Alessandro Strumia, talk at

#### "Is SUSY alive and well?"



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### Was SUSY alive and well?



## And for a good reason



## 1988: naturalness bounds on sparticles

"Barbieri was grilling a kidney without enough fire, a student was going around excited like a stock operator, telling 'chargino 200! gluino 400!"



Rattazzi expressed doubts but Barbieri said "bischero this a delicacy, eat it". So "it tasted weird, but my advisor thought it was gourmet, I ate it".

## 1990: LEP1 data speak

![](_page_4_Figure_1.jpeg)

'Neutralino' was synonymous of Dark Matter

## **1998: SK, LEP and cosmology speak**

SUSY needs  $Z_{2,3}$  and universal  $\tilde{m}$  to get  $B, L, L_i$  that come for free in the SM. Even so, if  $\tilde{m} \sim M_Z$ , p decay at dimension 5 is problematic in minimal SU(5).

"LEP2 experiments pose a serious naturalness problem for supersymmetric models". [hep-ph/9811386]. "About 95% of CMSSM parameter space is excluded" i.e. [hep-ph/9904247]

p(no SUSY at LEP|CMSSM) = 5%

FT = Bayesian **probability** of numerical accidents. "The well known naturalness problem of the Fermi scale has gained a pure low energy aspect".

![](_page_5_Picture_5.jpeg)

"The cosmological constant poses another serious unsolved problem, also related to power divergences". [hep-ph/0007265]

## The CMSSM

Use dimension-less ratios as parameters and fix the SUSY scale from

$$M_Z^2 \approx 0.2m_0^2 + 0.7M_3^2 - 2\mu^2 = (91 \,\text{GeV})^2 \times (\frac{M_3}{110 \,\text{GeV}})^2 + \cdots$$

![](_page_6_Figure_3.jpeg)

Main worry: so many sparticles at LHC that disentangling them will be hard.

## 2010: LHC data speak

![](_page_7_Figure_1.jpeg)

## The CMSSM

![](_page_8_Figure_1.jpeg)

Survives only only close to the critical line v = 0

Even one loop stop corrections start to be unnaturally big.

## LHC data speak badly about SUSY

![](_page_9_Picture_1.jpeg)

(Same message to workers on naturalness, diphoton...)

### **Dark Matter?**

![](_page_10_Figure_1.jpeg)

and 'Minimal Dark Matter' limit: higgsino at 1.1 TeV, wino at 2.5 TeV...

# Higgs at 125 GeV?

The MSSM prediction for  $\lambda$  i.e.  $M_h$  can now be computed in a simpler way:

1) Weak scale: SM at 2 loops; 2) 3 loop RGE running up to  $\tilde{m}$ ; 3)  $\lambda = \frac{1}{8}g^2c_{2\beta}^2 + 1$ -2 loop SUSY.

Sparticle corrections to  $y_t, g_3$  can be neglected.

Result: predicted  $M_h$  gets lower and more precise,  $\delta M_h \sim 1 \text{ GeV}$  Quasi-natural SUSY,  $\tan \beta = 20$ 

![](_page_11_Figure_6.jpeg)

Multi-TeV stop, huger tuning

## **Options: USUSY or USUSY**

160 Ugly Unnatural U  $\tan\beta = 50$ Split SUSY  $\tan\beta = 4$  $\sim$  TeV  $\gg$  TeV  $\tan\beta = 2$  $\tilde{m}$ 150  $\tan\beta = 1$ Naturalness No Bad Higgs mass  $m_h$  in GeV Higgs mass Bad Good 140 Dark Matter Bad Good Unification Bad Good High-Scale SUSY 130 Flavour Bad  $\sim$ Good Experimentally favored Models Bad Good 120 TeV signals Yes No Makes sense? Bah An\*\*\*pic? 110  $10^{6}$  $10^{8}$  $10^{10}$  $10^{12}$  $10^{14}$  $10^{16}$  $10^{18}$  $10^{4}$ Supersymmetry breaking scale in GeV

Predicted range for the Higgs mass

For example mini-split:  $\tilde{m} \sim 4\pi M_{1,2,3}$  with  $\tilde{H}$  or  $\tilde{W}$  as Minimal DM.

