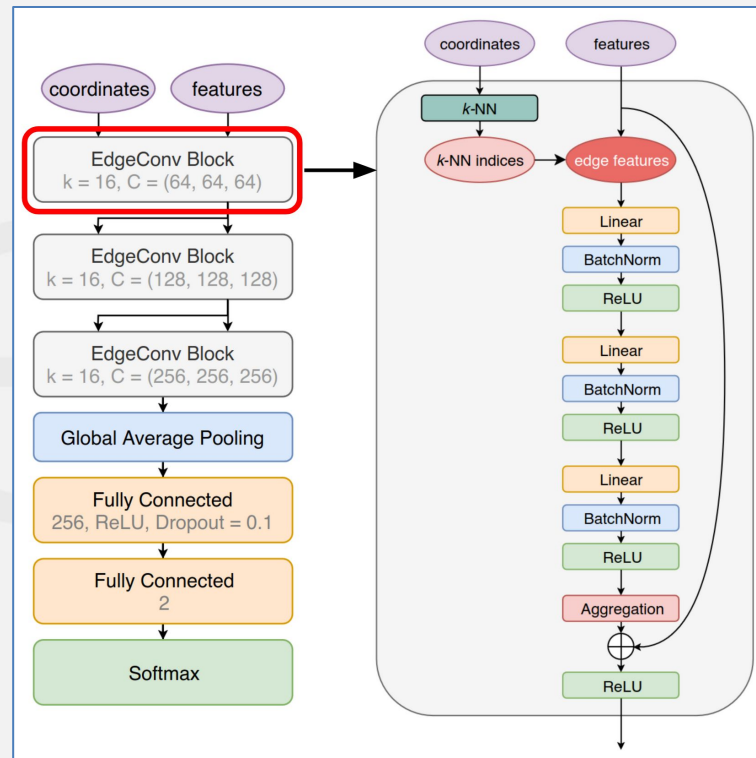
- “It is preferable not to shape the problem to the tool, but the tool to the problem” [My ML professor].
- GNNs using graphs as input **capture the irregular geometries** of our events with **no underlying assumption of on the geometry**.
- Using **nodes** with inputs:
 - Position where the hit happens (x, y, z)
 - The direction of the PMT collecting the hit (dx, dy, dz)
 - The time when the hit happens (t)
 - The time that the PMT is collecting over 3 PE (**ToT**)
- Connected to its N-nearest neighbors in an euclidean space.



ParticleNeT Model

- **ParticleNeT is a GNN** that was originally designed to jets at LHC.
- **Adapted and used in ORCA** under the name of ORCANEt since some years ago.
- **3 convolutional blocks** connected after to some **average pooling** to summarize the learned information and connect to a **fully connected perceptron** to learn more complex relations between node features.
- **Features a hierarchical structure**, early layers focused on learning low level features while deeper layers will capture more abstract, global patterns.
- Model with **521k trainable parameters***.

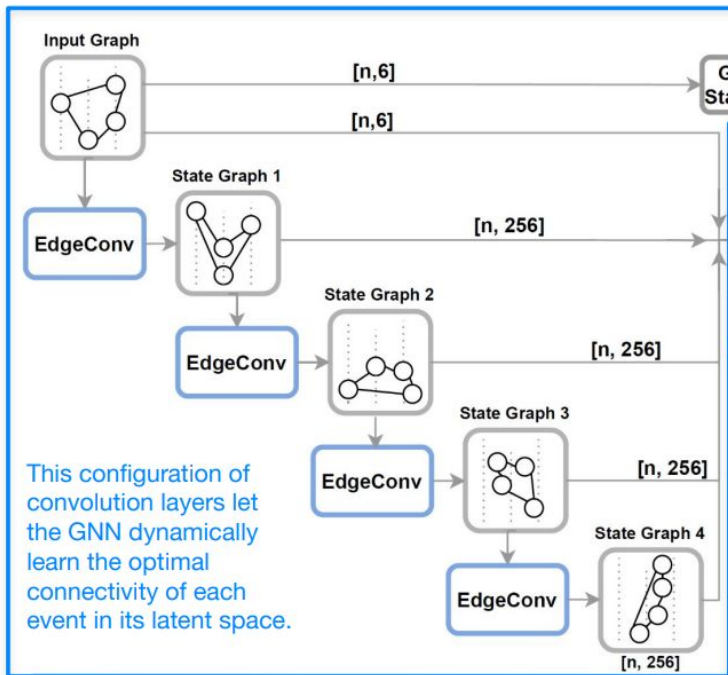
(*) Original model of ParticleNeT uses “mean” as the only node aggregation. Here it has been changed to [“mean”, “max”, “min”, “sum”] changing the number of trainable parameters of the model from ~370k to this 512k.



