Holography and Pseudospectra

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EXCELENCIA SEVERO OCHOA

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D. Arean, D. Fariña, K.L. JHEP 12 (2023) 187 • e-Print: 2307.08751 [hep-th]
D. Fariña, K.L., P. Romeu, P. Saura-Bastida arXiv:2407.06104 [hep-th]
D. Arean, D. Fariña, K.L Front.in Phys. 12 (2024) 1460268 • arXiv: 2407.04372 [hep-th]

Open Quantum Gravity, IFT Madrid, 06.02.2025

Pseudospectra and the AdS/CFT correspondence

- Normal Modes
- Black Holes and Quasinormal modes
- AdS(/CFT)
- Pseudospectra
- Pseudospectra of QNMs and CMMs
- Summary and Outlook

Normal Modes

Eigen modes of string:

$$\frac{d^2\Phi(x)}{dx^2} + \lambda\Phi(x) = 0$$

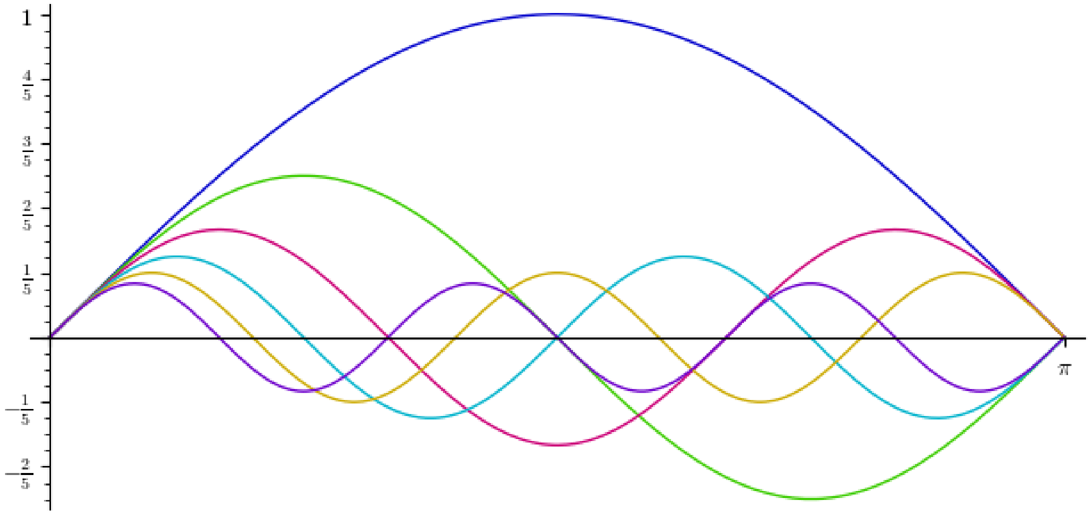
Boundary conditions:

$$\Phi(0) = \Phi(\pi) = 0$$

Hermitian operator:

$$\mathcal{L} = \frac{d^2}{dx^2} \quad , \quad \mathcal{L}^\dagger = \mathcal{L}$$

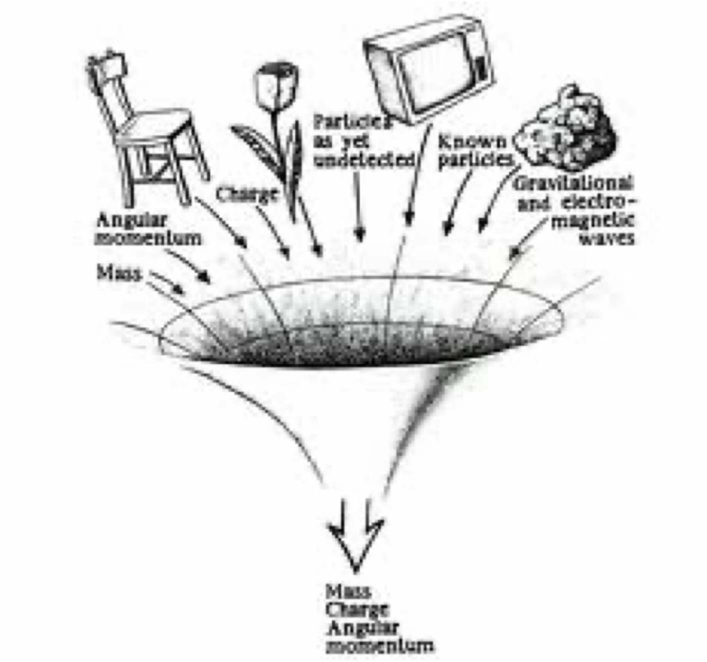
$$\langle \Psi, \Phi \rangle = \int_0^\pi dx \bar{\Psi}\Phi \qquad \langle \Psi, \mathcal{L}\Phi \rangle = \langle \mathcal{L}\Psi, \Phi \rangle$$



