

Numerical Stochastic Perturbation Theory in the Schrödinger Functional: Opportunities & Challenges

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Overview

- What is NSPT and what is it useful for?
- Teaming up NSPT and the Schrödinger Functional.

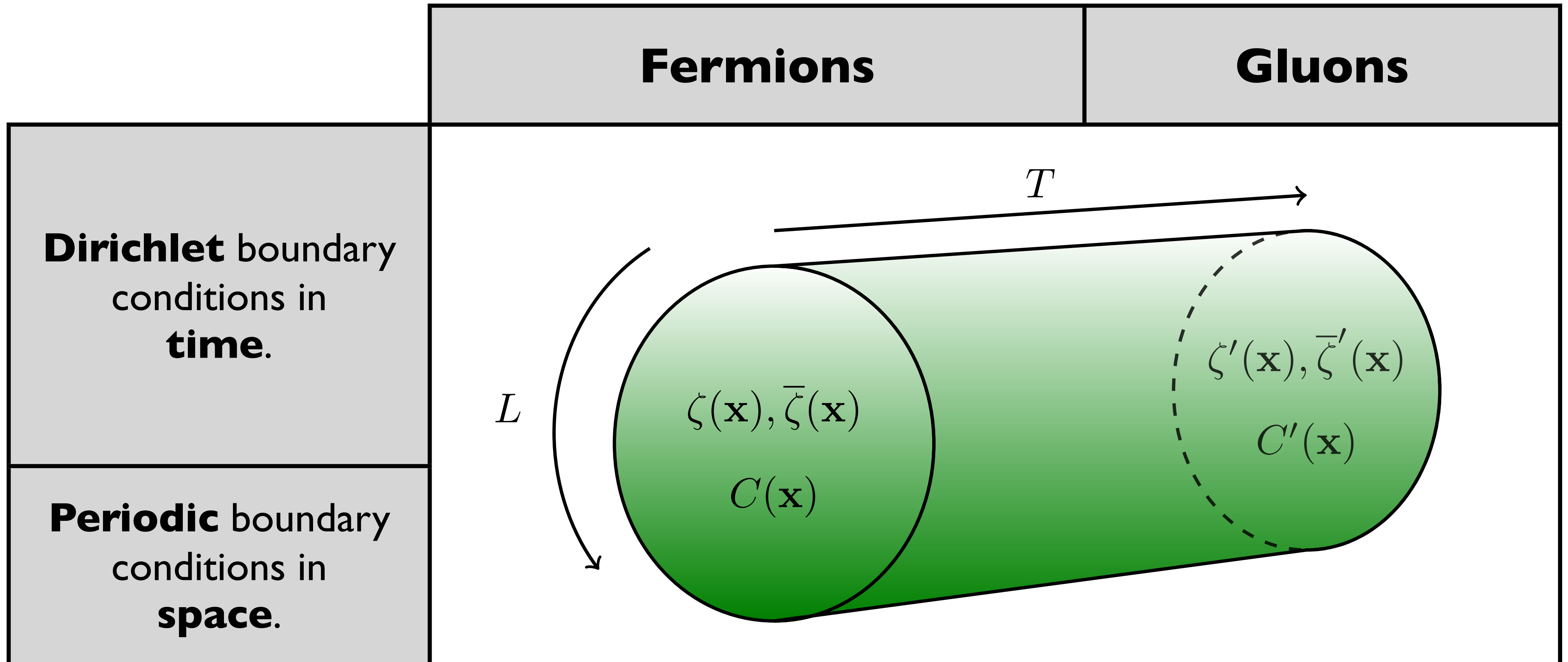
The Schrödinger Functional

	Fermions	Gluons
Dirichlet boundary conditions in time.	$\zeta(\mathbf{x}) = \sum_y \tilde{K}(\mathbf{x}, y) \psi(y)$ $\bar{\zeta}(\mathbf{x}) = \sum_y \bar{\psi}(y) K(y, \mathbf{x})$ $P_+ \psi(x) _{x_0=0} = \bar{\psi}(x) P_- _{x_0=0} = 0$	$U_k(x) _{x_0=0} = e^{C(\mathbf{x})}$ $U_k(x) _{x_0=T} = e^{C'(\mathbf{x})}$
Periodic boundary conditions in space.	$\psi(x + \hat{k}L) = e^{i\theta_k} \psi(x)$	$U_\mu(x + \hat{k}L) = U_\mu(x)$