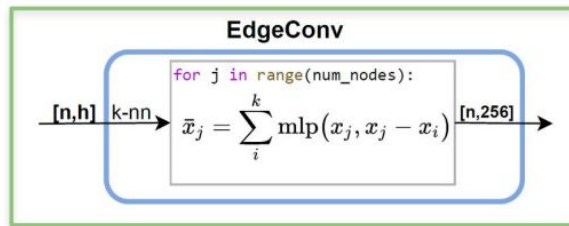
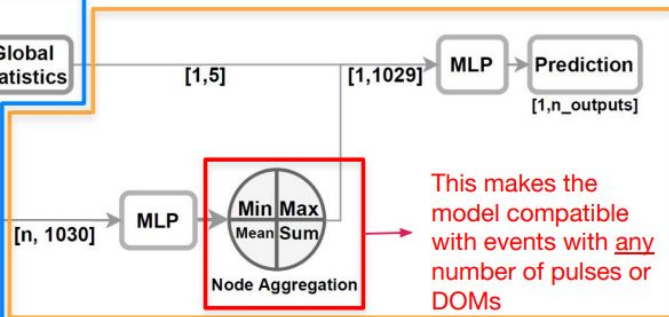
More complex GNNs - DYNEDGE



Graph Convolutional Layers



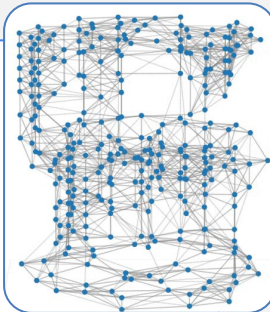
Classical Neural Networks



Our Choice in Convolution

(<https://arxiv.org/pdf/1801.07829.pdf>)

Captures globally relevant features in local areas by considering the difference of a node and the neighbours it's connected to. Only the neighbouring nodes contributes to the convolution!