$$\Phi_n(x) = \sin(nx)$$
$$\lambda = n^2$$

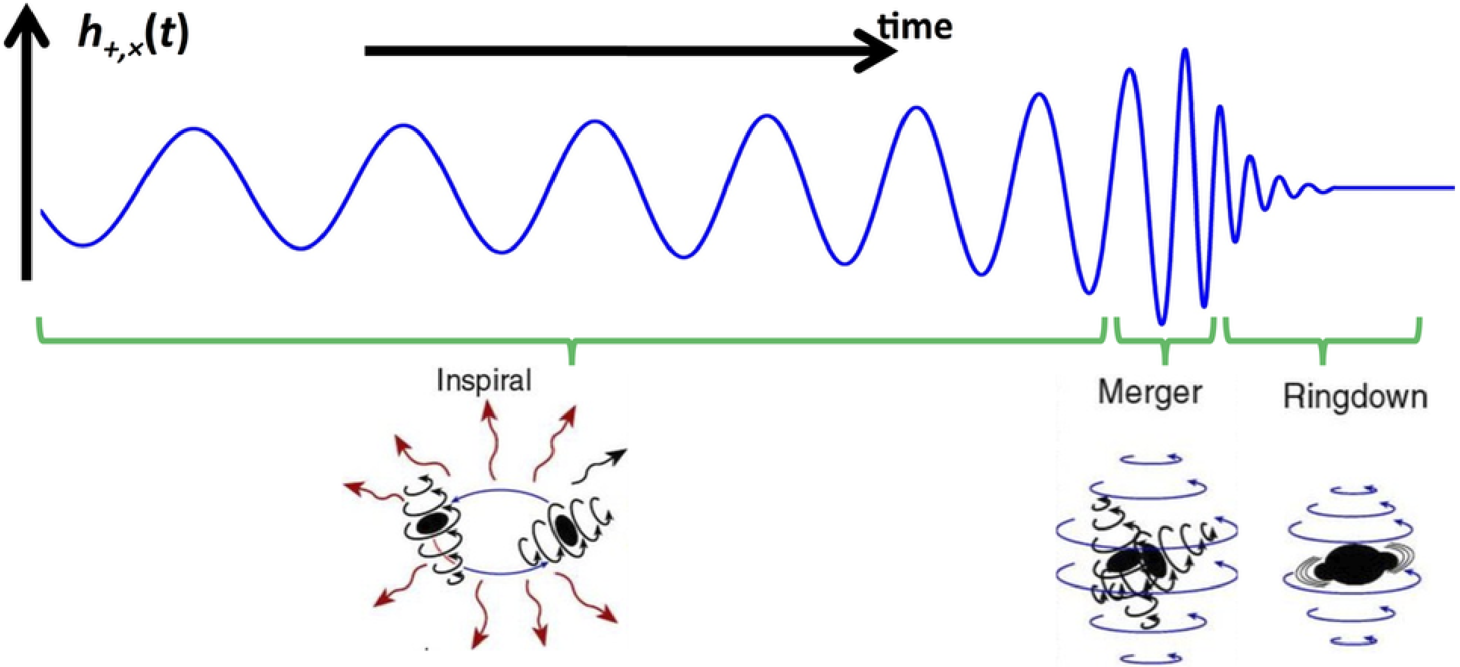
Quasi-Normal Modes

- Black Hole



- ➔ Black Holes no Hair
- ➔ Swallow everything
- ➔ Fate of a perturbation
  - ➔ Either fall into BH
  - ➔ Or radiate off to infinity
  - ➔ Perturbation eventually dies off

Quasi-Normal Modes



[https://www.soundsofspacetime.org/]

Quasi-Normal Modes

How to compute Quasi Normal Modes:

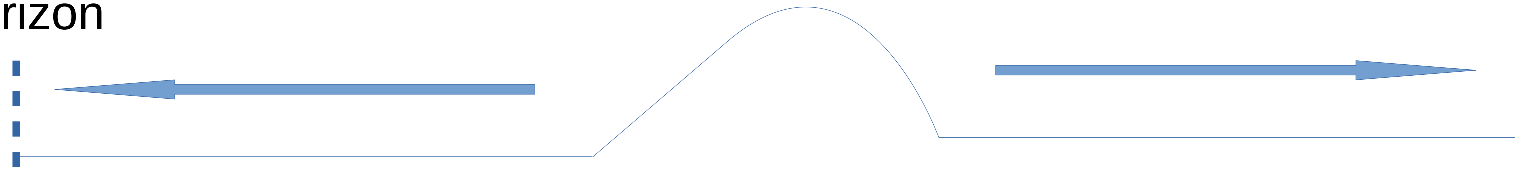
$$ds^2 = -f(r)dt^2 + \frac{dr^2}{f(r)} + r^2 d\Omega^2$$

$$\Phi = \phi(r) e^{-i\omega t} Y_{lm}(\Omega)$$

$$\phi'' + \frac{f'}{f}\phi + \left(\frac{\omega^2}{f^2} - \frac{l(l+1)}{r^2 f} + \frac{f'}{rf}\right)\phi = 0$$

“Outgoing” boundary conditions:

Horizon



$$\Phi \propto e^{-i\omega(t+r_\star)}$$

$$dr_* = \frac{dr}{f}$$

$$\Phi \propto e^{-i\omega(t-r_\star)}$$

[Vishveshwara]
[Press]

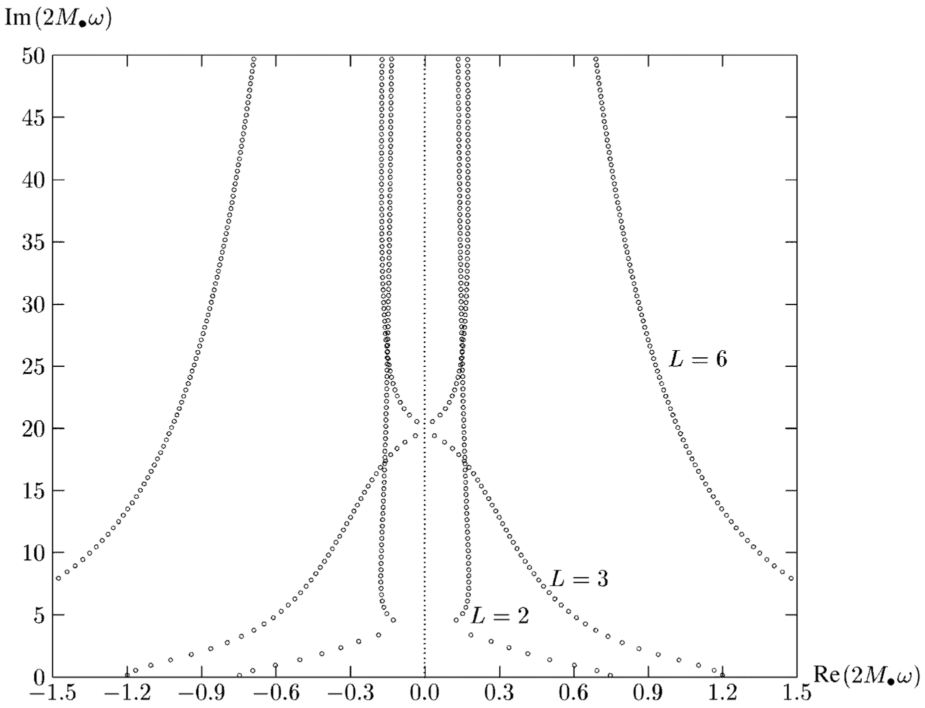
Quasi-Normal Modes

Leaky boundary conditions lead to complex frequencies

[Leaver, Prod. Roy. Soc. A402 (285)]

$$\omega_n = \Omega_n - i\Gamma_n$$

Oscillation   Damping



"Black Hole spectroscopy"

[Berti, Cardoso, Will, PRD 73 (2006)]

Reviews: [Nollert, "TOPICAL REVIEW", Class.Quant.Grav. 16 (1999) R159-R216]
[Berti, Cardoso, Starinets, Class.Quant.Grav. 26 (2009) 163001]

Holographic QNMs

Field in AdS black hole: $\Phi(r,t,\vec{x}) = e^{-i\omega t + i\vec{k}.\vec{x}}\phi_{\omega,\vec{k}}(r)$

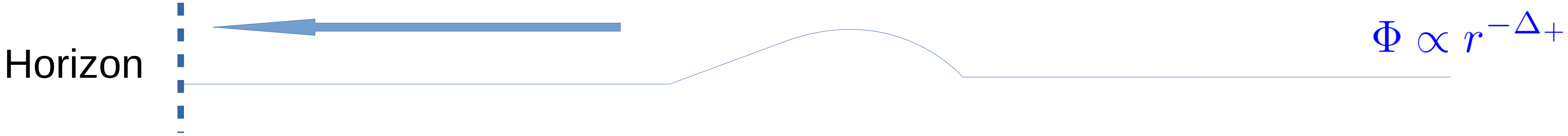
Boundary condition Horizon: $\phi_{\omega,\vec{k}} \propto e^{-i\omega(t+r_*)}$

Boundary condition boundary: $\phi \approx A(\omega,\vec{k})r^{-\Delta_-}(1+\dots) + B(\omega,\vec{k})r^{-\Delta_+}(1+\dots)$

Retarded Green's function:

$$G_R(\omega,\vec{k}) = K\frac{B(\omega,\vec{k})}{A(\omega,\vec{k})}$$

[Horowitz, Hubeny],
[Birmingham, Sachs, Solodhukin]
[Kovtun, Son, Starinets]



Holographic QNMs

Example: Scalar field in BTZ black hole

$$G_R(\omega,k)=\frac{(\omega^2-k^2)}{4\pi^2}\left[\psi\left(1-i\frac{\omega-k}{4\pi T}\right)+\psi\left(1-i\frac{\omega+k}{4\pi T}\right)\right]$$

$$\omega_n=\pm k-i4(n+1)$$ Exact spectrum of QNMs !