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The Schrödinger Functional



Perturbation Theory

In **perturbation theory**, we want to obtain **expansions** like this one:

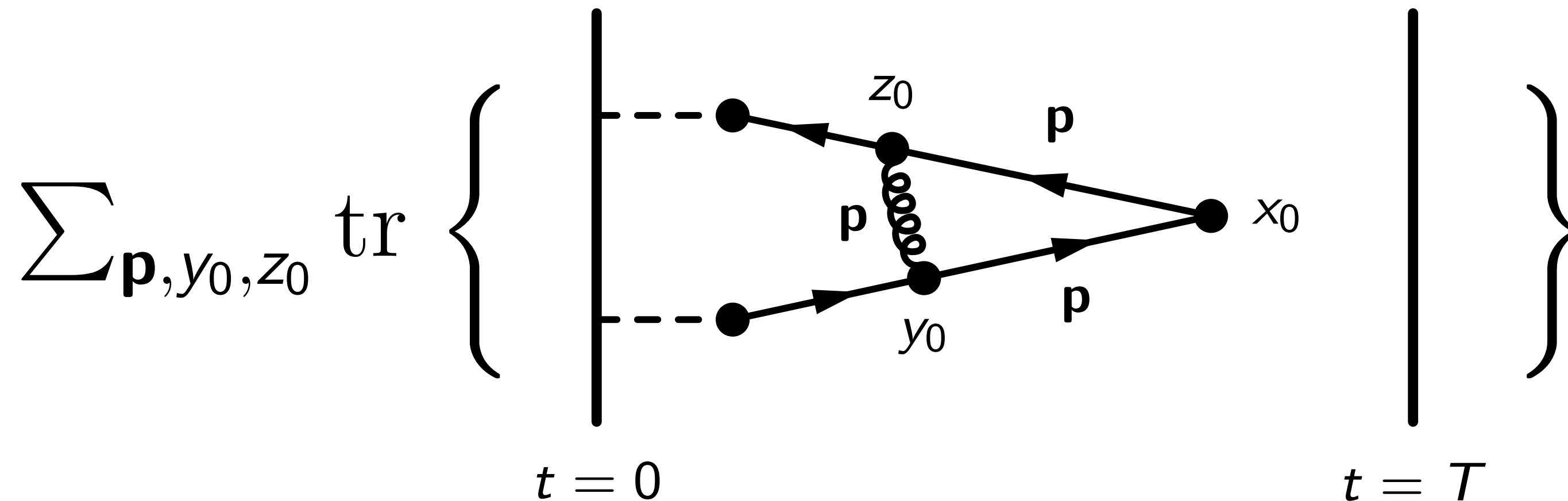
$$\langle O[\phi] \rangle = \frac{1}{Z} \int \mathcal{D}\phi O[\phi] \exp \{ S_0[\phi] + \alpha S_1[\phi] + \dots \} = O^{(0)} + \alpha O^{(1)} + \dots$$

Usually, one calculates $O^{(i)}$
using **Feynman diagrams** and **rules** extracted from

$$S[\phi] = S_0[\phi] + \alpha S_1[\phi] + \dots$$

PT in the SF

In the Schrödinger Functional one would like to **avoid** PT!



- Big number of diagrams even at low orders
- Background field makes Feynman rules complicated

Numerical Stochastic PT avoids both!

Stochastic Quantization I

We want to calculate an **expectation value**

$$\langle O[\phi] \rangle = \frac{1}{Z} \int \mathcal{D}\phi O[\phi] e^{-S[\phi]}$$

Introduce a new d.o.f., the **stochastic time** t .

The evolution in stochastic time is given by the **Langevin Equation**,

$$\dot{\phi}_\eta(x; t) = -\partial_{\phi_\eta(x; t)} S[\phi] + \eta(x; t)$$

With **Gaussian noise** η .