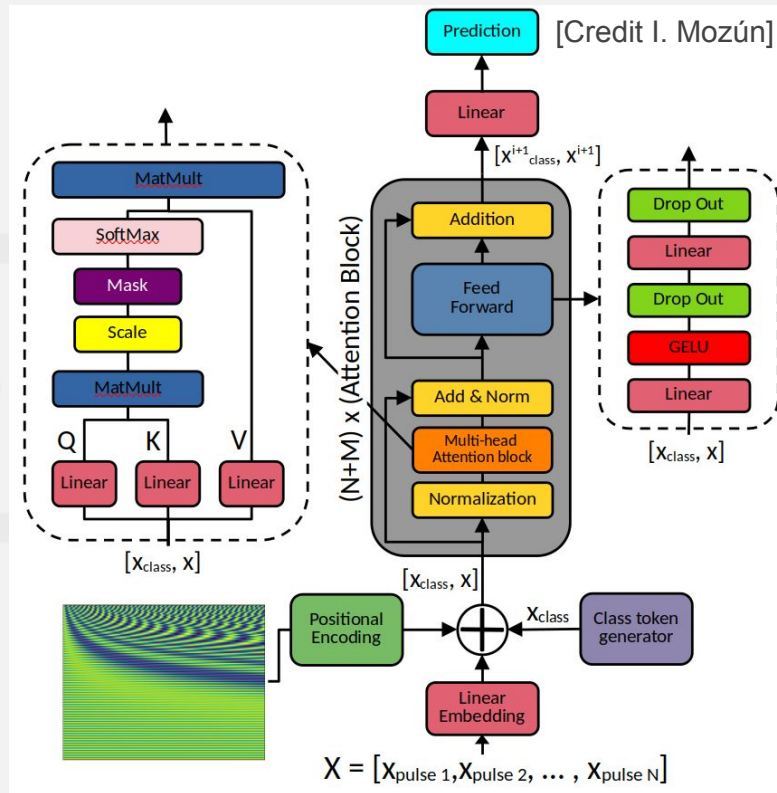


1.4M trainable parameters

Transformers in KM3NeT

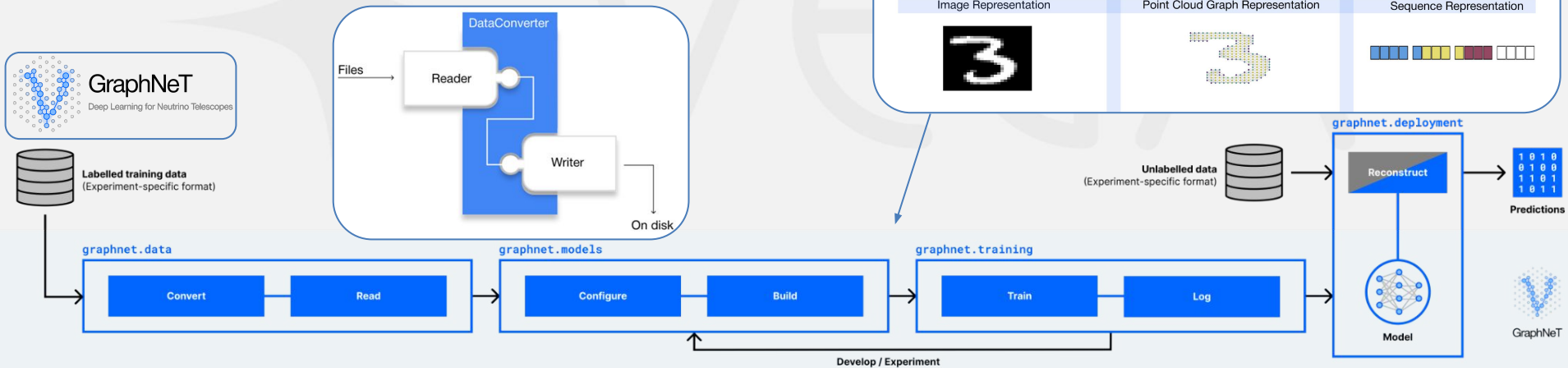
- Transformer model are **inspired in Natural Language Processing** tasks. They are starting to overtake but they have a **lot of learnable parameters that require a lot of data** for an optimal performance.
- The model arranges **light pulses in tokens** with information **[pos, dir, t, ToT]** in a sequence and learns relationships among them to then perform different tasks.
- Model with **1.6M trainable parameters**.
- Offers potential for **transfer learning**: fine-tuning pre-trained models for different detector configurations.

More details about the model and transfer learning in the follow up talk by Iván Mozún!



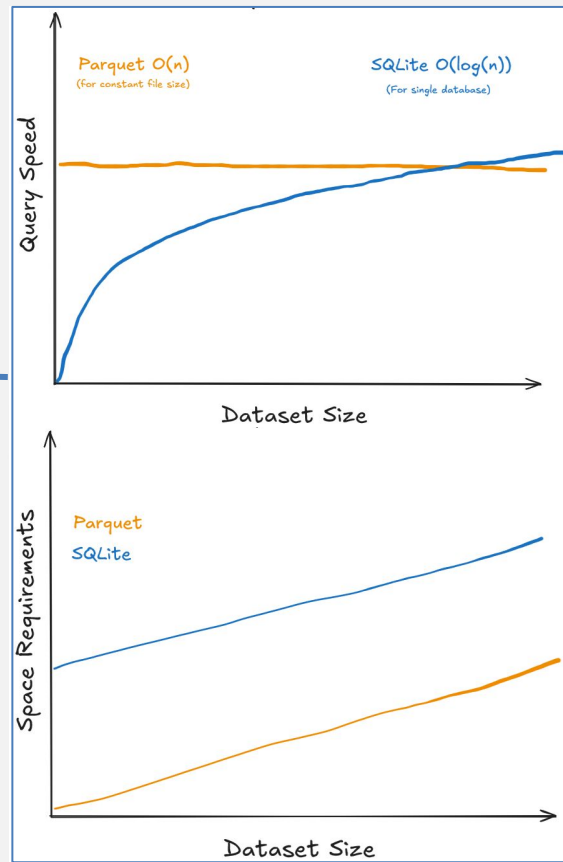
Comparing models in KM3NeT

- **GraphNeT** is a common framework for DL projects.
- It has a **modular structure** that makes it very **easy to embed your detector and your models** in the software.
- Dedicated [instructions](#) to include your detector.
- **Many models implemented.** ParticleNeT, DYNEDGE, RNNs and top 3 kaggle solution (see description [here](#))
- Working on **SQLite** or **Parquet**.