In general no exact solution, e.g. scalar in AdS_5:

$$\phi''+\left(\frac{5}{r}+\frac{f'}{f}\right)\phi'+\frac{\omega^2-f^2\vec{k}}{r^2 f}\phi=0$$

"Christmas tree"

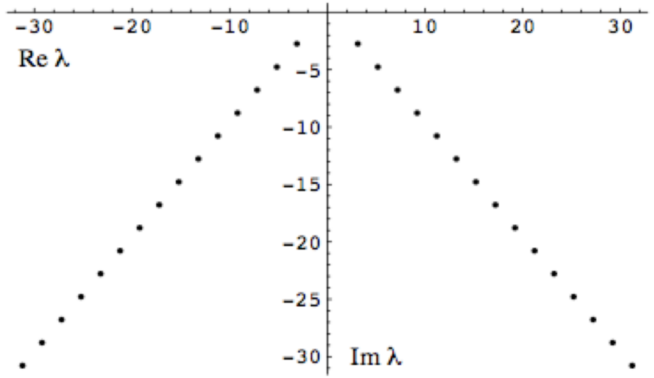


Fig. 1: The lowest 15 quasinormal frequencies in the complex λ-plane for $q=0$.

[Starinets]

Holographic QNMs

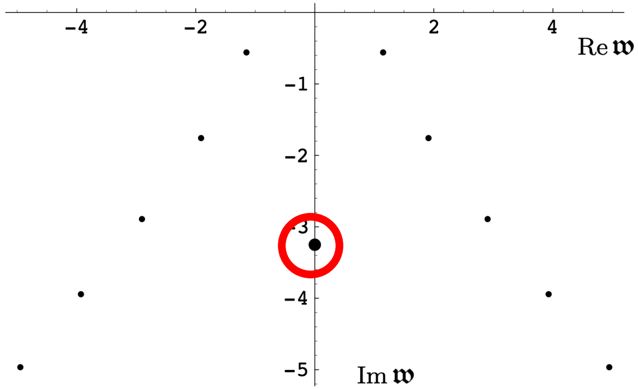
Gauge fields: new ingredient gauge symmetry: conserved current $\partial_\mu J^\mu = 0$

$$\frac{d}{dt}Q = 0 \qquad\qquad Q = \int d^3x J^0$$

2 channels:

- Transverse is like scalar
- Longitudinal new: diffusion

[Son, Starinets]

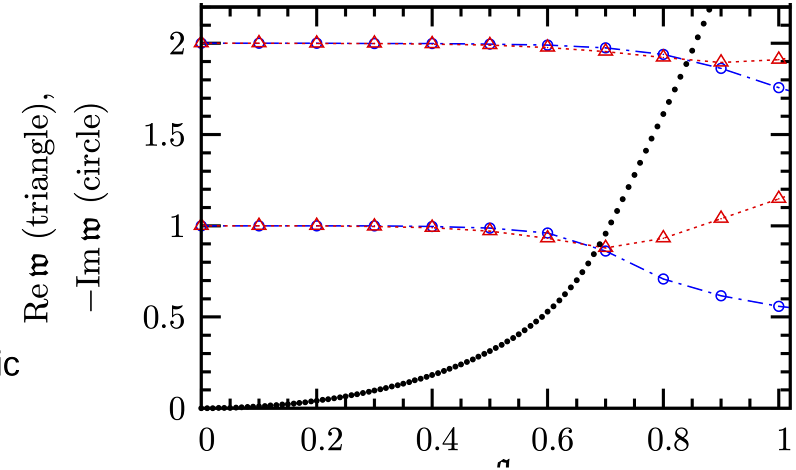


$$\omega_{\text{hydro}}(k) = -iDk^2$$

$$\omega = i\sum_{n=1}^{\infty} a_i k^{2n}$$

Problematic Causality!

[Amado, Hoyos, K.L., Montero]



Holographic QNMs

Metric: 3 channels:

- Tensor (Spin 2) is like scalar δg_{xy}
- Vector (Spin 1) is like vector → Shear channel $\delta g_{tx}, \delta g_{zx}$
- Scalar (Spin 0) new → Sound channel $\delta g_{tt}, \delta g_{zz}, \delta g_{tz}$

Shear = momentum diffusion $\omega(k) = -i\frac{\eta}{\epsilon + p}k^2 + O(k^2)$

$\eta = \frac{1}{4\pi}$ Universality of shear viscosity

[Policastro,Son,Starinets]
[Kovtun,Son,Starinets]

Sound = energy conservation $\omega(k) = \pm v_s k - i\gamma k^2 + O(k^3)$ $v_s^2 = \frac{\partial \epsilon}{\partial p} = 1d = \frac{1}{d}$

Hydrodynamics: $T^{\mu\nu} = (\epsilon + p)u^\mu u^\nu + p\eta^{\mu\nu} - \eta P_T^{\mu\alpha} P_T^{\nu\beta}(\partial_\alpha u_\beta + \partial_\beta u_\alpha - \frac{1}{d}\eta_{\alpha\beta}\partial.u)$

$\partial_\mu T^{\mu\nu} = 0$

Holographic QNMs

Superconducing QNMs:

- Holographic “Superconductor”
- Spontaneous U(1) symmetry breaking
- Goldstone mode = 2nd sound

[Gubser]
[Hartnoll, Herzog, Herzog]

$$S = \int d^4x\sqrt{-g}\left[R + \Lambda - \frac{1}{4}F^2 - |D\Phi|^2 - m^2|\Phi|^2\right]$$

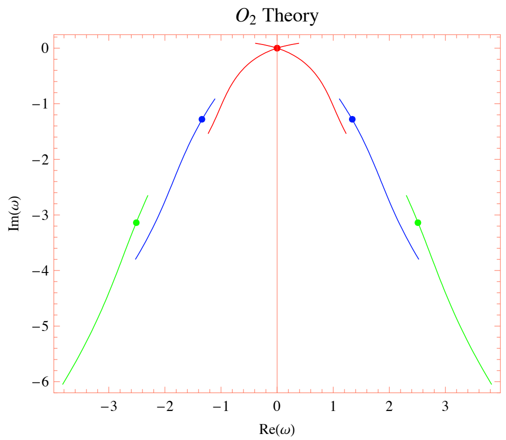
Temperature and chemical potential: $4\pi T = f'(r_h)$

$$\mu = A_t(\infty) - A_t(r_h)$$

Unbroken phase T<T$_c$

Growing mode T>T$_c$ $\Im(\omega) > 0$

[Amado,Kaminski, K.L.]