Stochastic Quantization 2

Defining the **'noise average'**

$$\langle O \rangle_\eta = \frac{1}{Z'} \int \mathcal{D}\eta \, O \, e^{-\frac{1}{4} \int d[z, \tau] \eta^2(z, \tau)},$$

One asserts that the **functional integral** can be calculated using

$$\langle O[\phi_\eta(x_1; t), \dots, \phi_\eta(x_n; t)] \rangle_\eta \xrightarrow{t \rightarrow \infty} \langle O[\phi(x_1), \dots, \phi(x_n)] \rangle$$

Stochastic Perturbation Theory

Split up the action into free and interacting parts

$$S[\phi] = S_0[\phi] + gS_1[\phi] + \dots$$

and formally write $\phi(x; t) = \sum_{i=0}^{\infty} g^i \phi^{(i)}(x; t)$

Using the Langevin equation one may now obtain $\langle O \rangle_{\eta} = \sum_{i=0}^{\infty} g^i \langle O \rangle_{\eta}^{(i)}$

By defining $(\phi + \varphi)^{(r)} = \phi^{(r)} + \varphi^{(r)}$, $(\phi\varphi)^{(r)} = \sum_{i=0}^r \phi^{(i)} \varphi^{(r-i)}$

Numerical Stochastic Perturbation Theory

- In NSPT, one integrates the perturbative Langevin Equation numerically.
- This is similar to hybrid MC methods c.f. Stefan's talk.
- However, there is no perturbative expression for an accept/reject step.
- Hence, one is stuck with a finite integration time τ and has to extrapolate $\tau \rightarrow 0$.

Stochastic Gauge Fixing

Zwanziger, 1981

In principle no GF is needed. However, if one looks at the Langevin eqn.,

$$\frac{\partial}{\partial t} A_{\mu}^a(\eta, x; t) = D_{\nu}^{ab} F_{\nu\mu}^b(\eta, x; t) + \eta_{\mu}^a(x; t)$$

One finds that for a solution in Fourier space

$$A_{\mu}^{(n)a}(k; t) = T_{\mu\nu}^{ab} \int_0^t ds e^{-k^2(t-s)} f_{\nu}^{(n)b}(k, s) + L_{\mu\nu}^{ab} \int_0^t ds f_{\nu}^{(n)b}(k, s)$$

The longitudinal component will diverge like a random walk. GF introduces a damping factor and stabilizes the simulation.

A similar statement holds for the gauge zero modes.

Gauge Fixing Pitfalls

The gauge fixing function in the SF **at the boundary** reads

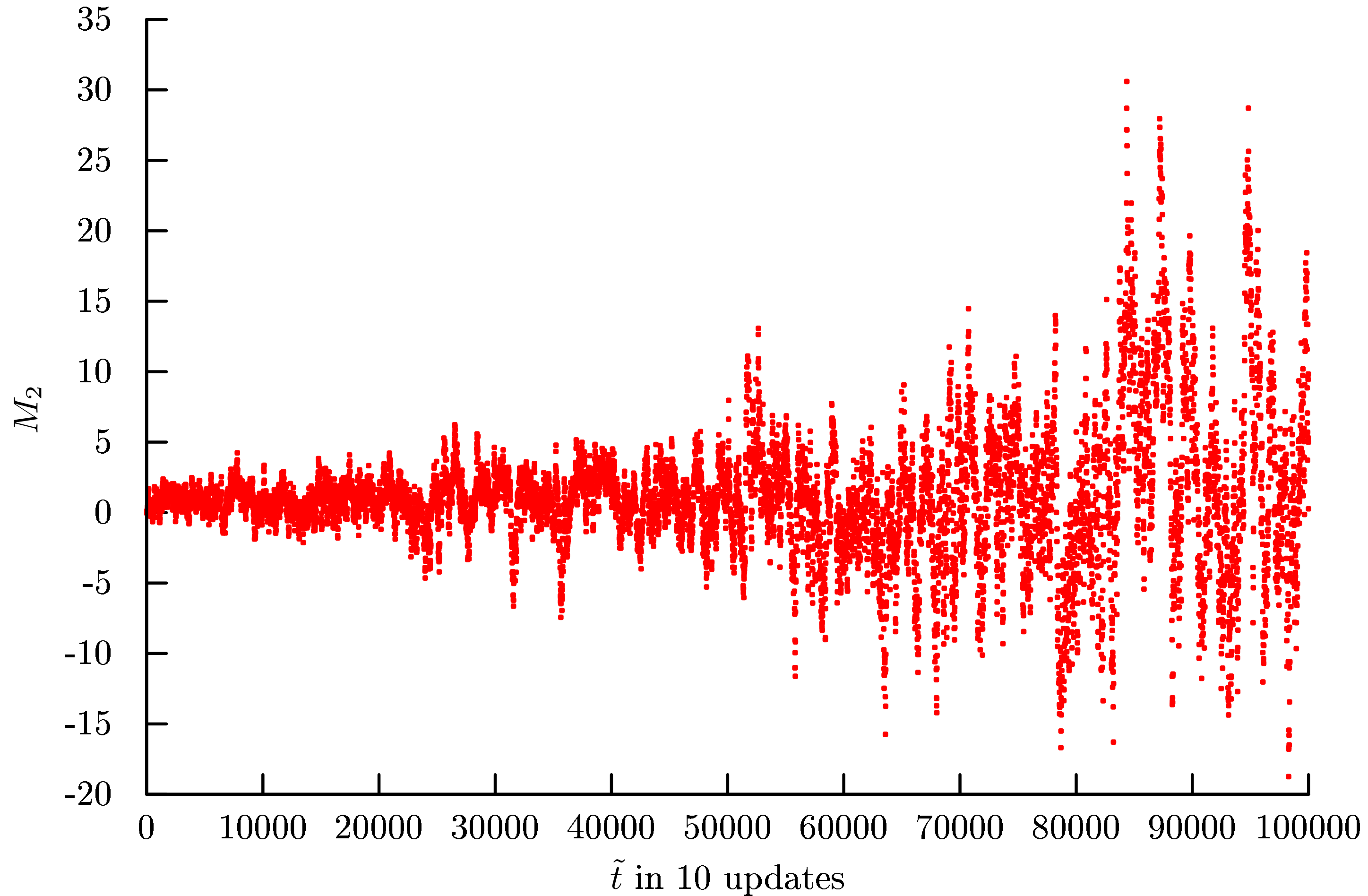
$$[d^* q(x)]_{ij} = \begin{cases} (a^2/L^3) \sum_{\mathbf{y}} [q_0(0, \mathbf{y})]_{ij} & \text{if } x_0 = 0, i = j \\ 0 & \text{else} \end{cases}$$