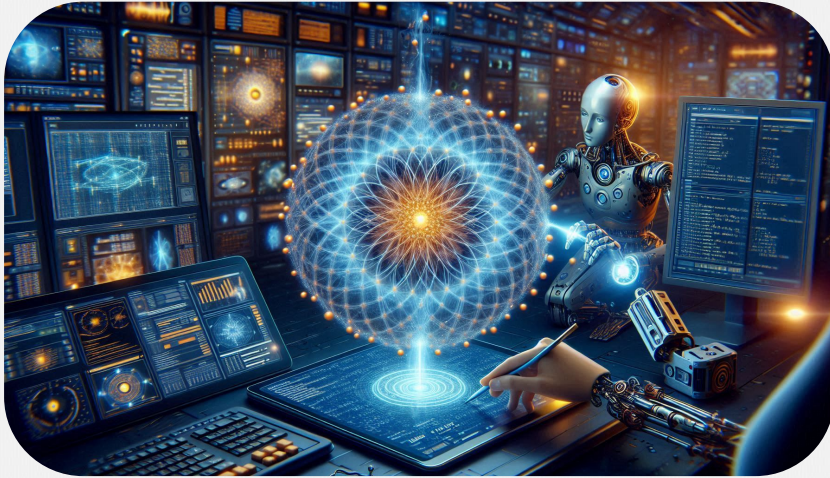
Comparing models in KM3NeT

- **GraphNeT** is a common framework for DL projects.
- It has a **modular structure** that makes it very **easy to embed your detector and your models** in the software.
- Dedicated [instructions](#) to include your detector.
- **Many models implemented.** ParticleNeT, DYNEDGE, RNNs and top 3 kaggle solution (see description [here](#))
- Working on **SQLite or Parquet.** ←



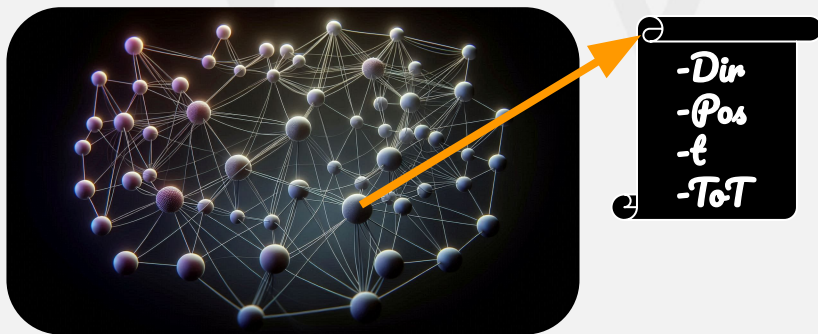
PART III

Preliminary model comparison



Comparing Models

- Comparison between **ParticleNeT**, **DYENDGE** and the **transformer** model training, validating and testing on the exact same events.
- Trained for: **classification between tracks and showers**, **reconstruction of the neutrino energy** and the incoming **direction of the neutrino**.
- Using electron (anti-)neutrino as showers and muon (anti-)neutrinos as tracks in the energy range 1-100 GeV.