QNMs

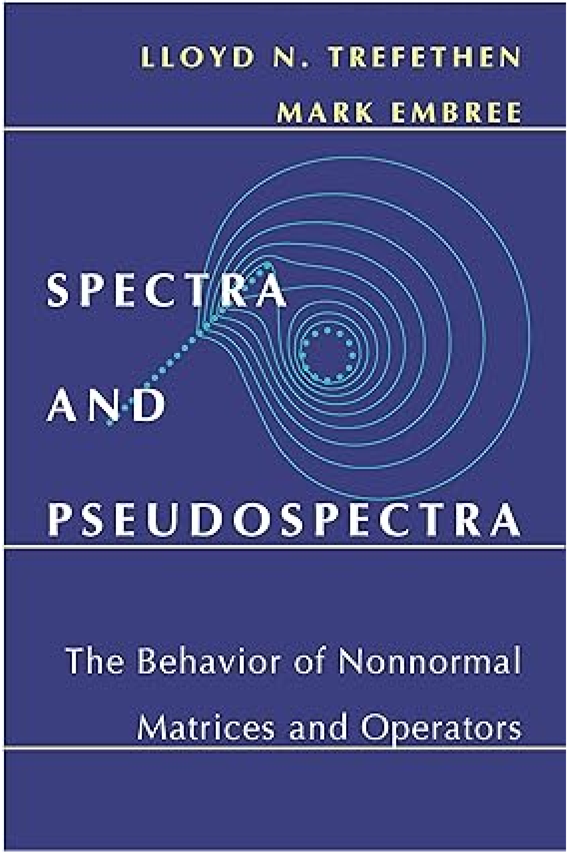
Quasinormal Modes are an essential ingredient to Gauge/Gravity duality!

Pseudospectra

- QNMs are eigenvectors of non-Hermitian operators
- No spectral theorem $\mathcal{O} \neq \sum_n |n\rangle \lambda_n \langle n|$
- QNMs are not complete (only late time)
- Eigenfunctions are singular in Schwarzschild coords
- Robustness and physical significance in question

Near Horizon $\phi(r) = e^{-i\omega r_*} = e^{-i\Omega r_* - \Gamma r_*} \quad , \quad r_* \to -\infty$

No Hilbert space interpretation



Pseudospectra

Resolvent: $\mathcal{R}(\mathcal{L}, z) = (\mathcal{L} - z)^{-1}$

Spectrum: $\sigma(\mathcal{L}) = \{z \in \mathbb{C} : "\mathcal{R}(\mathcal{L}, z) = \infty"\}$

Eigenvalues: $\mathcal{L} u_n = \lambda_n u_n$

Operator norm: $||\mathcal{L}|| = \sup_{u \in H} \frac{||\mathcal{L}u||}{||u||}$ *(Max. SVD)*

Definitions of **ε-Pseudospectra:**

1) Resolvent norm $\sigma_\epsilon(\mathcal{L}) = \{z \in \mathbb{C} \ : \ ||\mathcal{R}(\mathcal{L}, z)|| > 1/\epsilon\}$

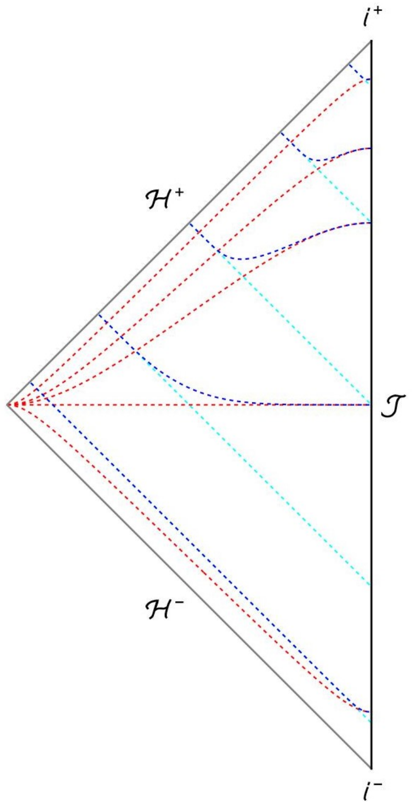
2) Perturbation $\sigma_\epsilon(\mathcal{L}) = \{z \in \mathbb{C}, \exists \delta\mathcal{L}, ||\delta\mathcal{L}|| < \epsilon : z \in \sigma(\mathcal{L} + \delta L)\}$

3) Pseudo eigenvector $\sigma_\epsilon(\mathcal{L}) = \{z \in \mathbb{C}, \exists u^\epsilon : ||(\mathcal{L} - z)u^\epsilon|| < \epsilon||u^\epsilon||\}$

Theorem: The 3 definitions are equivalent

Pseudospectra

How to deal with the QNM problems: chose better coordinates!



Schwarzschild coordinates (worst)

$$ds^2 = r^2\left[-f(r)dt^2 + d\vec{x}^2\right] + \frac{dr^2}{r^2 f(r)}$$

$$\phi(r) \propto (r - r_h)^{-i\omega/2}$$

Infalling Eddington-Finkelstein

$$dv = dt + \frac{dr}{r^2 f(r)}$$

$$ds^2 = r^2\left[-f(r)dv^2 + d\vec{x}^2\right] + 2dvdr$$

$\phi(r)$ regular at $r = r_h$

“Regular”

$$\tau = v - (1 - r_h/r)$$

$$ds^2 = r^2\left[-fd\tau^2 + d\vec{x}^2\right] + 2(1-f)d\tau dr + (2-f)\frac{dr^2}{r^2}$$

[Warnick]

Pseudospectra

Condition number: $$\kappa_i = \frac{||v_i||||u_i||}{|\langle v_i, u_i \rangle|}$$

Right eigenvector: $\mathcal{L}u_i = \lambda_i u_i$

Left eigenvector: $\mathcal{L}^\dagger v_i = \lambda_i^* v_i$

Perturbation: $||\delta\mathcal{L}|| = \epsilon$

Perturbed eigenvalue: $(\mathcal{L} + \delta\mathcal{L})u_i(\epsilon) = \lambda(\epsilon)u_i(\epsilon)$

$$|\lambda(\epsilon)_i - \lambda_i| \leq \epsilon\kappa_i$$

Def "small": Let d_{min} be the minimal distance between disconnected regions in the spectrum.

$\delta\mathcal{L}$, $||\delta\mathcal{L}|| = \epsilon$ is small if

$$\frac{\epsilon}{d_{min}} \ll 1$$

Pseudospectra

We need a physically motivated norm: **Energy** !