It acts on the fluctuation field, writing

$$U_\mu(x) = \exp\{g_0 a q_\mu(x)\} V_\mu(x)$$

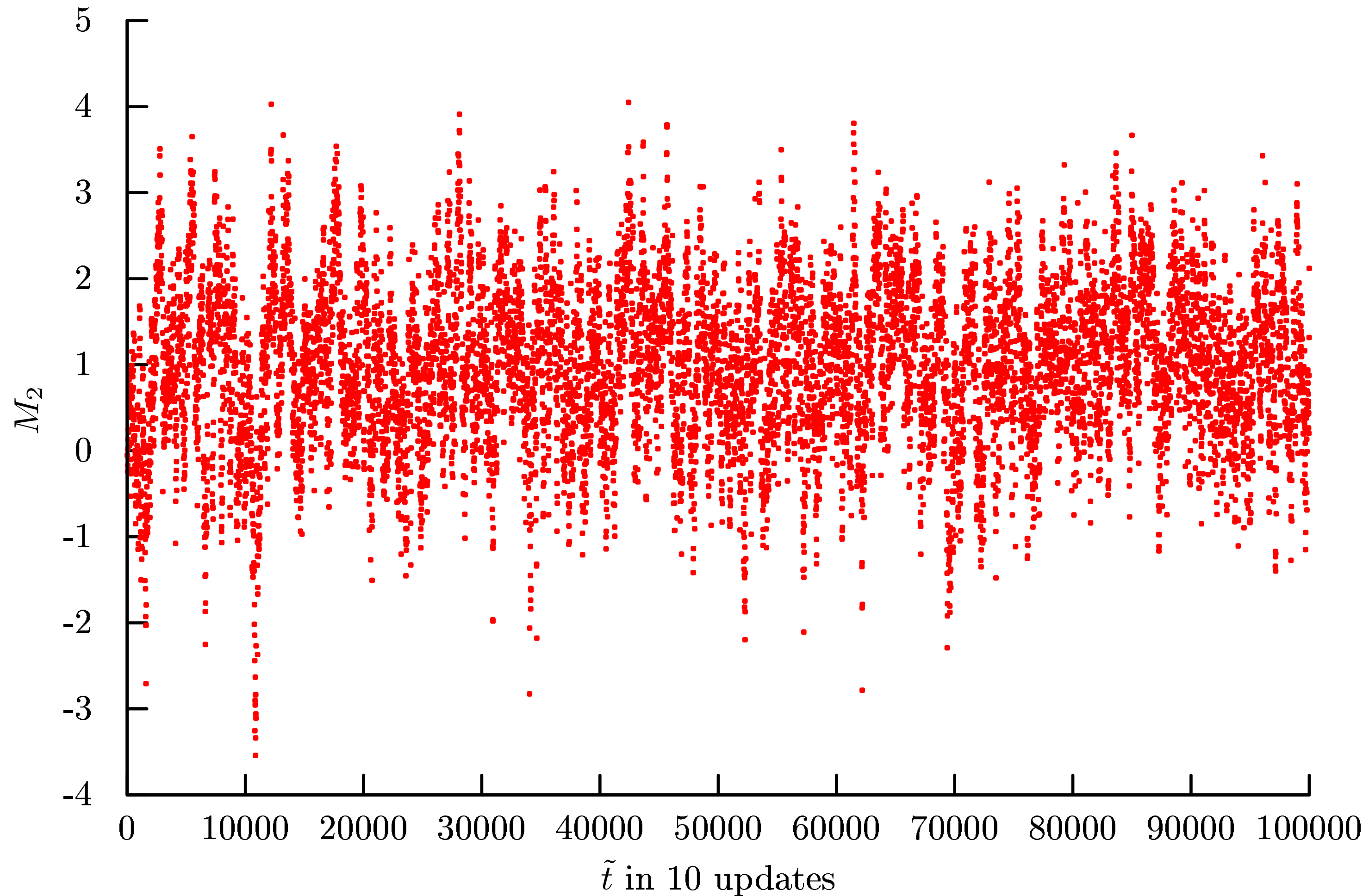
This amounts to suppressing **spatial zero modes** at the boundary.