	Tracks	Showers
Training	1.92M	1.92M
Validation	480k	480k
Test	278k	278k

	Trainable parameters
ParticleNeT	521k
Dynedge	1.4M
Transformer	1.6M

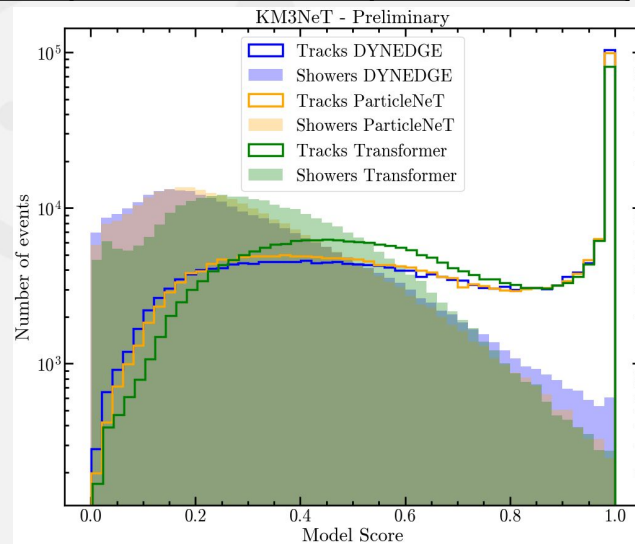
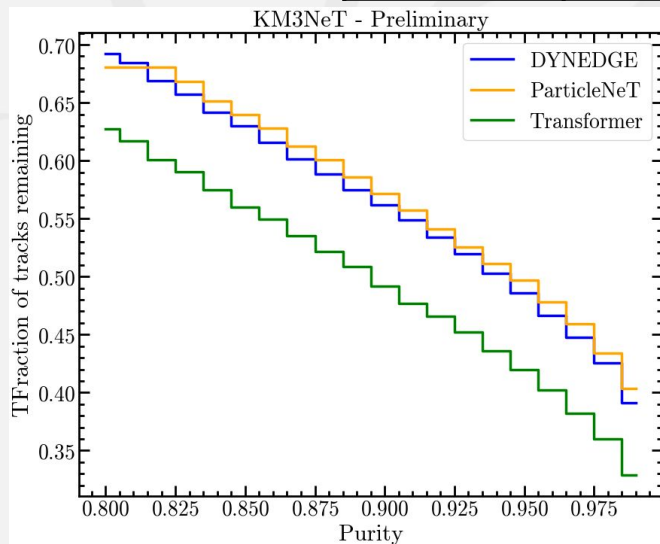
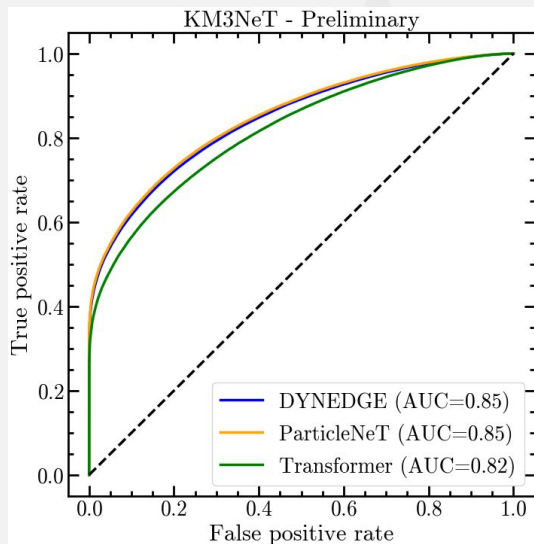
ParticleNeT and DYNEDGE trained 1 GPU Tesla V100-SXM2-32GB

Transformer trained in 8 GPUs NVIDIA A100-SXM4-80GB

Track-Shower Classification

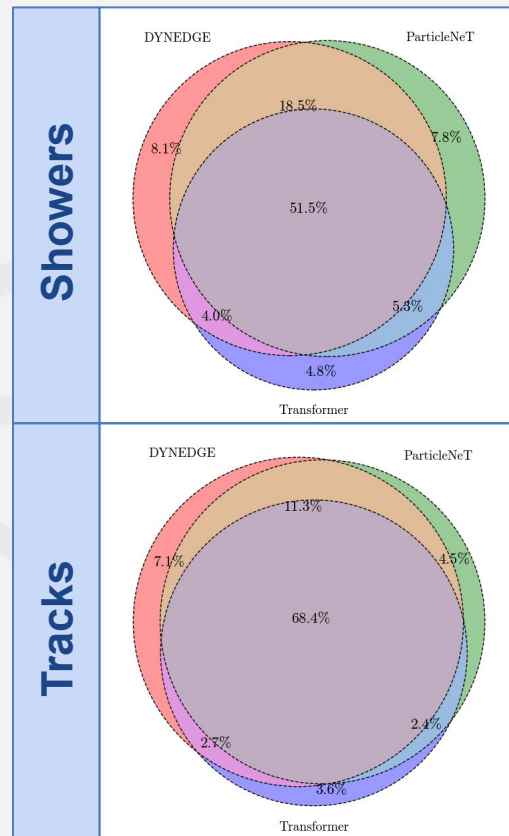
- Transformer performing worse DYNEDGE or ParticleNeT.
- The **simplest model is performing the best**.
Lack of training events to extract the full potential of DYNEDGE or the transformer?