Energycurrent: $J = t^{\mu} T_{\mu\nu} dx^{\mu}$ $E[\Phi] = \int_{\Sigma_t} \star J$

Schwarzschild:	$E = \frac{1}{2}\int d^3x dr \left[r^2 f (\Phi')^2 + \frac{(\partial_t \Phi)^2}{r^2 f} + m^2 \Phi^2 + \frac{(\vec{\partial}\Phi)^2}{r^2} \right]$	$\frac{dE}{dt} = 0$
Infalling EF:	$E = \frac{1}{2}\int d^3x dr r^3 \left[r^2 f (\Phi')^2 + m^2 \Phi^2 + \frac{(\vec{\partial}\Phi)^2}{r^2} \right]$	$\frac{dE}{dv} = -r_h^3 \int_{r=r_h} (\partial_v \Phi)^2$
Regular:	$E = \frac{1}{2}\int d^3x dr r^3 \left[r^2 f (\Phi')^2 + m^2 \Phi^2 + \frac{(\vec{\partial}\Phi)^2}{r^2} + (2-f)(\partial_\tau \Phi)^2 \right]$	$\frac{dE}{d\tau} = -r_h^3 \int_{r=r_h} (\partial_\tau \Phi)^2$

Pseudospectra

- We want to re-write the wave equation as a standard eigenvalue problem
- Only possible in regular coordinates

$$\psi = \partial_\tau \Phi \qquad \Psi = \begin{pmatrix} \Phi \\ \psi \end{pmatrix} \qquad \mathcal{L} = i \begin{pmatrix} 0 & 1 \\ L_1 & L_2 \end{pmatrix} \Rightarrow \mathcal{L}\Psi = \omega \Psi$$

- Compactify radial coordinate $\rho = 1 - \dfrac{r_h}{r}$

$$L_1 = [f(\rho) - 2]^{-1} \left[\frac{m^2 l^2}{(1-\rho)^2} + \mathbf{q}^2 - (1-\rho)^3 \left(\frac{f(\rho)}{(1-\rho)^3} \right)' \partial_\rho - f(\rho)\partial_\rho^2 \right]$$

$$L_2 = [f(\rho) - 2]^{-1} \left[(1-\rho)^3 \left(\frac{f(\rho)-1}{(1-\rho)^3} \right)' + 2\,(f(\rho)-1)\,\partial_\rho \right]$$

- Adjoint operator in energy norm $\mathcal{L}^\dagger = \mathcal{L} + \begin{pmatrix} 0 & 0 \\ 0 & -i\,\delta(\rho) \end{pmatrix}$ $\qquad \dfrac{d}{d\tau} E = -\bar{\psi}\psi|_{\rho=0}$

AdS in IEF [Cownden, Pantelidou, Zilhao] e-Print: 2312.08352

Pseudospectra

- No exact solutions → numerical methods
- Pseudospectral methods

- Differential operator becomes a (N+1)x(N+1) matrix D $F'(\rho_j) = D_{jk} F(\rho_j)$
- Boundary conditions: delete rows and columns corresponding to $\rho = 1$
  regularity corresponds to no boundary condition at $\rho = 0$
- Resolvent norm becomes maximal svd $||(\mathcal{L} - \omega \mathbf{1})^{-1}||_E \approx \inf(\mathrm{sv}((L - \omega \mathbf{1})_{N\times N})$
- Energy norm becomes a metric 2Nx2N matrix

$$E \approx \bar{u}_k^* G_E^{km} u_m \quad , \quad u \approx (\phi(\rho_j), \psi(\rho_j))^T$$

Pseudospectra

A toy example: $A = \begin{pmatrix} -1 & 0 \\ -50 & -2 \end{pmatrix}$ Eigenvalues: $\lambda_1 = -1 \ , \ \lambda_2 = -2$

1. ℓ_2 norm: $||u|| = [\bar{u}.u]^{1/2}$

$$\kappa_1 = \kappa_2 = \sqrt{2501} \approx 50$$

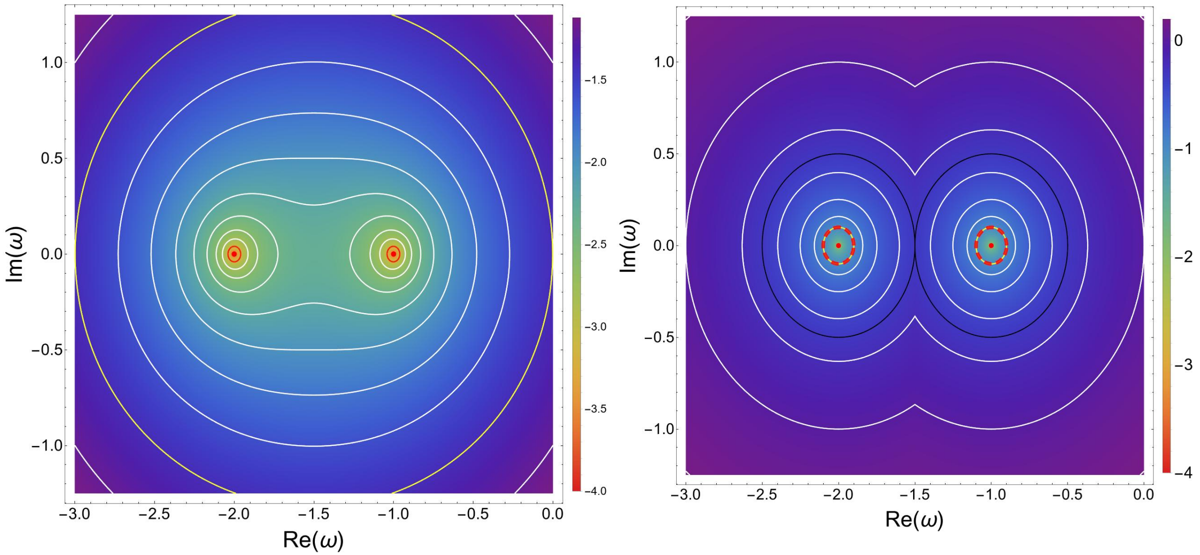
1. G-norm: $||u||_G = [\bar{u}.G.u]^{1/2}$

$$A^\dagger = \left[(G.A.G^{-1})^T\right]^* = A$$

$$G = \begin{pmatrix} 20000 & 50 \\ 50 & 1 \end{pmatrix}$$

$$\kappa_1 = \kappa_2 = 1$$

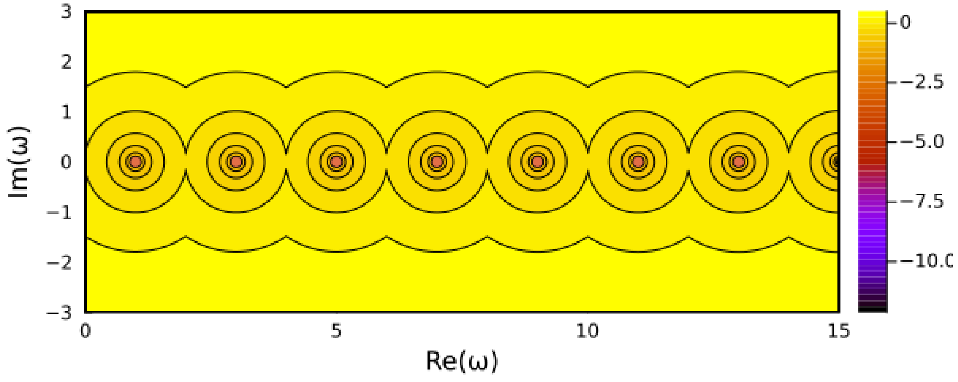
Contour maps of $\log ||A - \omega \mathbf{1}||$