Bad History



Incorrect gauge fixing leads to a slow increase of noise (which can be confusing).

Good History



Correct gauge fixing leads to a drastic decrease of the noise.

Advantages of NSPT

- No Feynman Rules, i.e. easier to implement various actions.
- No Feynman Diagrams, i.e. higher order are only a matter of CPU time.
- NSPT codes can benefit from non-perturbative ones ...
- ... and vice versa!

No Free Lunch

- Way more numerical effort than common PT.
- Stochastic noise makes extraction of logs difficult.

But NSPT has many applications, e.g. improvement:

$$\mathcal{O}_I(a/L) = \frac{\mathcal{O}(a/L)}{1 + \delta(a/L)}$$

$$\delta(a/L) = \frac{\mathcal{O}(a/L) - \mathcal{O}(0)}{\mathcal{O}(0)} = \delta^{(0)}(a/L) + g_0^2 \delta^{(1)}(a/L) + \dots$$

The SF coupling

The SF introduces an external scale $T = cL$ (usually $c = 1$).

This enables us to give a precise definition for a coupling

$$\bar{g}(\mu = 1/L)$$

Furthermore one can calculate a discrete beta-function

$$u = \bar{g}^2(L), \quad u' = \bar{g}^2(sL), \quad \sigma(s, u) = u'$$

On the lattice: $\Sigma(s, u, a/L) = u'$

Defining the SF coupling

The usual choice for the background field is 'defined' by setting

$$C = i/L \operatorname{diag}(\phi_1, \phi_2, \phi_3), \quad C' = i/L \operatorname{diag}(\phi'_1, \phi'_2, \phi'_3)$$

Where ϕ_i, ϕ'_i depend on two parameters, ν and η .

$$\text{Defining as usual } e^{-\Gamma} = \int D[u] e^{-S[U]}$$

one can define a coupling through $\left. \frac{\partial \Gamma}{\partial \eta} \right|_{\eta=\nu=0} = \frac{k}{\bar{g}^2}$

$O(a)$ improvement

To implement $O(a)$ improvement, one defines

$$S[U] = \frac{1}{g^2} \sum_p w(p) \operatorname{tr}[1 - U(p)]$$

For plaquettes at the boundary, containing a spatial link, one sets

$$w(p) = c_t(g_0) = 1 + c_t^{(1)} g^2 + \dots \quad \text{else} \quad w(p) = 1$$

The Coupling Order by Order

We then write the coupling order by order,

$$\bar{g}^2(L) = g_0^2 + m_1(L/a)g_0^4 + m_2(L/a)g_0^6 + \dots$$

with

$$m_1 = m_1^a + c_t^{(1)} m_1^b,$$
$$m_2 - m_1^2 = m_2^a + c_t^{(1)} m_2^b + [c_t^{(1)}]^2 m_2^c + c_t^{(2)} m_2^d$$

We use this as a check for our quenched code!