	N-epochs	Loss function	Time per epoch
ParticleNeT	38	BinaryCrossEntropy	~26min
DYNEDGE	28	BinaryCrossEntropy	~28min
Transformer	52	BinaryCrossEntropy	~15min



Track-Shower Classification

- Check if the models are targeting different events. If, for instance, one model finds it easy to classify a track as a track while the other struggles, while for a different track the opposite happens, **then the combination of both predictions** might be valuable.
- Here it has been considered events classified as a shower for a model those with a score smaller than 0.3 and tracks the ones with a score greater than 0.7.
- Most of the events are understood the same way by the three models.
- Tried a BDT with features: the three scores + the three energy reconstructions (see in a moment) but very little improvement in the classification (AUC=0.86).

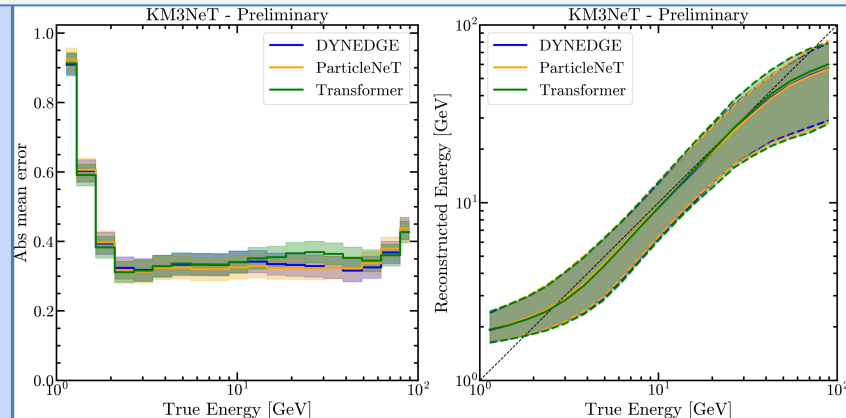


Energy Reconstruction

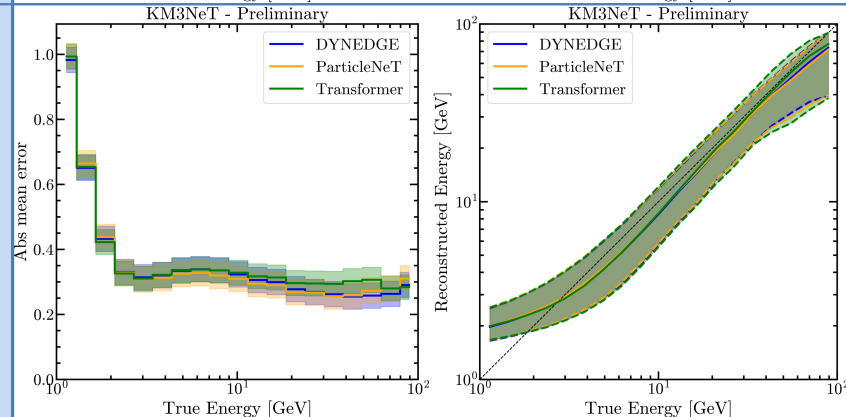
- Very **similar performances of the three models** in the entire energy range.
- For tracks and showers all the three models **struggle with events below 3 GeV**.
- For tracks **over 60 GeV the three models have difficulties** in the prediction as well.

	N-epochs	Loss function	Time per epoch
ParticleNeT	29	LogCosh	~22mins
DYNEDGE	22	LogCosh	~23mins
Transformer	58	LogCosh	~15mins