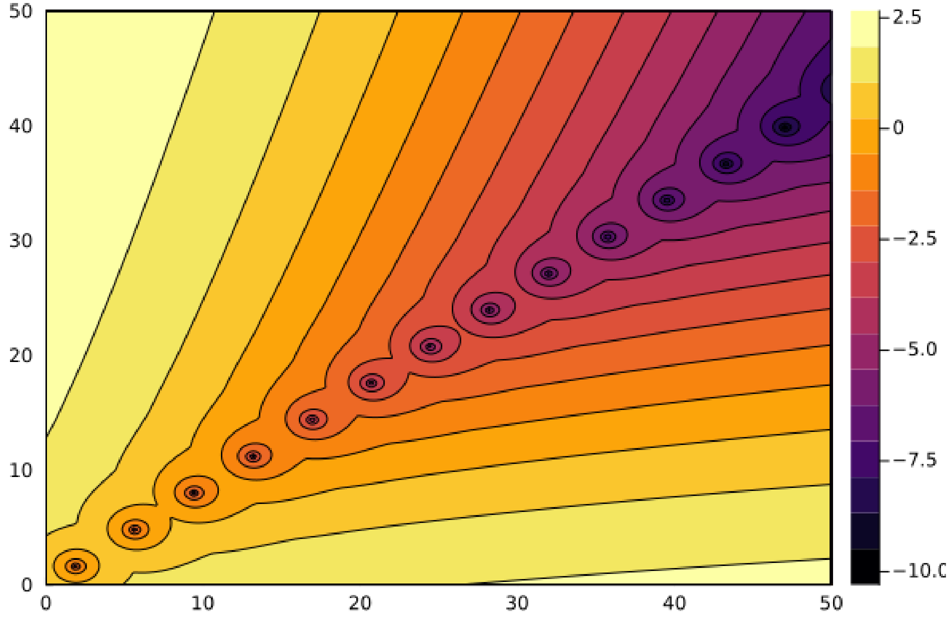
Pseudospectra

Harmonic Oscillator: $$-\frac{d^2\phi}{dx^2} + c\,x^2\,\phi = \omega\phi \qquad G_E = \int dx\bar{\phi}\phi$$

c=1



c=1+3i



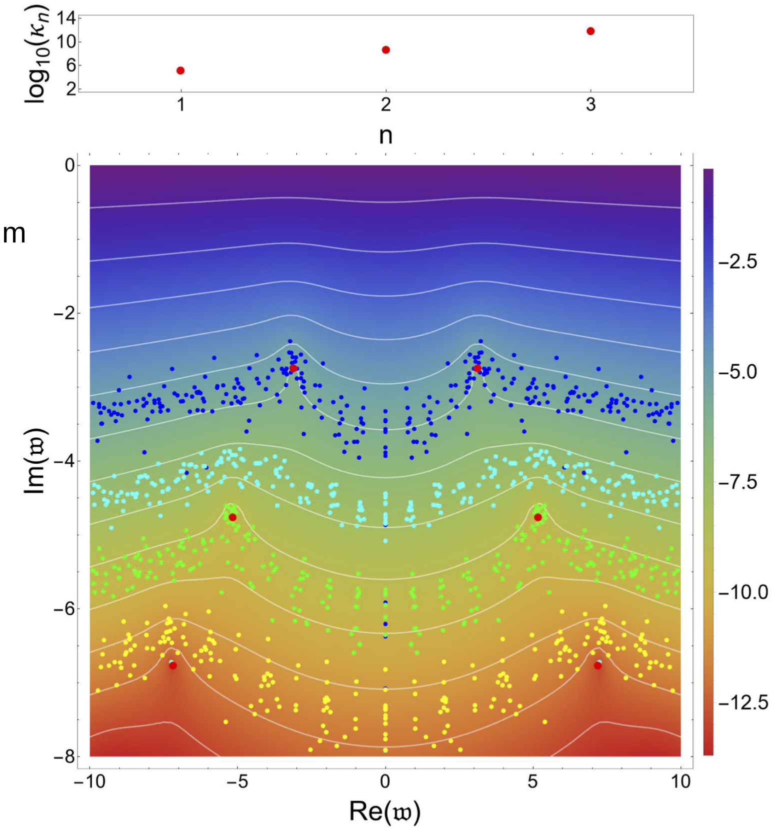
QNM-Pseudospectra

Pseudospectra of massless scalar in AdS_5:

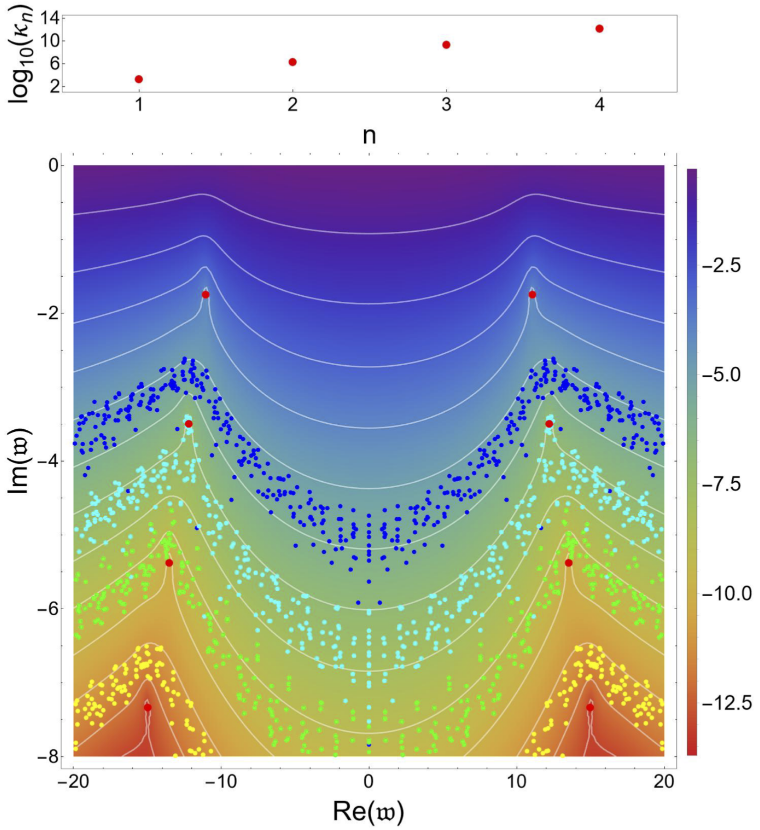
"Selective" Pseudospectrum
Local random
Potential perturbations

$||V_{\mathrm{rand}}||$

- 10^{-1}
- 10^{-3}
- 10^{-5}
- 10^{-7}



(a) $m^2 l^2 = 0$, $\mathfrak{q} = 0$.