## **Does Unnatural SUSY make sense?**

Does the anthropic multiverse justify Unnatural SUSY?

- $m_p \ll M_{\text{Pl}}$  allows systems with  $N \sim M_{\text{Pl}}^3 / m_p^3 \gg 1$  particles.
- $y_d v \approx \alpha_{\rm em} \Lambda_{\rm QCD}$  allows chemistry.

#### But natural solutions exist, difficult to argue that multiverse avoids them.

If we live in a multiverse with many low-energy SUSY vacua, the likely outcome is again natural SUSY with  $\tilde{m} \sim M_Z$  (e.g. mini-split SUSY with  $M_3 \sim m_p$ ).

If we live in a multiverse with many Planck-scale SUSY vacua, the likely lowenergy physics is

- an anthropically acceptable alternative to the SM that does not involve an unnaturally light Higgs scalar;
- or (even within the Standard Model) a smaller y or a smaller  $M_{\text{Pl}}$ .

To argue differently one needs to add ad hoc counting or DM restrictions.

Keep searching alternatives to anthropic nirvana

## Subtle is the Lord

What is going on? We are confused but nature is surely following some logic

![](_page_14_Picture_2.jpeg)

## Data speak and tell Standardissimo Model

We now have all SM parameters, let's assume SM and see what happens.

Facts:

- 1: SM can be extrapolated above  $M_{\rm PI}$ .
- 2:  $\lambda(M_{\mathsf{Pl}}) \approx 0$  at  $\approx 2\sigma$ .
- 3:  $\beta(\lambda)$  vanishes around  $M_{\text{Pl}}$ .

![](_page_15_Figure_6.jpeg)

## Scalarphobic vs scalarfriendly

Scalarphobic theorists believe that scalars are unnatural because  $\delta M_h \sim g_{SM} \Lambda$ . But power divergences give no physical effect. In quantum mechanics it's better to stick to observables, without adding realism. Maybe scalarphobic theorists over-interpret equations, as happened with the æther: "wave  $\Rightarrow$  medium".

#### Scalarfriendly theorists can try new roads:

**Finite naturalness**. Upper bounds on new physics from naturalness of physical corrections:  $\delta M_h \sim g_{\text{new}} M_{\text{new}}$  at 1/2/3 loops. Allows SM + DM + neutrino masses + baryogenesis + inflation + axions. No GUT, no string. **Dynamical generation of the weak scale**.  $\int dE E = 0$  if physics is dimensionless. Simple models where a vev ( $\lambda < 0$ ) or a condensate ( $g \rightarrow \infty$ ) generates  $M_h$  and DM. Even models with 0 new parameters: predict  $m_h^2 < 0$ ,  $M_{\text{DM}}$ ,  $\Omega_{\text{DM}}$ . **Gravity**:  $\delta M_h \sim E^2/M_{\text{Pl}}$ : natural at  $E \lesssim \sqrt{M_h M_{\text{Pl}}}$ . New physics there could be the spin 2 negative-norm (?) graviton of agravity = dimension-less renormalizable gravity, where  $\delta M_h \sim g_{\text{gravity}}^2 M_{\text{Pl}}$ .

Dynamical generation of the Planck scale.

**Inflation**. Dimension-less theories allow super-Planckian vevs and give quasi-flat potentials:

 $\epsilon, \eta \sim \beta, \qquad P_R \sim M_h / \bar{M}_{\rm Pl}.$ 

**Total Asymptotic Freedom**. Theories valid up to infinite energy with all couplings  $g, y, \lambda$ flowing to zero. No cut-off, predictions, e.g.  $g_Y = 0, M_t = 186 \text{ GeV}, M_\tau = 0$  in the SM. Weak-scale extensions of the SM into SU(3)<sup>3</sup>.

Fundamental models of composite Higgs.

![](_page_17_Figure_4.jpeg)

![](_page_17_Picture_5.jpeg)

## Is SUSY well?

No Giving up naturalness maybe better than giving up the rest

![](_page_18_Figure_2.jpeg)

## Is SUSY popular?

Yes

![](_page_19_Figure_2.jpeg)

year

### Is SUSY alive?

"Ibis redibis non morieris in bello"