Tracks



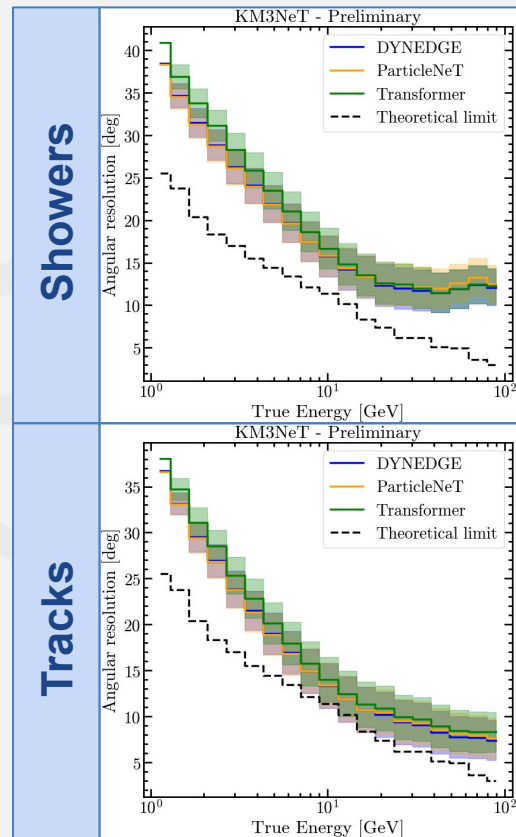
Showers



Direction Reconstruction

- **Better performance of ParticleNeT and DYNEDGE** over the transformer.
- Can reach **resolutions below 15°** for tracks and for showers over **10 GeV**.
- VonFisherMises loss function outputs **as well a measurement of how certain the model is of a prediction**. Can be used to achieve a sample of over 30% of the events but with angular resolutions close to the theoretical limit. (See [this talk](#) at NPML2024)

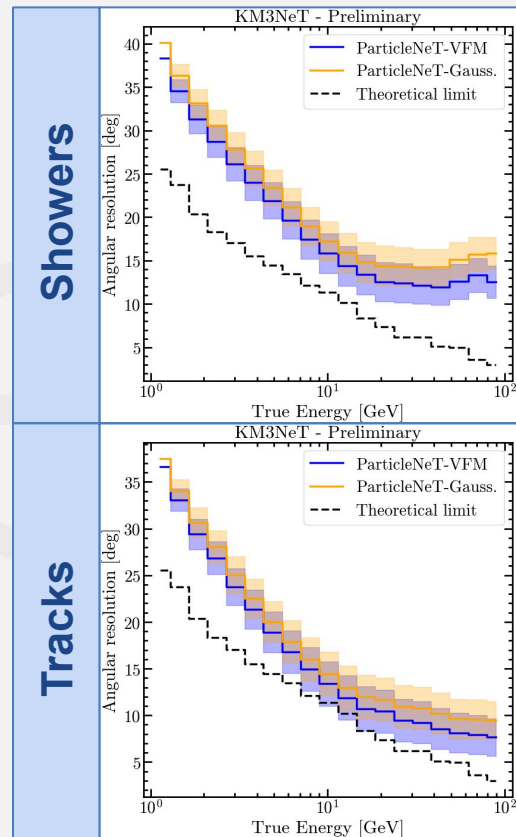
	N-epochs	Loss function	Time per epoch
ParticleNeT	45	VonfisherMises	~27mins
DYNEDGE	25	VonFisherMises	~31mins
Transformer	56	GaussianNegativeLogLikelihood	~28mins



Trying new loss functions?

- **ParticleNeT** in KM3NeT and the **transformer** model used **GaussianNegativeLogLikelihood** (GNLL) as loss function for direction reconstructions while **DYENDGE** for IceCube direction reconstruction used **VonFisherMises** (VFM). Can any of the models perform better by using a different loss function?
- **ParticleNeT** trained with the exact same hyperparameters shows an **improvement when changing from the GNLL to the VFM** in every energy bin.
- Transformer might enhance its performance as well (under study).