(b) $m^2 l^2 = 0$, $\mathfrak{q} = 10$.

QNM-Pseudospectra

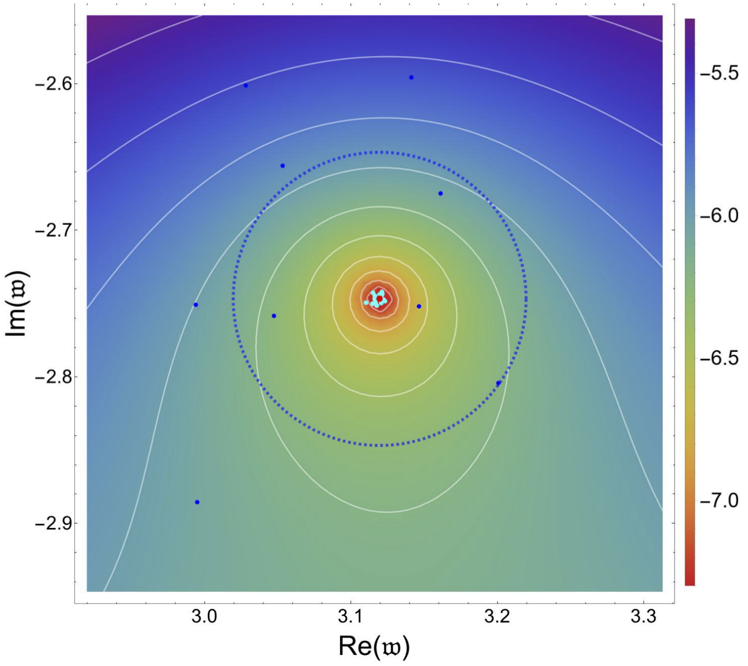
Zoom into first QNM:

Circle of stability

$|\lambda(\epsilon)_i - \lambda_i| \leq \epsilon \kappa_i$

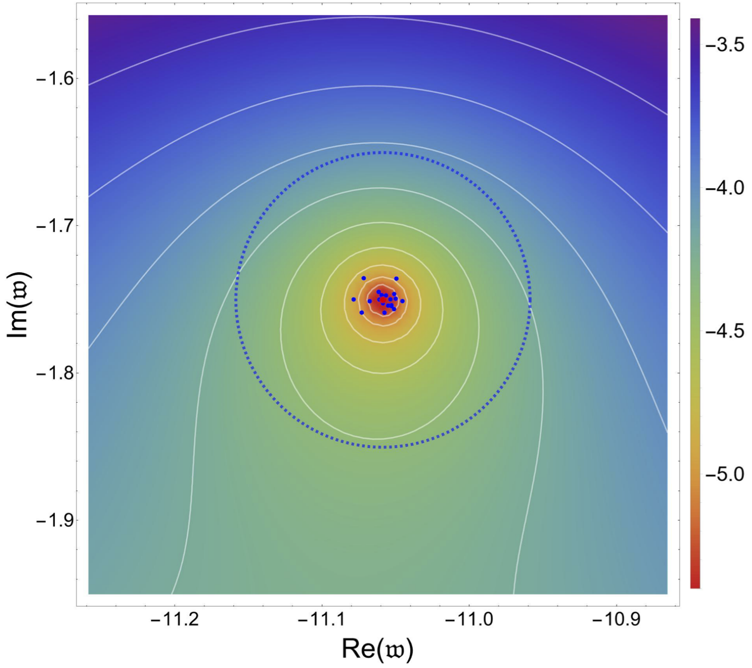
● 10^{-1}
● 10^{-3}

Restricted spectral instability



(a) $m^2 l^2 = 0$, $\mathfrak{q} = 0$.

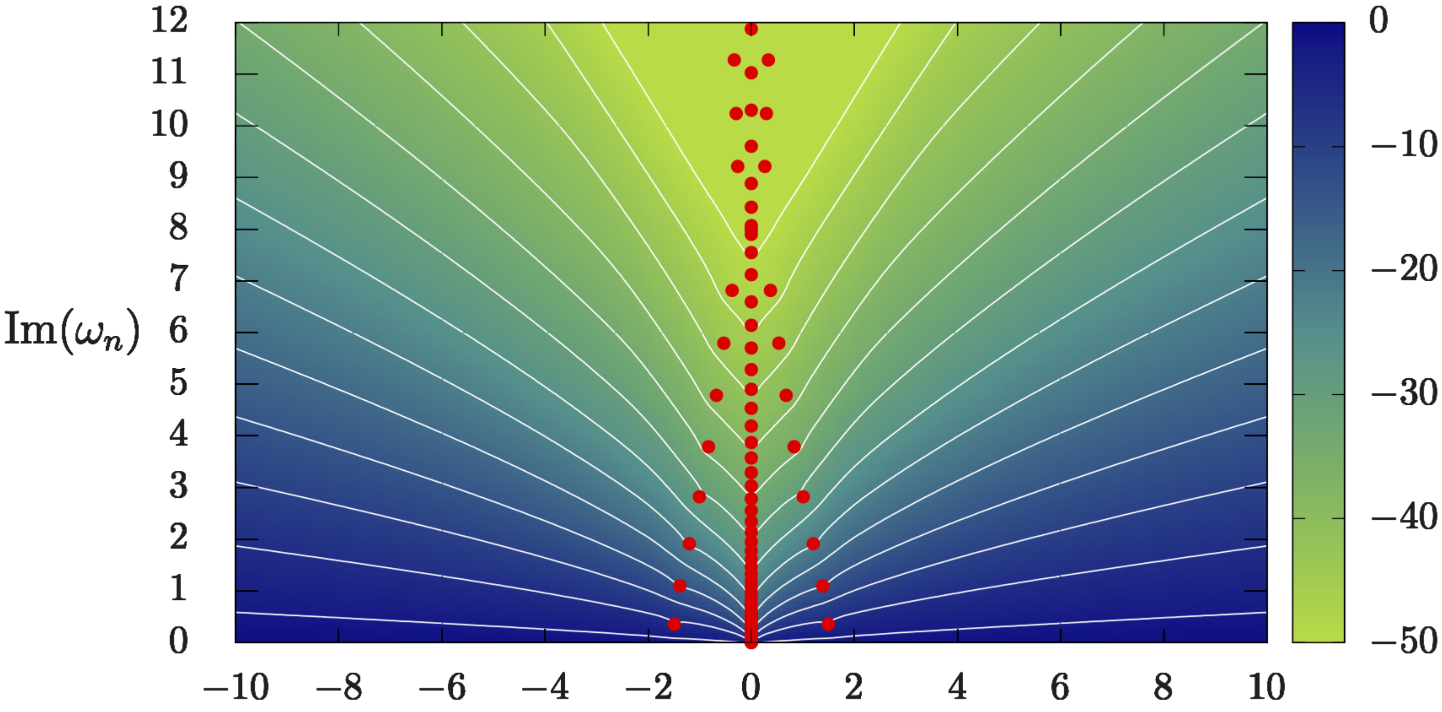
Restricted spectral stability



(b) $m^2 l^2 = 0$, $\mathfrak{q} = 10$.