Data Analysis

- We use a custom implementation of the method presented by Ulli Wolff (CPC 156, 2004) to control autocorrelation and determine the errors.
- The extrapolation $\tau \rightarrow 0$ was done using a bootstrap analysis.
- We performed a cross check using binned Jackknife samples.

Our Data

- Did **test runs** on $L/a = 4, 8, 12$
- Coupling is known to be hard to measure
- We experience
 - Lots of **noise**
 - Slow **thermalization**
 - Long **autocorrelation** times

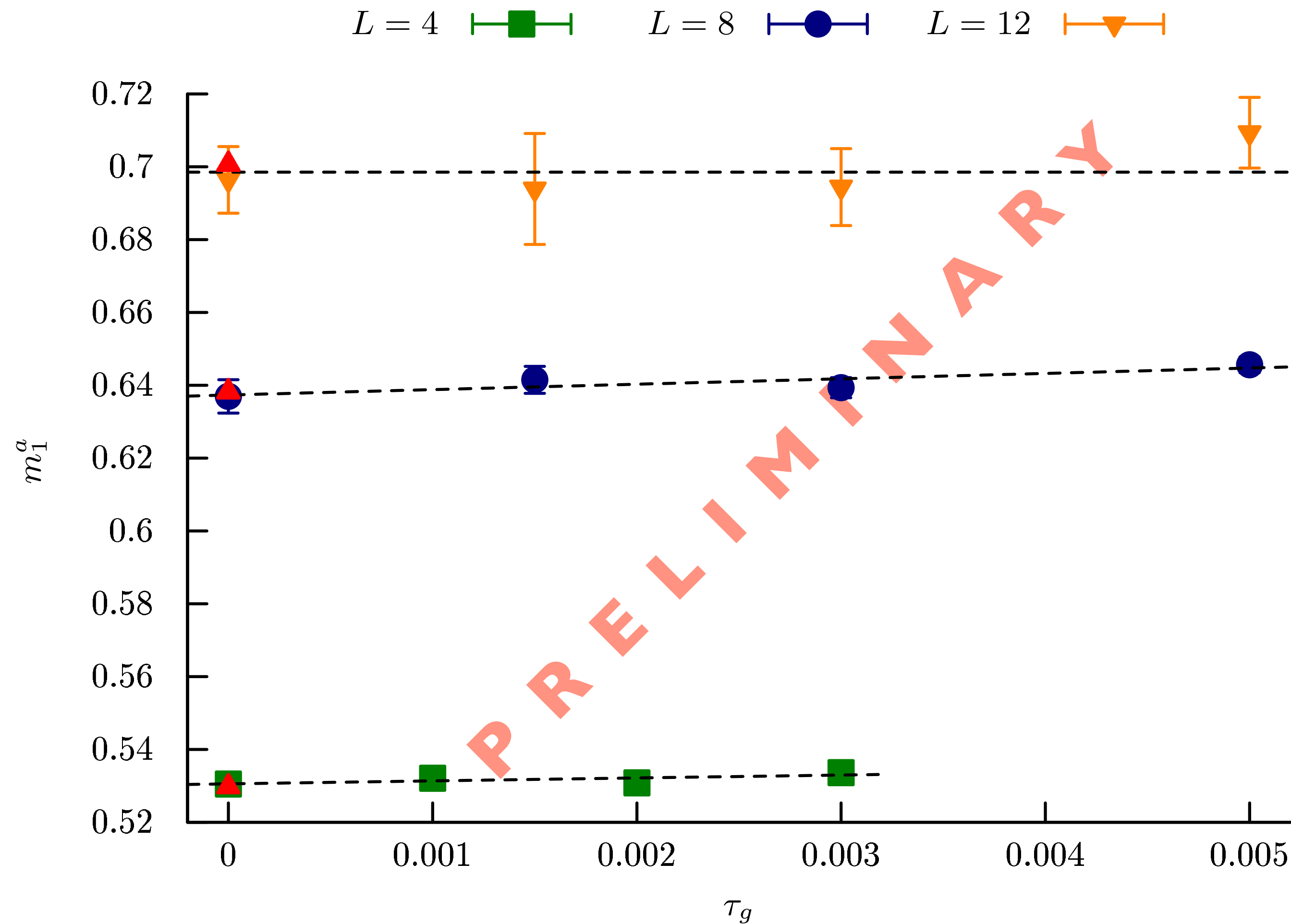
Simulation Details

- $L = 4$ on TURING (Univ. of Milan Bicocca)
- $L = 8, 12$ on FERMI (CINECA)