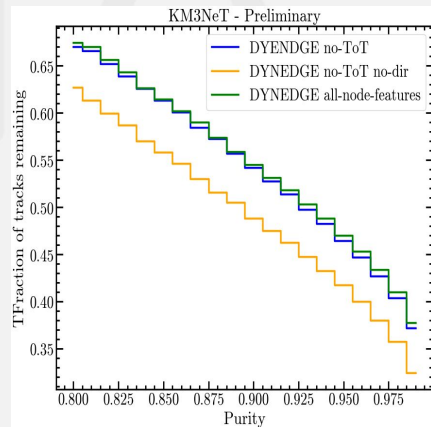
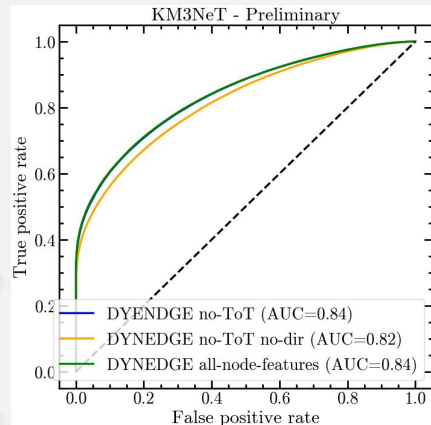
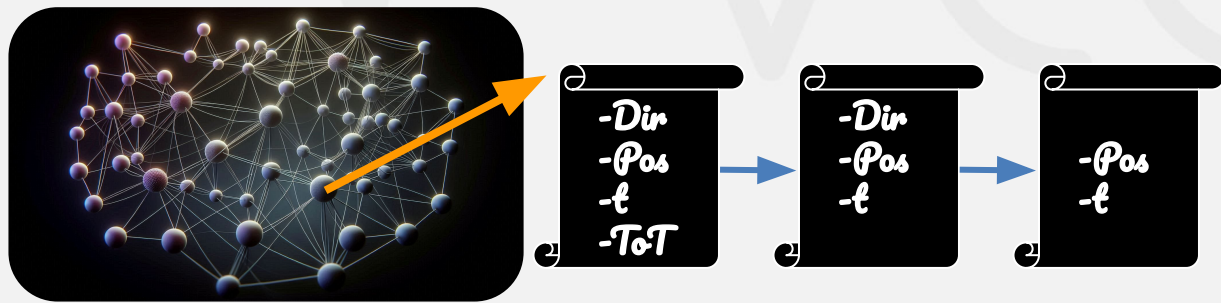
	N-epochs	Time per epoch
ParticleNeT Gauss.	29	~28mins
ParticleNeT VFM	45	~27mins



Node features studies

- How well the model **performs with less node information?**
- ToT does not affect** the performance but **PMT direction does.**
- Need to check per task.
(work in progress).

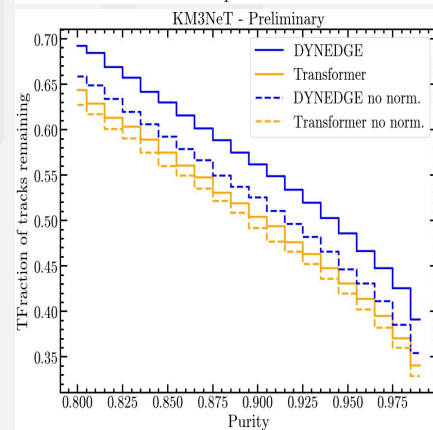
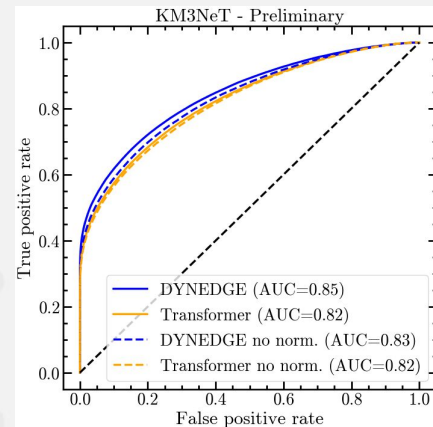
	N-epochs	Time per epoch
DYN. all	26	~27mins
DYN. no-ToT	26	~22mins
DYN. no_ToT,dir	38	~20mins



Node Inputs Normalization Effect

- In principle, **the model should be able to find out the range of values for each feature and learn how to deal with it.**
- Providing a **normalization so that the features are in the range [-10, 10]** (as in [here](#)) **enhances the performance and improves the training times.**
- Need to study and optimize the normalization? Normalizing x, y and z so that they lay within -10 and 10 can not distort the geometry? Is this possible that this also affects to the choice of neighbours?
- For sure it affects, but the optimization is work in progress.

	N-epochs	Time per epoch
DYNEDGE (Norm / No-norm)	26 / 30	21 mins / 24mins
Transformer (Norm / No-norm)	32 / 52	15.3mins / 14.9mins



Conclusions

- **Deep learning** techniques are **very suitable techniques to study neutrino events in KM3NeT** detector as they capture information that with BDT or likelihood based methods for study are missing.
- **Need to perform further comparisons** and then study how well is the agreement between the predictions on simulation and data.
- It is key to see the **adaptability of models** for different detector configurations as we are working on a growing project. **See next talk about Transfer Learning!**
- IceCube, KM3NeT, P-One, Baikal-GVD...very similar detectors. Would it be nice to study how well the techniques used in one detector work in another. (work in progress, **see paper soon**)



Thank you!

Models trained in:



Appendix - TS score over E, Zen and Az