QNM Pseudospectra

In asymptotically flat space: [Jaramillo, Macedo, Al Sheik, PRX 11 (2021) 3, 031003 • e-Print: 2004.06434]



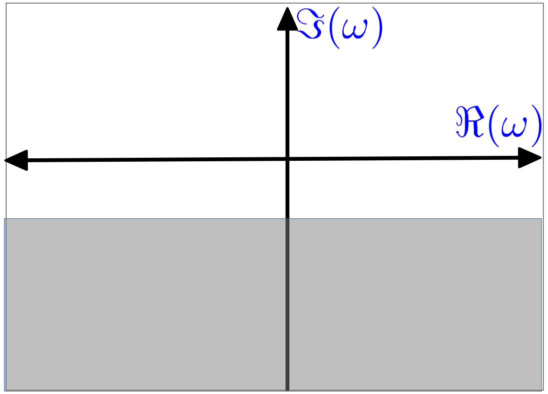
Review: [Destounis, Duque] e-Print: 2308.16227

QNM-Pseudospectra

Caveats: Ingoing modes are integrable in the energy norm:

$$\phi_{\text{in}} \approx \rho^{i\omega} = \rho^{i\Omega}\rho^{\Gamma}$$

$$|\phi'_{\text{in}}|^2 \approx \rho^{2\Gamma-1} \qquad \Gamma = -\Im(\omega) > \frac{1}{2}$$



- Hilbert space of square integrable functions with energy norm
  In this sense the ingoing solutions belong to the spectrum
- Pseudospectrum in IEF N → infinity limit converges to this limit
- Math: Sobolev norm, Hilbert space $H^{(k)}$
  Physics: higher derivative theories!

$$||\Phi||^2 = \int \sum_{m=0}^{k} |D^m \Phi|^2$$

- In regular coords resolvent seems to be divergent in general
- Finite N provides a natural cutoff → what is the optimal N?
  e.g stretched Horizon, Planck length away from horizon
- Universality? Holographic interpretation?

[Warnick: CMP. 333 (2015) 2, 959-1035 • e-Print: 1306.5760 [gr-qc]]

[Boyanov, Cardoso, Destounis,Jaramillo, Macedo] e-Print: 2312.11998 [gr-qc]

Pseudospectra of CMMs

$$G_R(\omega,\vec{k}) = K\frac{B(\omega,\vec{k})}{A(\omega,\vec{k})}$$

Look for poles with complex k and real ω

$$\Phi \propto e^{-i\omega t + ikx} e^{-\kappa|x|}$$

Absorbtion length: $L = \frac{1}{\kappa}$

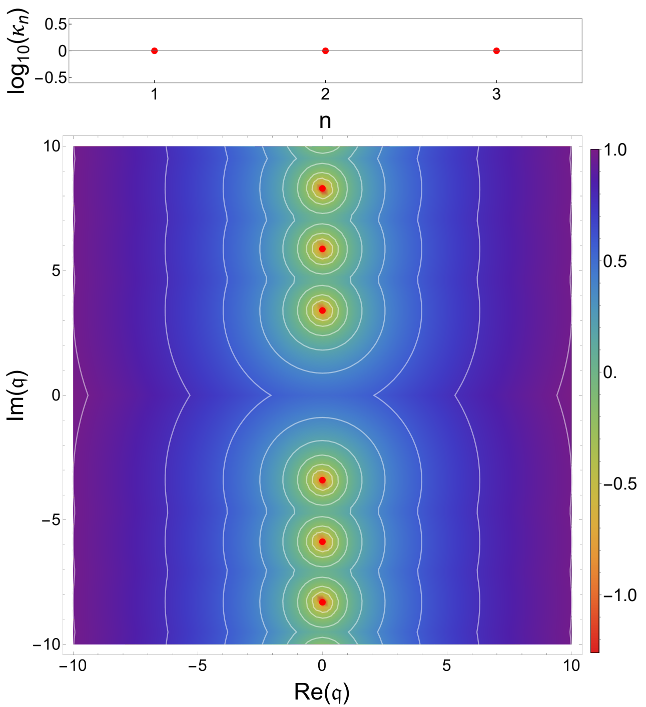
Norm = time averages energy density:

$$\bar{\varrho}_E[\phi] = \int \frac{d\rho}{(1-\rho)^3}\left\{ f\partial_\rho\phi\partial_\rho\bar{\phi} + z_h^2\partial_3\phi\partial_3\bar{\phi} + z_h^2(2-f)\omega^2\phi\bar{\phi} + \frac{m^2l^2}{(1-\rho)^2}\phi\bar{\phi}\right\}$$

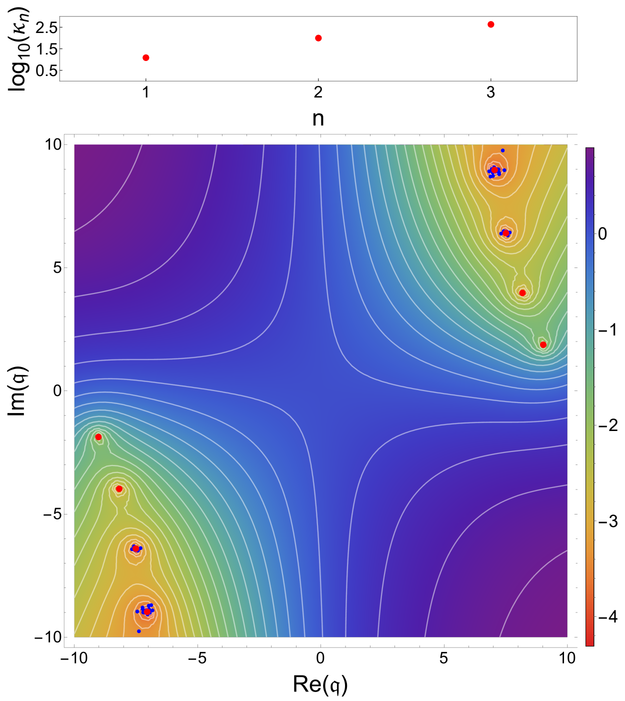
In asymptotically flat space: Regge poles or Complex Angular Momentum

Pseudospectra of CMMs

$\omega = 0$



$\omega = 10\pi T$



Summary

- Quasinormal Modes are central to black hole physics
- Subject to spectral instability
- Choice of norm is important → Energy norm
- QNMs or continuous spectrum? → Norm dependence
- CMMs do not suffer from this problem, well defined
- Pseudospectra of CAMs for asymptotically flat black holes?
- Physical significance: Hammer vs. Antenna