This is an ideal test case to **check core performances!**

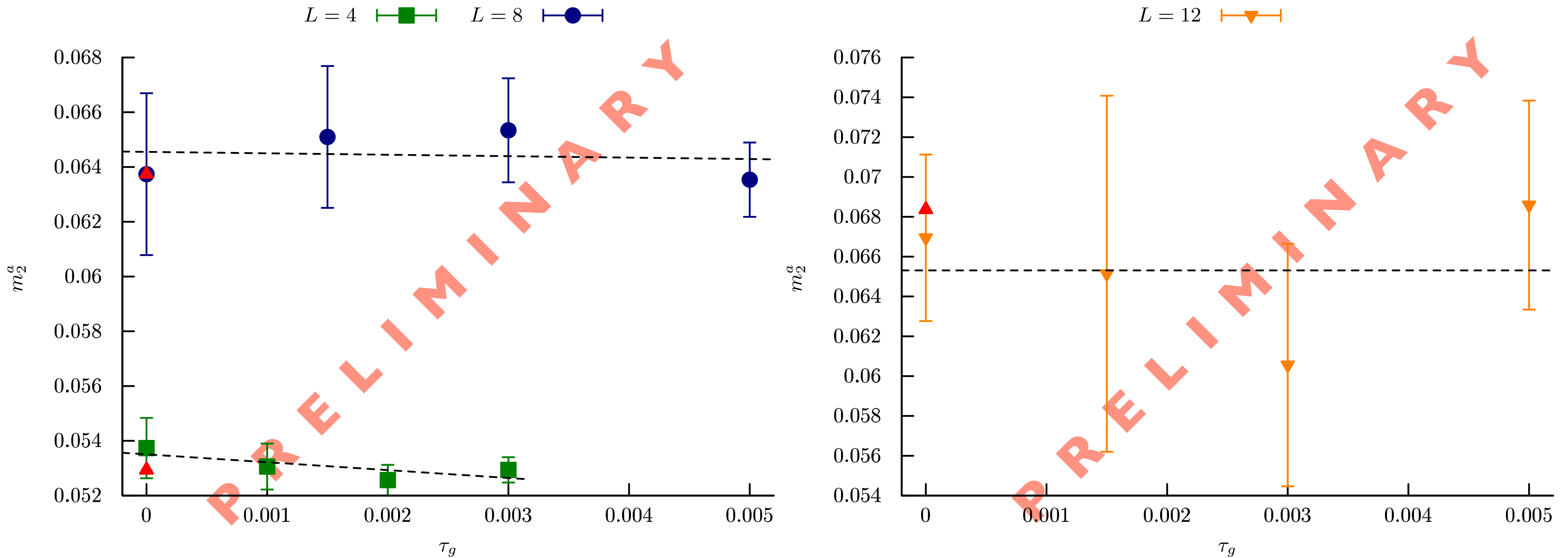
	$L = 12$			$L = 8$			$L = 4$		
τ	0.0015	0.003	0.005	0.0015	0.003	0.005	0.001	0.002	0.003
N_{eff}	270(50)	540(80)	800(100)	1.6(2)k	3.3(2)k	6.3(3)k	6.6(3)k	14.9(5)k	21.8(6)k
CORE-H	2k			1.6k			33		

Results - One Loop



Known results (red) from: Bode, Weisz, Wolff: hep-lat/9809175

Results - Two Loops



Known results (red) form: Bode, Weisz, Wolff: hep-lat/9809175

Conclusions

- Quenched NSPT for the SF is implemented.
- Applications: Cut-off effects.
- Next Step: Fermions!

A word about logs.

An observable with at most a logarithmic divergence looks like this

$$f(I) = \sum_{n=0}^{\infty} \frac{a_n + b_n \log I}{I^n} \quad I = L/a$$

M. Lüscher, P.Weisz, 1996,

One can extract the coefficients using successive fits.

A. Bode, P.Weisz, U.Wolff, 2000.

But: One requires many data points with high precision.